



# Update on MB Measurements in ATLAS

Regina Kwee

on behalf of the ATLAS Collaboration

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# Introduction

- **Last** presented in this WG: inclusive measurements in
  - phase space regions at all 3 energies  
 $n_{\text{ch}} \geq 1, |\eta| < 2.5, p_{\text{T}} > 500 \text{ MeV}$  at  $\sqrt{s} = 0.9, 2.36, 7 \text{ TeV}$
  - most inclusive phase spaces  
 $n_{\text{ch}} \geq 2, |\eta| < 2.5, p_{\text{T}} > 100 \text{ MeV}$  at  $\sqrt{s} = 0.9, 7 \text{ TeV}$
- **Today** will highlight the latest, final results:
  - larger  $n_{\text{ch}}$  ranges, finalization of systematics
  - new correction method for correction of  $\langle p_{\text{T}} \rangle$  vs  $n_{\text{ch}}$
  - **studied new phase spaces**
    - diffraction-limited for tuning
    - hard-component
    - common phase spaces



# Overview

Phase spaces studied at 900 GeV and 7 TeV:

- $N_{ch} \geq 1$ ,  $p_T > 500$  MeV,  $|\eta| < 2.5$  (1<sup>st</sup> paper, 900 GeV / PLHC, 7 TeV a.k.a. 1.0 / 1.5)
- $N_{ch} \geq 6$ ,  $p_T > 500$  MeV,  $|\eta| < 2.5$  (AMBT1 tuning, a.k.a. 1.5T)
- $N_{ch} \geq 2$ ,  $p_T > 100$  MeV,  $|\eta| < 2.5$  (ICHEP, a.k.a. 2.0)
- $N_{ch} \geq 20$ ,  $p_T > 100$  MeV,  $|\eta| < 2.5$  (new)
- $N_{ch} \geq 1$ ,  $p_T > 2.5$  GeV,  $|\eta| < 2.5$  (new)
- $N_{ch} \geq 1$ ,  $p_T > 500$  MeV,  $|\eta| < 0.8$  (new)
- $N_{ch} \geq 1$ ,  $p_T > 1.0$  GeV,  $|\eta| < 0.8$  (new)

2nd MB Paper:  
arXiv 1012.5104  
Now in New  
Journal of Physics

ATLAS-CONF-  
2010-101

no change in selection criteria since last presentation  
and same analysed dataset



# Analysis Strategy

- Charged-particle multiplicity measurements:

$$\frac{1}{N_{ev}} \cdot \frac{dN_{ch}}{d\eta}, \quad \frac{1}{N_{ev}} \cdot \frac{1}{2\pi p_T} \cdot \frac{d^2N_{ch}}{d\eta dp_T}, \quad \frac{1}{N_{ev}} \cdot \frac{dN_{ev}}{dn_{ch}} \quad \text{and} \quad \langle p_T \rangle \text{ vs. } n_{ch}$$

- Correct for trigger and vertex efficiency, event-wise

$$w_{ev}(n_{sel}^{BS}) = \frac{1}{\epsilon_{trig}(n_{sel}^{BS})} \cdot \frac{1}{\epsilon_{vtx}(n_{sel}^{BS}, x)}$$

$n_{sel}^{BS}$  = # of tracks  
passing track selection  
criteria wrt beamspot

- Tracking efficiency correction on track-by-track basis

$$w_{trk}(p_T, \eta) = \frac{1}{\epsilon_{trk}(p_T, \eta)} \cdot (1 - f_{nonp}(p_T)) \cdot (1 - f_{okr}(p_T, \eta))$$

- Correct for events which have  $N_{sel}=0,1$  but  $n_{ch} \geq 2$

$$1 / (1 - (1 - \epsilon_{trk})^{n_{ch}} - n_{ch} \cdot \epsilon_{trk} \cdot (1 - \epsilon_{trk})^{(n_{ch}-1)})$$

- $N_{ch}$  and  $p_T$  distributions corrected using 1D Bayesian unfolding.



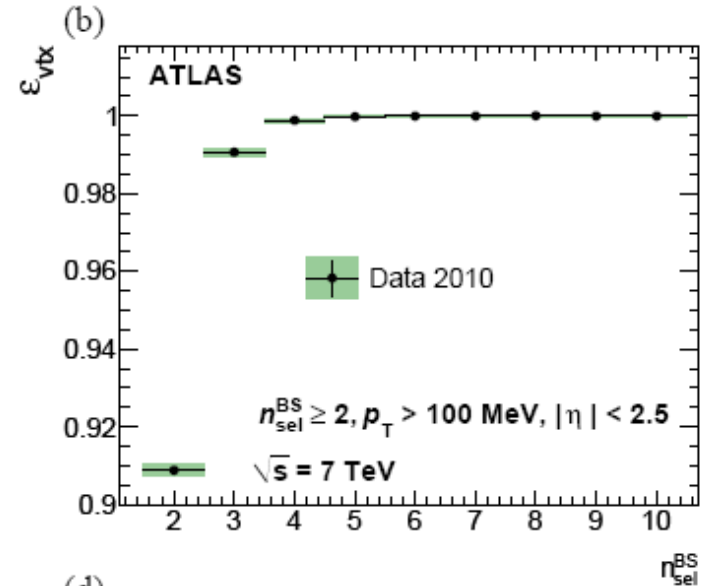
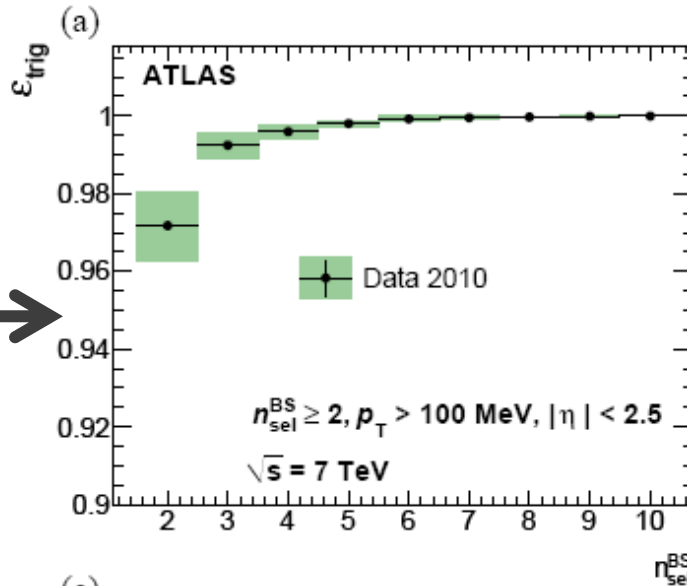
# Systematic Uncertainties

- Systematic uncertainties were evaluated for each final distribution
- Dominant uncertainty comes from the tracking efficiency due to the material description
- New systematic considered for the  $n_{ch}$ -unfolding
  - accounts for the MC  $p_T$  spectrum dependence of the unfolding matrix
- A different method is now (since ICHEP) used to obtain high  $p_T$  systematics due to
  - badly reconstructed high  $p_T$  tracks (very low  $p_T$  tracks get reconstructed at high  $p_T$  tracks due to wrong track extrapolation)
  - track resolution

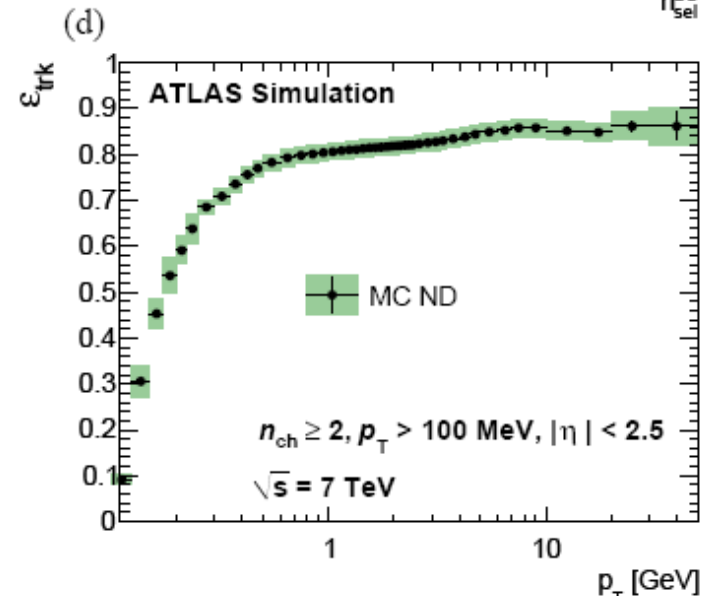
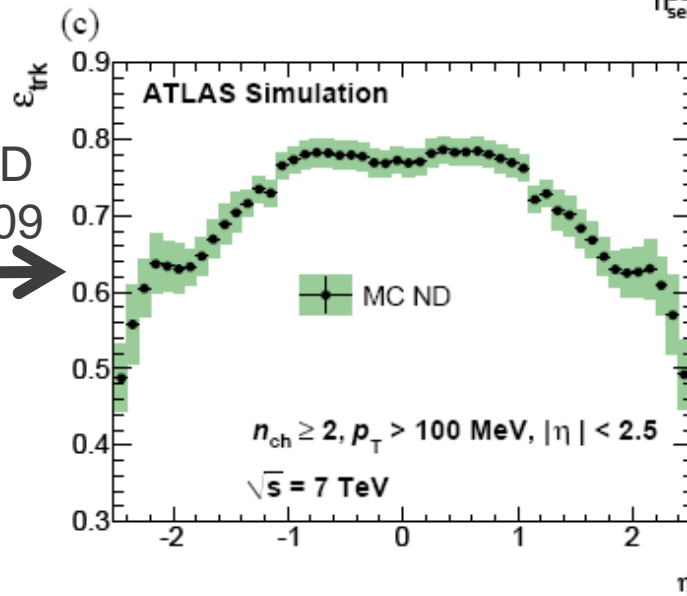


