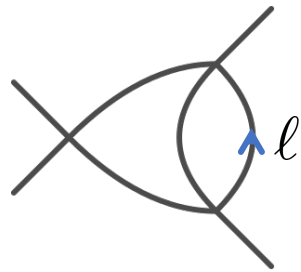


Three things we've learned about infrared divergences this year

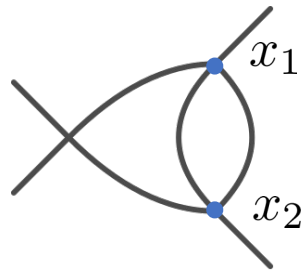
Sebastian Mizera (IAS)

together with Nima Arkani-Hamed (IAS)
and Aaron Hillman (Princeton)

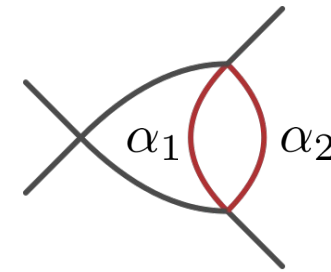
We have multiple ways of thinking about UV divergences



loop momentum
blowing up



space-time points
colliding



worldline segments
shrinking

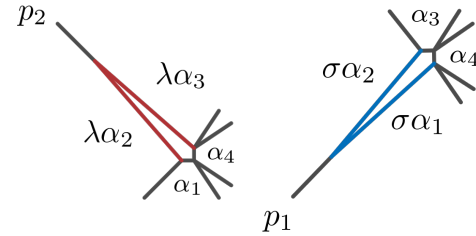


Fascinating connections with algebraic geometry, topology,
and combinatorics, but also new computational tools

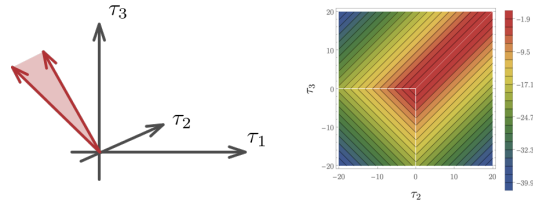
**On the other hand, most of the intuition for IR divergences
comes from the loop momentum space**

Question:
**What is the meaning of IR divergences
in Schwinger parameters?**

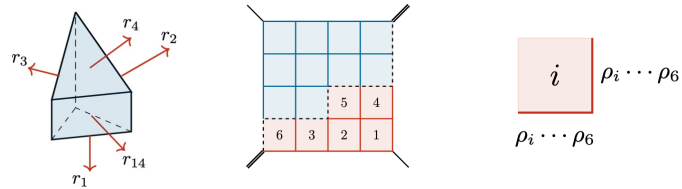
1. Worldline interpretation



2. Divergences as redundancies



3. Computational scheme



Feynman integrals in Schwinger parametrization

(concrete examples later on)

$$\int_0^\infty \frac{d^E \alpha}{\mathcal{U}^{D/2}} \mathcal{N} e^{i\mathcal{F}/\mathcal{U}}$$

Schwinger parameters

Space-time dimension $D = 4 - 2\varepsilon$

(Translating back to the loop momenta: $l_a^\mu = \frac{1}{\mathcal{U}} \sum_{\text{spanning trees } T} p_{T,a}^\mu \prod_{e \notin T} \alpha_e$)

Momentum flowing through edge a along T

The integrand $\frac{d^E \alpha}{\mathcal{U}^{D/2}} \mathcal{N} e^{i\mathcal{F}/\mathcal{U}}$ **features Symanzik polynomials**
 (concrete examples later on)

$$\mathcal{U} = \sum_{\substack{\text{spanning} \\ \text{trees } T}} \prod_{e \notin T} \alpha_e$$

$$\mathcal{F} = \sum_{\substack{\text{spanning} \\ \text{2-trees } T_L \sqcup T_R}} p_{T_L}^2 \prod_{e \notin T_L, T_R} \alpha_e - \mathcal{U} \sum_{e=1}^E m_e^2 \alpha_e$$

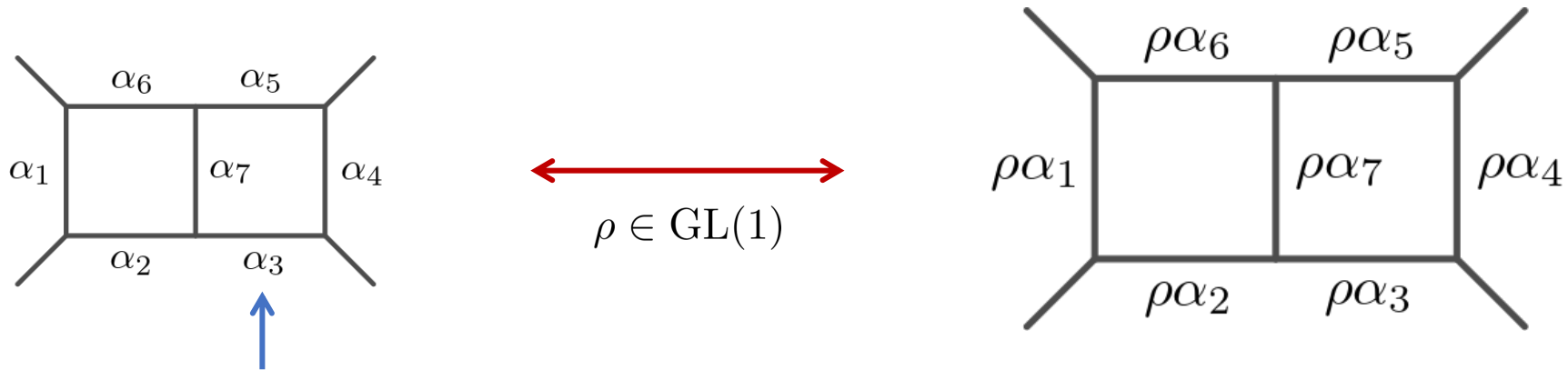
$$\mathcal{N} = \dots = 1$$



In this talk,
 but can be easily included

Overall GL(1) redundancy

(reparametrization invariance)



