

# EXPLORING BSM MODELS AT LONG BASELINE NEUTRINO EXPERIMENTS

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# NEUTRINO OSCILLATIONS AT LONG BASELINE EXPERIMENTS

$$i \frac{d}{dt} \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \left[ \frac{1}{2E_\nu} U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} U^\dagger + A_{CC} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} \right] \begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix}$$

- GeV neutrinos,  $\sim 10^3$  Km baseline  $\longrightarrow$  Atmospheric oscillation maxima  $\Delta m_{31}^2 L/4E \sim (2n + 1)\pi/2$
- Well known muon neutrino beam  $\longrightarrow$  Possibility of spectral analysis, small electronic neutrino contamination
- Possibility to look at different oscillation channels

$$\nu_\mu \rightarrow \nu_e$$

$(\theta_{13}, \theta_{23}, \delta)$

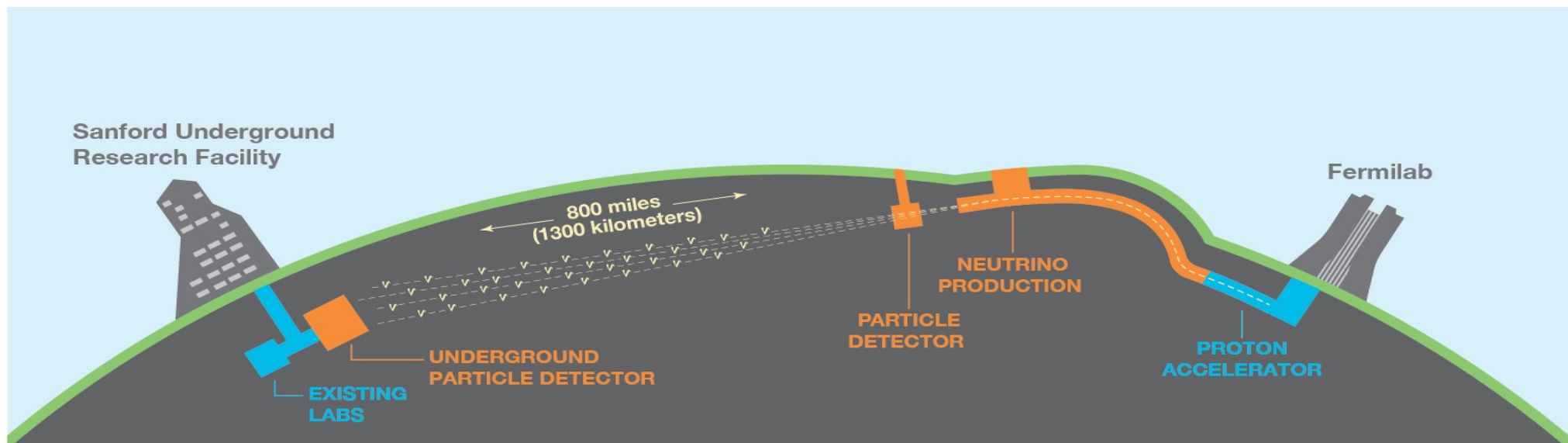
$$\nu_\mu \rightarrow \nu_\mu$$

$(\theta_{23}, \Delta m_{31}^2)$

$$\nu_\mu \rightarrow \nu_\tau$$

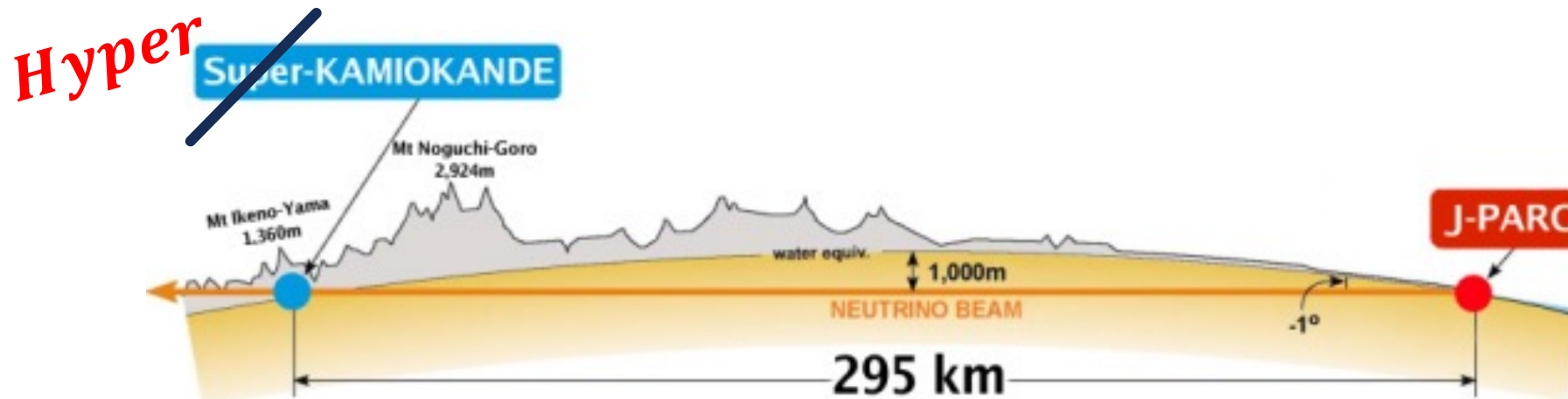
BSM  
 $E_\nu > 3 \text{ GeV}$

# FUTURE LBL EXPERIMENTS: DUNE



- 2.5 GeV peaked neutrinos,  $L=1200$  km **FIRST OSCILLATION MAXIMUM**
- On axis, broad band beam **L/E SCAN**
- All channels accessible (also NC), LAr-TPC detectors **GREAT PRECISION ON PARAMETERS AND MH**
- Matter effects **FAKE CP VIOLATION, STILL 70% CP SENSITIVITY COVERAGE AT  $3\sigma$**

# FUTURE LBL EXPERIMENTS:T2HK



- 0.6 GeV peaked neutrinos,  $L=295$  km **FIRST OSCILLATION MAXIMUM**
- Off-axis, narrow band beam **PRECISION MEASUREMENTS AT THE FIRST MAXIMUM**
- Unprecedented statistics on electron appearance (Cherenkov detector) **GREAT PRECISION ON PARAMETERS AND  $MH$**
- Almost no matter effects **GOOD SENSITIVITY TO CPV**

## FUTURE LBL EXPERIMENTS:T2HKK



- Second proposed detector in Korea. **SECOND OSCILLATION MAXIMUM**
- Same beam and same detector of the T2HK **INCREASED STATISTICS**
- Improvement of CPV sensitivity due to the clean environment provided by the second oscillation maximum



# STUDY OF BSM MODELS AT DUNE AND T2HKK



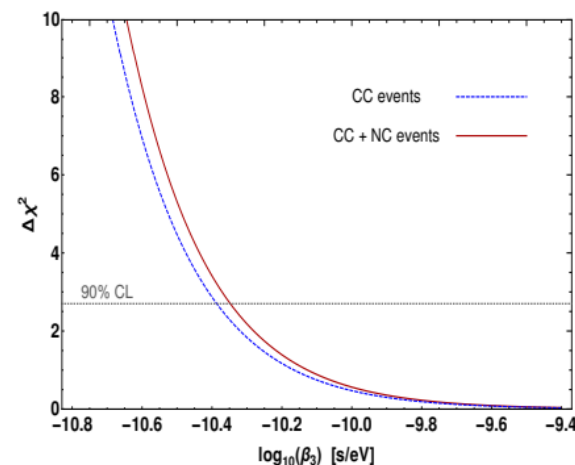
# INVISIBLE DECAY MODEL AT DUNE (GIARNETTI, GHOSHAL, MELONI 2003.09012)

**Invisible decay:** one of the neutrinos decays into invisible particles, e.g. sterile neutrino states

$$H = U \left[ \frac{1}{2E} \begin{pmatrix} 0 & 0 & 0 \\ 0 & \Delta m_{21}^2 & 0 \\ 0 & 0 & \Delta m_{31}^2 \end{pmatrix} - i \frac{1}{2\beta_3 E} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix} \right] U^\dagger + \begin{pmatrix} A & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

- All the CC channels and the **NC channel** are sensitive to the decay parameter

