

Completing the standard model

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GOBIERNO
DE ESPAÑA

MINISTERIO
DE CIENCIA
E INNOVACIÓN



GENERALITAT
VALENCIANA
Conselleria d'Educació,
Investigació, Cultura i Esport



FLASY 2022 | 9th Workshop on Flavour Symmetries and
Consequences in Accelerators and Cosmology

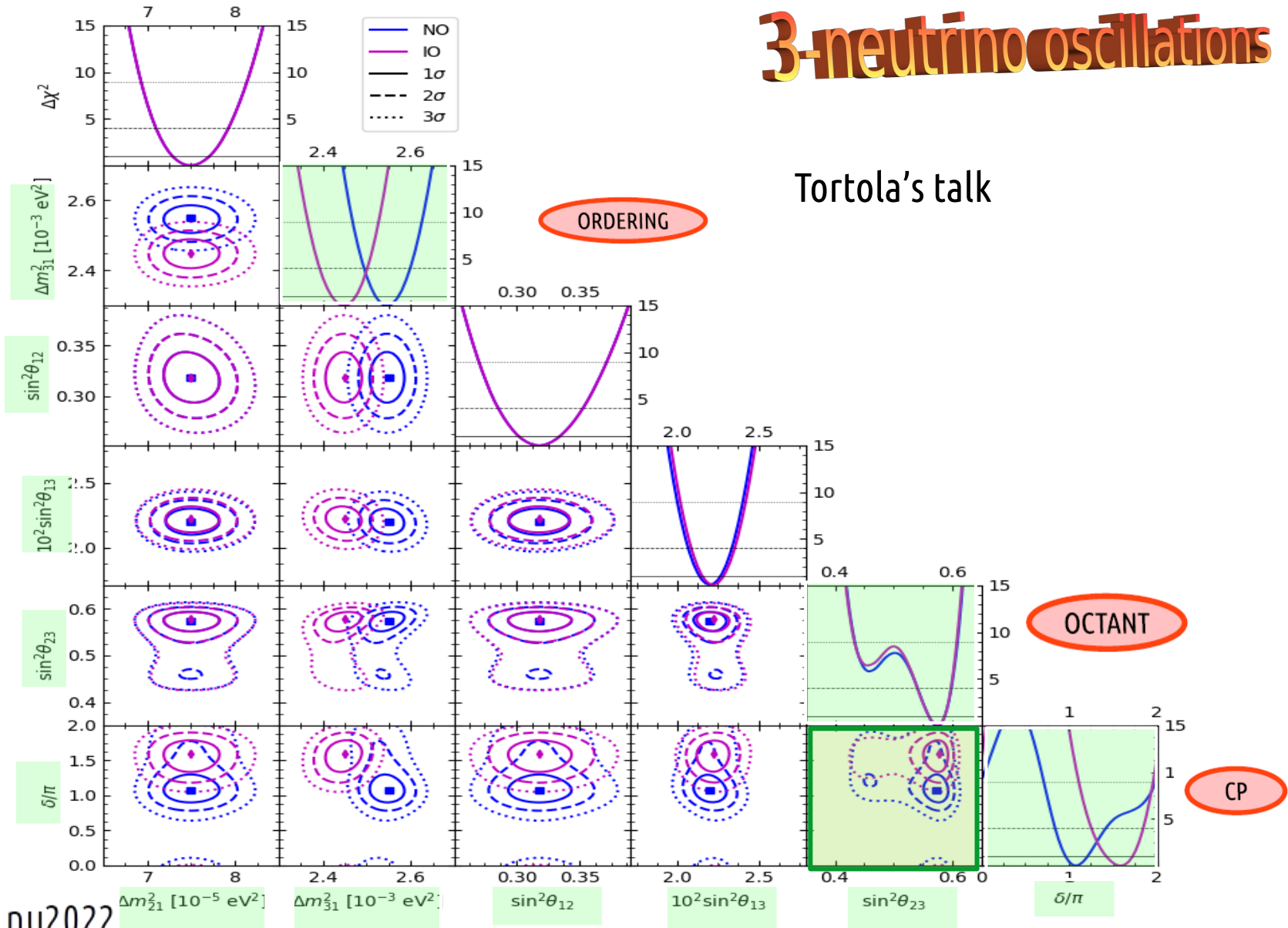
June 27 - July 1, 2022

IST Congress Centre, Instituto Superior Técnico
Lisbon, Portugal

Organised by Centro de Física Teórica de Partículas (CTFP)

3-neutrino oscillations

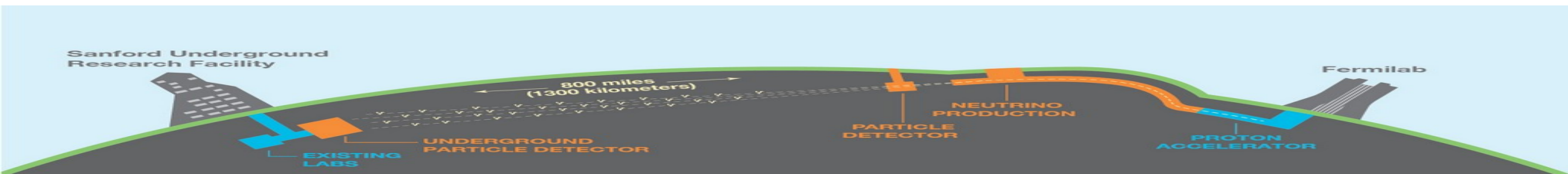
Tortola's talk



Updates from nu2022

$$\sin^2 2\theta_{13} = 0.0853^{+0.0024}_{-0.0024} \quad (2.8\% \text{ precision})$$

DUNE Hyper-K



Leptonic CPV reviews

Nunokawa, Parke, JV

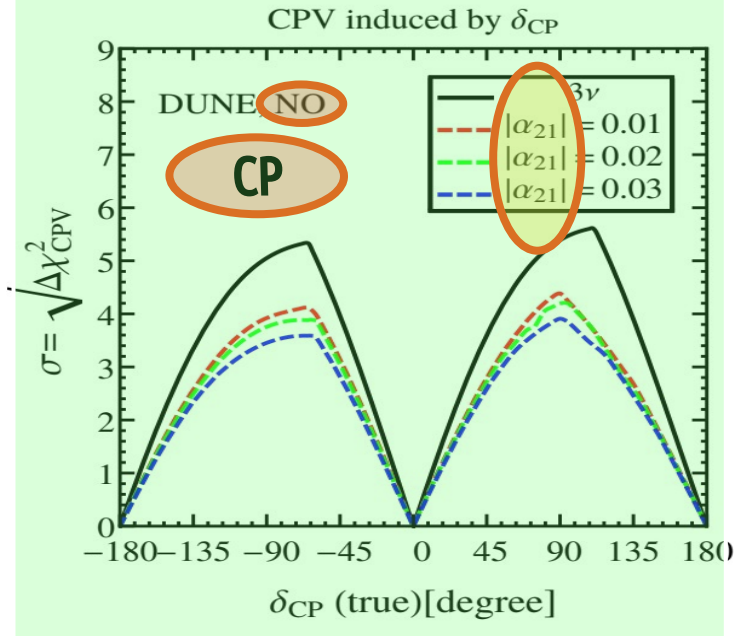
Prog.Part.Nucl.Phys. 60 (2008) 338

Branco, Felipe, Joaquim,

Rev.Mod.Phys. 84 (2012) 515

CPV

DUNE



PhysRevLett117(2016)061804

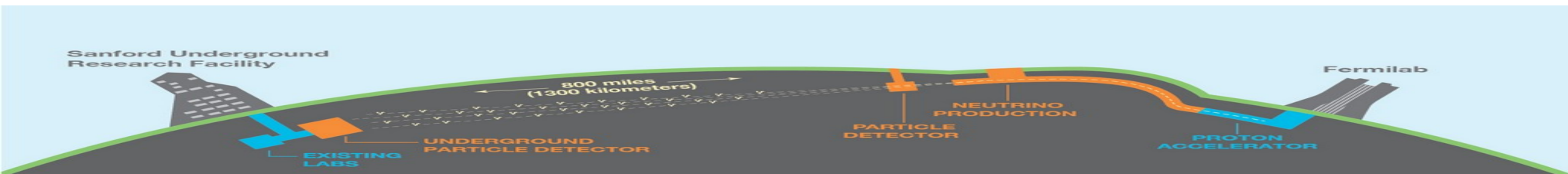
New J.Phys. 19 (2017) 9, 093005

PhysRevD97 (2018) 095026

2008.12769

DUNE Hyper-K

Expected CP discovery Sensitivity: standard 3-nu vs Unitarity violation



Leptonic CPV reviews

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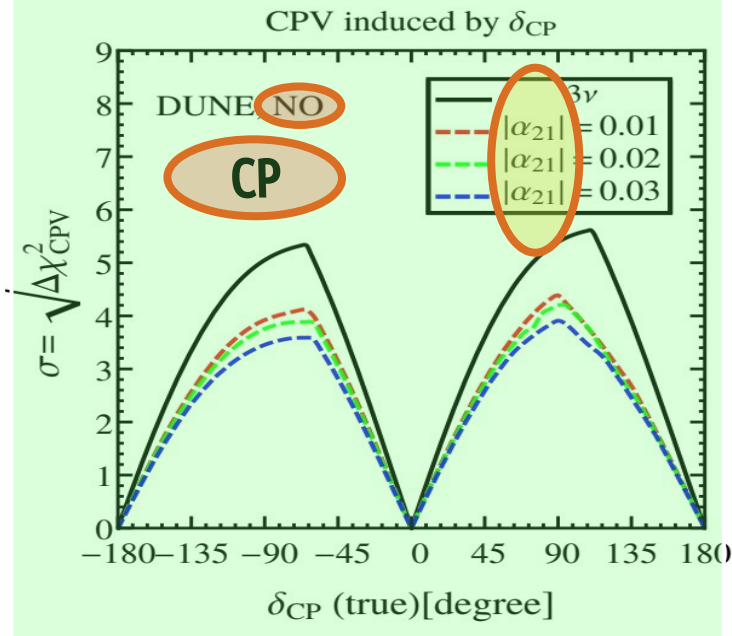
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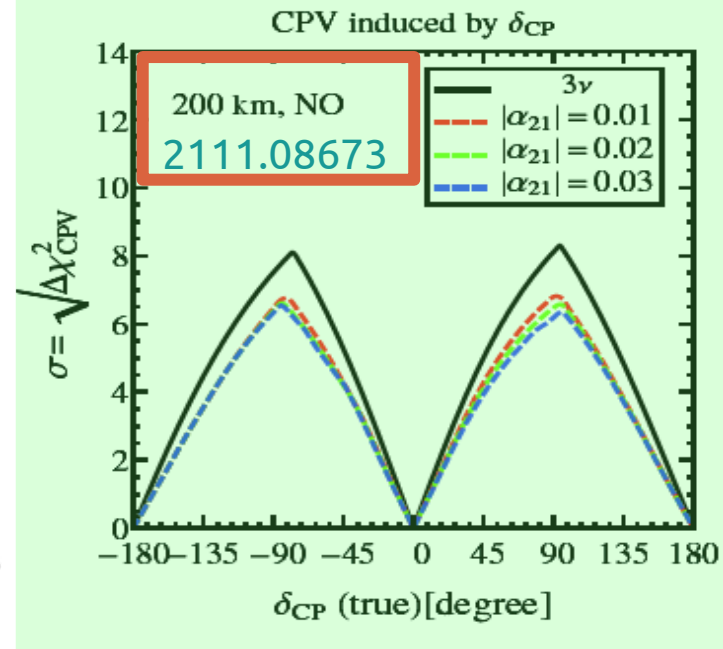
Rev.Mod.Phys. 84 (2012) 515

