

# Scotogenic $A_5 \rightarrow A_4$ Dirac Neutrinos

Ernest Ma

Physics and Astronomy Department  
University of California  
Riverside, CA 92521, USA

# Contents

- Two Useful Ideas
- Tetrahedron and Pentatope
- Model Particle Content
- Seesaw Charged Leptons
- Scotogenic Dirac Neutrino Mass
- Freeze-In Dark Matter
- Concluding Remarks

## Two Useful Ideas

(A) Neutrino masses exist because of dark matter, i.e. **scotogenically**. [Ma, PRD 73, 077301 (2006)]

For **Majorana** neutrinos, dark parity may be related to lepton parity:  $\pi_D = (-1)^{L+2j}$ , where  $j$  is the intrinsic spin of the particle. [Ma, PRL 115, 011801 (2015)]

For **Dirac** neutrinos, dark number may be related to lepton number:  $D = L - (2j)_{[mod\ 2]}$ .  
[Ma, PLB 809, 135736 (2020)]

(B) Non-Abelian discrete flavor symmetries such as  $A_4$  allow all charged-lepton masses to be different, and yet predict a pattern for neutrino masses and mixing.

[Ma/Rajasekaran, PRD 64, 113012 (2001)]

Up to 2012, the focus of model building was

**Tribimaximal**:  $\theta_{13} = 0$ ,  $\sin^2 \theta_{23} = 1/2$ ,  $\sin^2 \theta_{12} = 1/3$ .

[Harrison/Perkins/Scott, PLB 530, 167 (2002)]

At present, the data are closer to

**Cobimaximal**:  $\theta_{13} \neq 0$ ,  $\sin^2 \theta_{23} = 1/2$ ,  $\delta_{CP} = \pm\pi/2$ .

[Babu/Ma/Valle, PLB 552, 207 (2003);

Grimus/Lavoura, PLB 579, 113 (2004)]

# Tetrahedron and Pentatope

In 3 dimensions, the points  $(1, 0, 0)$ ,  $(0, 1, 0)$ ,  $(0, 0, 1)$ ,  $(1, 1, 1)$  form a **tetrahedron** with the symmetry  $A_4$ .

It has the irreducible representations  $\underline{1}$ ,  $\underline{1}'$ ,  $\underline{1}''$ ,  $\underline{3}$ .

In the decomposition  $\underline{3} \times \underline{3} = \underline{1} + \underline{1}' + \underline{1}'' + \underline{3} + \underline{3}$ ,

$$U_\omega = \frac{1}{\sqrt{3}} \begin{pmatrix} 1 & 1 & 1 \\ 1 & \omega & \omega^2 \\ 1 & \omega^2 & \omega \end{pmatrix}$$

is obtained with  $\omega = \exp(2\pi i/3) = -1/2 + i\sqrt{3}/2$ .

[Cabibbo, PLB 72, 333 (1978); Wolfenstein, PRD 18, 958 (1978)]

In most applications, the charged-lepton mass matrix is diagonalized on the left by  $U_\omega$ , then the PMNS matrix is obtained from multiplying  $U_\omega^\dagger$  to the unitary matrix  $U_\nu$  diagonalizing the neutrino mass matrix also on the left. If

$$U_\nu = \begin{pmatrix} 0 & 1 & 0 \\ 1/\sqrt{2} & 0 & -i/\sqrt{2} \\ 1/\sqrt{2} & 0 & i/\sqrt{2} \end{pmatrix},$$

then **tribimaximal** mixing occurs. [Ma, PRD 70, 031901R (2004)] If  $U_\nu$  is an orthogonal matrix, then **cobimaximal** mixing occurs. [Fukuura/Miura/Takasugi/Yoshimura, PRD 61, 073002 (2000)]

In 4 dimensions, the points  $(2, 0, 0, 0)$ ,  $(0, 2, 0, 0)$ ,  $(0, 0, 2, 0)$ ,  $(0, 0, 0, 2)$ ,  $(\varphi, \varphi, \varphi, \varphi)$ , where  $\varphi = (1 + \sqrt{5})/2 = 1.618$  is the golden ratio, form a **pentatope** with the symmetry  $A_5$ .