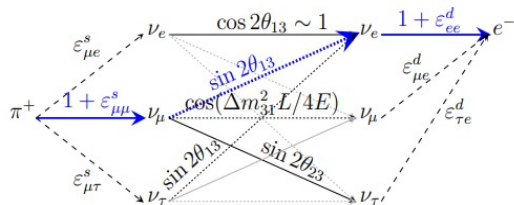
$$\sum_{\alpha}^{e,\mu,\tau} P_{\mu\alpha} = 1 + (e^{-\frac{L}{\beta_3 E}} - 1) \cos^2 \theta_{13} \sin^2 \theta_{23} \neq 1$$



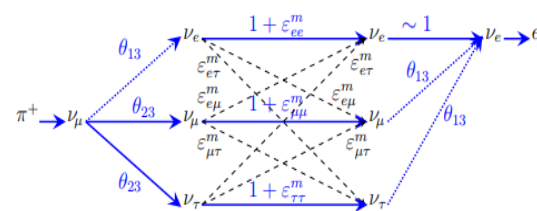
**New constraint from  
DUNE  $5.1 \times 10^{-11}$  s/eV**

# NON STANDARD INTERACTIONS

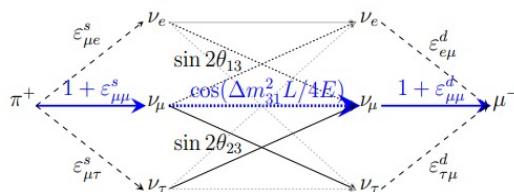
- Source NSI:** occur when neutrinos are produced by hadrons decays at the accelerator level
- Propagation NSI:** occur while neutrinos travel in the matter
- Detector NSI:** occur while neutrinos are detected



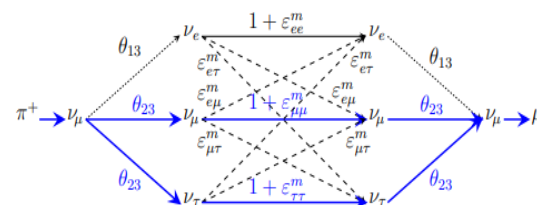
(a)



(a)



(b)