Schwinger parameters
(\sim lengths of the edges)

Modding out by the GL(1) redundancy

UV divergence factored out, $d = E - LD/2$

$$\int_0^\infty \frac{d^E \alpha}{\mathcal{U}^{D/2}} e^{i\mathcal{F}/\mathcal{U}} = \# \Gamma(d) \int_0^\infty \frac{d^E \alpha}{\underbrace{\text{GL}(1)}_{\text{GL}(1)}} \frac{1}{\mathcal{U}^{D/2-d} \mathcal{F}^d}$$

Integrate out ρ

Fixing the GL(1) gauge, e.g.,

$$\sum_{e=1}^E \alpha_e = 1 \quad \text{“Feynman parameters”}$$

$$\alpha_E = 1 \quad \text{This talk}$$

Big picture:
Singularities as saddle points of the worldline action $\mathcal{V} = \mathcal{F}/\mathcal{U}$

$$\alpha_e \partial_{\alpha_e} \mathcal{V} = 0 \quad \text{for} \quad e = 1, 2, \dots, E$$

How many constraints on the external kinematics?

Codimension-0

(any value of external kinematics)

UV/IR
divergences

[this talk]

Codimension-1

(e.g. $s = 4m^2$)

Normal/anomalous
thresholds

[related talks/posters by
Cordova, Correia, Hannesdottir,
He, Henn, McLeod, Pokraka,
Zhiboedov, ...]

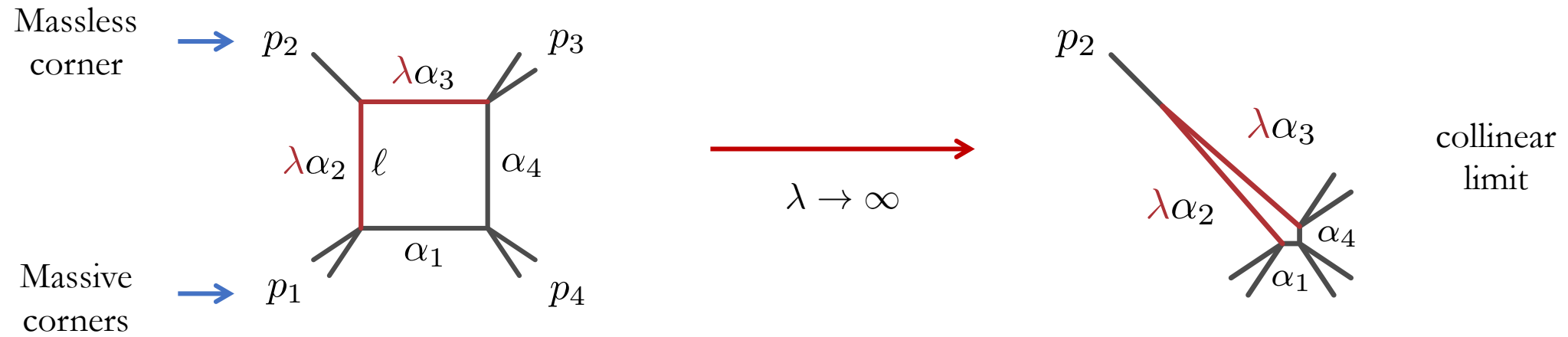
Codimension-2 and higher

(e.g. $s = t = 0$)

Absence puts
interesting constraints

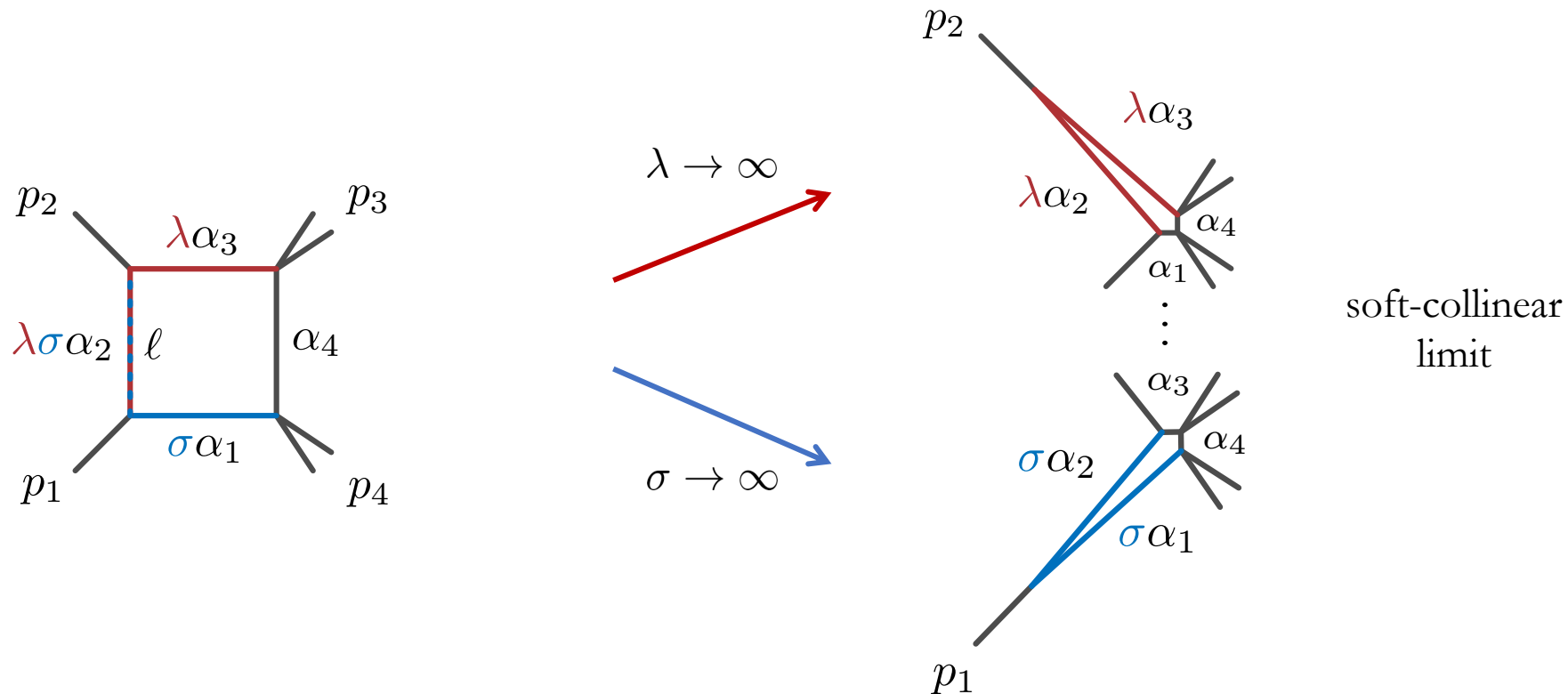
[SM, Telen '21]
[Hannesdottir, SM '22]