# Event and Track Level Corrections

both  
from data



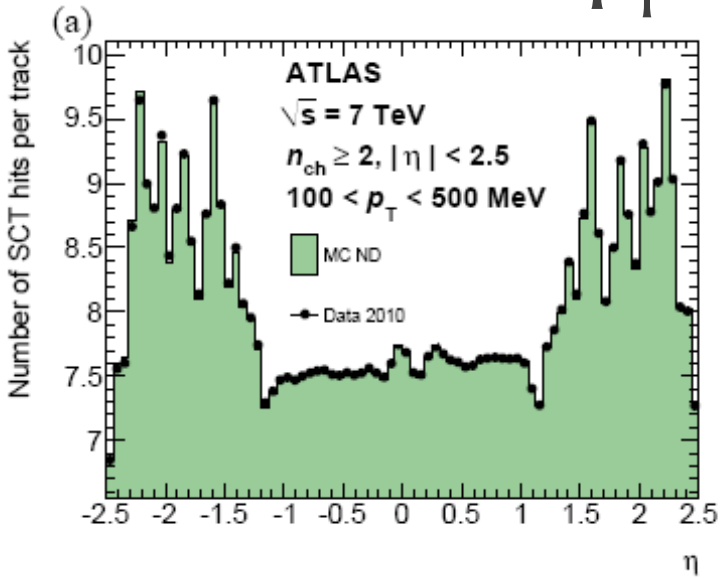
from MC ND  
ATLASMC09



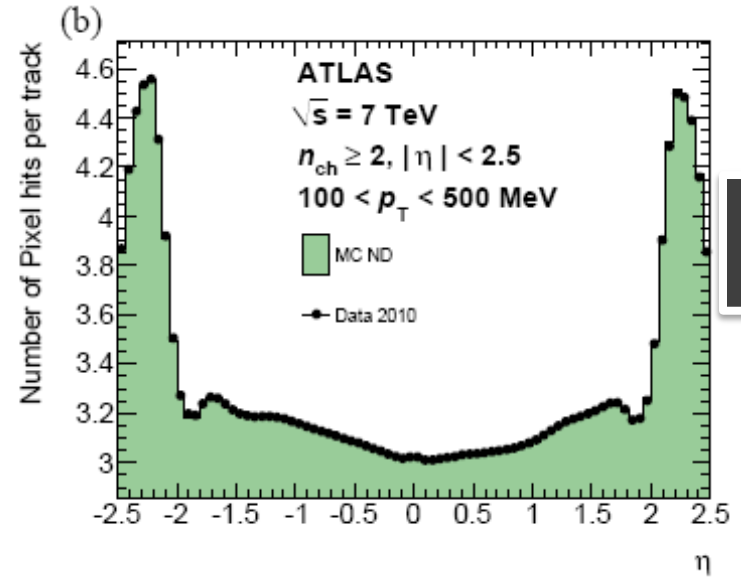


# Data and MC comparison for tracks

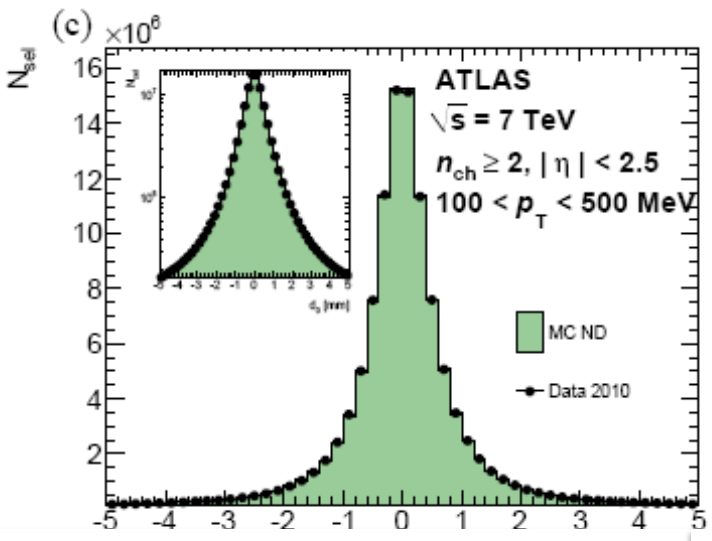
$100 < p_T < 500 \text{ MeV}$



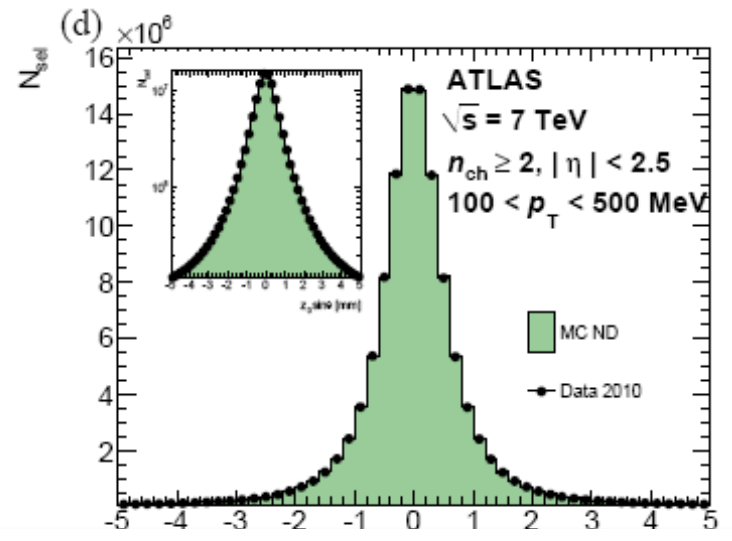
hits  
in SCT



hits  
in Pixel



transverse impact parameter  $d_0$  [mm]



longitudinal impact parameter  $z_0 \sin\theta$  [mm]



# Results

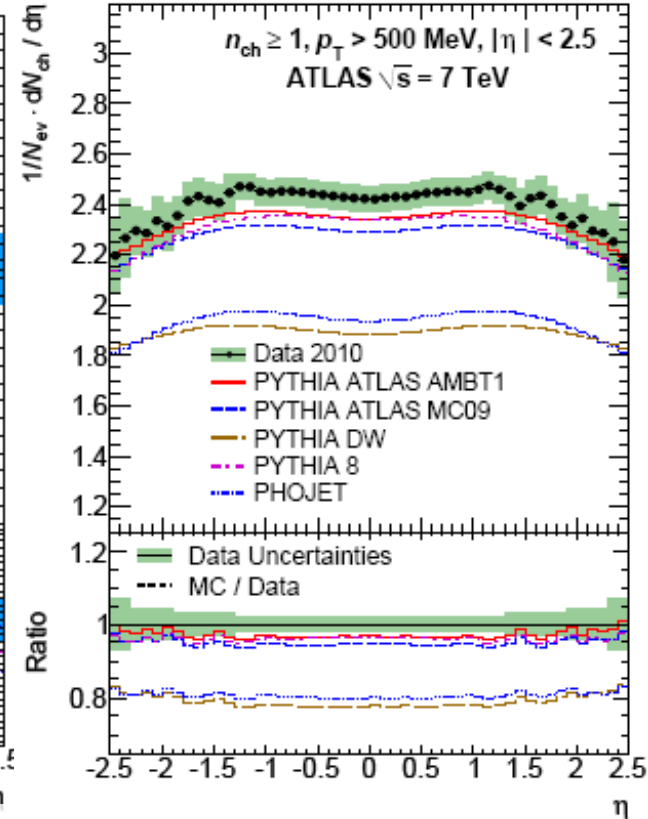
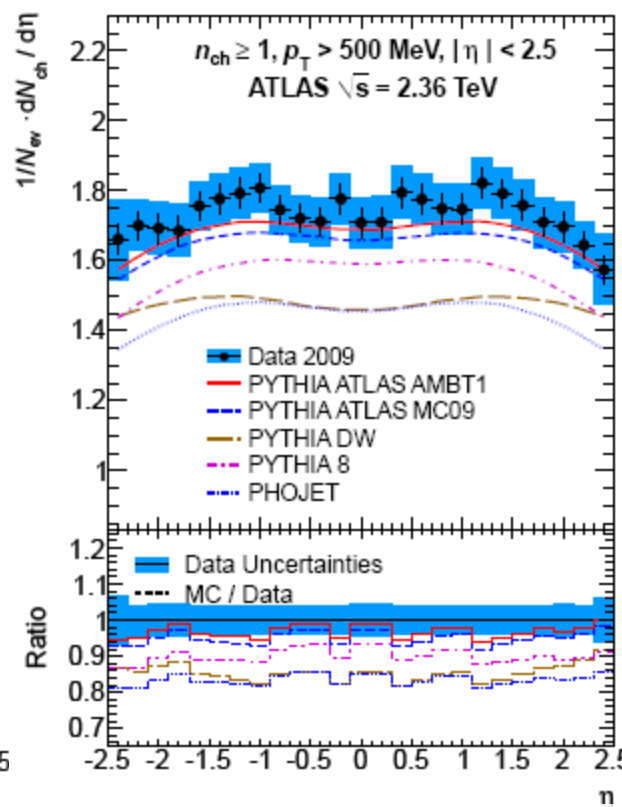
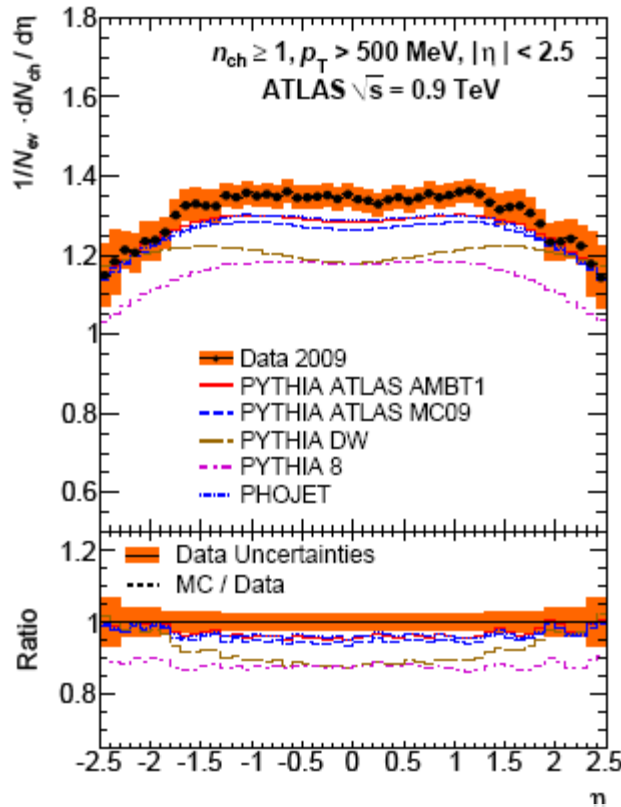
Show only selected results, full results are in the paper.





