

Unitarity and clock dependence in quantum cosmology

Lucía Menéndez-Pidal

University of Nottingham

Work in collaboration with Steffen Gielen

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Nottingham
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- 1 Introduction
- 2 Our model
- 3 Quantisation(s) and unitarity
- 4 Effects on dynamics
- 5 Conclusion

What is the problem of time (POT)?

Diffeomorphism invariance \implies GR is a **constrained** system:

$$\mathcal{S}_{EH} = \int dt d^3x \left[p^{ab} \dot{h}_{ab} - N \mathcal{H}_{\perp}^g - N^a \mathcal{H}_a^g \right]$$

N and N^a (lapse and shift functions) are Lagrange multipliers. Hence,

$$\mathcal{H}_{\perp}^g = 0, \quad \mathcal{H}_a^g = 0$$

When we quantise we find

$$\hat{\mathcal{H}}_{\perp}^g \Psi = 0, \quad \hat{\mathcal{H}}_a^g \Psi = 0$$

There is no external time variable. How do we make observables evolve?

The model we studied

Ingredients

- flat FLRW: $ds^2 = -N(\tau)^2 d\tau^2 + a(\tau)^2(dx^2 + dy^2 + dz^2)$
- a cosmological constant Λ from **unimodular gravity**
- free massless **scalar field** ϕ

$$S_{PUM} = V_0 \int_{\mathbb{R}} d\tau \left\{ \frac{3\dot{a}^2 a}{N\kappa} - Na^3 \frac{\Lambda}{\kappa} + \Lambda \dot{T} + \frac{a^3}{2N} \dot{\phi}^2 \right\}$$

Change of variables:

$$v \propto a^3, \quad \lambda \propto \Lambda, \quad t \propto T, \quad \varphi \propto \phi,$$

where $\{t, \lambda\} = 1$. **The cosmological constant is a constant of motion.**

$$\mathcal{H} = \tilde{N} \left[-\pi_v^2 + \frac{\pi_\varphi^2}{v^2} + \lambda \right] \xrightarrow{\text{constraint}} \mathcal{C} = -\pi_v^2 + \frac{\pi_\varphi^2}{v^2} + \lambda = 0$$



Wheeler–DeWitt equation

$$\hat{C}\Psi = 0 \implies \left(\hbar^2 \frac{\partial^2}{\partial v^2} + \frac{\hbar^2}{v} \frac{\partial}{\partial v} - \frac{\hbar^2}{v^2} \frac{\partial^2}{\partial \varphi^2} - i\hbar \frac{\partial}{\partial t} \right) \Psi(v, \varphi, t) = 0$$

(WdW)

POT: how to extract time evolution from this equation?

Answer: A \rightarrow choose an internal variable to play the rôle of time parameter (relational quantisation), and B \rightarrow Dirac quantisation (ask me about that in the breaks :D)

Answer A: choose a clock

t and φ are good clocks classically, hence both can a priori be good clocks for the quantum theory

The t -clock theory

$$i\hbar\partial_t\Psi = - \underbrace{\left(\hbar^2\partial_v^2 + \frac{\hbar^2}{v}\partial_v - \frac{\hbar^2}{v^2}\partial_\varphi^2 \right)}_{\hat{H}}\Psi$$

Schr. eq. \implies Schr. inner product

$$\langle\Psi|\Phi\rangle_t = \int d\varphi dv v\bar{\Psi}\Phi$$

The φ -clock theory

Multiply (WdW) by v^2

$$\hbar^2\partial_\varphi^2\Psi = \underbrace{\left(\hbar^2(v\partial_v)^2 - iv^2\hbar\partial_t \right)}_{\hat{G}}\Psi$$

K.G. eq. \implies K.G. inner product

$$\langle\Psi|\Phi\rangle_\varphi = i \int dt \frac{dv}{v} (\bar{\Psi}\partial_\varphi\Phi - \Phi\partial_\varphi\bar{\Psi})$$



choose a clock (continued)

\hat{H} and \hat{G} are not self-adjoint \implies boundary conditions have to be imposed to preserve unitarity

t-clock theory

$$\int d\varphi [v\bar{\Psi}\partial_v\Phi - v\Phi\partial_v\bar{\Psi}]_{v=0} = 0$$