CPV

DUNE



ESSnuSB



PhysRevLett117(2016)061804

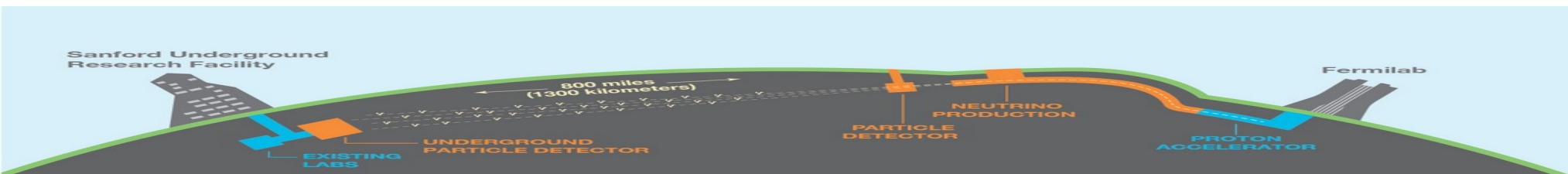
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Branco, Felipe, Joaquim,

Rev.Mod.Phys. 84 (2012) 515

@jwvalle3

TBM interpretation

Harrison,
Scott
& Perkins
2002

$$\begin{bmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{bmatrix}$$

θ_{13}



CP

systematic revamping

Chen et al

Phys.Lett. B753 (2016) 644

Phys.Rev. D94 (2016) 033002

JHEP 1807 (2018) 077

Phys.Lett. B792 (2019) 461

Phys.Rev. D99 (2019) 075005

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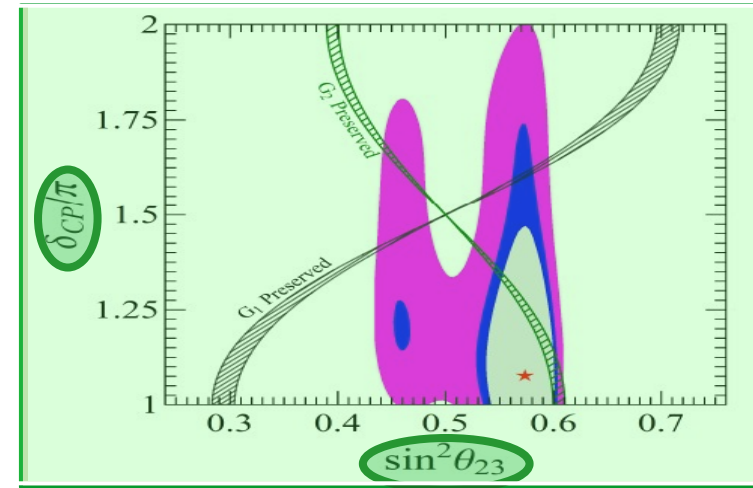
Chen et al
Phys.Lett. B753 (2016) 644
Phys.Rev. D94 (2016) 033002
JHEP 1807 (2018) 077
Phys.Lett. B792 (2019) 461
Phys.Rev. D99 (2019) 075005

Phys.Rev.D98(2018)055019

$$\sin^2 \theta_{12} \cos^2 \theta_{13} = \frac{1}{3} :$$

$$\tan 2\theta_{23} \cos \delta_{CP} = \frac{\cos 2\theta_{13}}{\sin \theta_{13} \sqrt{2 - 3 \sin^2 \theta_{13}}}$$

an example



Bi-Large lepton mixing pattern

$$\begin{bmatrix} 1 - \frac{1}{2}\lambda^2 & -\lambda e^{i\phi} & A\lambda^3 e^{i\phi} \\ \lambda e^{-i\phi} & 1 - \frac{1}{2}\lambda^2 & -A\lambda^2 \\ 0 & A\lambda^2 & 1 \end{bmatrix} \begin{bmatrix} 1 - \frac{5\lambda^2}{2} & 2\lambda & -\lambda \\ -2\lambda + 3\lambda^2 & 1 - \frac{13\lambda^2}{2} & 3\lambda \\ \lambda + 6\lambda^2 & -3\lambda + 2\lambda^2 & 1 - 5\lambda^2 \end{bmatrix}$$

$\sin \theta_{12}^{\text{CKM}} = \lambda$ and $\sin \theta_{23}^{\text{CKM}} = A\lambda^2$, where $\lambda = 0.22453 \pm 0.00044$, $A = 0.836 \pm 0.015$

Largest Q-mixing similar to smallest L-mixing
Cabibbo angle as universal seed for flavor
mixing

Phys.Rev. D86 (2012) 051301

Phys.Rev.D87 (2013) 053013

Phys.Lett. B748 (2015) 1-4

predicting solar & atm

Bi-Large lepton mixing pattern

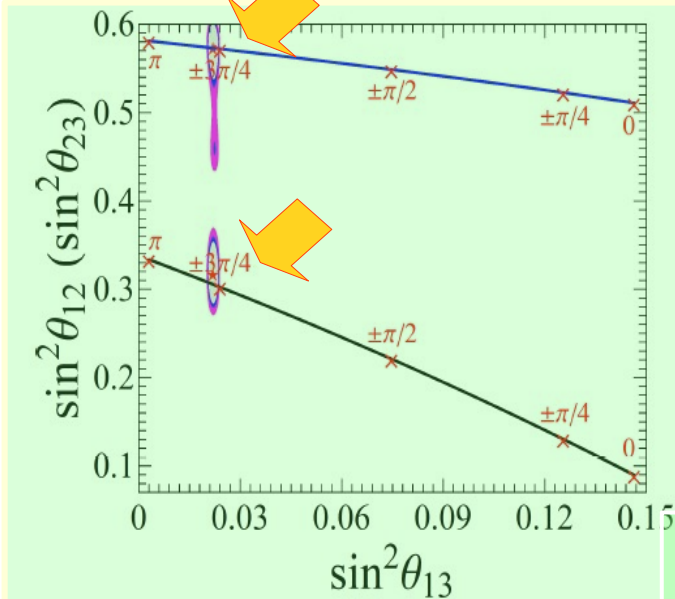
$$\begin{bmatrix} 1 - \frac{1}{2}\lambda^2 & -\lambda e^{i\phi} & A\lambda^3 e^{i\phi} \\ \lambda e^{-i\phi} & 1 - \frac{1}{2}\lambda^2 & -A\lambda^2 \\ 0 & A\lambda^2 & 1 \end{bmatrix} \begin{bmatrix} 1 - \frac{5\lambda^2}{2} & 2\lambda & -\lambda \\ -2\lambda + 3\lambda^2 & 1 - \frac{13\lambda^2}{2} & 3\lambda \\ \lambda + 6\lambda^2 & -3\lambda + 2\lambda^2 & 1 - 5\lambda^2 \end{bmatrix}$$

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predicting solar & atm



looser realization
Phys.Lett.B 796 (2019) 162

Many other patterns, e.g. trimaximal,
most can be probed at DUNE

e.g. Phys.Rev.D97(2018)095025

From Phys.Lett. B792 (2019) 461

neutrinoless double beta decay

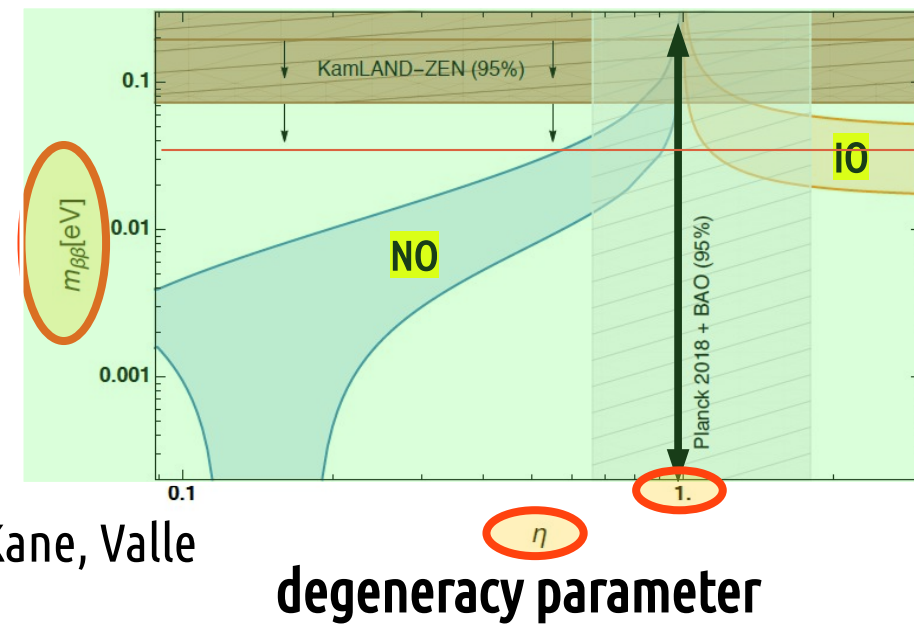
$$\left| \sum_j U_{ej}^2 m_j \right| = \left| c_{12}^2 c_{13}^2 m_1 + s_{12}^2 c_{13}^2 m_2 e^{2i\phi_{12}} + s_{13}^2 m_3 e^{2i\phi_{13}} \right|$$

Schechter & JV PRD22 (1980) 2227

Rodejohann, JV Phys.Rev. D84 (2011) 073011

➤ **Quasi-degenerate strongly-disfavored**

Lattanzi, Gerbino, Freese, Kane, Valle
JHEP 10 (2020) 213



neutrinoless double beta decay

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Schechter & JV PRD22 (1980) 2227
 Rodejohann, JV Phys.Rev. D84 (2011) 073011

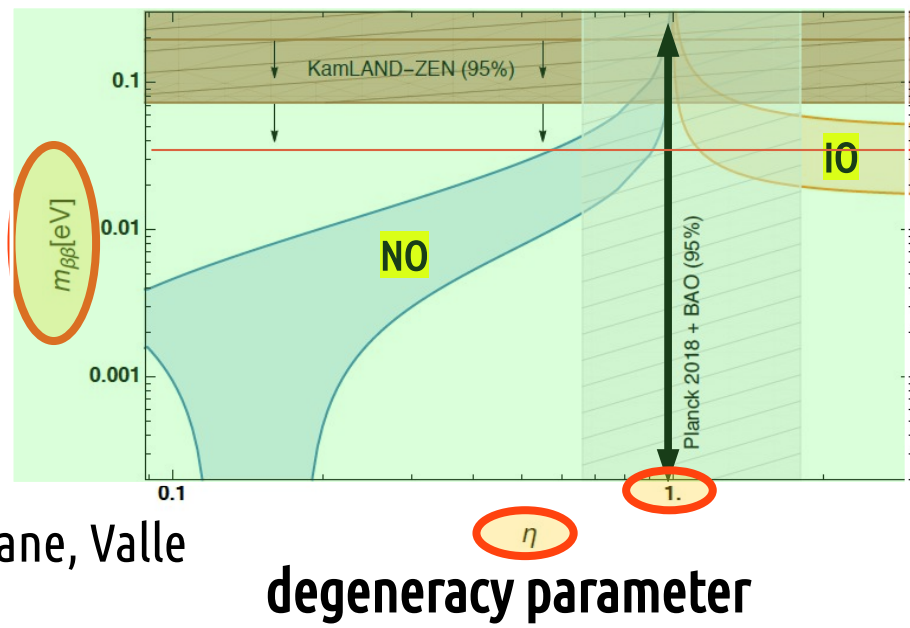
➤ **Quasi-degenerate strongly-disfavored**