$$1/N_{ev} \frac{dn_{ch}}{d\eta}$$

phase space at all 3 energies

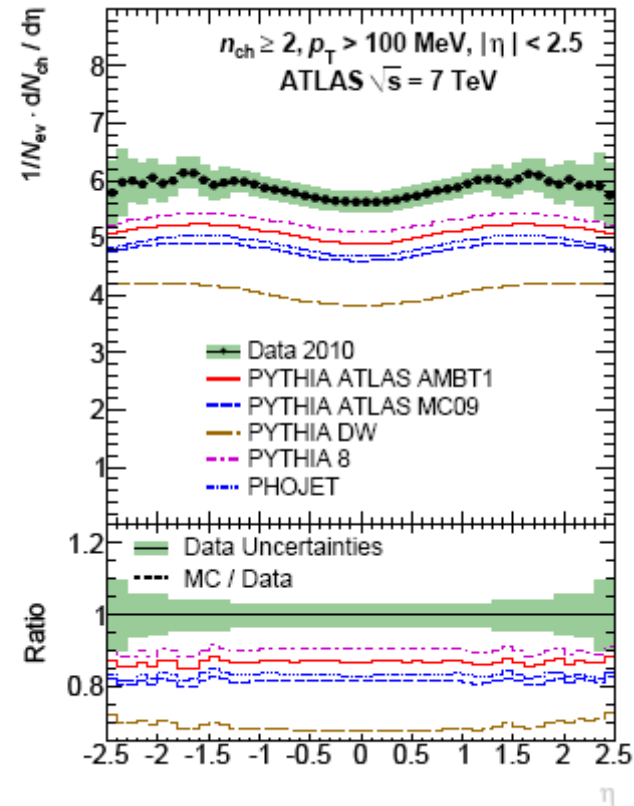
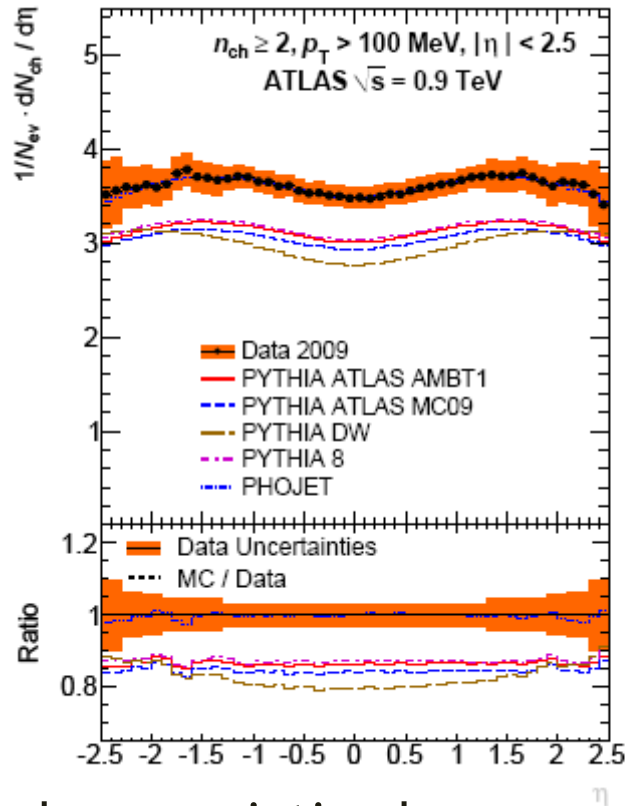


- Density is roughly constant within  $|\eta| < 1$
- Little shape variation between the models except for DW (has a more pronounced dip)



$$1/N_{ev} \frac{dn_{ch}}{d\eta}$$

most inclusive phase space

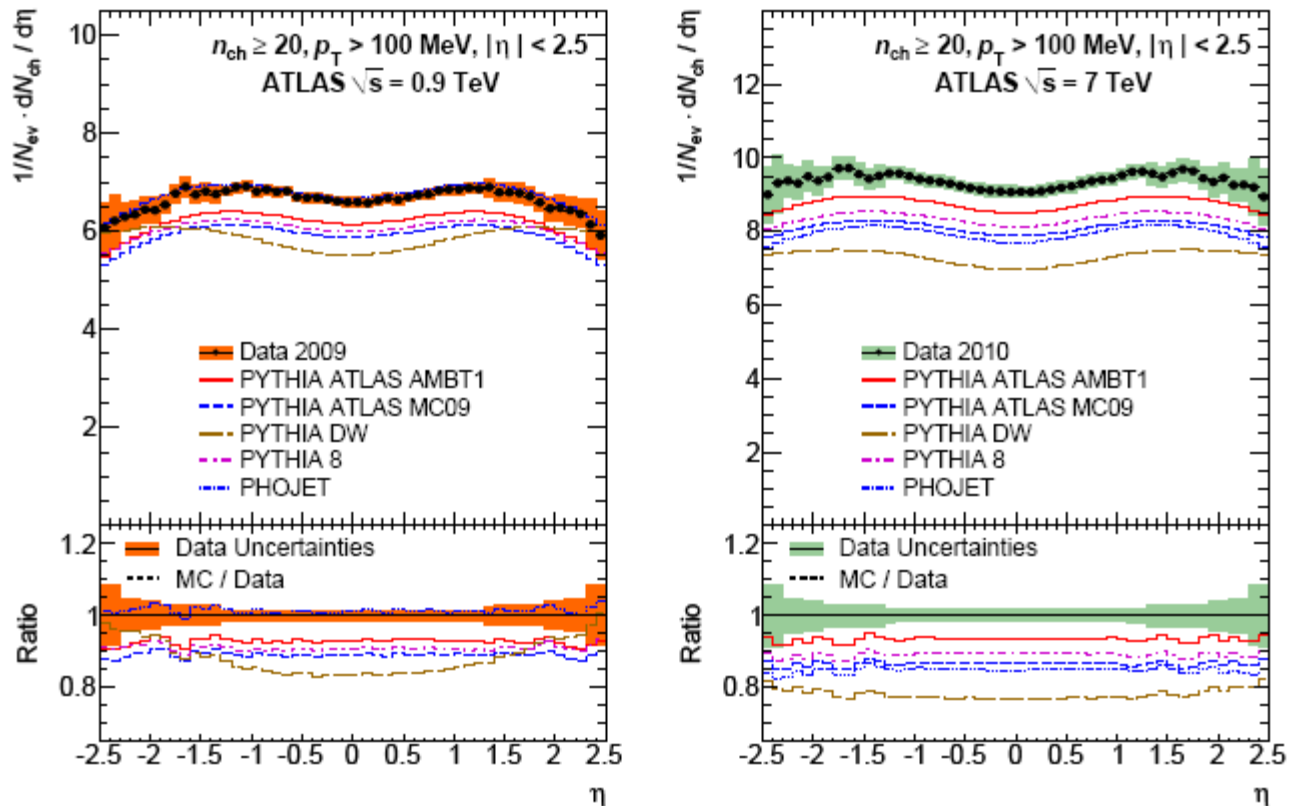


- Even less shape variation here
- Strong variation on the normalisation by all models, none reproduce the data, not even the ATLAS tune AMBT1



$$1/N_{ev} \frac{dn_{ch}}{d\eta}$$

diffractive-limited phase space with  $n_{ch} \geq 20$

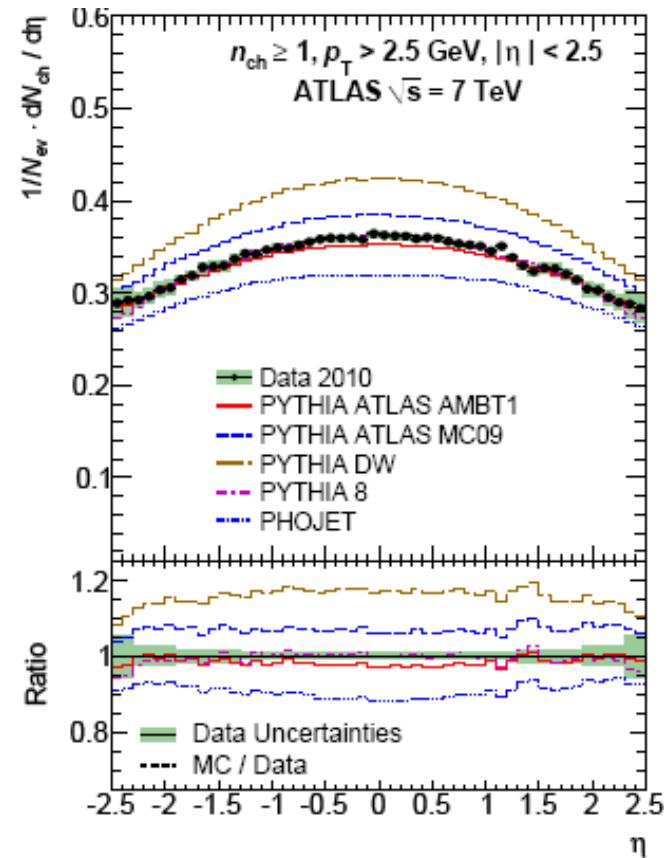
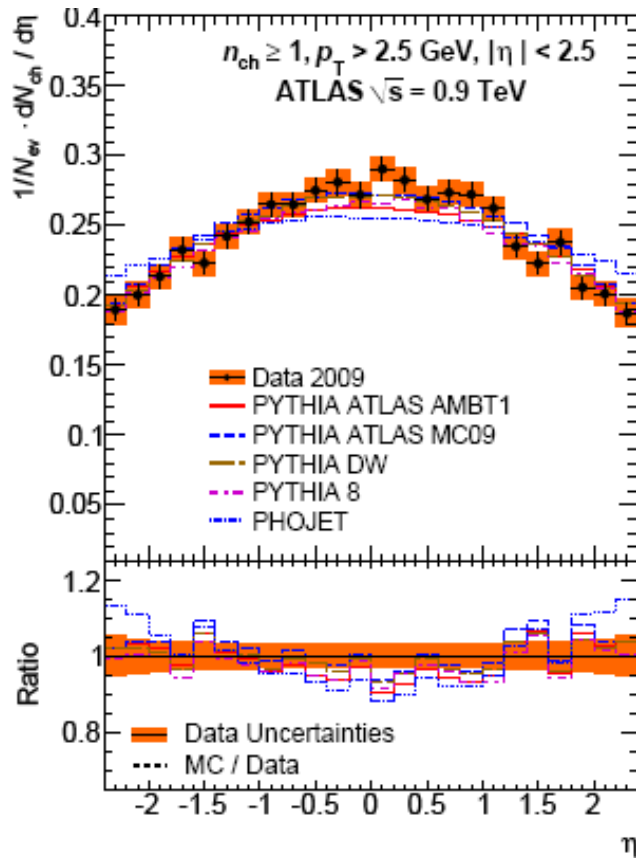