φ -clock theory

$$\int dt [v\bar{\Psi}\partial_v\Phi - v\Phi\partial_v\bar{\Psi}]^{v=\infty} = 0$$

Remark:

- In the *t*-theory, the boundary condition is at $v = 0$ (classical singularity)
- In the φ -theory, the boundary condition is at $v = \infty$



Singularity resolution

expectation values = $\langle \Psi_{sc} | v(t) | \Psi_{sc} \rangle = \langle v(t) \rangle_{\Psi_{sc}}$ and
 $\langle \Psi_{sc} | v(\varphi) | \Psi_{sc} \rangle = \langle v(\varphi) \rangle_{\Psi_{sc}}$

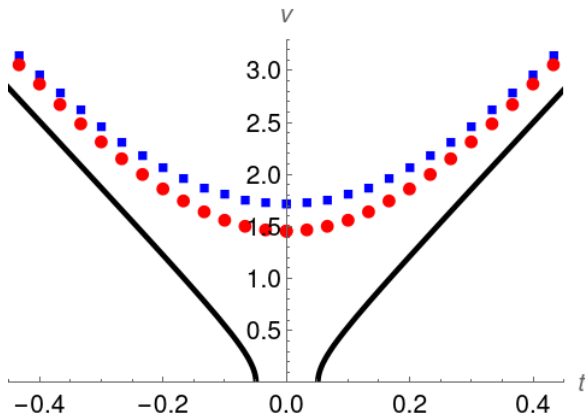


Figure: $\langle v(t) \rangle_{\Psi_{sc}}$ for different semiclassical states

Quantum recollapse

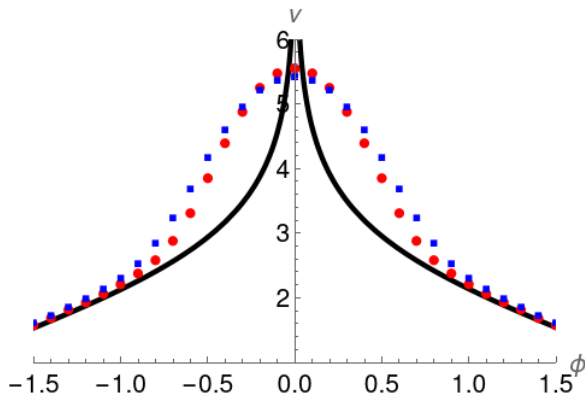


Figure: $\langle v(\varphi) \rangle_{\psi_{sc}}$ for different semiclassical states

A few things to take from this talk

- The problem of time has many nuances
- Quantum cosmology is a good testing ground
- Unitarity requirements are what lead to different theories and,
- They are present in different quantisation schemes
- Singularity resolution is not a feature of the theory, it is a feature of the clock \implies clock choices should be given more attention when studying these models
- Path integral quantisation?



Thank you!

Contact: lucia.menendez-pidal@nottingham.ac.uk

Classical solutions

For $\lambda > 0$ the classical solutions are

$$v(t) = \sqrt{-\frac{\pi_\varphi^2}{\lambda} + 4\lambda(t - t_0)^2}, \quad \varphi(t) = \frac{1}{2} \log \left| \frac{\pi_\varphi - 2\lambda(t - t_0)}{\pi_\varphi + 2\lambda(t - t_0)} \right| + \varphi_0$$

$$v(\varphi) = \frac{|\pi_\varphi|}{\sqrt{\lambda} |\sinh(\varphi - \varphi_0)|}, \quad t(\varphi) = -\frac{\pi_\varphi}{2\lambda} \coth(\varphi - \varphi_0) + t_0$$

$$t(v) = t_0 - \operatorname{sgn}(\pi_v) \frac{1}{2} \sqrt{\frac{v^2}{\lambda} + \frac{\pi_\varphi^2}{\lambda^2}}, \quad \varphi(v) = \varphi_0 + \log \left| \frac{\pi_\varphi}{\sqrt{\lambda} v} + \sqrt{\frac{\pi_\varphi^2}{\lambda v^2} + 1} \right|$$