[Everett/Stuart, PRD 79, 085005 (2009)].

It has the irreducible representations  $\underline{1}$ ,  $\underline{3}$ ,  $\underline{3}'$ ,  $\underline{4}$ ,  $\underline{5}$ .

Their multiplication rules are  $3 \times 3 = 1 + 3 + 5$ ,

$3' \times 3' = 1 + 3' + 5$ ,  $3 \times 3' = 4 + 5$ ,  $3 \times 4 = 3' + 4 + 5$ ,

$3' \times 4 = 3 + 4 + 5$ ,  $4 \times 4 = 1 + 3 + 3' + 4 + 5$ , etc.

The decompositions to  $A_4$  representations are  $1 \sim 1$ ,

$3 \sim 3$ ,  $3' \sim 3$ ,  $4 \sim 3 + 1$ ,  $5 \sim 3 + 1' + 1''$ .

# Model Particle Content

Consider the following particle content:

fermion/scalar	$SU(2)_L \times U(1)_Y$	$A_5$	$A_4$	$L$	$D = L - 2j$
$L_L = (\nu, l)_L$	$(2, -1/2)$	3	3	1	0
$l_R$	$(1, -1)$	3	3	1	0
$\nu_R$	$(1, 0)$	$3'$	3	1	0
$E_{L,R}$	$(1, -1)$	3	3	1	0
$\Phi = (\phi^+, \phi^0)$	$(2, 1/2)$	1	1	0	0
$N_{L,R}$	$(1, 0)$	$4$	3,1	0	$-1$
$\eta = (\eta^0, \eta^-)$	$(2, -1/2)$	$3'$	3	1	1
$\zeta^0$	$(1, 0)$	3	3	1	1
$\zeta^-$	$(1, -1)$	$3'$	3	1	1



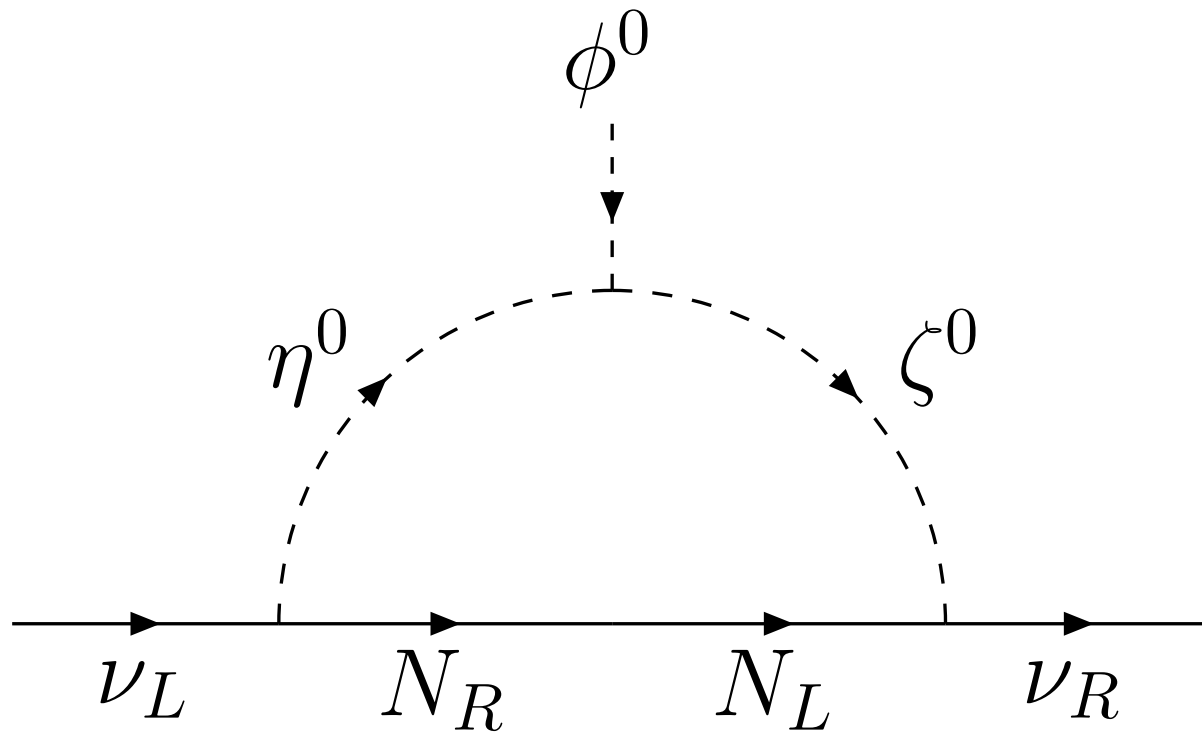
There is only one Higgs doublet  $\Phi$  as in the Standard Model (SM). Only renormalizable terms are allowed in this model, i.e. it is ultraviolet complete. All dimension-4 terms must obey  $A_5$ , whereas  $A_4$  and its breaking occur through soft terms.