- Situation is not significantly different when requiring 20 charged particles or more, only phojet matches data at  $\sqrt{s} = 900 \text{ GeV}$



$$1/N_{ev} \frac{dn_{ch}}{d\eta}$$

phase space with „hard“-component

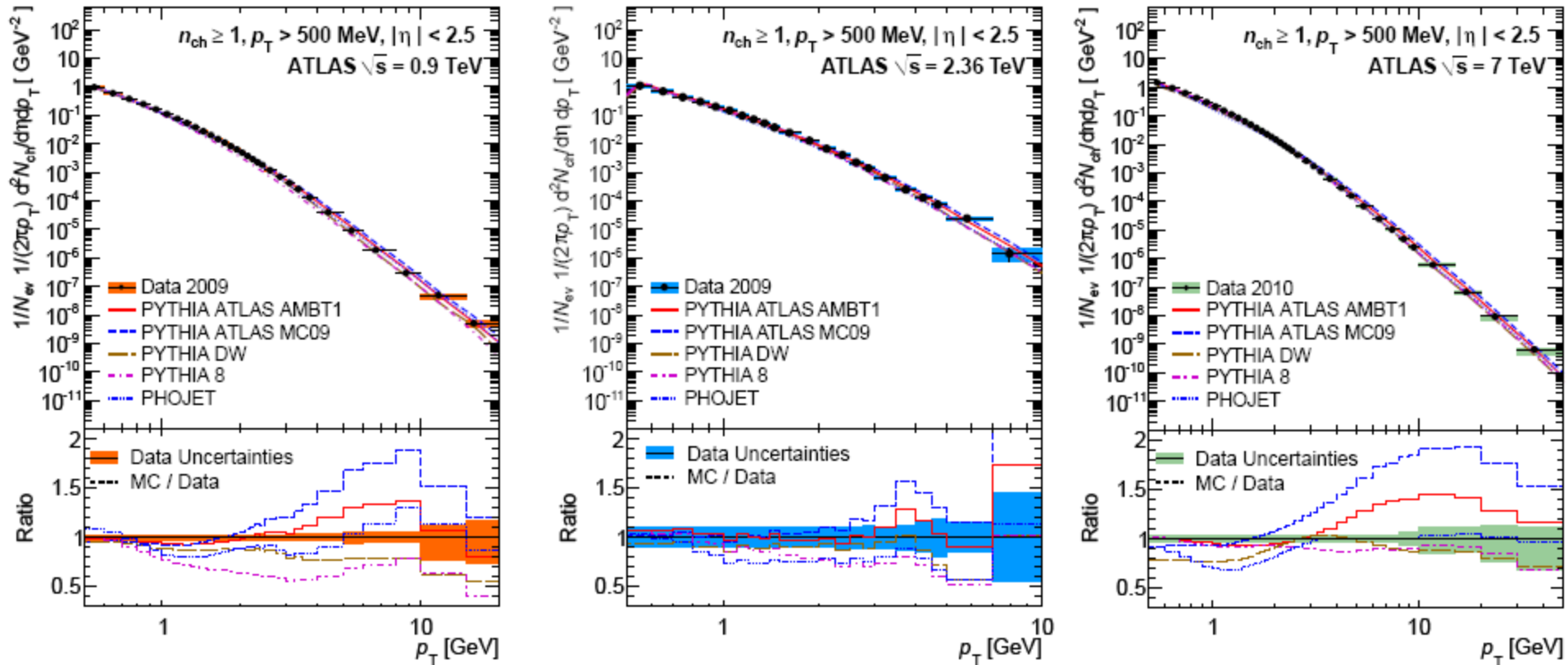


- Achieve much better description of the data at  $\sqrt{s} = 900 \text{ GeV}$ , only AMBT1 and pythia8 at  $\sqrt{s} = 7 \text{ TeV}$  match the data



$$1/N_{ev} \frac{1}{2\pi p_T} \frac{d^2N_{ch}}{d\eta dp_T}$$

phase space at all 3 energies

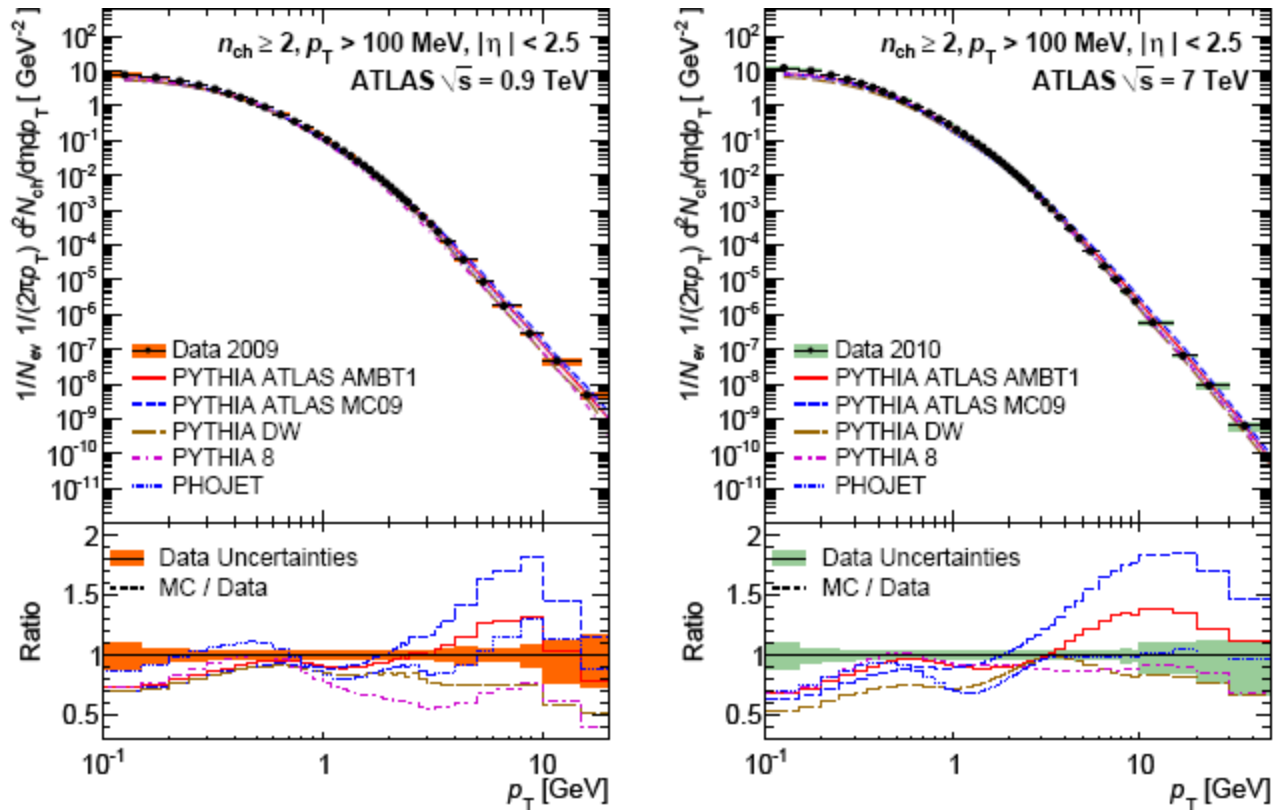


- Not described by any of the models, especially above 1 GeV
- pythia8 comes quite close to the data at  $\sqrt{s} = 7$  TeV



$$\frac{1}{N_{ev}} \frac{1}{2\pi p_T} \frac{d^2 N_{ch}}{d\eta dp_T}$$

most inclusive phase space

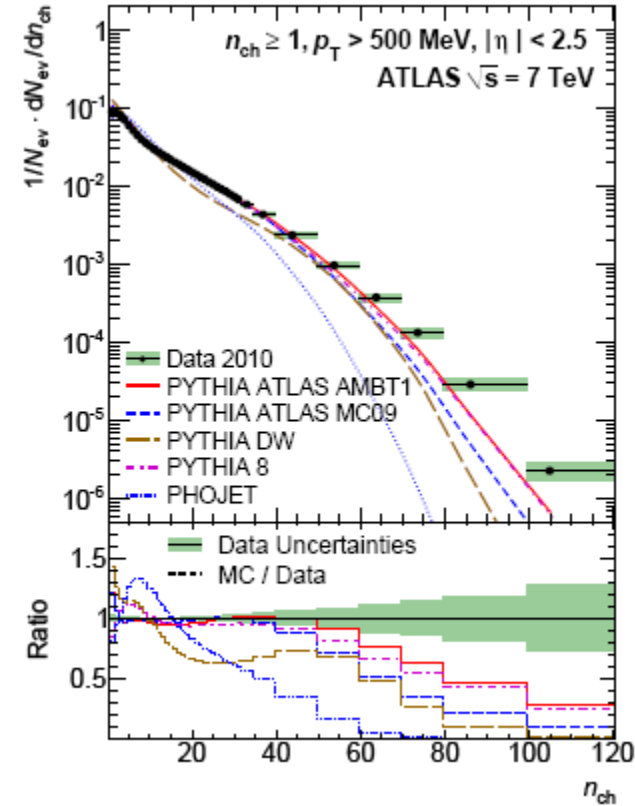
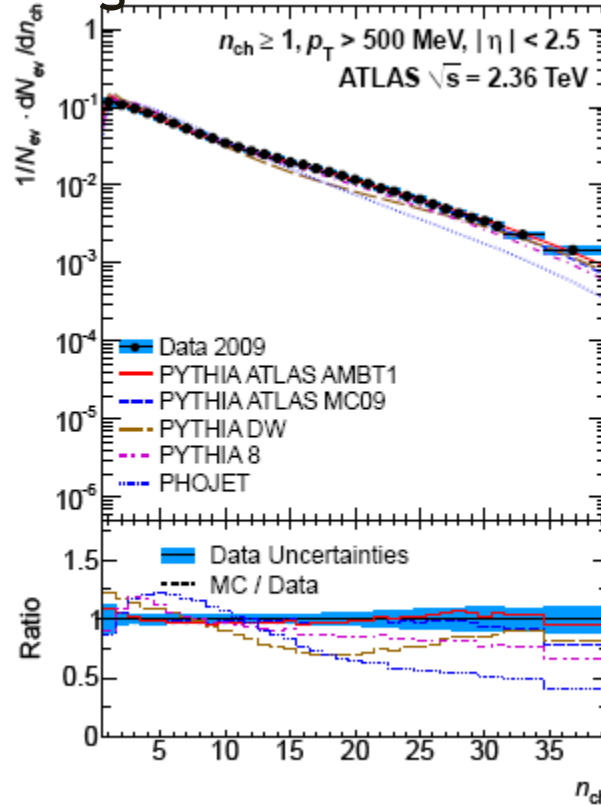
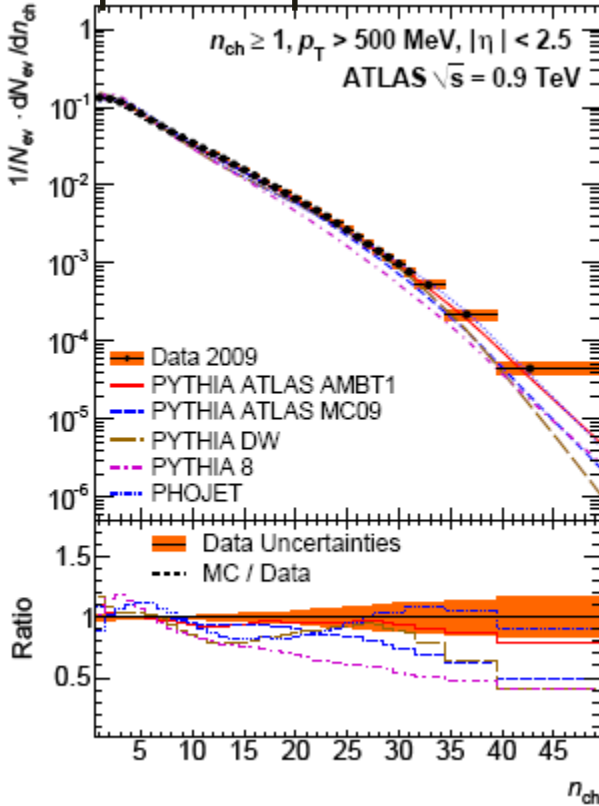