Simplest example of an IR divergence



Cross-check:
$$\ell^\mu = -p_2^\mu \frac{\alpha_3}{\alpha_2 + \alpha_3} + \mathcal{O}(\lambda^{-1})$$

Next-to-simplest example of an IR divergence

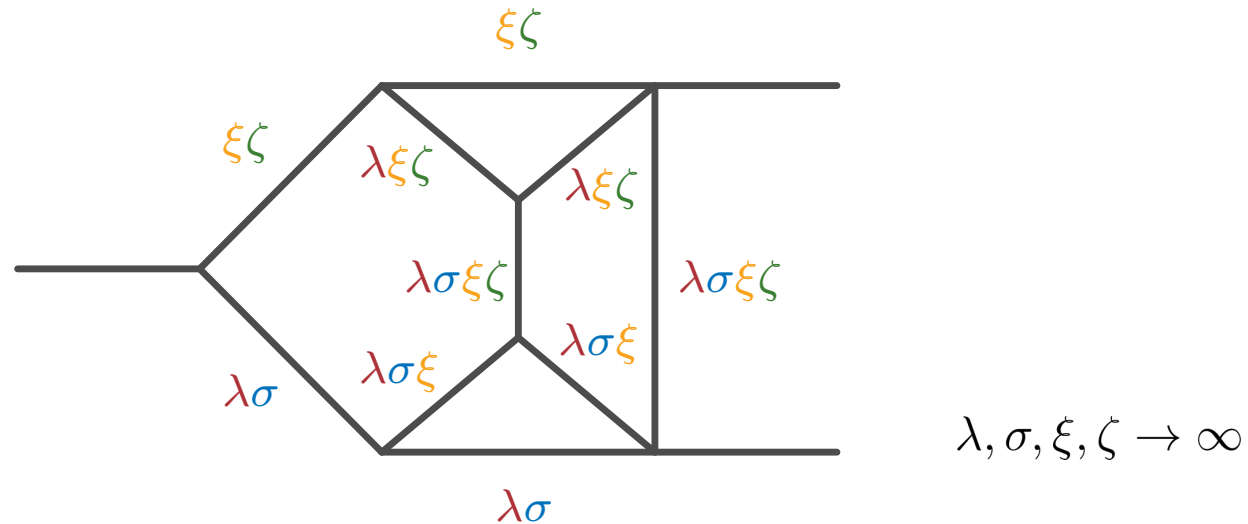


Cross-check:
$$\ell^\mu = \frac{p_1^\mu \sigma \alpha_1 - p_2^\mu \lambda \alpha_3}{\sigma \alpha_1 + \lambda \sigma \alpha_2 + \lambda \alpha_3} + \mathcal{O}(\lambda^{-1}, \sigma^{-1})$$

Related to
[Yellespur '19]

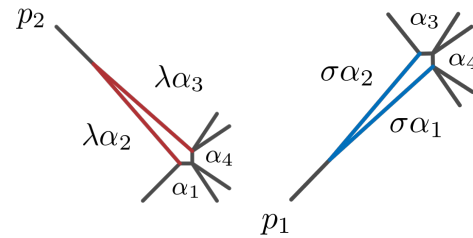
Worldline scalings can become quite involved

(e.g., four-loop contribution to the QCD cusp anomalous dimension)



We'll see that they can be entirely classified for any diagram

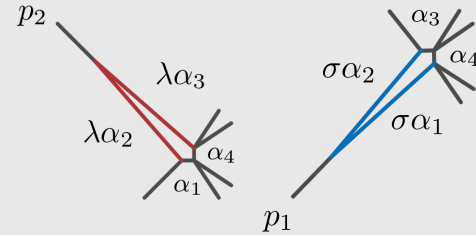
1. Worldline interpretation



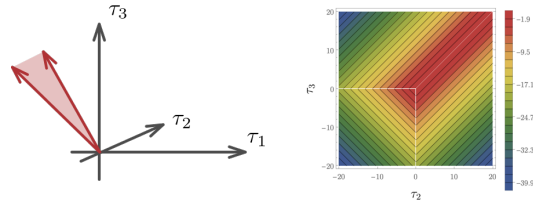
IR divergences come from worldline segments expanding at different rates