**Allowed:**  $\bar{l}_R \Phi^\dagger L_L, \bar{E}_R \Phi^\dagger L_L (3 \times 1 \times 3);$

$\bar{L}_L \eta N_R (3 \times 3' \times 4); \bar{\nu}_R \zeta^0 N_L (3' \times 3 \times 4).$

**Forbidden:**  $\bar{\nu}_R \tilde{\Phi}^\dagger L_L (3' \times 1 \times 3); \bar{N}_R \tilde{\Phi}^\dagger L_L (4 \times 1 \times 3).$

The soft trilinear term  $\Phi \eta \zeta^0$  ( $1 \times 3' \times 3$ ) breaks  $A_5$ , but not  $A_4$ . Hence  $\nu_L$  cannot couple to  $\nu_R$  at tree level because of  $A_5$ , but is linked to it in one loop under  $A_4$ .



# Seesaw Charged Leptons

Since both  $l_R$  and  $E_R$  transform as  $\underline{3}$  under  $A_5$ , the latter may be defined to be the one which couples to  $l_L$  through  $\phi^0$  and the former does not. Once chosen this way, both  $l_R$  and  $E_R$  will couple to  $E_L$ . Hence the  $6 \times 6$  mass matrix linking  $(l, E)_L$  to  $(l, E)_R$  is

$$\mathcal{M} = \begin{pmatrix} 0 & M_{lE} \\ M_{El} & M_{EE} \end{pmatrix}.$$

Under  $A_5$ , all the  $3 \times 3$  entries  $M_{lE}$ ,  $M_{El}$ ,  $M_{EE}$  are proportional to the identity, with  $M_{lE}$  coming from  $\langle \phi^0 \rangle$ .

It is now assumed that the soft term  $M_{El}$  breaks  $A_5$  in a way compatible with  $A_4$ , using  $E \sim 3$  and  $l_R \sim 3$  and a gauge singlet flavon  $\sim 3$ . [Ma, MPLA 21, 2931 (2006)]

$$M_{El} = \begin{pmatrix} h_1 v' & h_2 v'' & h_3 v'' \\ h_3 v'' & h_1 v' & h_2 v'' \\ h_2 v'' & h_3 v'' & h_1 v' \end{pmatrix}$$

This magically decomposes to  $U_\omega M_{diag} U_\omega^\dagger$ , with eigenvalues  $h_1 v' + (h_2 + h_3)v''$ ,  $h_1 v' + (h_2 \omega + h_3 \omega^2)v''$ ,  $h_1 v' + (h_2 \omega^2 + h_3 \omega)v''$ . The seesaw  $3 \times 3$  charged-lepton mass matrix is  $M_{ll} = M_{lE} M_{EE}^{-1} M_{El}$ , as desired.