- $\sqrt{s} = 900 \text{ GeV}$  phojet describes the data best over the whole range, even though the agreement is not excellent
- generally low  $p_T$  particles are underpredicted, more significant at  $\sqrt{s} = 7 \text{ TeV}$



$$1/N_{ev} \frac{dN_{ev}}{dn_{ch}}$$

phase space at all 3 energies

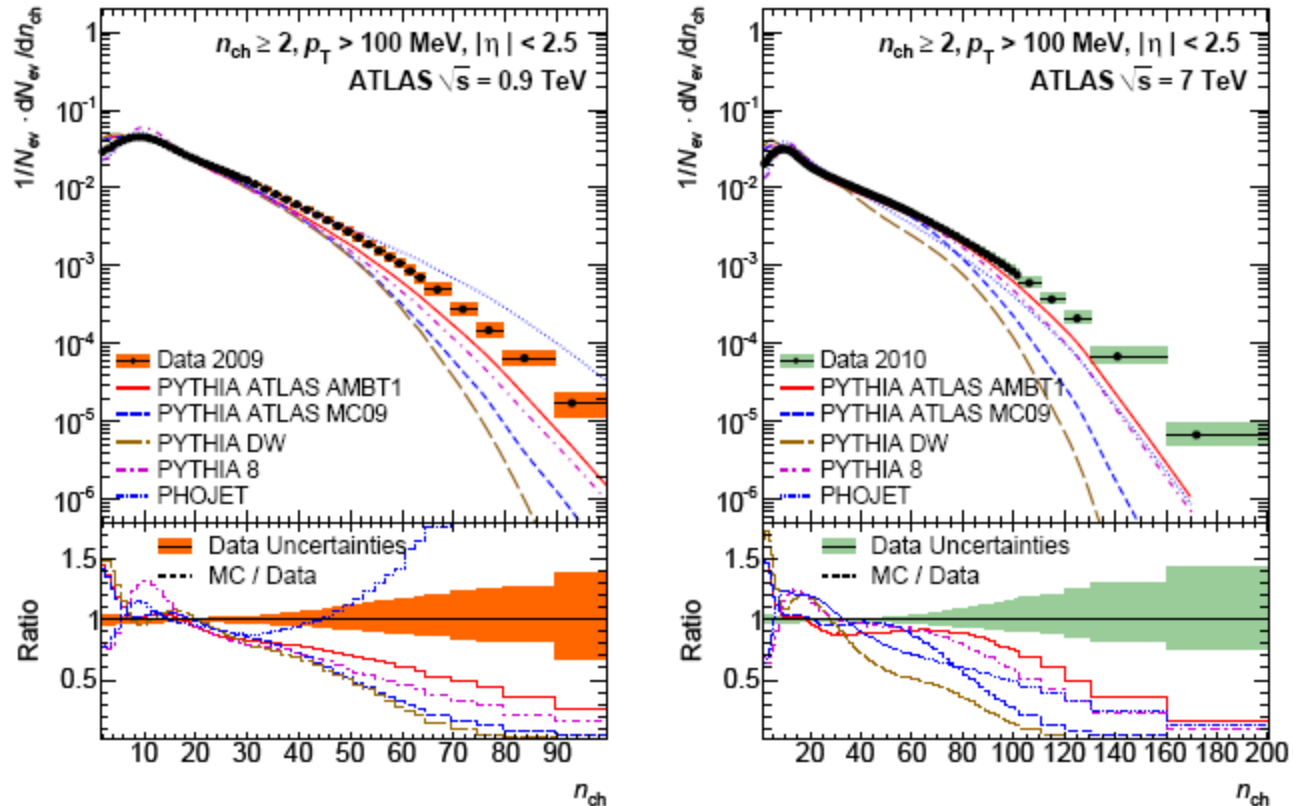


- Models overpredict low-multiplicity events, compensated by underprediction of high-multiplicity events
- phojet models distribution reasonably well but only at  $\sqrt{s} = 900 \text{ GeV}$



$$1/N_{ev} \frac{dN_{ev}}{dn_{ch}}$$

most inclusive phase space



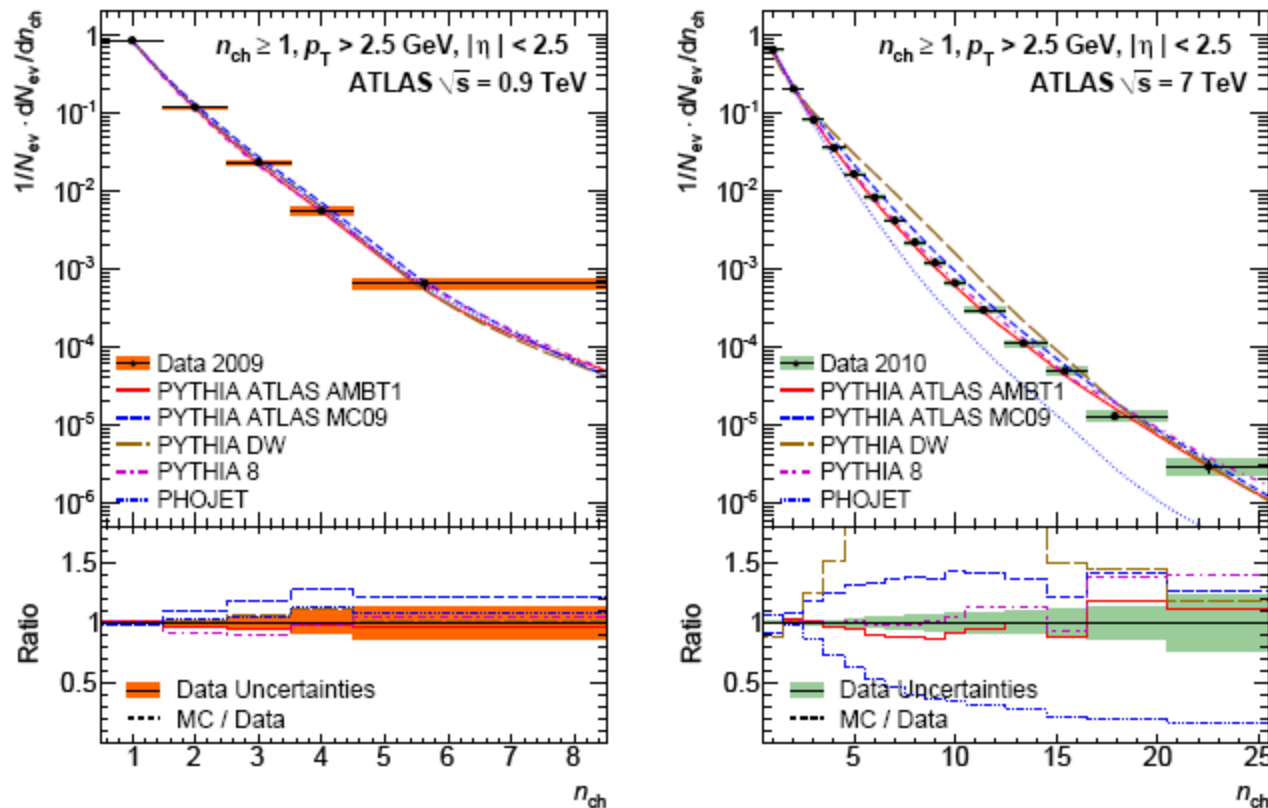
- Variations of low and high multiplicity events are even larger here





$$1/N_{ev} \frac{dN_{ev}}{dn_{ch}}$$

phase space with hard component



- Fairly good at  $\sqrt{s} = 900$  GeV, but large difference between the tunes/models  $\sqrt{s} = 7$  TeV : all but AMBT1 and pythia8 are very different to the data



## $\langle p_T \rangle$ vs $n_{ch}$

„Old“ method: correction in 2 steps

- bin-by-bin correction of  $\langle p_{T\ reco} \rangle$  to  $\langle p_T \rangle$
- $n_{ch}$  unfolding

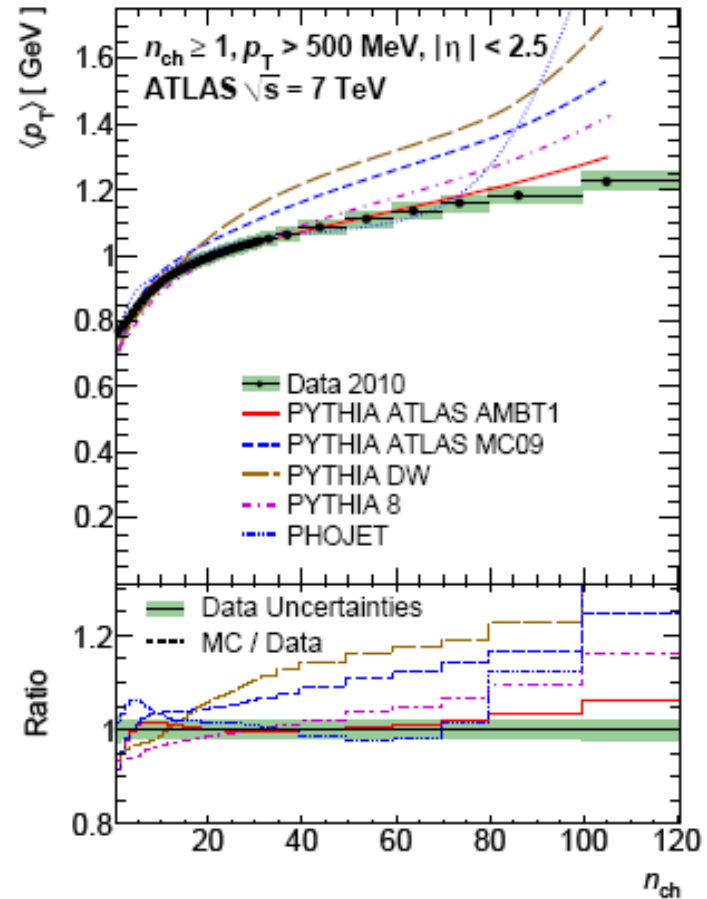
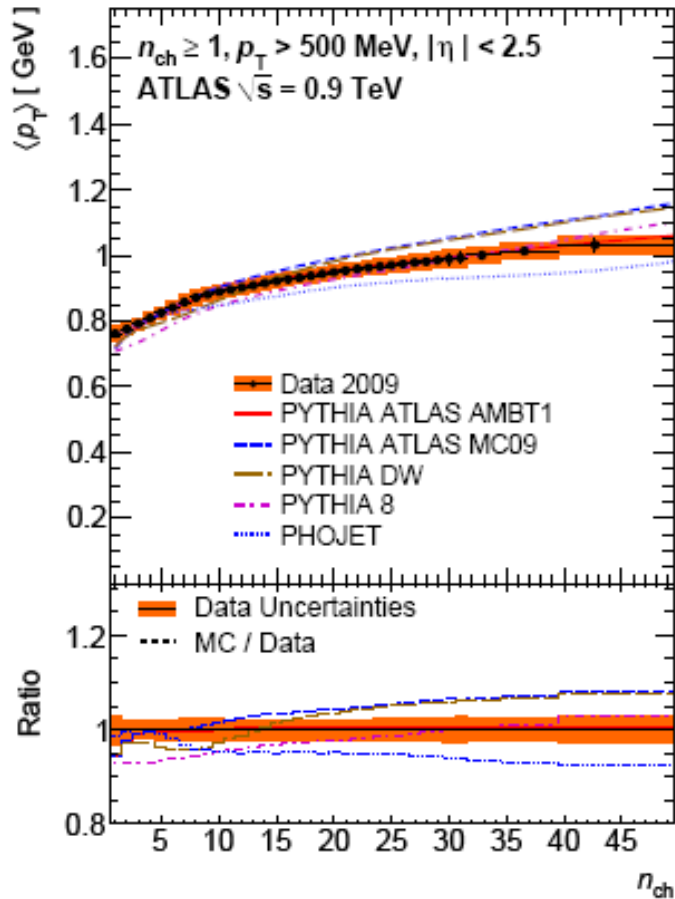
New method was developed to remove dependence of MC  $p_T$  spectrum