Lattanzi, Gerbino, Freese, Kane, Valle
 JHEP 10 (2020) 213

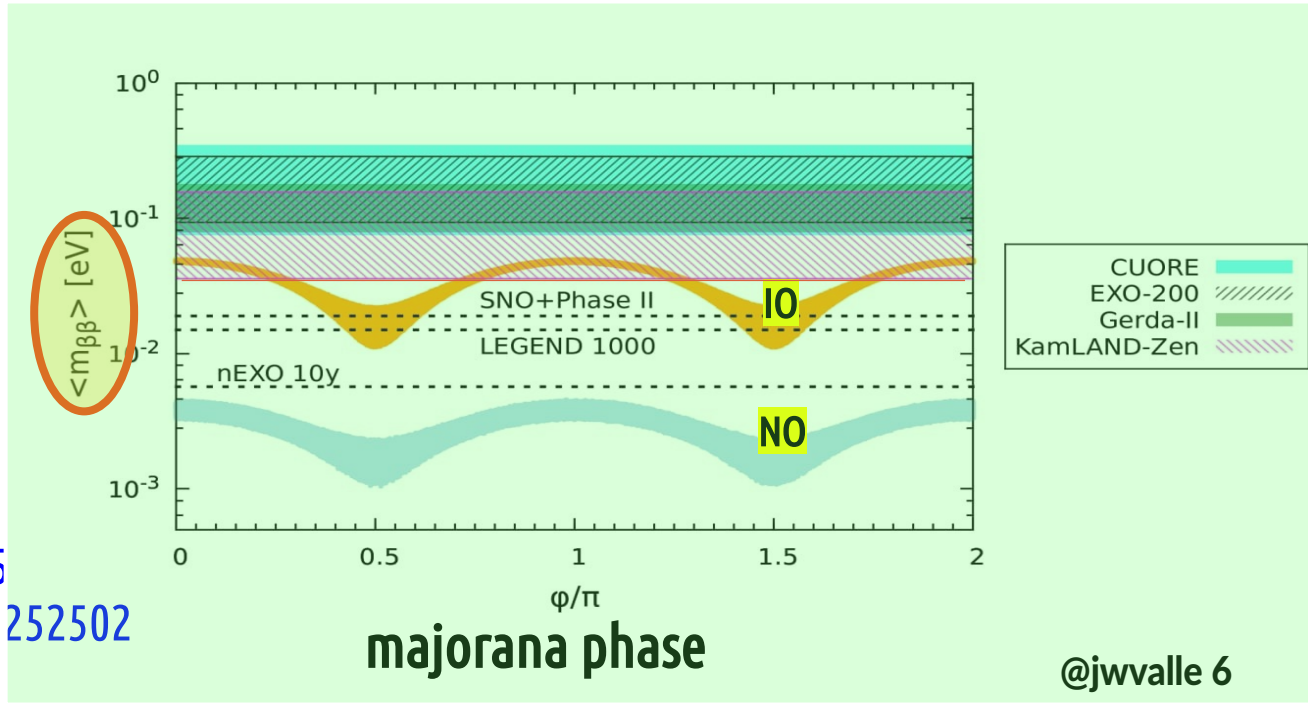
➤ **One-massless neutrino**

Reig et al Phys.Lett. B790 (2019)303
 Barreiros, Felipe & Joaquim JHEP01 (2019) 223
 Mandal et al PLB789 (2019) 132
 Avila et al Eur.Phys.J.C 80 (2020) 10, 908

Agostini et al. Science 365 (2019) 1445
 Final Gerda Phys.Rev.Lett. 125(2020) 252502



degeneracy parameter



majorana phase

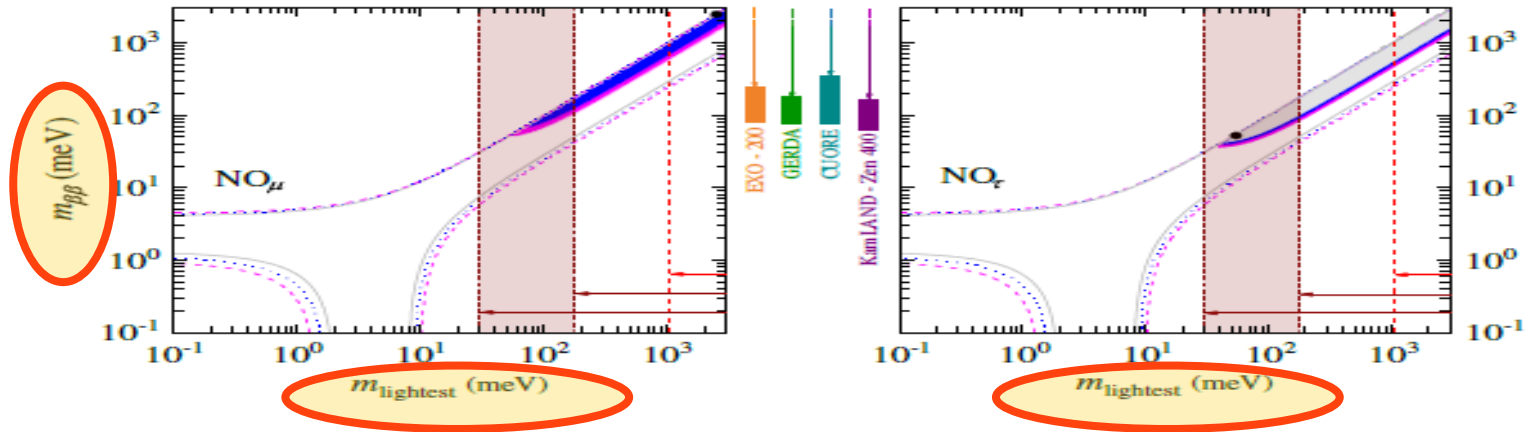
3-massive case

Lower bounds from family symmetries

Dorame et al PhysRevD86(2012)056001

Dorame et al Nucl.Phys.B 861 (2012) 259-270

King et al Phys.Lett. B 724 (2013) 68-72 etc



From Barreiros et al [JHEP04\(2021\)249](#)

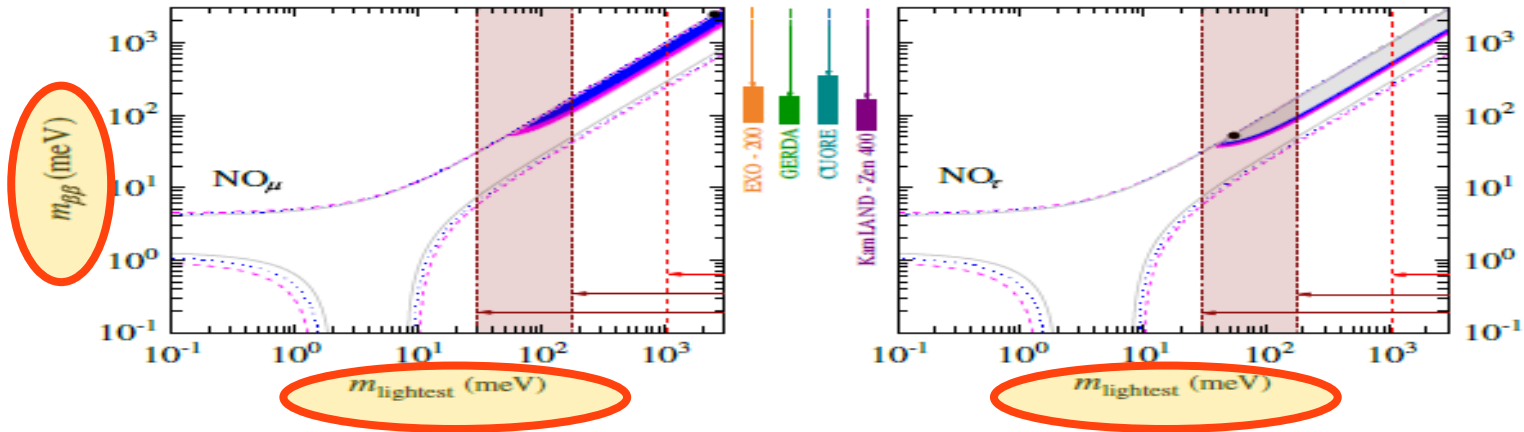
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Dorame et al PhysRevD86(2012)056001

Dorame et al Nucl.Phys.B 861 (2012) 259-270

King et al Phys.Lett. B 724 (2013) 68-72 etc

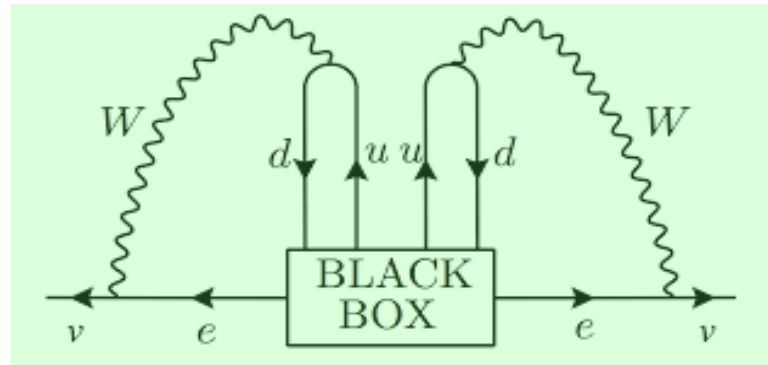


From Barreiros et al JHEP04(2021)249

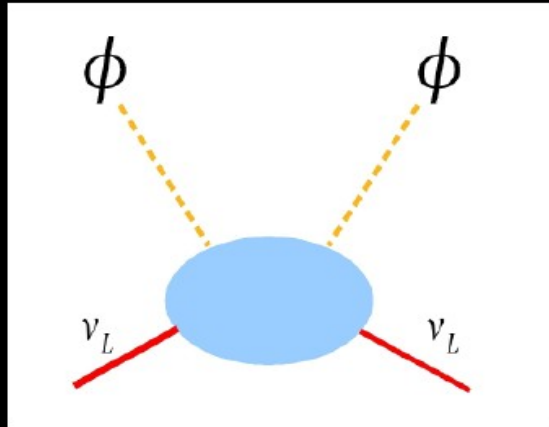
Schechter, Valle PhysRev D25 (1982) 2951

Duerr, Lindner, Merle JHEP06(2011)091

B.J.P. Jones 2108.09364 (TASI 2020)