## Scotogenic Dirac Neutrino Mass

Let  $\psi_1^0 = \eta^0 \cos \theta - \zeta^0 \sin \theta$ ,  $\psi_2^0 = \eta^0 \sin \theta + \zeta^0 \cos \theta$ , with masses  $m_{1,2}$ , then the radiative Dirac neutrino mass matrix  $(\mathcal{M}_\nu)_{ij}$  is [using  $F(x, y) = x \ln(x/y)/(x - y)$ ]

$$\frac{\sin \theta \cos \theta}{16\pi^2} \sum_{k, k', a} f_{ika}^L f_{k'k4}^\phi f_{k'ja}^R [F(m_2^2, M_a^2) - F(m_1^2, M_a^2)].$$

The  $f$  couplings come from  $3 \times 3' \rightarrow 4$  of  $A_5$ . Let  $(a_1, a_2, a_3) \sim 3$ ,  $(b_1, b_2, b_3) \sim 3'$ , then the four components of  $N$  are  $3^{-1/2}[\varphi^{-1}a_3b_2 - \varphi a_1b_3, \varphi a_3b_1 + \varphi^{-1}a_2b_3, -\varphi^{-1}a_1b_1 + \varphi a_2b_2, a_2b_1 - a_1b_2 + a_3b_3]$ .

If  $M_{1,2,3,4}$  are all equal, then all three Dirac neutrino masses are equal as well. Assume instead that  $N_4$  is very heavy (breaking  $A_5 \rightarrow A_4$ ), and all entries of the  $3 \times 3$  mass matrix spanning  $N_{1,2,3}$  are much smaller than  $m_{1,2}$ . In that case,  $\mathcal{M}_\nu$  reduces to [Ma, PLB 717, 235 (2012)]

$$\frac{\sin \theta \cos \theta \ln(m_2^2/m_1^2)}{16\pi^2} \sum_{k,k',a,b} f_{ika}^L f_{k'k4}^\phi f_{k'jb}^R M_{ab}.$$

$$\propto \tilde{\mathcal{M}} = \begin{pmatrix} \varphi^2 M_{12} & M_{11} + M_{33} & \varphi^{-2} M_{32} \\ M_{22} + M_{33} & \varphi^{-2} M_{21} & \varphi^2 M_{31} \\ \varphi^{-2} M_{13} & \varphi^2 M_{23} & M_{11} + M_{22} \end{pmatrix}.$$

To obtain **cobimaximal** mixing,  $\tilde{\mathcal{M}}$  is assumed to be diagonalized by an orthogonal matrix  $\mathcal{O}$  on the left. This means that  $\tilde{\mathcal{M}}\tilde{\mathcal{M}}^\dagger$  should be real. Using

$$U_{l\nu} = \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13} \\ (s_{12} - ic_{12}s_{13})/\sqrt{2} & (-c_{12} - is_{12}s_{13})/\sqrt{2} & ic_{13}/\sqrt{2} \\ (s_{12} + ic_{12}s_{13})/\sqrt{2} & (-c_{12} + is_{12}s_{13})/\sqrt{2} & -ic_{13}/\sqrt{2} \end{pmatrix},$$

and the central values  $s_{13} = 0.148$ ,  $c_{13} = 0.989$ ,  
 $s_{12} = 0.554$ ,  $c_{12} = 0.832$ ,

$$\mathcal{O} = U_\omega U_{l\nu} = \begin{pmatrix} 0.927 & -0.363 & 0.085 \\ 0.336 & 0.714 & -0.614 \\ 0.162 & 0.598 & 0.785 \end{pmatrix}.$$

As a numerical example, assume first a reduction of parameters:  $\varphi^2(M_{12}, M_{31}, M_{23}) = \varphi^{-2}(M_{21}, M_{13}, M_{32})$ , then  $\mathcal{M}_\nu$  is of the form

$$\mathcal{M}_\nu = \begin{pmatrix} y_1 & x_1 & y_3 \\ x_2 & y_1 & y_2 \\ y_2 & y_3 & x_3 \end{pmatrix}, \text{ with } \mathcal{M}_\nu \mathcal{M}_\nu^\dagger = \mathcal{O} \begin{pmatrix} m_1^2 & 0 & 0 \\ 0 & m_2^2 & 0 \\ 0 & 0 & m_3^2 \end{pmatrix} \mathcal{O}^T,$$