- Correct numerator and denominator separately, assume
  - tracking efficiency depends only on  $p_T$  and  $\eta$
  - $p_T$  spectrum is the same in the  $n_{sel}$  bin and the  $n_{ch}$  bin the event migrates to
- Limited validity is visible in closure tests
  - considered as systematic uncertainty (estimated using different MC samples)

→ Change of mean values from 2% up to 4% at higher  $n_{ch}$  values in the most inclusive phase space. Understood as  $p_T$  spectrum difference between data and MC.



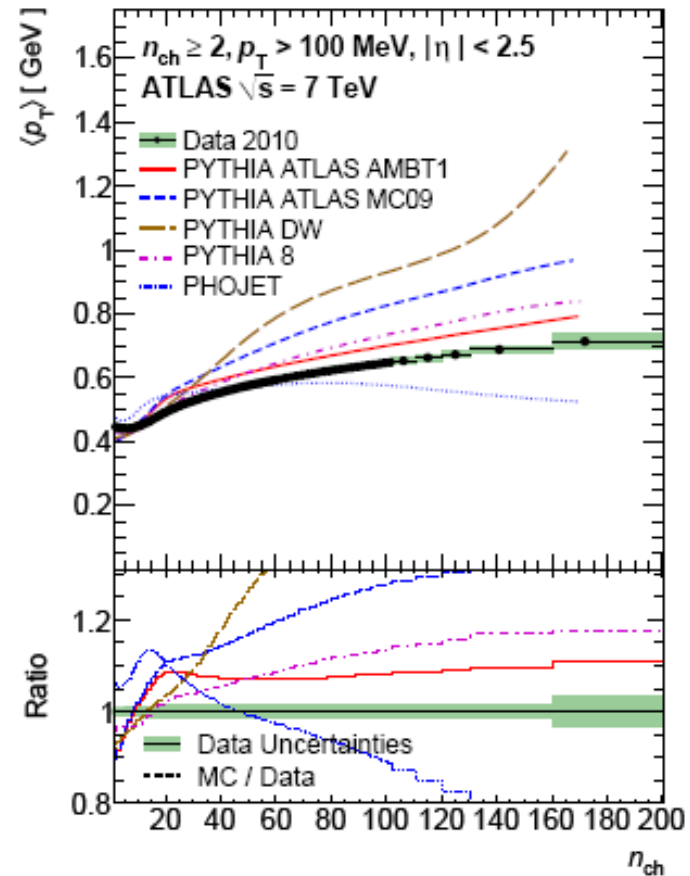
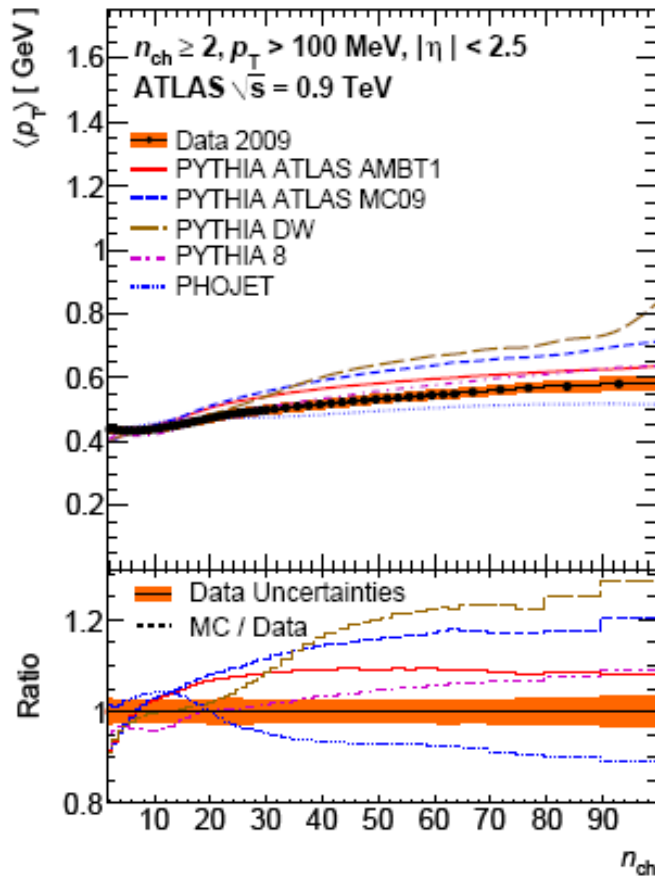
# $\langle p_T \rangle$ vs. $n_{ch}$



- AMBT1 tune yields the closest description of the data



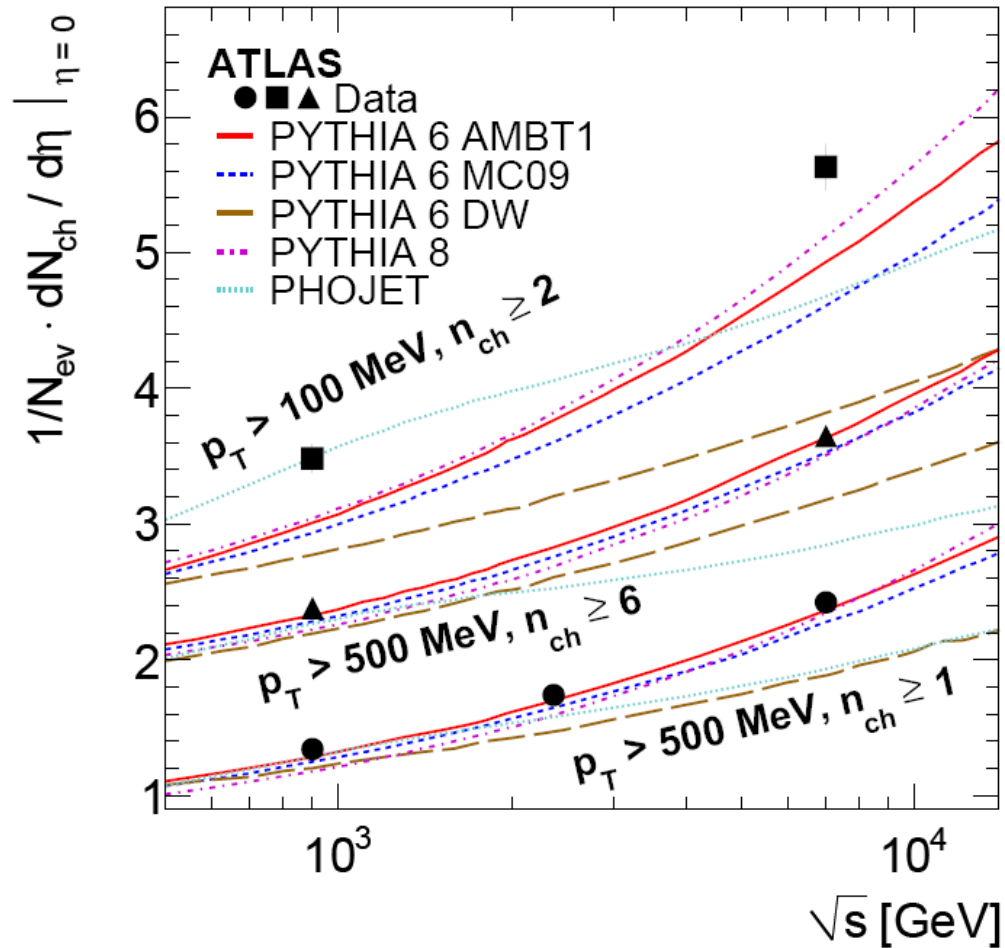
# $\langle p_T \rangle$ vs. $n_{ch}$



- Still true in the most inclusive phase space, but discrepancy gets larger
- Measurements are very precise



# $1/N_{ev} dN_{ch}/d\eta$ at $\eta = 0$



- The harder the  $p_T$  cut the better the agreement.
- Only data points at  $\sqrt{s} = 900 \text{ GeV}$  is well matched by phojet .