Origin of neutrino mass

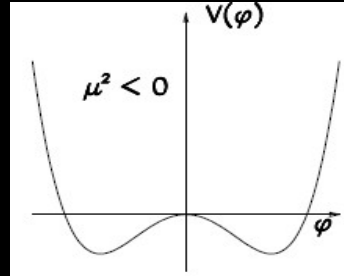


Origin of neutrino mass

stability

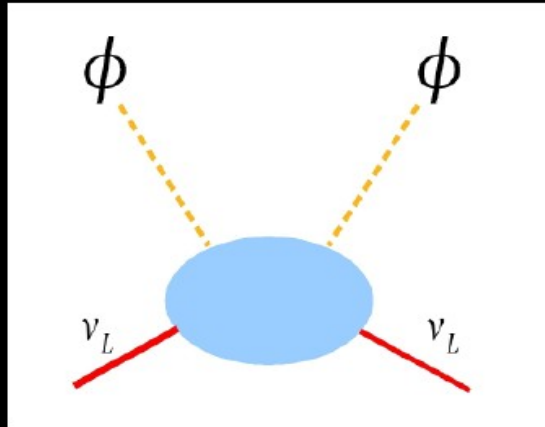
SEESAW
dynamics

$$v_3 v_1 \sim v_2^2$$



Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029

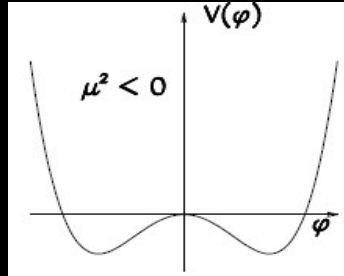


Origin of neutrino mass

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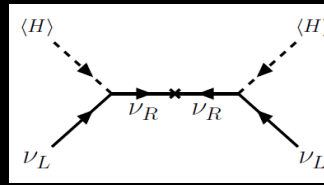
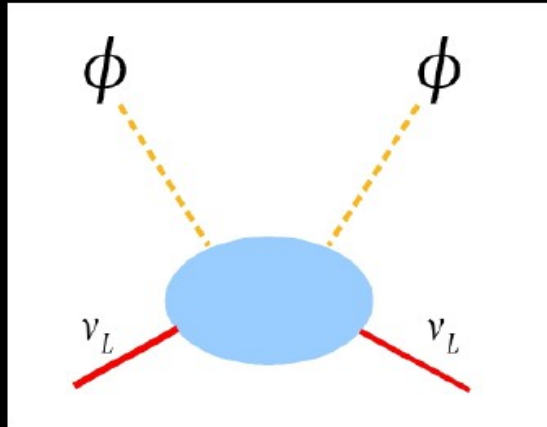
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stability



Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029



TYPE I

Minkowski 77

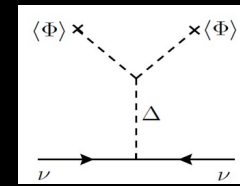
Gellman Ramond Slansky 80

Glashow, Yanagida 79

Mohapatra Senjanovic 80

Lazarides Shafi Weterrich 81

Schechter-Valle 80 & 82



TYPE II

Schechter-Valle 80 & 82

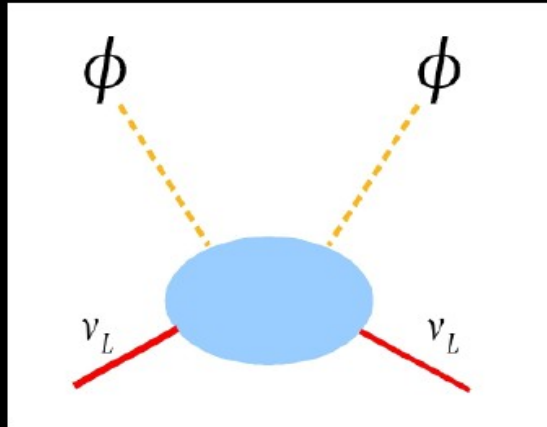
Origin of neutrino mass

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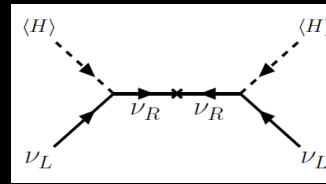
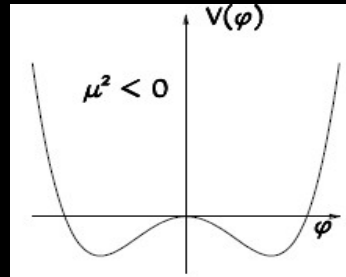
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Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029



stability



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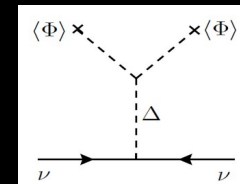
Schechter-Valle 80 & 82

L-R seesaw

of Rs = # Ls

SM seesaw

of singlets arbitrary



TYPE II

Schechter-Valle 80 & 82

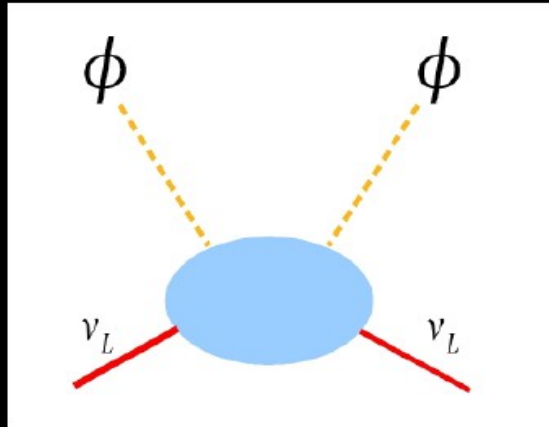
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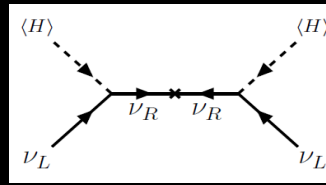
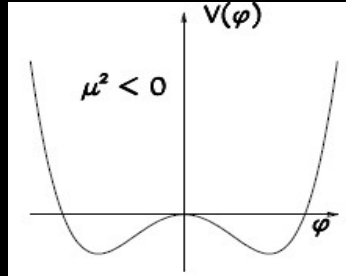
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Phys.Rev.D 101 (2020) 115030

JHEP03(2021)212 & JHEP07(2021) 029



stability



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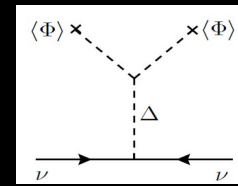
of singlets arbitrary

■ MISSING PARTNER

(3,2) min viable type1 seesaw

(3,1) scoto-seesaw template

$$m_{\beta\beta}$$



TYPE II

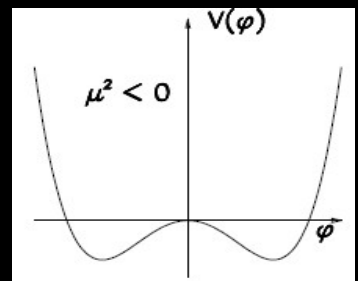
Schechter-Valle 80 & 82

Origin of neutrino mass

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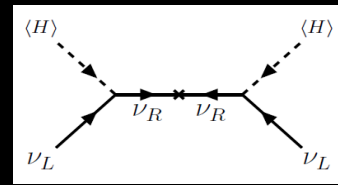
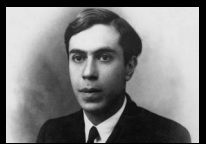
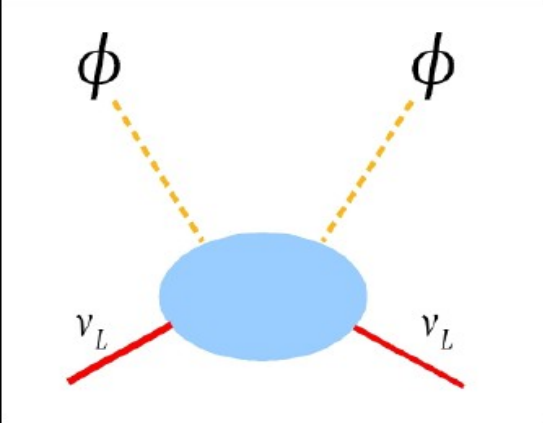
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stability



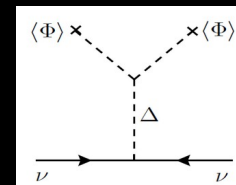
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JHEP03(2021)212 & JHEP07(2021) 029



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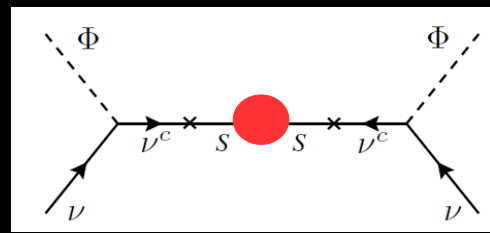
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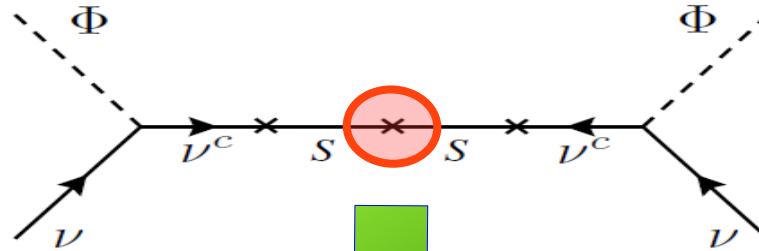
■ LOW-SCALE Type1 SEESAW (3,6) ISS & LSS



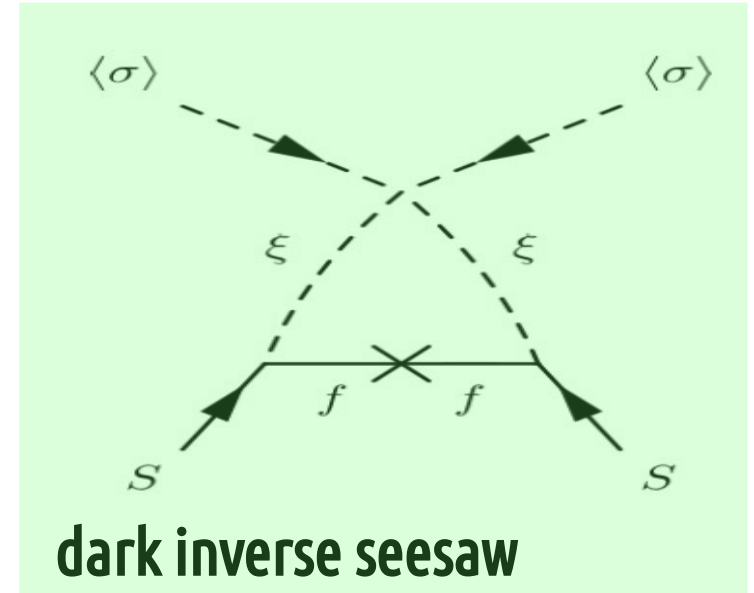
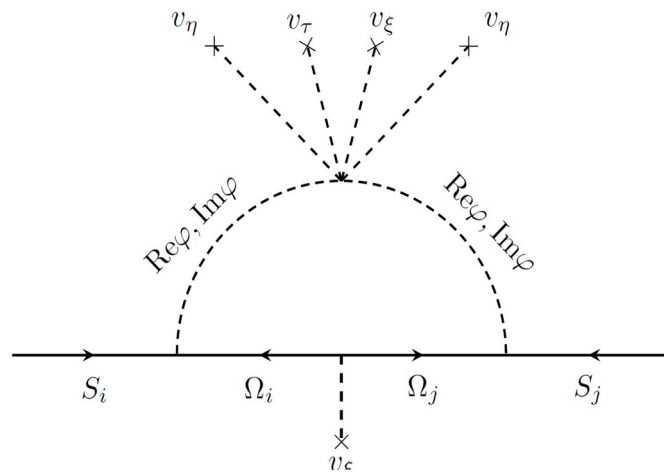
Mohapatra,Valle 86