(cf. UV divergences)

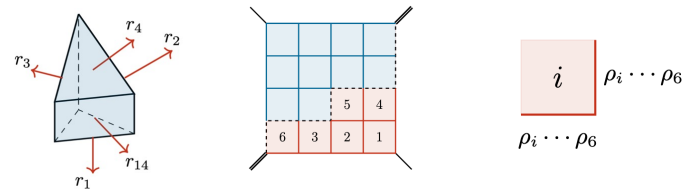
1. Worldline interpretation



2. Divergences as redundancies



3. Computational scheme



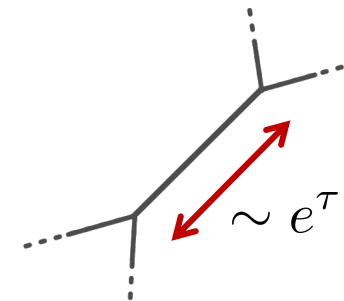
First change the variables

$$\alpha_e = \exp(\tau_e) \quad \text{for every edge } e$$

Integration measure:

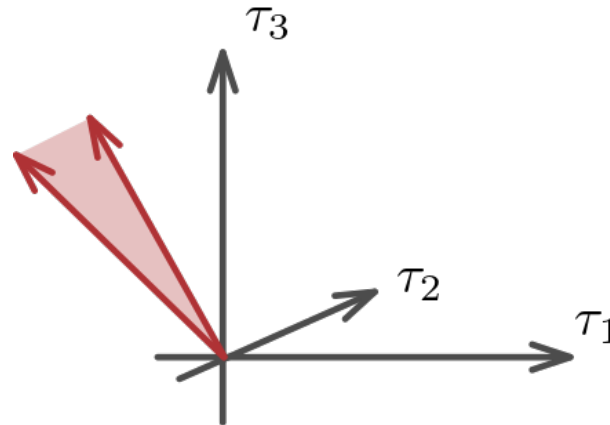
$$\int_0^\infty \frac{d^E \alpha / \prod_{e=1}^E \alpha_e}{\text{GL}(1)} (\dots) = \int_{-\infty}^\infty \frac{d^E \tau}{\text{GL}(1)} (\dots)$$

All divergences at infinity: $\tau_e \rightarrow \begin{cases} +\infty & \text{IR} \\ -\infty & \text{UV} \end{cases}$



UV/IR divergences comes from approaching infinity from different directions

(avoid kinematic singularities with $\tau_e \rightarrow \tau_e + i\varepsilon\partial_{\alpha_e}\mathcal{V}$) [SM '21]



We can linearize this problem using tropical geometry!

(Other applications of tropical geometry: [Tourkine, Cachazo, Early, Guevara, Sepulveda, Borges, Umbert, Panzer, Borinsky, Arkani-Hamed, He, Lam, Spradlin, Salvatori, Lukowski, Parisi, Williams, Drummond, Foster, Gurdogan, Kalousios, Papathanasiou, Henke, Eberhardt, SM, ...])

Tropical approximation of the integrand

$$\frac{d^{\mathbb{E}}\tau}{\text{GL}(1)} \left(\frac{e^{\sum_{e=1}^{\mathbb{E}}\tau_e}}{\mathcal{U}^{D/2-d}\mathcal{F}^d} \right) \approx \frac{d^{\mathbb{E}}\tau}{\text{GL}(1)} e^{\text{Trop}}$$



Exponentially
accurate at infinity

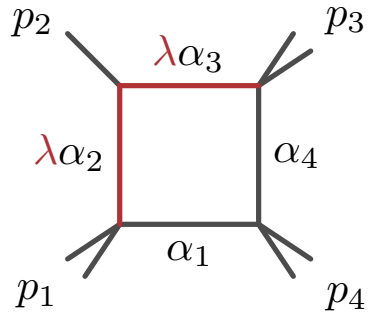
where

$$\text{Trop} = \sum_{e=1}^{\mathbb{E}}\tau_e - (D/2 - d) \max(\mathcal{U}) - d \max(\mathcal{F})$$