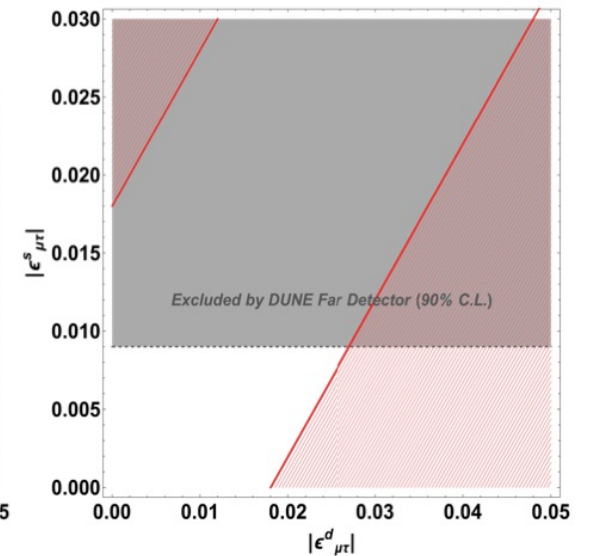
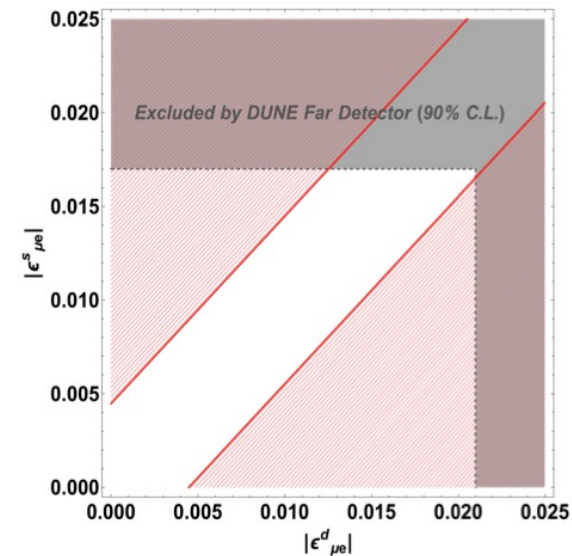
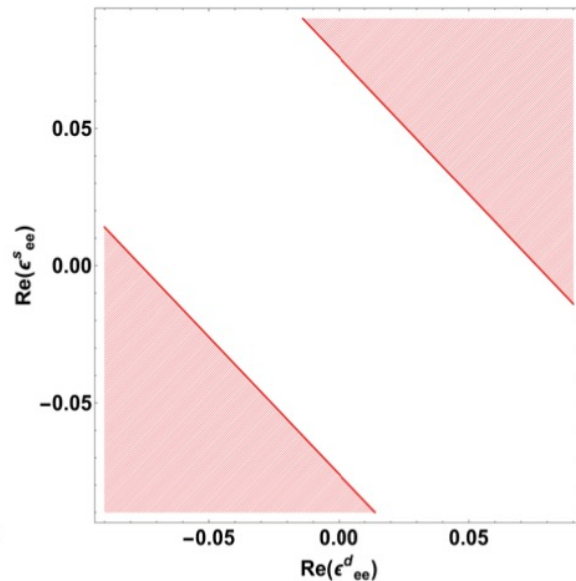
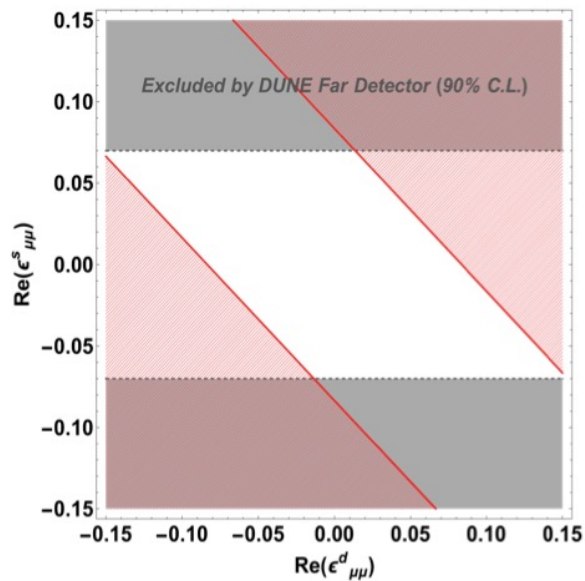




# SOURCE AND DETECTOR NSI AT DUNE ND (GIARNETTI, MELONI 2005.10272)

$$P_{\alpha\alpha} = 1 + 2|\varepsilon_{\alpha\alpha}^s| \cos \Phi_{\alpha\alpha}^s + 2|\varepsilon_{\alpha\alpha}^d| \cos \Phi_{\alpha\alpha}^d$$

$$P_{\alpha\beta} = |\varepsilon_{\alpha\beta}^s|^2 + |\varepsilon_{\alpha\beta}^d|^2 + 2|\varepsilon_{\alpha\beta}^s||\varepsilon_{\alpha\beta}^d| \cos(\Phi_{\alpha\beta}^s - \Phi_{\alpha\beta}^d)$$