# Extrapolation to $p_T = 0$

- For comparison purposes, extrapolate  $p_T$  spectrum down to  $p_T = 0$
- Estimate number of charged-particles with  $p_T < 100$  MeV with 3 different methods
- Each methods yields a correction factor that is applied to the mean charged-particle multiplicity in bins of  $\eta$  averaged over  $|\eta| < 2.5$
- Differences taken as syst. uncertainty

| Method | cf at $\sqrt{s} = 900$ GeV | cf at $\sqrt{s} = 7$ TeV |
|--------|----------------------------|--------------------------|
| 1      | 1.065                      | 1.063                    |
| 2      | 1.068                      | 1.065                    |
| 3      | 1.055                      | 1.051                    |

final result

|  |  |
|--|--|
| $1.065 \pm 0.011$ (tot)<br>at $\sqrt{s} = 900$ GeV | $1.063 \pm 0.014$ (tot)<br>at $\sqrt{s} = 7$ TeV |
|--|--|

- [Method 1](#): Use a modified Tsallis function to fit the  $p_T$  spectrum (accounts also for fact that  $\eta$  is measured while  $y$  is used)
- [Method 2](#): Assume  $p_T$  distribution is flat in the lower  $p_T$  region. Extract correction factor from first measured bin  $100 < p_T < 150$  MeV
- [Method 3](#): Use AMBT1





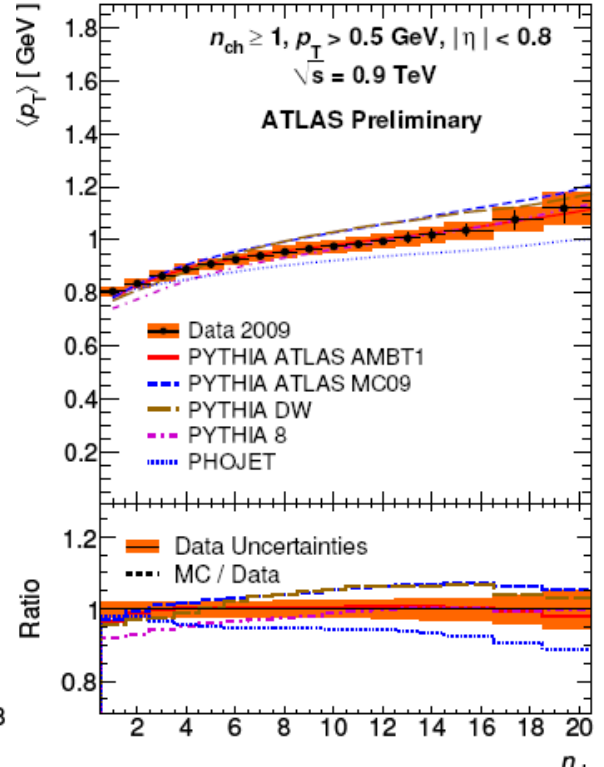
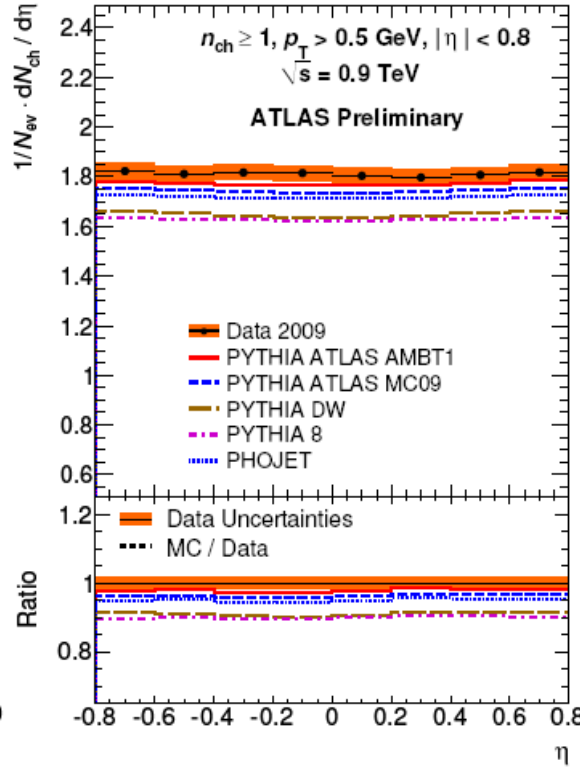
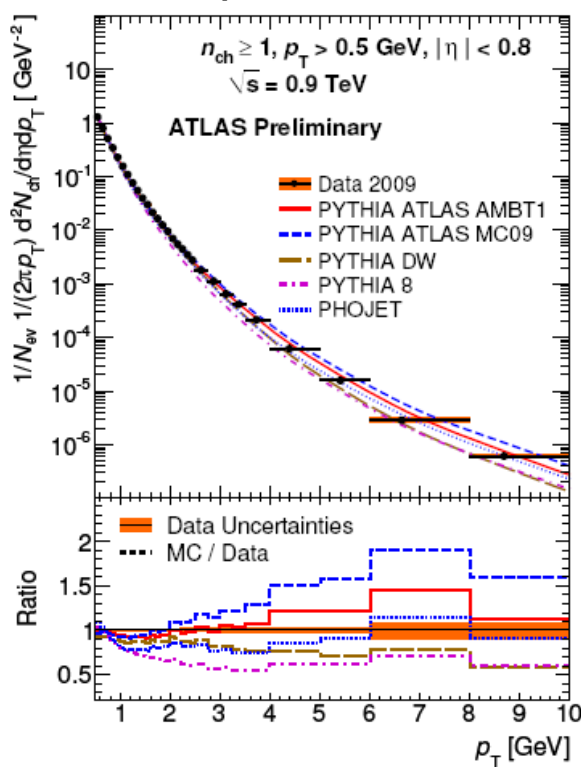
# MB Common Plots

ATLAS-CONF-  
2010-101

- All MB distributions analysed in „common“ phase spaces with the same analysis procedure

$n_{\text{ch}} \geq 1, |\eta| < 0.8, p_{\text{T}} > 0.5, 1 \text{ GeV}$  at  $\sqrt{s} = 0.9, 7 \text{ TeV}$

examples at  $\sqrt{s} = 900 \text{ GeV}$





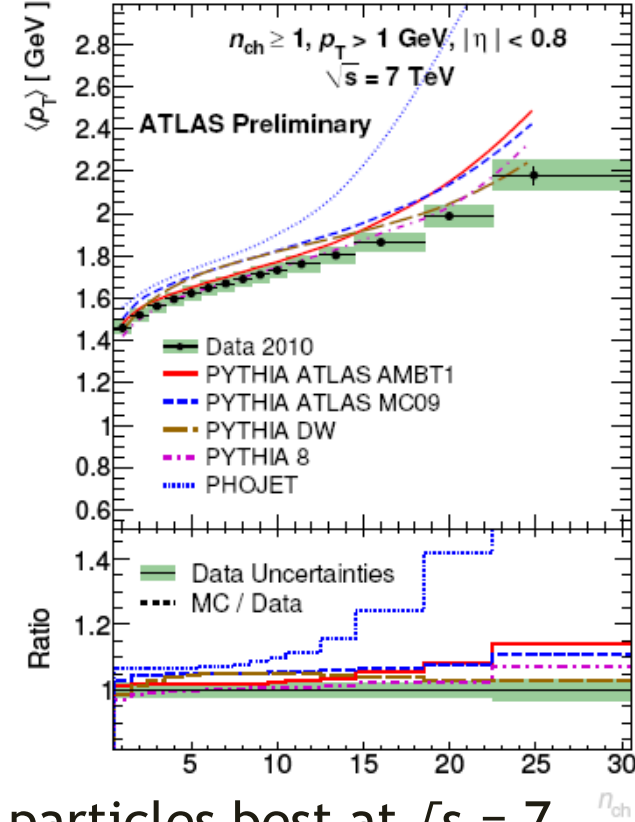
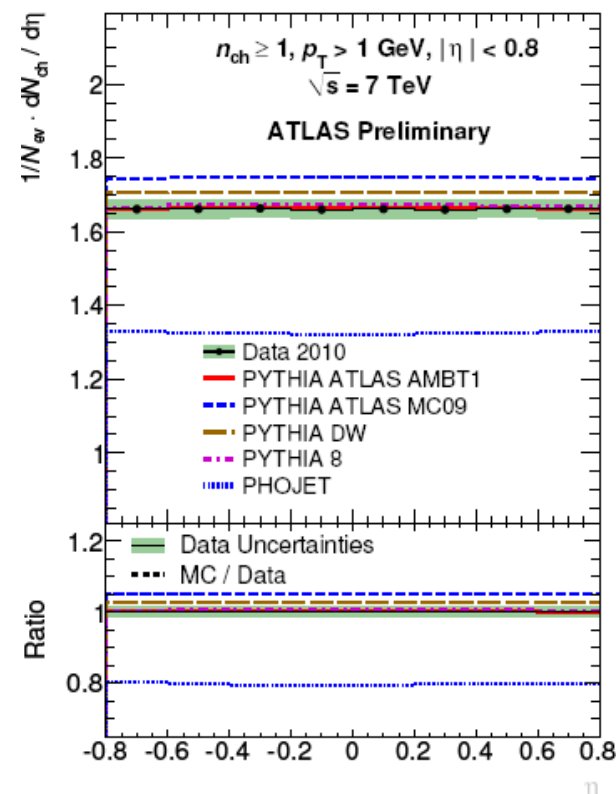
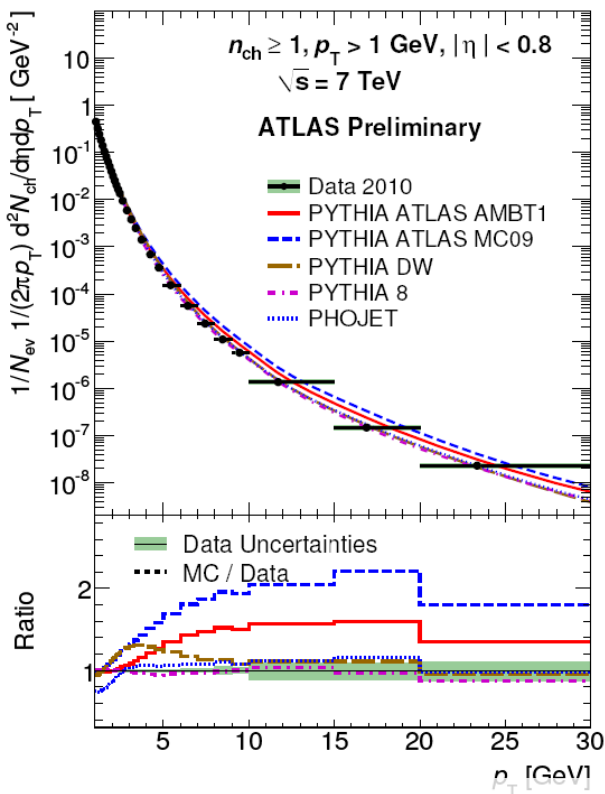


# MB Common Plots

ATLAS-CONF-2010-101

- Examples for other phase space

$n_{ch} \geq 1, |\eta| < 0.8, p_T > 1 \text{ GeV}$  at  $\sqrt{s} = 0.9, 7 \text{ TeV}$



- AMBT1 and pythia8 describe the central charged particles best at  $\sqrt{s} = 7 \text{ TeV}$ , but pythia8 provides a better description of the high  $p_T$  tails