Keep only the leading monomials

Example: collinear divergence in D=4

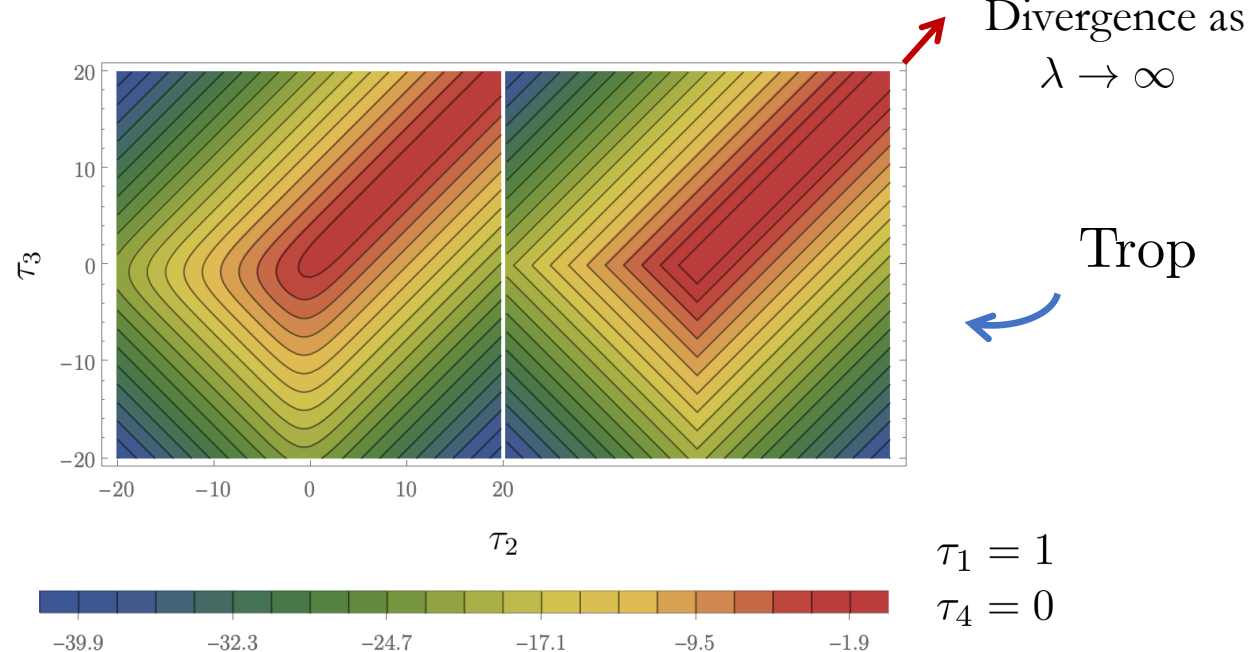


$$\mathcal{U} = e^{\tau_1} + e^{\tau_2} + e^{\tau_3} + e^{\tau_4}$$

$$\mathcal{F} = s e^{\tau_1 + \tau_3} + t e^{\tau_2 + \tau_4} + p_1^2 e^{\tau_1 + \tau_2} + p_3^2 e^{\tau_3 + \tau_4} + p_4^2 e^{\tau_4 + \tau_1}$$

$$\text{Trop} = \tau_1 + \tau_2 + \tau_3 + \tau_4 - 2 \max(\tau_1 + \tau_3, \tau_2 + \tau_4, \tau_1 + \tau_2, \tau_3 + \tau_4, \tau_4 + \tau_1)$$

$$\log \left(\frac{e^{\tau_1 + \tau_2 + \tau_3 + \tau_4}}{\mathcal{U}^0 \mathcal{F}^2} \right)$$



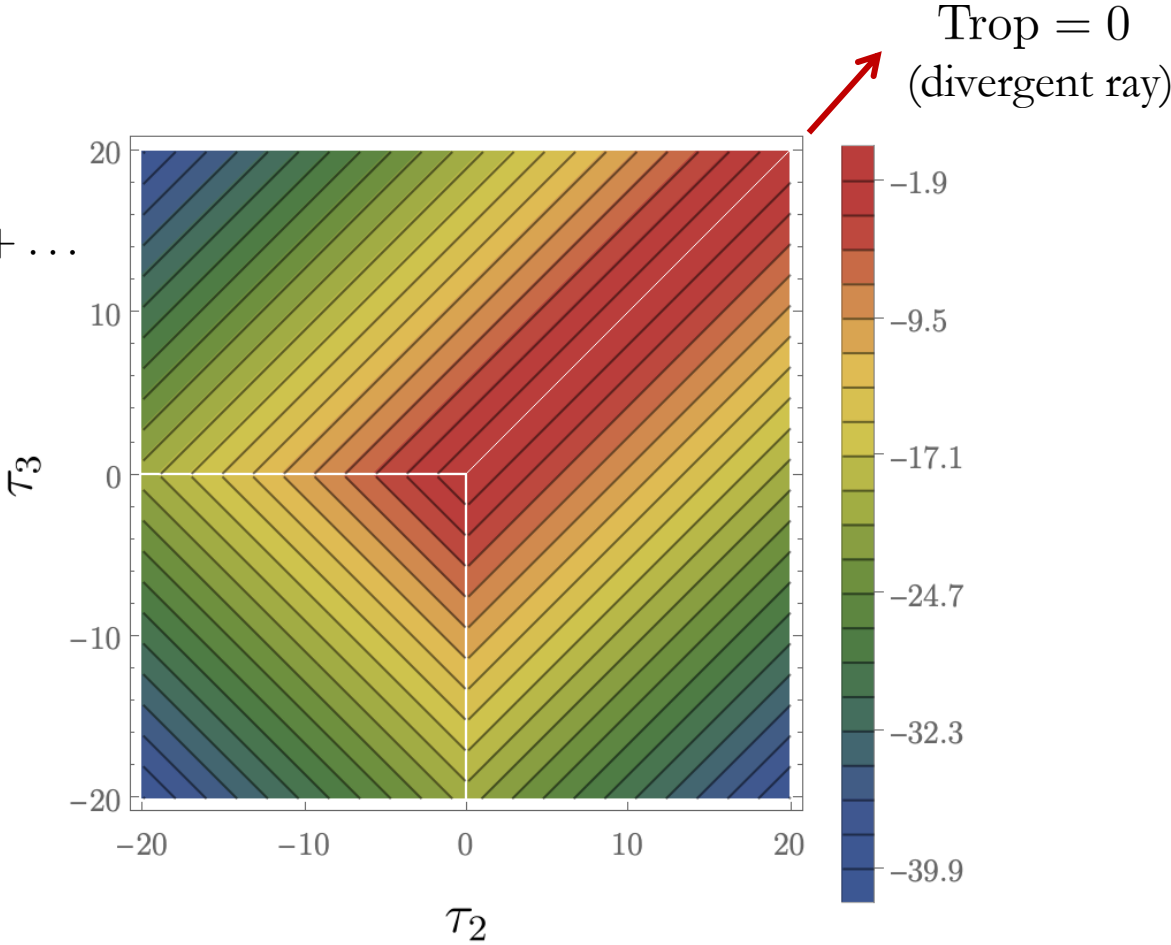
Enhanced symmetry from $GL(1)$ to $GL(1)^2$ \implies divergence

$$\frac{d^4\alpha}{(s\alpha_1\alpha_3 + t\alpha_2\alpha_4 + p_1^2\alpha_1\alpha_2 + p_3^2\alpha_3\alpha_4)^2} + \dots$$