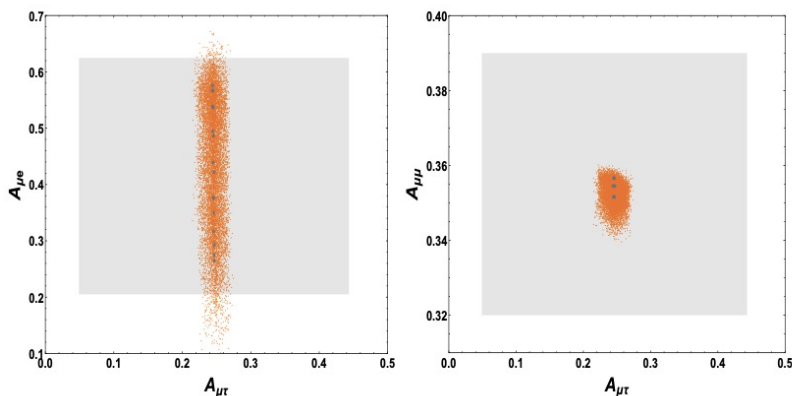


# CP VIOLATION WITH NSI AT DUNE (GIARNETTI, MELONI 2106.00030)

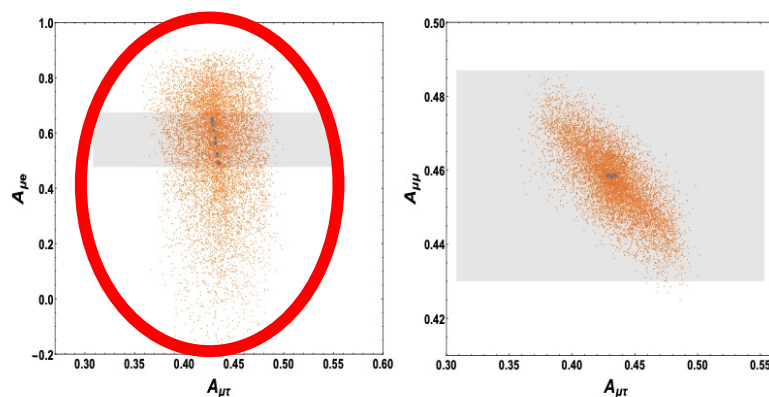
$$A_{\alpha\beta} \equiv \frac{P(\nu_\alpha \rightarrow \nu_\beta) - P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta)}{P(\nu_\alpha \rightarrow \nu_\beta) + P(\bar{\nu}_\alpha \rightarrow \bar{\nu}_\beta)}$$

$\delta_{CP}$  induced

Matter induced



Standard 2.5 GeV flux



High energy flux

## Search for hints of new physics:

- Compute the asymmetries in the standard model with their uncertainties
- Compute the asymmetries in NSI model with parameters in their allowed ranges
- Check whether some sets of new parameters generate anomalous asymmetries