- Akhmedov et al Phys.Rev.D53 (1996) 2752
- PhysLettB368 (1996) 270
- Malinsky et al PhysRevLett95(2005)161801

doubly protected inverse seesaw



radiative
inverse seesaw



L-R scheme

Cárcamo Hernández et al JHEP 1902 (2019) 065

dark inverse seesaw

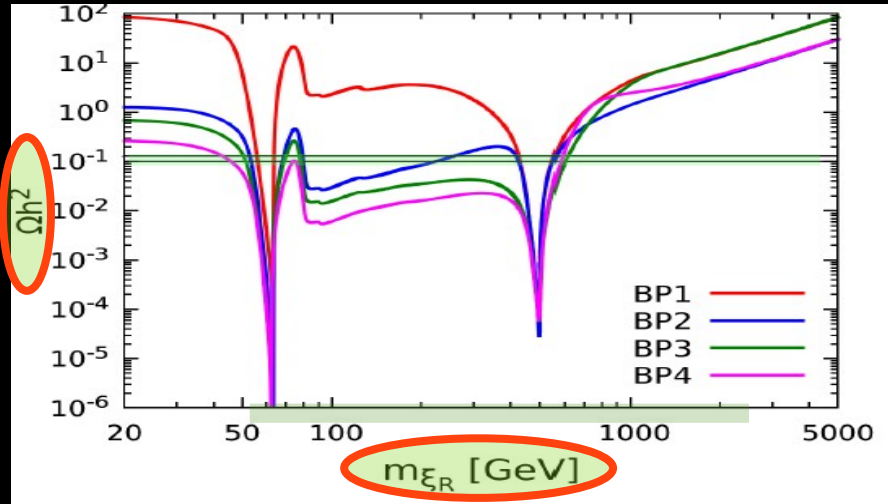
Mandal et al Phys.Lett.B821 (2021) 136609



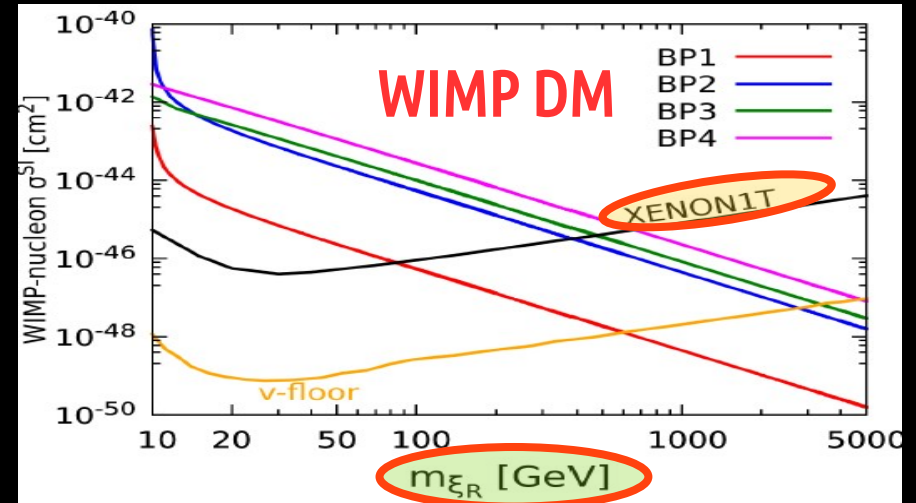
dark inverse typeI seesaw mechanism

LambdaCDM

Phys.Lett.B 821 (2021) 136609



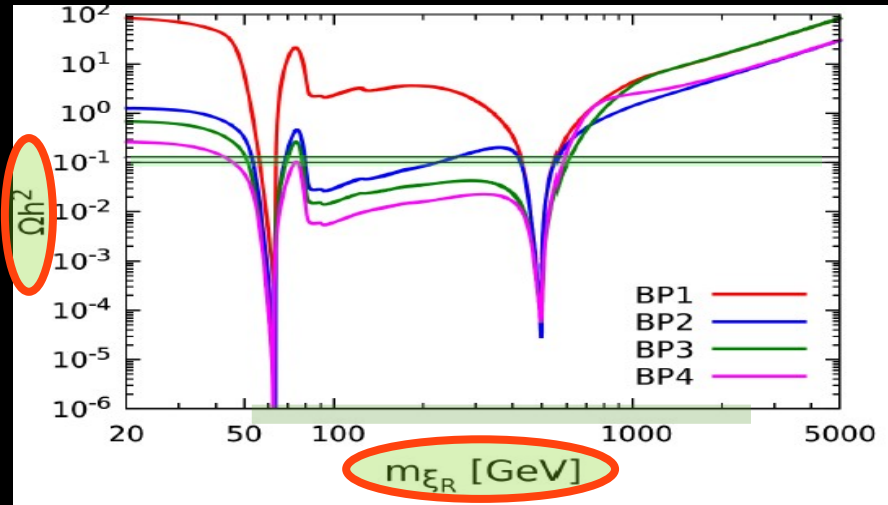
Xenon1T PhysRevLett.121.111302



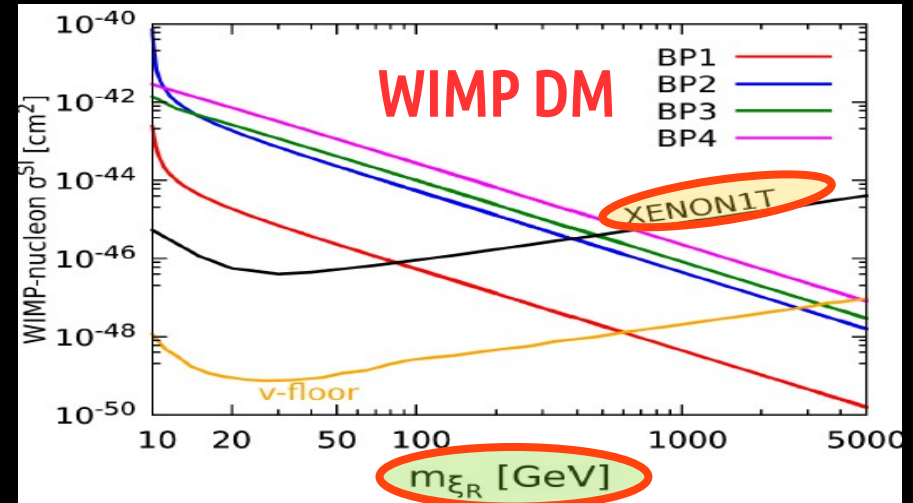
dark inverse type I seesaw mechanism

LambdaCDM

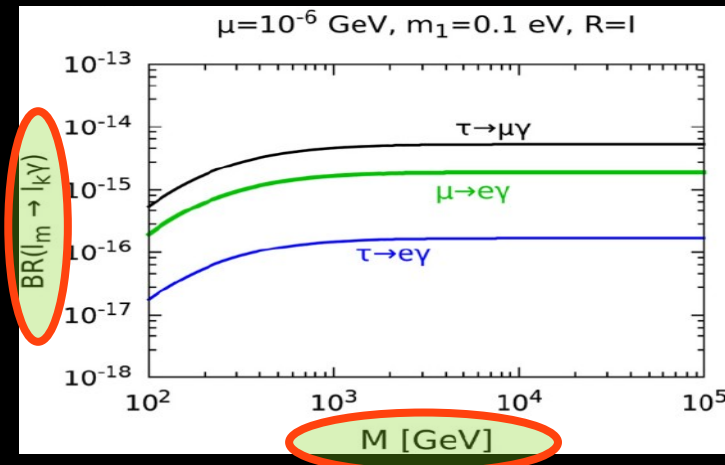
Phys.Lett.B 821 (2021) 136609



Xenon1T PhysRevLett.121.111302



e.g. large cLFV from inverse type I seesaw
Mandal et al
Phys.Lett.B 821 (2021) 136609
(larger values possible)



neutrino properties from colliders

current oscillation data
can reconstruct **type-II**
so that it can be tested
at high-energies