Invariant under

$$GL(1) : \alpha_e \rightarrow \rho\alpha_e$$

$$GL(1) : (\alpha_2, \alpha_3) \rightarrow \lambda(\alpha_2, \alpha_3)$$

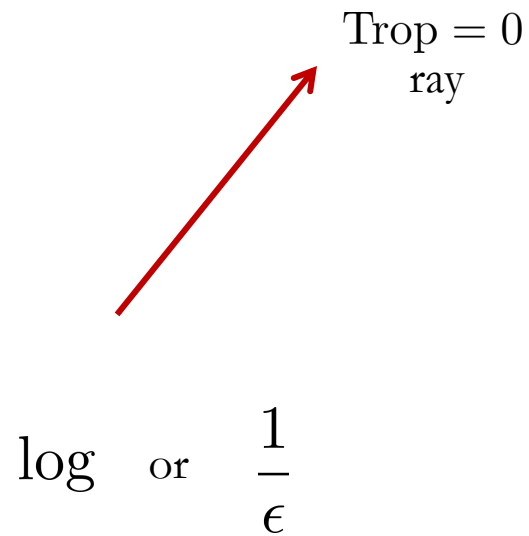


What kind of a divergence?

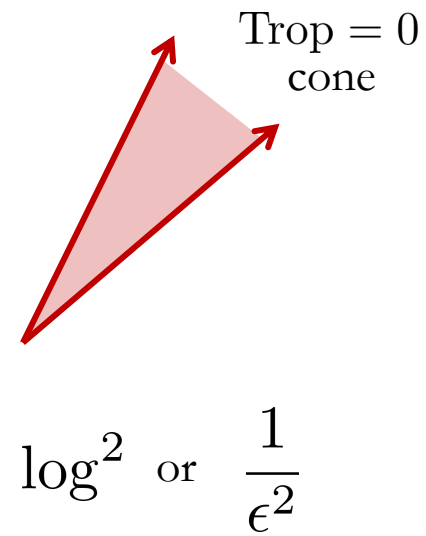
$$\int_0^\infty \frac{d^E \alpha / \prod_{e=1}^E \alpha_e}{\text{GL}(1)} e^{\text{Trop}}$$

$$\text{Trop} \rightarrow \begin{cases} > 0 & \text{power-law} \\ = 0 & \text{logarithmic} \\ < 0 & \text{finite} \end{cases}$$

How bad of a divergence?

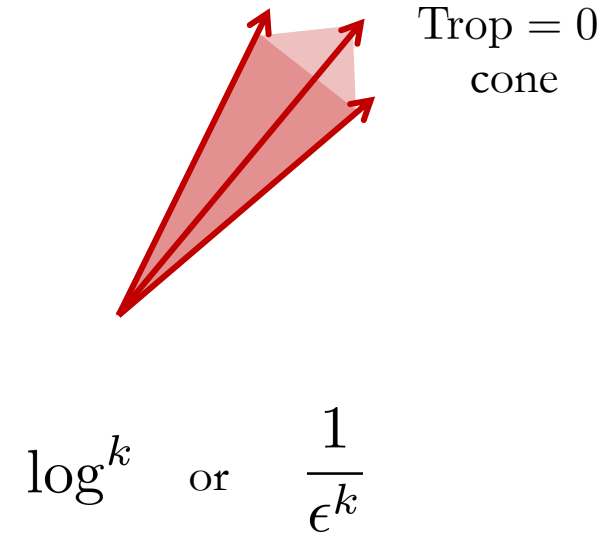


e.g., collinear



e.g., soft-collinear

...



generalizations
 $k \leq 2L$

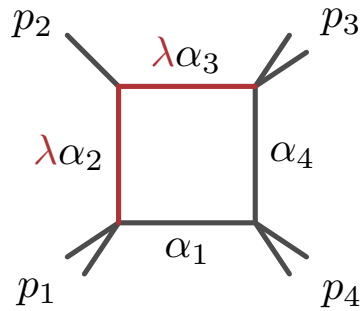
Modding out by the enhanced $GL(1)^{k+1}$ redundancy

(specialize to logarithmic from now on)

$$\Gamma(d) \int_0^\infty \frac{d^E \alpha}{GL(1)} \frac{1}{\mathcal{U}^{D/2-d} \mathcal{F}^d} = \Gamma(d) \sum_{\text{divergent cones } C} \text{vol}(C) \int_0^\infty \frac{d^E \alpha}{GL(1)^{k+1}} \frac{1}{\mathcal{U}_C^{D/2-d} \mathcal{F}_C^d} + \mathcal{O}(\epsilon^{-k+1})$$