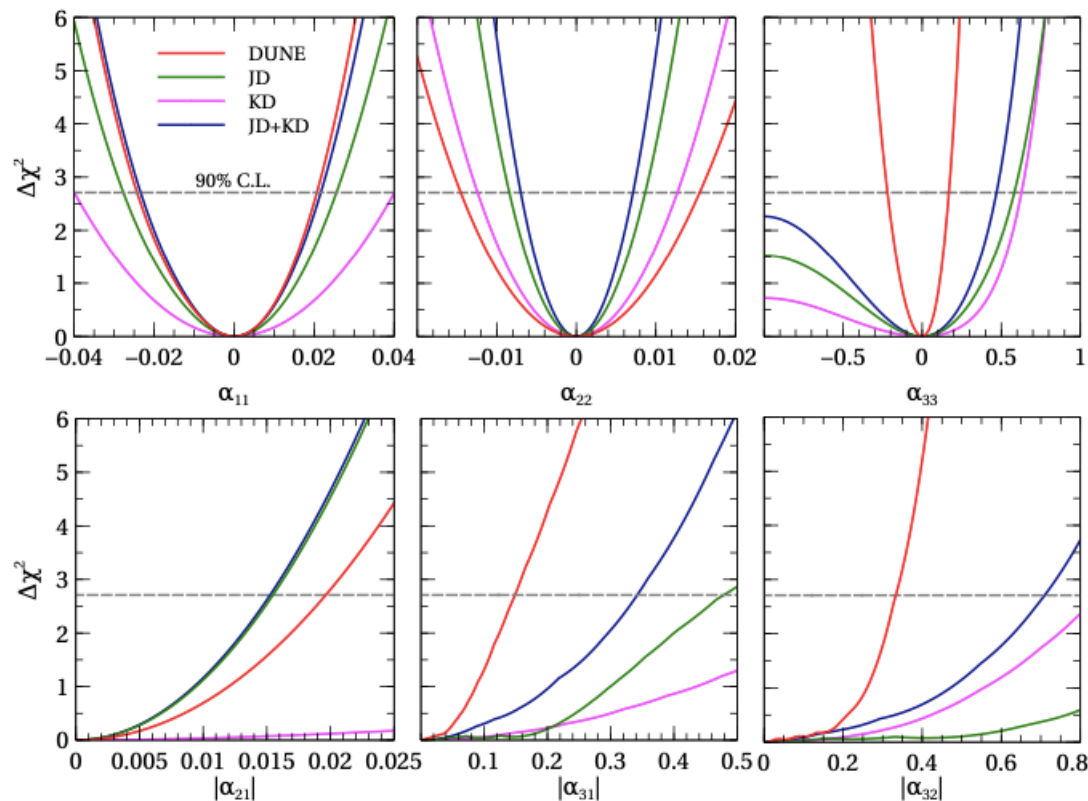
# NON UNITARITY OF THE PMNS MATRIX

$$N = (1 + \alpha)U_{PMNS} . \quad \alpha = \begin{pmatrix} \alpha_{11} & 0 & 0 \\ |\alpha_{21}|e^{i\phi_{21}} & \alpha_{22} & 0 \\ |\alpha_{31}|e^{i\phi_{31}} & |\alpha_{32}|e^{i\phi_{32}} & \alpha_{33} \end{pmatrix}$$

Different models which explain the introduce new particles in the standard model in the leptonic sector. In these cases, the neutrino mixing matrix is no longer 3x3.