Miranda et al Phys.Rev.D105 (2022) 095020

seesaw mediator produced in
@ e^+e^- / pp collisions



Miranda et al PLB 829 (2022) 137110

neutrino properties from colliders

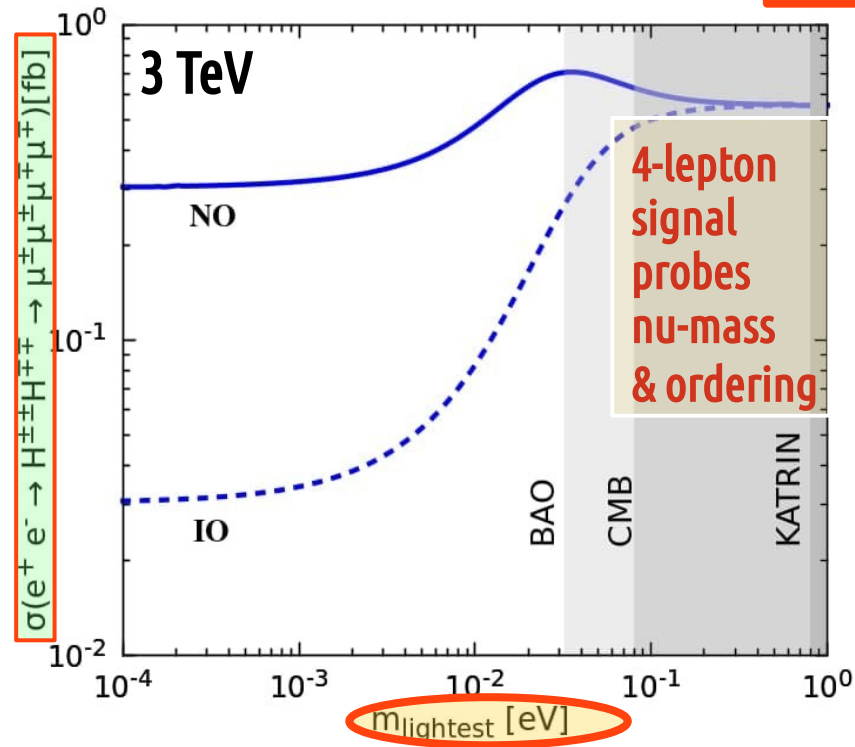
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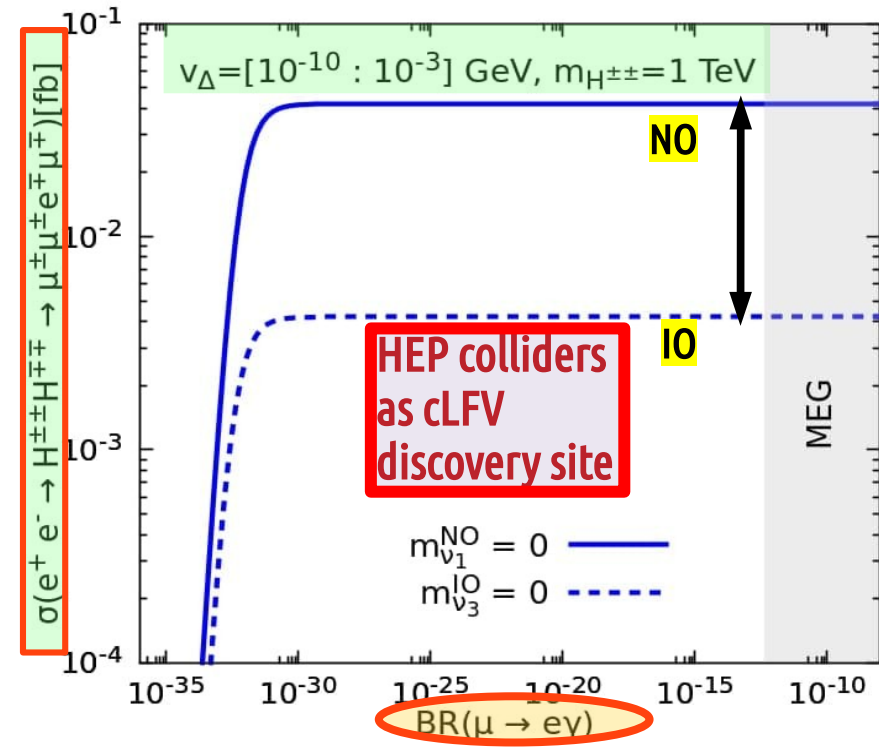
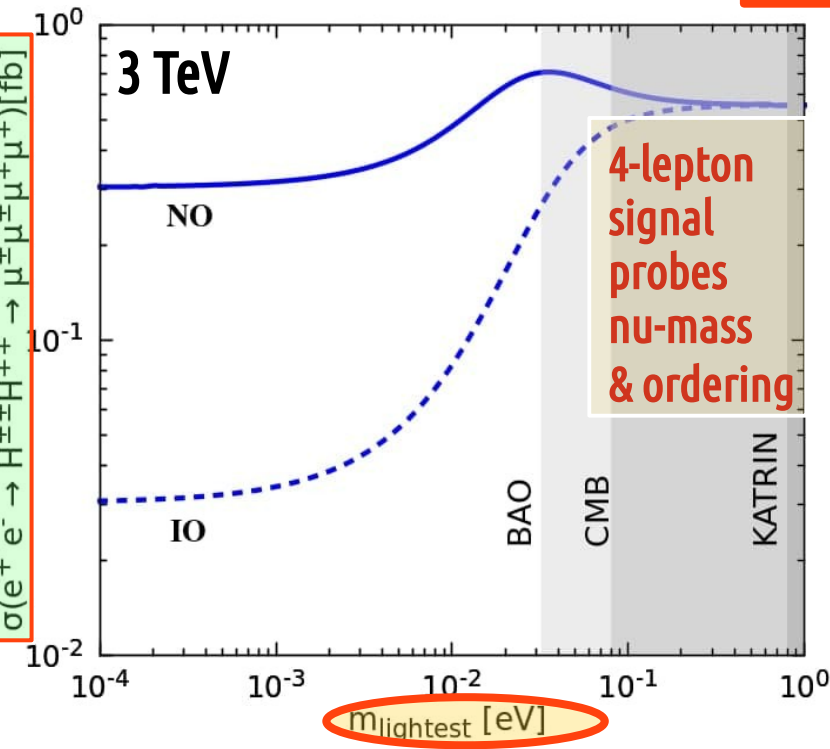
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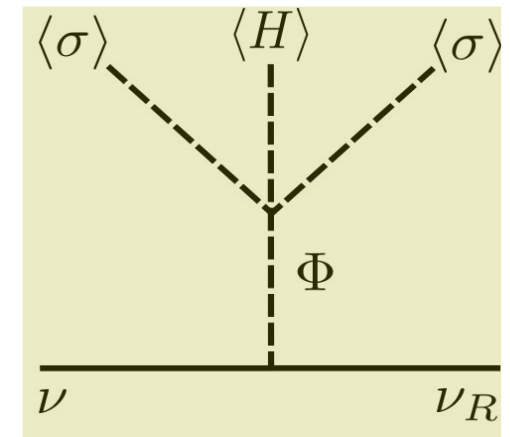
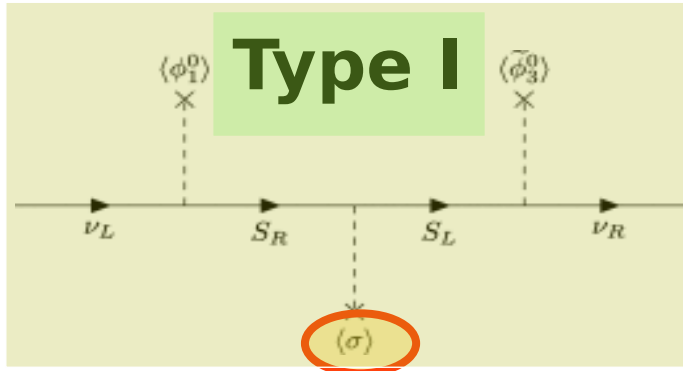
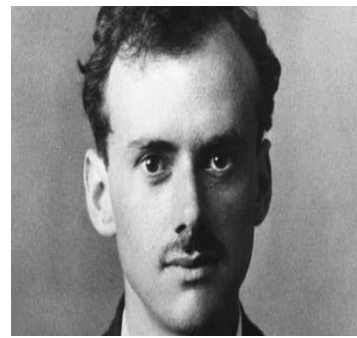
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Miranda et al PLB 829 (2022) 137110



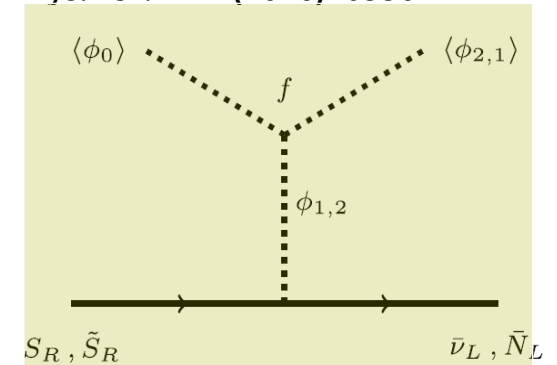
Seesawing a la



Type II

Phys.Lett. B762 (2016) 162-165

Phys.Rev. D94 (2016) 033012



Phys.Lett. B761 (2016) 431-436

Phys.Lett. B767 (2017) 209-213

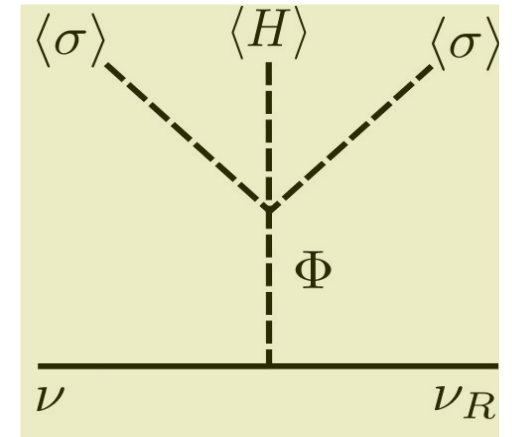
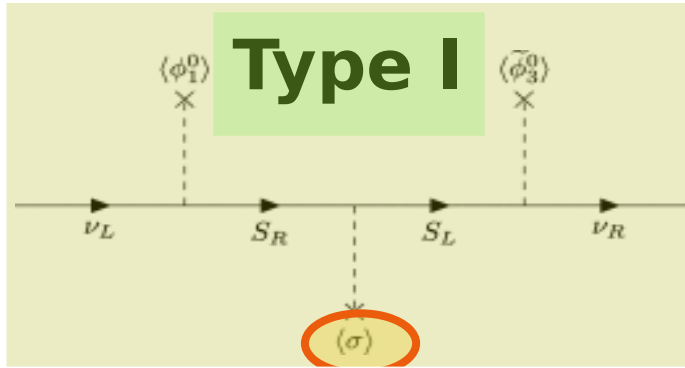
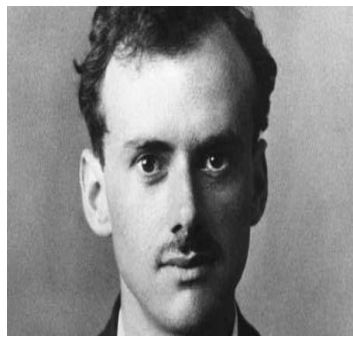
Phys.Rev. D98 (2018) 035009

Phys.Lett. B781 (2018) 122-128

Addazi et al Phys.Lett. B759 (2016) 471-478

Phys.Lett. B755 (2016) 363-366

Seesawing a la



Type II

**symmetry protecting small neutrino mass
+ Diracness**

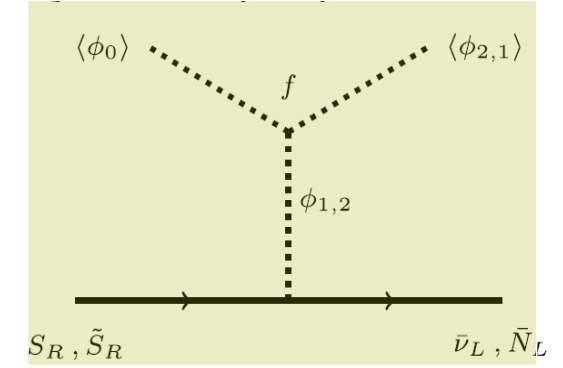
Peccei-Quinn symmetry [Phys.Lett.B 810 \(2020\) 135829](#)

$$m_\nu^D \simeq \frac{y^{\nu_1} (y^S)^{-1} (y^{\nu_2})^T}{\sqrt{2}} \frac{v \langle W \rangle}{v \langle \sigma \rangle}$$

← SU3L
← PQ

[Phys.Lett. B762 \(2016\) 162-165](#)

[Phys.Rev. D94 \(2016\) 033012](#)



[Phys.Lett. B761 \(2016\) 431-436](#)

[Phys.Lett. B767 \(2017\) 209-213](#)

[Phys.Rev. D98 \(2018\) 035009](#)

[Phys.Lett. B781 \(2018\) 122-128](#)

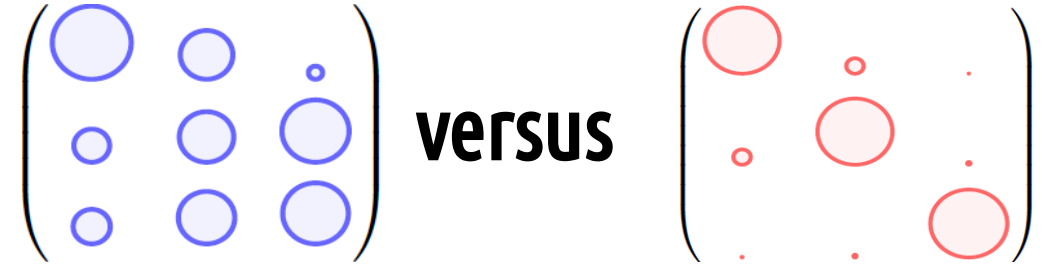
[Addazi et al Phys.Lett. B759 \(2016\) 471-478](#)