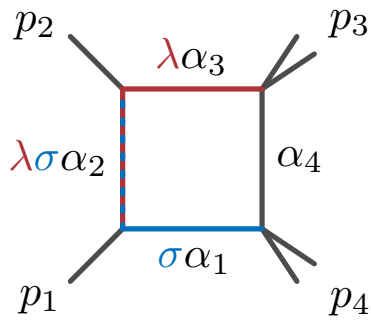
Volume of the cone $\propto \frac{1}{\epsilon^k}$
↓
↑ Fix the $GL(1)^{k+1}$ redundancy ↑ Simplified Symanzik polynomials
↑

Example: collinear divergence in $D=4-2\epsilon$



$$\begin{aligned}
 & \text{Fix } \alpha_3 = \alpha_4 = 1 && \text{Set } \epsilon = 0 \\
 & \text{---} && \text{---} \\
 & = -\frac{1}{\epsilon} \int \frac{d^4\alpha}{\text{GL}(1)^2} \frac{(\alpha_1 + \alpha_2 + \alpha_3 + \alpha_4)^{2\epsilon}}{(s\alpha_1\alpha_3 + t\alpha_2\alpha_4 + p_1^2\alpha_1\alpha_2 + p_3^2\alpha_3\alpha_4 + \cancel{p_4^2\alpha_4\alpha_1})^{2-\epsilon}} + \mathcal{O}(\epsilon^0) \\
 & = -\frac{1}{\epsilon} \int_0^\infty \frac{d\alpha_1 d\alpha_2}{(s\alpha_1 + t\alpha_2 + p_1^2\alpha_1\alpha_2 + p_3^2)^2} + \mathcal{O}(\epsilon^0) \\
 & = -\frac{1}{\epsilon} \frac{\log\left(\frac{st}{p_1^2 p_3^2}\right)}{st - p_1^2 p_3^2} + \mathcal{O}(\epsilon^0)
 \end{aligned}$$

Example: soft-collinear divergence in $D=4-2\epsilon$



$$\begin{aligned}
 & \text{Fix } \alpha_2 = \alpha_3 = \alpha_4 = 1 && \text{Set } \epsilon = 0 \\
 & = \frac{1}{\epsilon^2} \int \frac{d^4\alpha}{\text{GL}(1)^3} \frac{(\alpha_1 + \alpha_2 + \alpha_3 + \alpha_4)^{2\epsilon}}{(s\alpha_1\alpha_3 + t\alpha_2\alpha_4 + \cancel{p_3^2\alpha_3\alpha_4} + \cancel{p_4^2\alpha_4\alpha_1})^{2-\epsilon}} + \mathcal{O}(\epsilon^{-1}) \\
 & = \frac{1}{\epsilon^2} \int_0^\infty \frac{d\alpha_1}{(s\alpha_1 + t)^2} + \mathcal{O}(\epsilon^{-1}) \\
 & = \frac{1}{\epsilon^2} \frac{1}{st} + \mathcal{O}(\epsilon^{-1})
 \end{aligned}$$

Example: cusp anomalous dimension at four loops

$$\begin{aligned}
 &= \frac{1}{144\epsilon^4} \int \frac{d^{10}\alpha}{\text{GL}(1)^5} \frac{1}{s^2} [(\alpha_1\alpha_5 + \alpha_3\alpha_5 + \alpha_1\alpha_7) (\alpha_2\alpha_6 + \alpha_4\alpha_6 + \alpha_2\alpha_8) \alpha_9 \\
 &\quad + (\alpha_1 + \alpha_3) (\alpha_2 + \alpha_4) (\alpha_5 + \alpha_7) (\alpha_6 + \alpha_8) \alpha_{10}]^{-2} + \mathcal{O}(\epsilon^{-3}) \\
 &= \frac{\pi^4}{5184s^2\epsilon^4} + \mathcal{O}(\epsilon^{-3})
 \end{aligned}$$

In agreement with [\[Henn, Smirnov, Smirnov, Steinhauser '16-20\]](#)

Rich literature on closely related topics

- Sector decomposition & expansion by regions

[Binoth, Heinrich, Bogner, Weinzierl, Kaneko, Ueda, Borowka, Jones, Kerner, Schlenk, Zirke, Jahn, Langer, Magerya, Poldaru, Villa, Pak, Smirnov, Smirnov, Ananthanarayan, Ramanan, Sarkar, Semenova, ...]

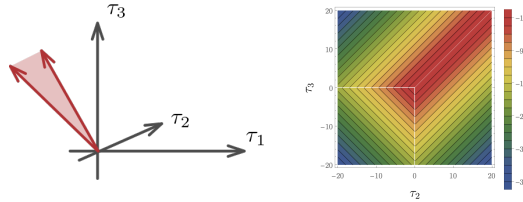
- Feynman integrals as GKZ hypergeometric functions

[enormous literature in the 20th century,
de la Cruz, Klausen, Tellander, Helmer, Feng, Chang, Chen, Zhang,
Abreu, Britto, Duhr, Gardi, Matthew, Mastrolia, Telen, SM, ...]