Thus, the PMNS matrix that we observe is only a submatrix of the complete one, losing its unitarity property.

# NU AT FUTURE DUNE, T2HK AND T2HKK (AGARWALLA, DAS, GIARNETTI, MELONI, 2111.00329)



-T2HKK has great limits on  $a_{22}$

-DUNE and T2HKK have comparable limits on  $a_{21}$  and  $a_{11}$

-Other parameters sensitivity dominated by DUNE due to bigger matter effects



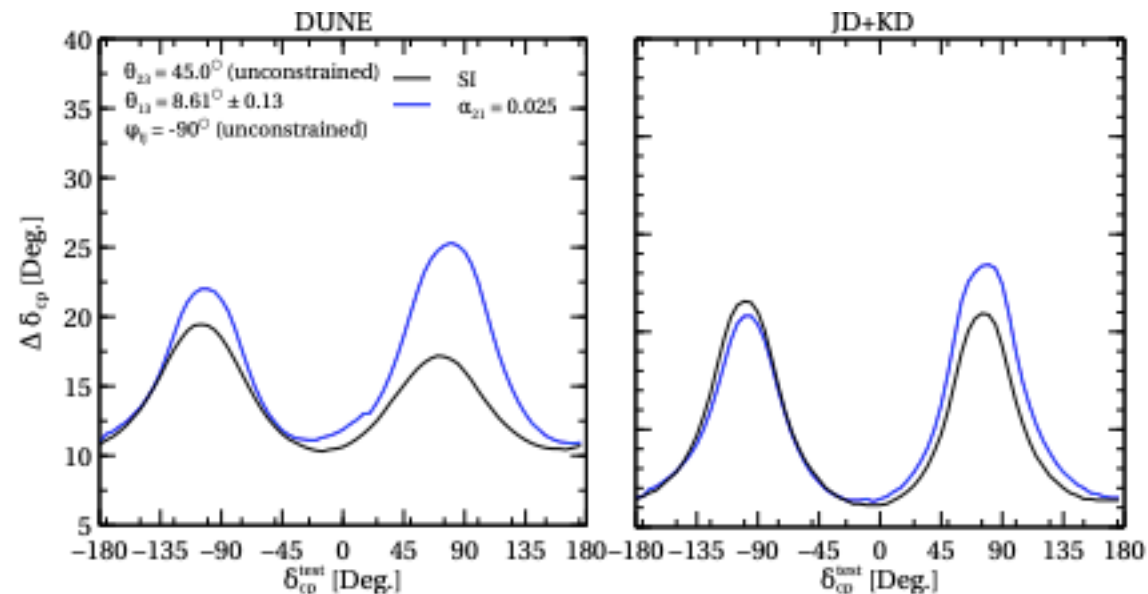
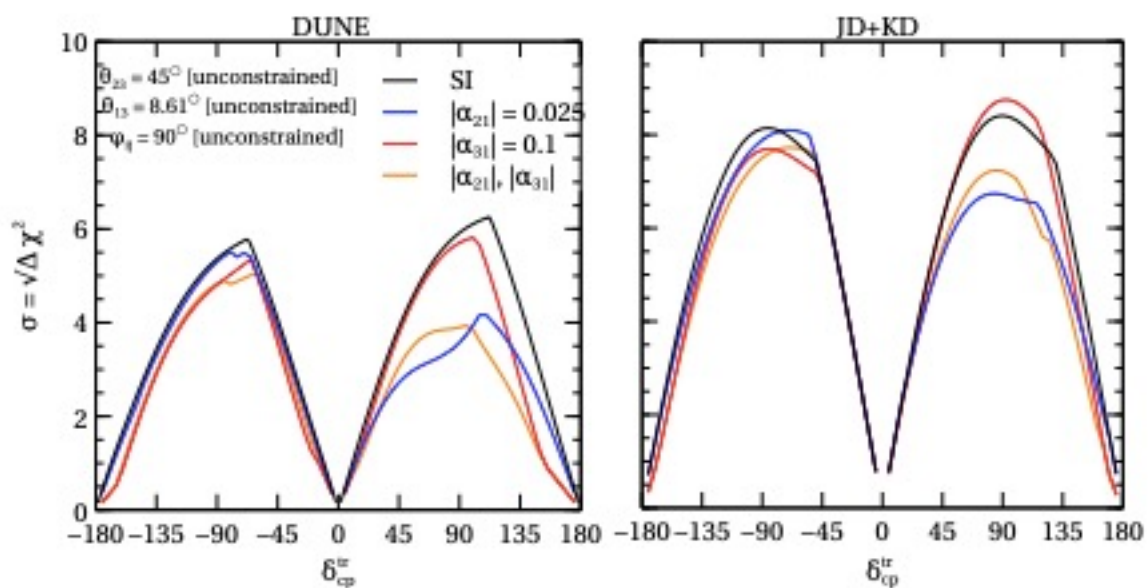
*However, if the flux is measured at the ND, sensitivity to  $a_{22}$  is lost.*