The big bang/big crunch singularity is at $t_{\text{sing}} = \frac{|\pi_\varphi|}{2\lambda}$, $\log \frac{v_{\text{sing}}}{v_0} = -\infty$,

$\varphi_{\text{sing}} = \pm\infty$

Spatial infinity is at $t_\infty = \pm\infty$, $\log \frac{v_\infty}{v_0} = \infty$, $\varphi_\infty = \varphi_0$

Answer B: Dirac quantisation

Freedom to multiply the WdW by a phase space function

$$\hat{\mathcal{C}}_1 \Psi = 0, \implies \left(\hbar^2 \frac{\partial^2}{\partial v^2} + \frac{\hbar^2}{v} \frac{\partial}{\partial v} - \frac{\hbar^2}{v^2} \frac{\partial^2}{\partial \varphi^2} - i\hbar \frac{\partial}{\partial t} \right) \Psi = 0 \quad (\text{WdW1})$$

$$\hat{\mathcal{C}}_2 \Psi = 0, \implies \left(\hbar^2 \left(v \frac{\partial}{\partial v} \right)^2 - \hbar^2 \frac{\partial^2}{\partial \varphi^2} - i\hbar v^2 \frac{\partial}{\partial t} \right) \Psi = 0 \quad (\text{WdW2})$$

where $\hat{\mathcal{C}}_2 = v^2 \hat{\mathcal{C}}_1$. This motivates different kinematical inner products

$$\langle \Psi | \Phi \rangle_{\text{kin}_1} = \int dt d\varphi dv \, v \bar{\Psi} \Phi, \quad \langle \Psi | \Phi \rangle_{\text{kin}_2} = \int dt d\varphi \frac{dv}{v} \bar{\Psi} \Phi$$

Demanding self-adjointness of $\hat{\mathcal{C}}_1$ and $\hat{\mathcal{C}}_2$ reduces to demanding self-adjointness of $\hat{\mathcal{H}}$ and $\hat{\mathcal{G}}$



Dirac quantisation extra

Demanding \hat{C}_1 and \hat{C}_2 to be self-adjoint with respect to their respective inner products corresponds to self-adjointness of the one dimensional operators

$$\hat{D}_1 = \hbar^2 \left(-\frac{\partial^2}{\partial v^2} - \frac{k^2 + \frac{1}{4}}{v^2} \right), \quad \hat{D}_2 = -\hbar^2 \left(v \frac{\partial}{\partial v} \right)^2 - \lambda v^2$$

Recall,

$$\hat{H} = \left(\hbar^2 \frac{\partial^2}{\partial v^2} + \frac{\hbar^2}{v} \frac{\partial}{\partial v} - \frac{\hbar^2}{v^2} \frac{\partial^2}{\partial \varphi^2} \right), \quad \hat{G} = \left(\hbar^2 \left(v \frac{\partial}{\partial v} \right)^2 - iv^2 \hbar \frac{\partial}{\partial t} \right)$$

Despite being two dimensional, study \hat{H} and \hat{G} reduces to analysing \hat{D}_1 and \hat{D}_2



Allowed wave functions in the different theories

$$\Psi_t \sim \int \frac{d\lambda}{2\pi\hbar} \frac{dk}{2\pi} e^{ik\varphi} e^{i\lambda \frac{t}{\hbar}} \alpha(k, \lambda) \operatorname{Re} \left[e^{i\vartheta(k) - i|k| \log \sqrt{\frac{\lambda}{\lambda_0}}} J_{i|k|} \left(\frac{\sqrt{\lambda}}{\hbar} v \right) \right]$$

$$\Psi_\varphi \sim \int \frac{d\lambda}{2\pi\hbar} \frac{dk}{2\pi} e^{ik\varphi} e^{i\lambda \frac{t}{\hbar}} \alpha(k, \lambda) \operatorname{Re} \left[\sqrt{\frac{\sinh((|k| - i\kappa_0(\lambda))\frac{\pi}{2})}{\sinh((|k| + i\kappa_0(\lambda))\frac{\pi}{2})}} J_{i|k|} \left(\frac{\sqrt{\lambda}}{\hbar} v \right) \right]$$