[Phys.Lett. B755 \(2016\) 363-366](#)



flavour legacy of oscillations

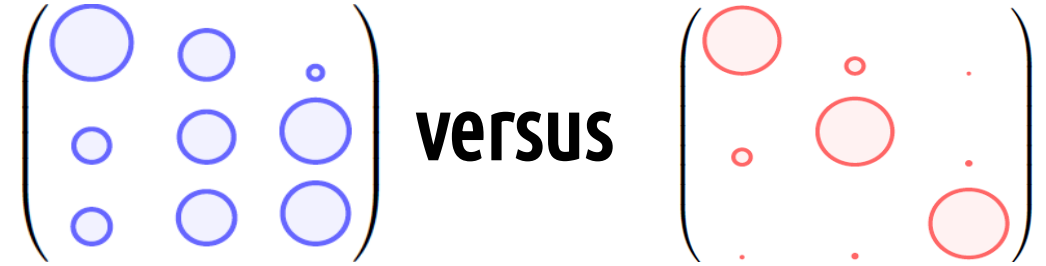
Q/L mixing pattern



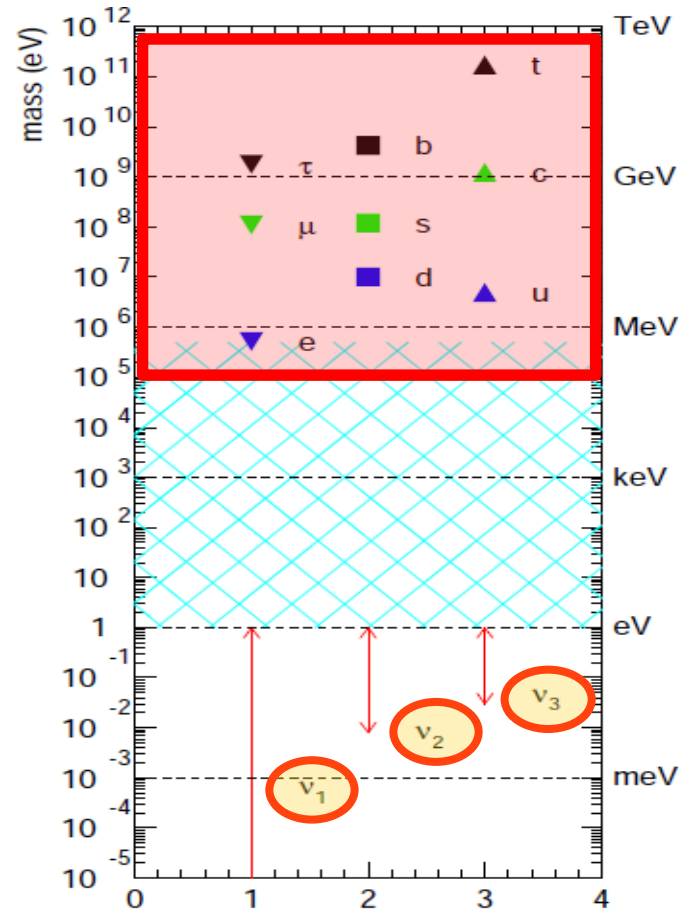


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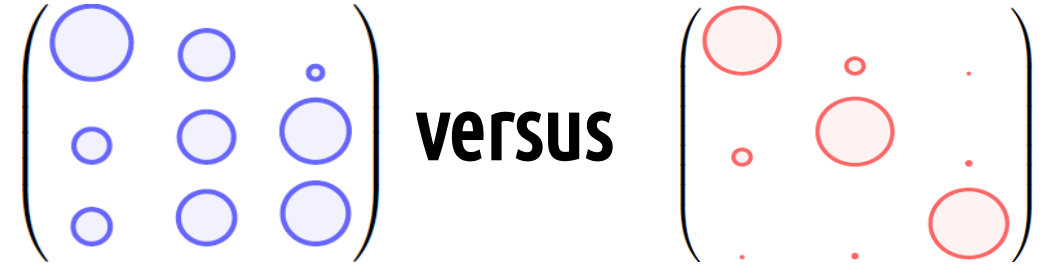
Q/L mass hierarchies





flavour legacy of oscillations

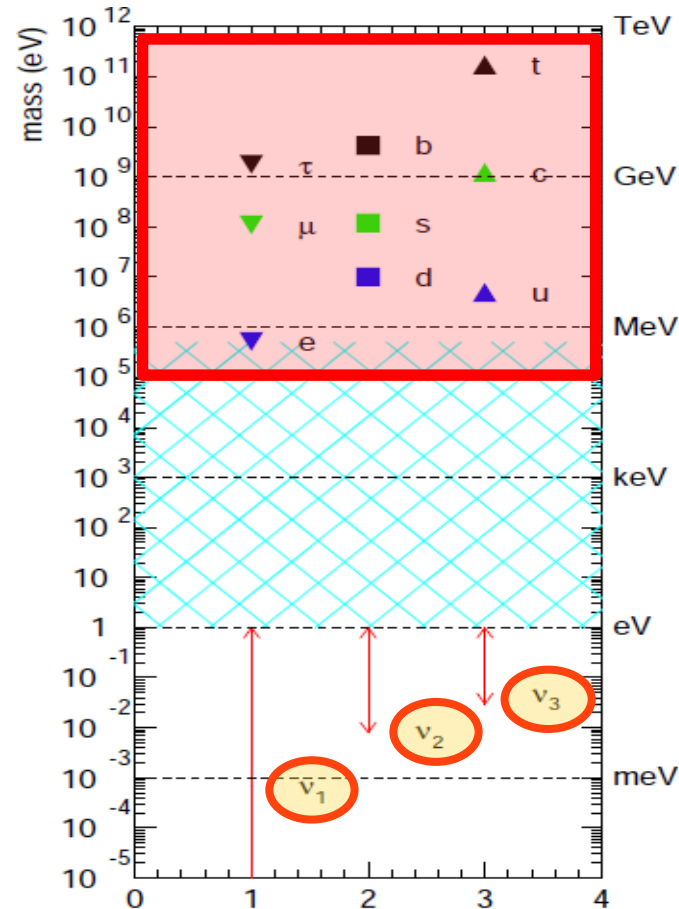
Q/L mixing pattern



Q/L mass hierarchies

Golden
Q-L mass
relation

$$\frac{m_\tau}{\sqrt{m_e m_\mu}} \approx \frac{m_b}{\sqrt{m_d m_s}}$$



Morisi et al	Phys.Rev. D84 (2011) 036003
King et al	Phys. Lett. B 724 (2013) 68
Morisi et al	Phys.Rev. D88 (2013) 036001
Bonilla et al	Phys.Lett. B742 (2015) 99
Reig, JV, Wilczek	Phys.Rev. D98 (2018) 095008

5D Warped flavour dynamics

Randall-Sundrum Phys.Rev.Lett. 83 (1999) 3370

- **mass hierarchies from geometry**

Arkani-Hamed & Schmaltz hep-ph/9903417

- **mixing angles from family symmetry**

5D Warped flavour dynamics

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■ mass hierarchies from geometry

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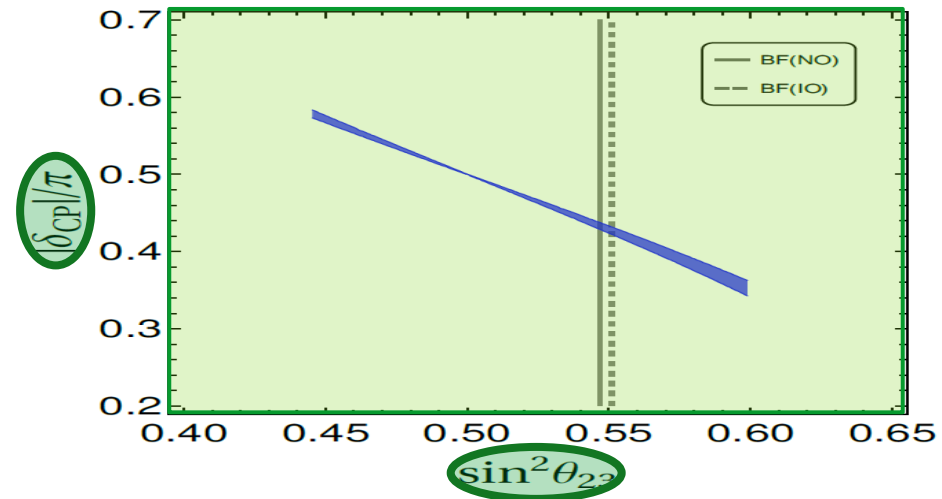
■ mixing angles from family symmetry

TM mixing pattern predicted from T'

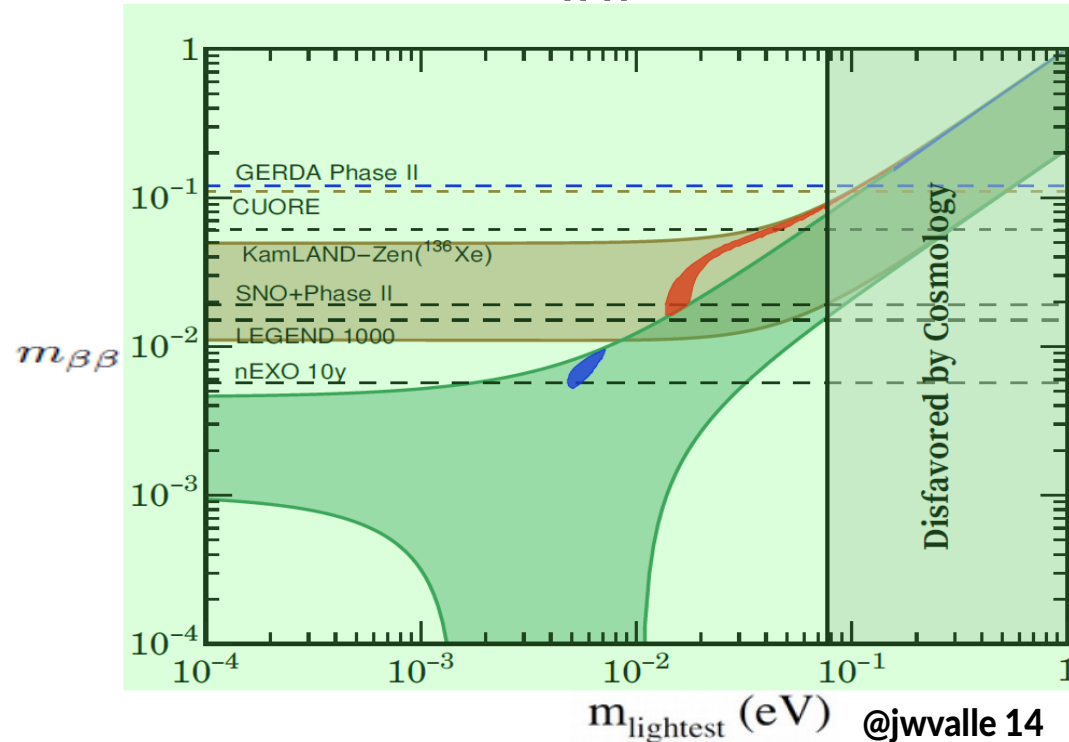
$$\cos^2 \theta_{12} \cos^2 \theta_{13} = \frac{2}{3} \quad \text{TM1 pattern}$$

$$\cos \delta_{CP} = \frac{(3 \cos 2\theta_{12} - 2) \cos 2\theta_{23}}{3 \sin 2\theta_{23} \sin 2\theta_{12} \sin \theta_{13}}$$

Chen et al Phys. Rev. D 102, 095014 (2020)



TM1



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Chen et al Phys. Rev. D 102, 095014 (2020)