# Conclusion

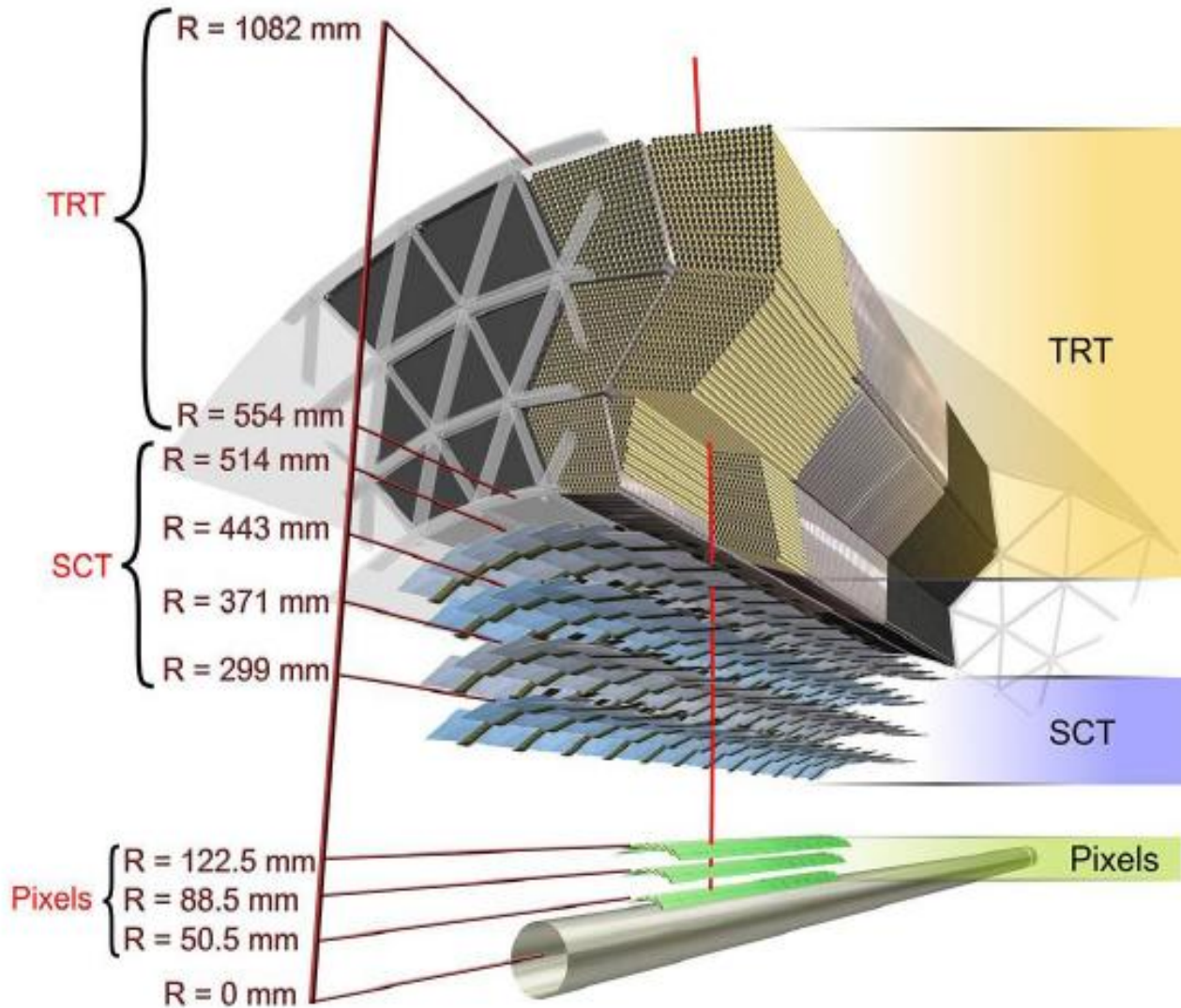
- Properties of charged-particles were studied in well defined phase space regions
- Data was corrected with minimal model dependence
- Precision of this analysis reveals clear differences between the MC models and the measured distributions
- In general agreement in models and data are better when diffractive contribution is smaller
- More analysis details in the paper
  - 2nd MB Paper: arXiv 1012.5104
  - see also public page:  
[https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/PAPERS/MinBias\\_02/](https://atlas.web.cern.ch/Atlas/GROUPS/PHYSICS/PAPERS/MinBias_02/)
  - HEPDATA <http://hepdata.cedar.ac.uk/view/p7918>



back up

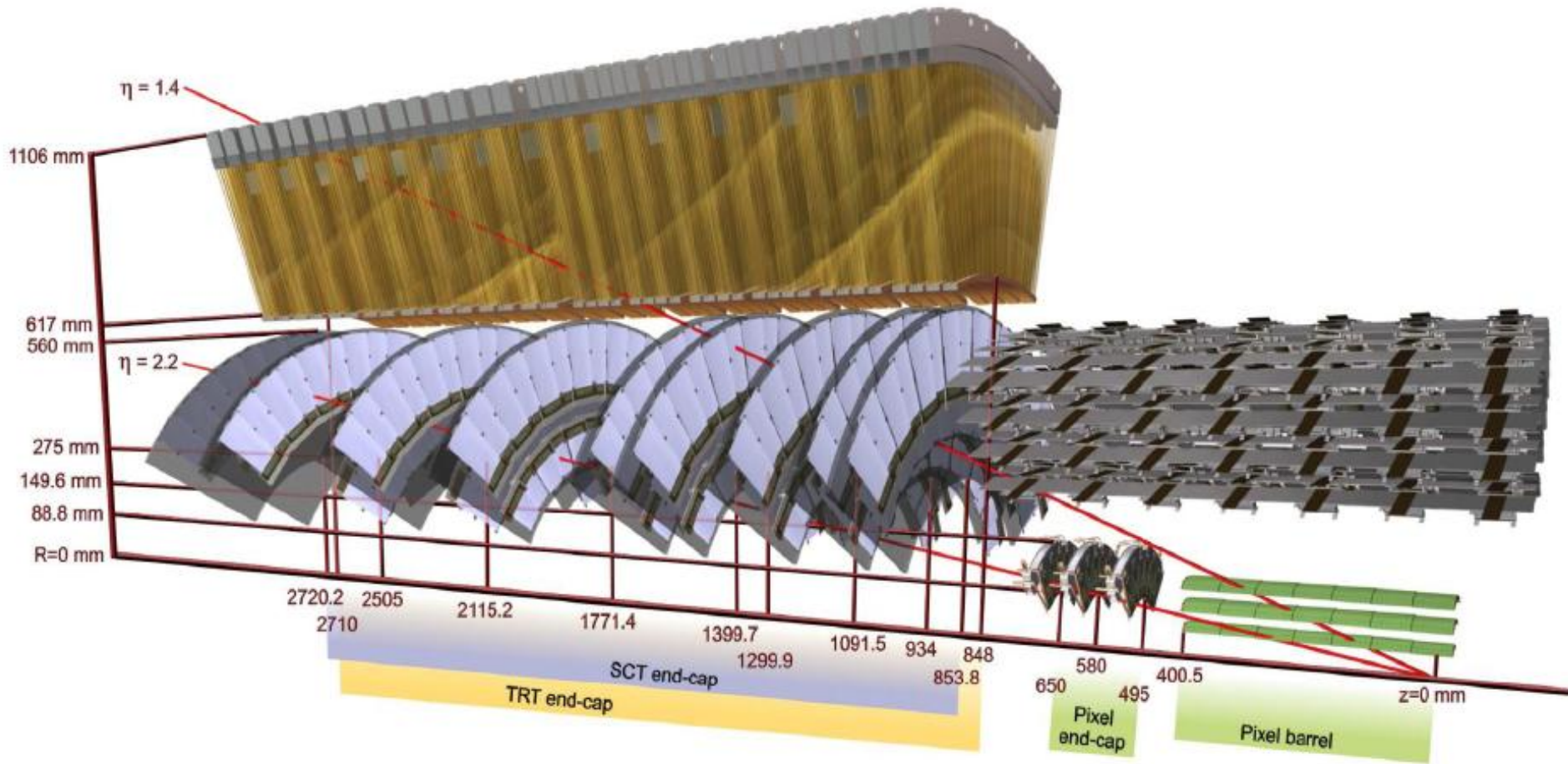


# ATLAS Inner Detector Barrel





# ATLAS Inner Detector Endcap





# Event Selection

## Requirements per event

- Single-arm trigger (MBTS\_1)
- at least 1 reconstructed vertex for beam background removal
- $n_{\text{ch}} \geq 1, 6, 20$
- pile-up removal: if a second reconstructed vertex has  $\geq 4$  tracks

## Track Quality criteria to reduce background to primary tracks

- hit in the innermost layer if expected
- 2, 4 or 6 hits in the silicon strip detector depending for  $p_{\text{T}} < 100, 200$  and  $\geq 300$  MeV (different for the analysis at  $\sqrt{s} = 2.36$  TeV)

## Kinematic phase space cuts on the track

- $p_{\text{T}} > 0.1, 0.5$  or  $2.5$  GeV
- $|\eta| < 2.5, 0.8$

Event sample:  
at  $\sqrt{s} = 7$  TeV:  $\sim 190 \mu\text{b}^{-1}$   
at  $\sqrt{s} = 900$  GeV:  $\sim 7 \mu\text{b}^{-1}$



# Systematic Uncertainty on Tracking Efficiency

| Systematic Uncertainty         | Size                              | Region  |
|--------------------------------|-----------------------------------|---|
| Material                       | $\pm 2 - 15\%$                    | decreases with $p_T$ , increases with $ \eta $                                    |
| $\chi^2$ prob. cut             | $\pm 10\%$                        | flat, only for $p_T > 10$ GeV   |
| Resolution                     | $\pm 5\%$<br>negligible<br>$-7\%$ | $100 < p_T < 150$ MeV<br>$0.15 < p_T < 10$ GeV<br>$p_T > 10$ GeV                  |
| Track Selection                | $\pm 1\%$                         | flat in $p_T$ and $\eta$  |
| Truth Matching                 | $\pm 1\%$                         | only for $\sqrt{s} = 2.36$ TeV Pixel Tracks                                       |
| Efficiency correction factor   | $\pm 4\%$                         | only for $\sqrt{s} = 2.36$ TeV ID Track   |
| Alignment and other high $p_T$ | $-3\%$ to $-30\%$                 | only for $p_T > 10$ GeV<br>averaged over $\eta$ , increases with increasing $p_T$ |



## $\langle p_T \rangle$ vs $n_{ch}$

**new method** was developed to remove dependence of MC  $p_T$  spectrum:  
correct separately 2 components

- **Numerator:** Sum  $p_T$  of all tracks in all events vs  $n_{ch}$
  - **Denominator:** Sum of all tracks in all events vs  $n_{ch}$
- apply track-by-track corrections  
(track weights, no dependence on MC  $p_T$  spectrum)
- use same  $n_{ch}$ -unfolding matrix for denominator as for  $1/N_{ev} dN_{ev}/dn_{ch}$  correction