**+** *Improved limits with DUNE  $\nu_\tau$  sample and Near detectors*

# NU AND CPV

$$P_{\mu e} \subset -2|\alpha_{21}|\tau \sin \Delta_{31} \sin(\delta_{CP} - \phi_{21} + \Delta_{31}) + f(\alpha_{31}, A)$$

Effect of NU on CP phase determination!  
(Future developments...)



# PROPAGATION VECTOR NSI VS SCALAR NSI

$$\mathcal{L}_{\text{NSI}}^{\text{eff}} = -2\sqrt{2}G_F \sum_{f,\alpha,\beta} \varepsilon_{\alpha\beta}^f (\bar{\nu}_\alpha \gamma_\rho \nu_\beta) (\bar{f} \gamma^\rho f) \longrightarrow H = \frac{1}{2E} \left[ U M^2 U^\dagger + A \begin{pmatrix} 1 + \varepsilon_{ee} & \varepsilon_{e\mu} & \varepsilon_{e\tau} \\ \varepsilon_{e\mu}^* & \varepsilon_{\mu\mu} & \varepsilon_{\mu\tau} \\ \varepsilon_{e\tau}^* & \varepsilon_{\mu\tau}^* & \varepsilon_{\tau\tau} \end{pmatrix} \right]$$

Vector Propagation NSI: coupled to matter potential, limits from oscillation experiments

$$\mathcal{L}_{\text{scalar NSI}}^{\text{eff}} = y_f Y_{\alpha\beta} (\bar{\nu}_\alpha \nu_\beta) (\bar{f} f) \longrightarrow \delta \tilde{M} = \sqrt{\Delta m_{31}^2} \begin{pmatrix} \eta_{ee} & \eta_{e\mu} & \eta_{e\tau} \\ \eta_{e\mu}^* & \eta_{\mu\mu} & \eta_{\mu\tau} \\ \eta_{e\tau}^* & \eta_{\mu\tau}^* & \eta_{\tau\tau} \end{pmatrix}$$

LBL???

Scalar NSI: modification of the mass matrix, oscillations sensitive to absolute mass scale

# CONCLUSIONS

- Oscillation measurements have now reached a good precision, bounds to new physics parameters are possible
- DUNE and T2HK (T2HKK) are two long baseline experiments which will be able to determine (hopefully) all the unknown oscillation observables
- Their great potential could be used to find bounds on new physics parameters which could be competitive with other non-oscillation bounds
- The complementarity between the two experiments is essential