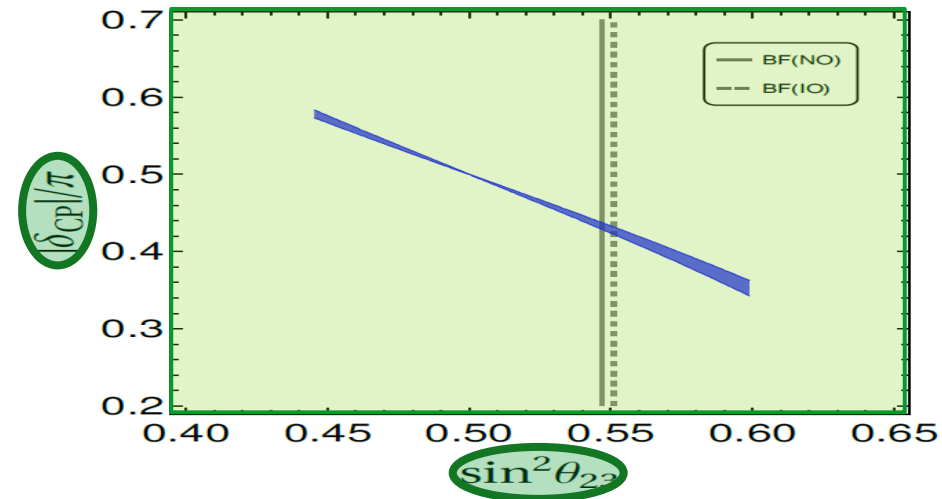
TM2

Another model with Dirac neutrinos

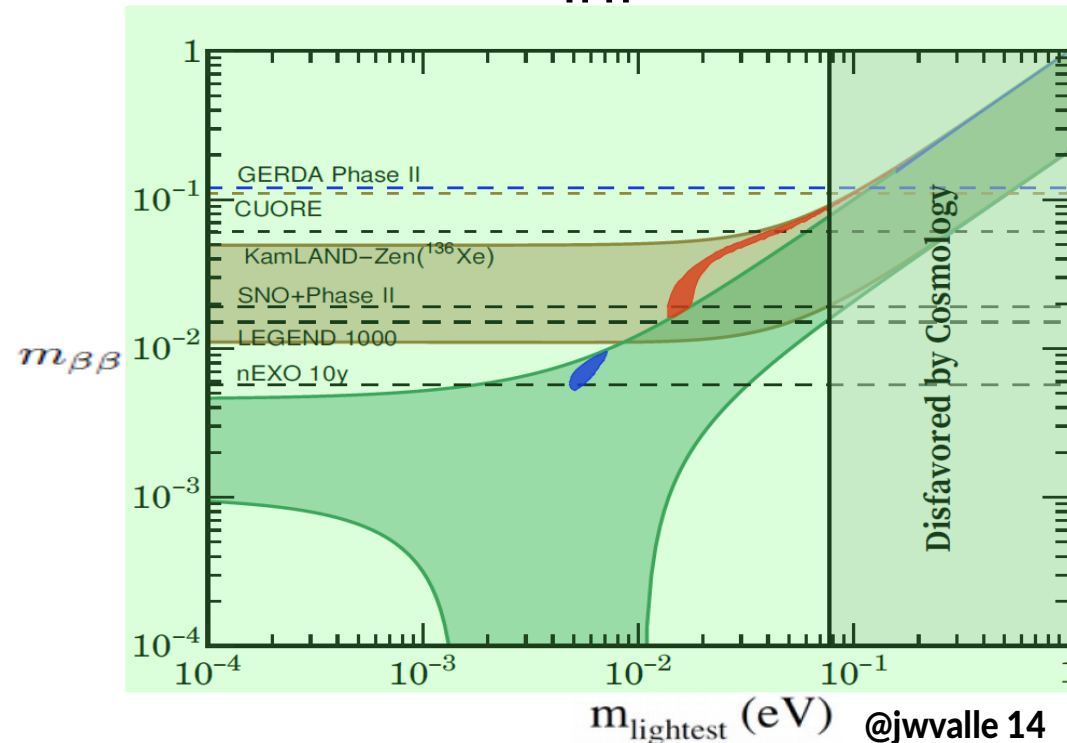
Chen et al JHEP01(2016)007

Phys. Rev. D95 (2017) 095030

Phys.Lett. B771 (2017) 524



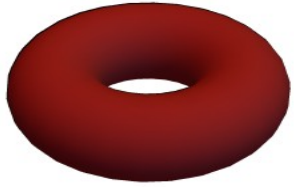
TM1



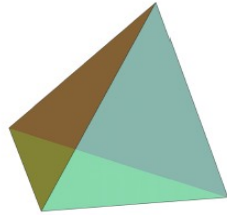
family symmetry from 6D orbifold

$$\mathcal{M} = \mathbb{M}^4 \times (\mathbb{T}^2 / \mathbb{Z}_2)$$

Phys.Lett.B 801 (2020) 135195
Phys.Rev.D 105 (2022) 055030



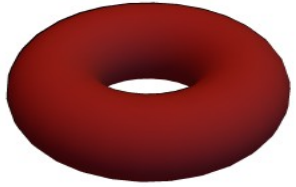
**A4 family symmetry
as emergent**



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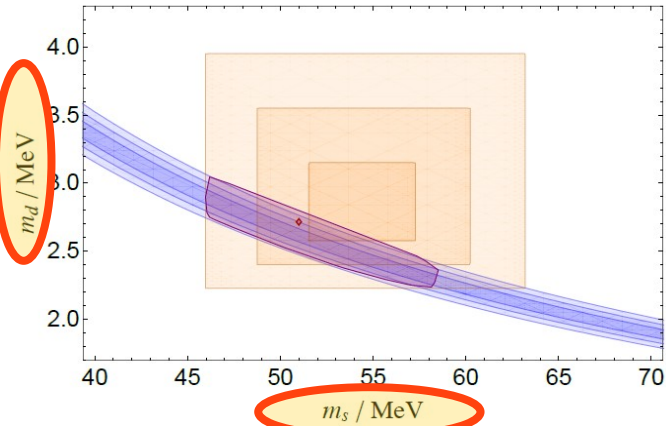
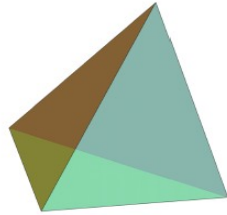
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Golden Q-L relation

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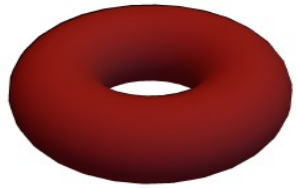


Good global fit of flavor observables

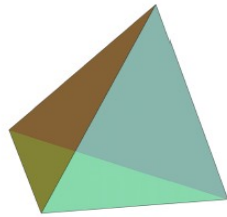
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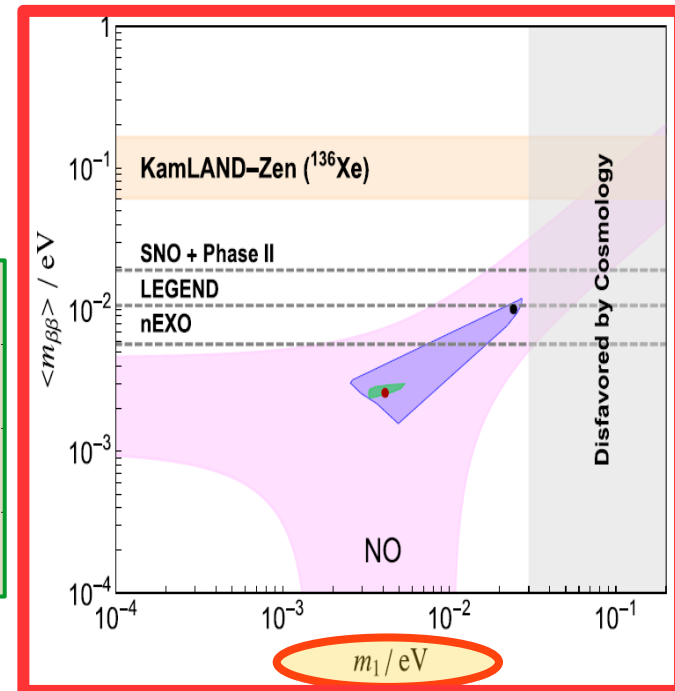
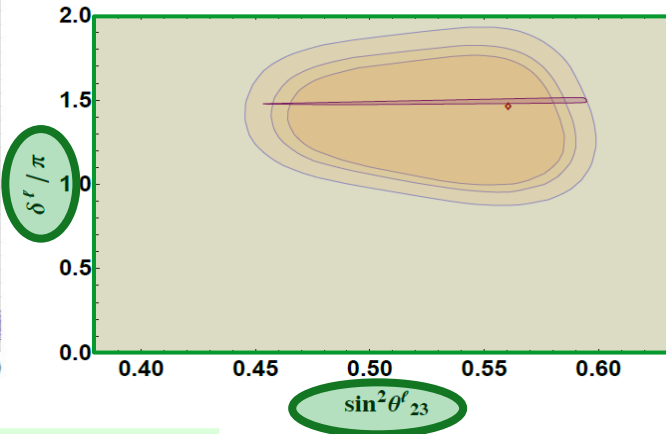
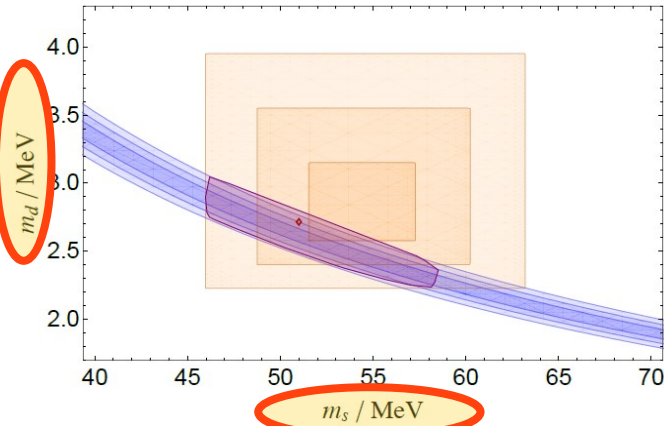


Sign of CP phase!

Phys.Rev.D 101 (2020) 11, 116012

Golden Q-L relation

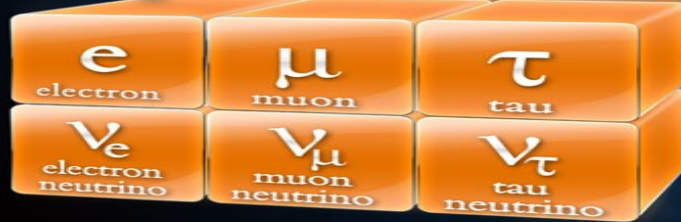
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Good global fit of flavor observables



Leptons



H
Higgs boson

Force Carriers



NEUTRINO MASS

DARK MATTER

FLAVOR PROBLEM

NU-COMPLETION of SM

Thank You



Leptons



H
Higgs boson

Force Carriers



NEUTRINO MASS

DARK MATTER

NU-COMPLETION of SM

FLAVOR PROBLEM



**EW vacuum stability from seesaw
neutrinos & DM
scoto-seesaw
dark inverse seesaw
Diracness from DM stability ...**

**Diracness from PQ symmetry
Warped flavordynamics
Family symmetry as emergent
cosmology from neutrinos
GWs from lepton # violation..**

@jwvalle 16