resulting in

$$\begin{pmatrix} x_1^2 + y_1^2 + y_3^2 \\ x_2^2 + y_1^2 + y_2^2 \\ x_3^2 + y_2^2 + y_3^2 \end{pmatrix} = \begin{pmatrix} 0.859 & 0.113 & 0.026 \\ 0.132 & 0.510 & 0.358 \\ 0.007 & 0.377 & 0.616 \end{pmatrix} \begin{pmatrix} m_1^2 \\ m_2^2 \\ m_3^2 \end{pmatrix},$$



and

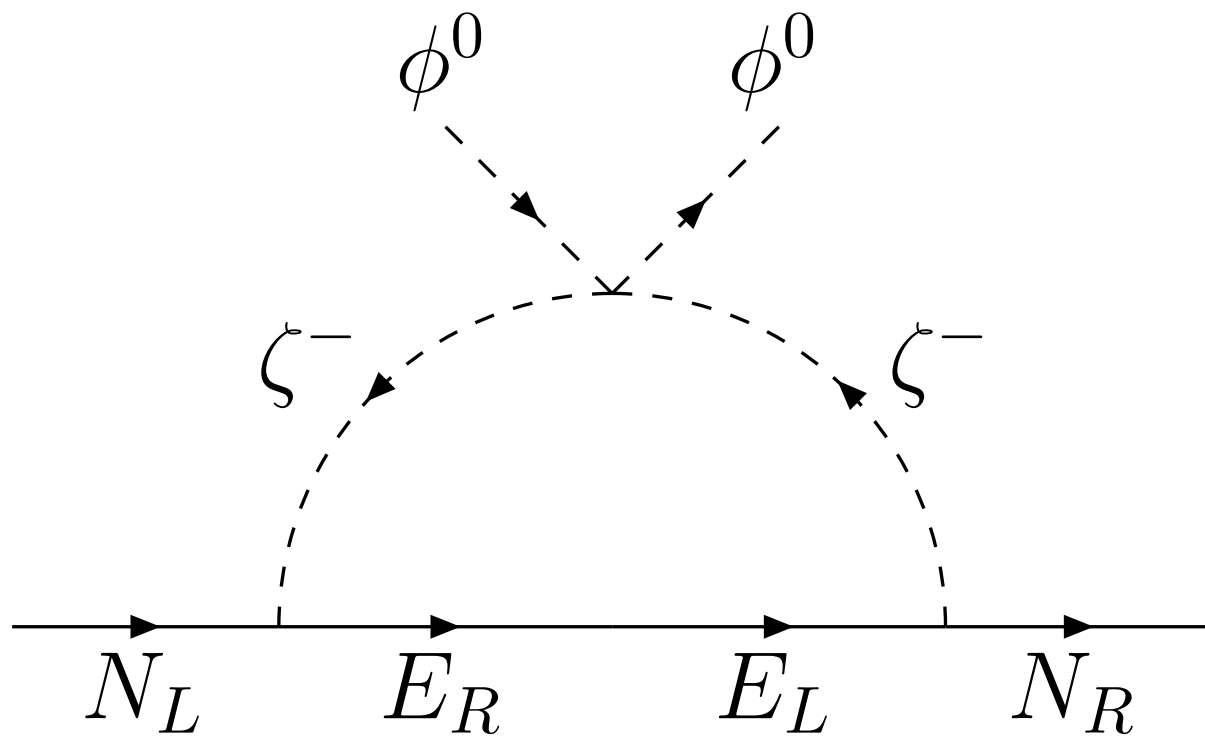
$$\begin{pmatrix} y_1(x_1 + x_2) + y_2y_3 \\ y_3(x_1 + x_3) + y_1y_2 \\ y_2(x_2 + x_3) + y_1y_3 \end{pmatrix} = \begin{pmatrix} -0.337 & 0.240 & 0.097 \\ 0.079 & -0.206 & 0.127 \\ -0.031 & -0.438 & 0.469 \end{pmatrix} \begin{pmatrix} m_1^2 \\ m_2^2 \\ m_3^2 \end{pmatrix}.$$

Assuming normal ordering of neutrino masses with  $\Delta m_{21}^2 = 7.53 \times 10^{-5} \text{ eV}^2$  and  $\Delta m_{32}^2 = 2.453 \times 10^{-3} \text{ eV}^2$ , the above 6 equations are solved for  $y_1 = 0$ , resulting in  $x_1 = 0.00597 \text{ eV}$ ,  $x_2 = 0.01334 \text{ eV}$ ,  $x_3 = 0.02711 \text{ eV}$ ,  $y_2 = 0.02850 \text{ eV}$ ,  $y_3 = 0.00924 \text{ eV}$ ,  $m_1 = 0.00684 \text{ eV}$ . Hence  $m_2 = 0.011 \text{ eV}$ ,  $m_3 = 0.051 \text{ eV}$ , with  $\sum m_\nu = 0.07 \text{ eV}$ , less than  $0.15 \text{ eV}$  from cosmology.

# Freeze-In Dark Matter

Defining  $D = L - (2j)_{[mod\ 2]}$ , the charged leptons and neutrinos have  $D = 0$ , whereas  $D = 1$  for  $\eta, \zeta$  and  $D = -1$  for  $N$ .

The dark neutral fermions  $N$  are presumably light, say of order GeV in mass, and they are weakly coupled to the leptons. They are not suitable for freeze-out dark matter, but with the help of the charged scalar gauge singlet  $\zeta^- \sim 3'$ , they may be rare decay products of the SM Higgs boson, with a lifetime long enough to be freeze-in dark matter.



The effective Higgs Yukawa coupling  $f_h$  to  $\bar{N}N$  is

$$f_h = \frac{\lambda_\zeta f_\zeta^L f_\zeta^R v m_E}{16\pi^2} \left[ \frac{1}{m_\zeta^2 - m_E^2} - \frac{m_E^2 \ln(m_\zeta^2/m_E^2)}{(m_\zeta^2 - m_N^2)^2} \right].$$

The decay rate of  $h$  to  $\bar{N}N$  is

$$\Gamma_h = \frac{f_h^2 m_h}{8\pi} \sqrt{1 - 4r^2} (1 - 2r^2), \quad r = m_N/m_h.$$

If the reheat temperature of the Universe is  $T_R \sim 1$  to 10 TeV, and  $f_h \sim 10^{-11}$ , then  $N$  has the correct relic density, as a feebly interacting massive particle (FIMP).

## Concluding Remarks

Flavor symmetries may relate Dirac neutrino masses and mixing to their dark matter counterparts through the **scotogenic** mechanism, using the symmetry  $A_5$  breaking to  $A_4$ . A numerical example of **cobimaximal** neutrino mixing fitting present data with normal ordering of neutrino masses is obtained, implying a  $3 \times 3$  dark fermion mass matrix of the form (in units of order GeV)

$$\mathcal{M}_N \sim \begin{pmatrix} 0.987 & 0 & 7.461 \\ 0 & 1.724 & 0.353 \\ 1.089 & 2.419 & -0.390 \end{pmatrix}.$$

Details of this presentation appear in PLB 829, 137104 (2022) [arXiv:2202.13031].

Further development:

The lepton number  $L$  is an input global symmetry assumed in this model. However, if the dark number  $D$  is gauged and is broken by 3 (or more) units, then  $D$  remains a global symmetry.

[Ma/Picek/Radovicic, PLB 726, 744 (2013)]

The linkage of  $N$  and  $\eta, \zeta$  to the leptons through  $A_5 \rightarrow A_4$  would again lead to  $D = L - (2j)_{[mod\ 2]}$ .

[Ma, arXiv:2203.12034]

# International Conference on **Neutrinos** and **Dark Matter** (**NuDM-2022**)

25-28 September 2022 in **Sharm El-Sheikh, Egypt**

[https://conferences.andromedapublisher.com/  
NuDM-2022/index.html](https://conferences.andromedapublisher.com/NuDM-2022/index.html)

email: [NuDM-2022@andromedapublisher.com](mailto:NuDM-2022@andromedapublisher.com)

**Your participation is most welcome!**