[poster by Datta]

We want to exploit the extra combinatorics of Symanzik polynomials

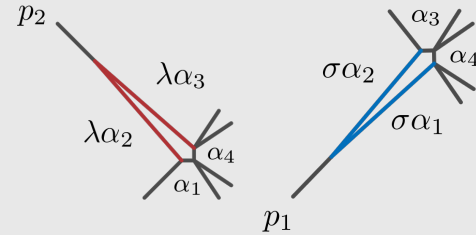
2. Divergences as redundancies



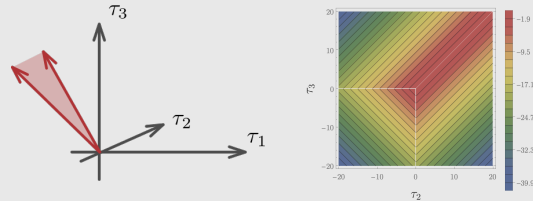
**Divergence is a $GL(1)^{k+1}$ redundancy
computed by fixing $k+1$ Schwinger parameters**

(cf. fixing gauge in gauge theory)

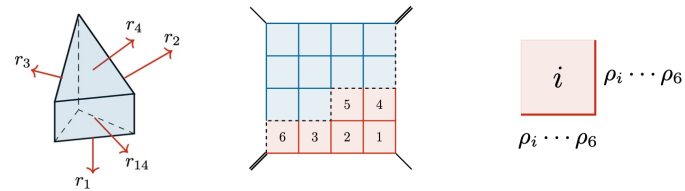
1. Worldline interpretation



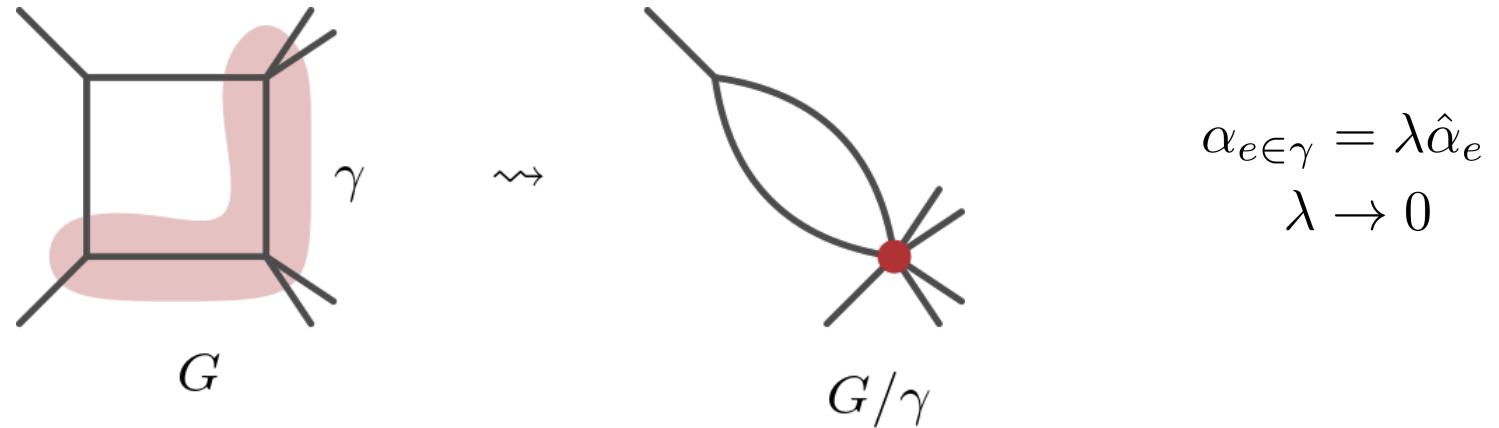
2. Divergences as redundancies



3. Computational scheme

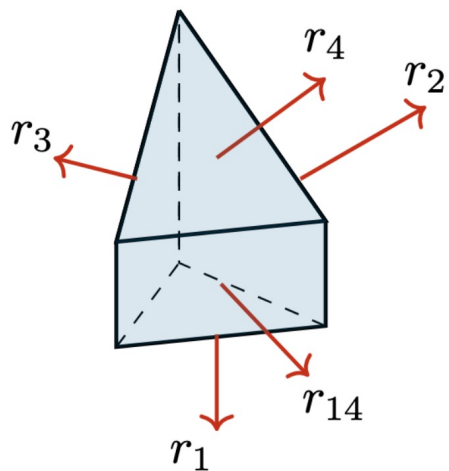


No need to do it “by hand” for every Feynman diagram



Rays generated by subdiagrams γ

Feynman polytopes



- Where are the divergences?
- What type are they?
- How do they nest/overlap?

$$\text{Trop}(\gamma) = \begin{cases} -d_\gamma & \text{if } \mathcal{F}_{G/\gamma} \neq 0 \\ d_{G/\gamma} & \text{if } \mathcal{F}_{G/\gamma} = 0 \end{cases}$$

UV-like
IR-like

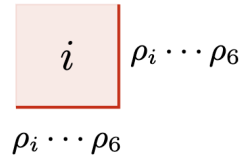
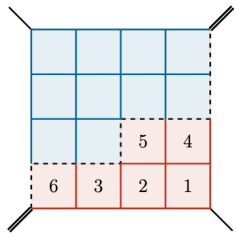
Previously studied in the UV case with generic kinematics

- Blow-ups of the Schwinger parameter space
[Bloch, Esnault, Kreimer, Brown, Schultka '05-19]
- Numerical computations of finite integrals
[Panzer, Borinsky '19-20]

IR singularities need non-generic kinematics
(Feynman polytopes are in general *not* generalized permutohedra)

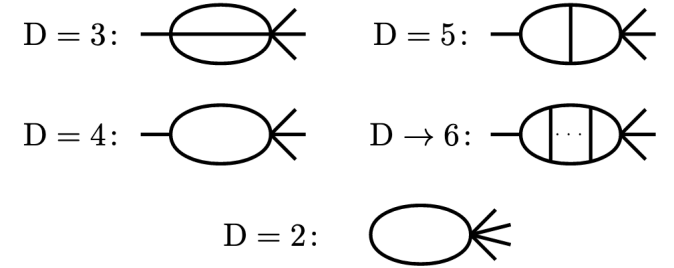
Lots of applications to IR divergences, including all-loop results

(details in [\[Arkani-Hamed, Hillman, SM '22\]](#))



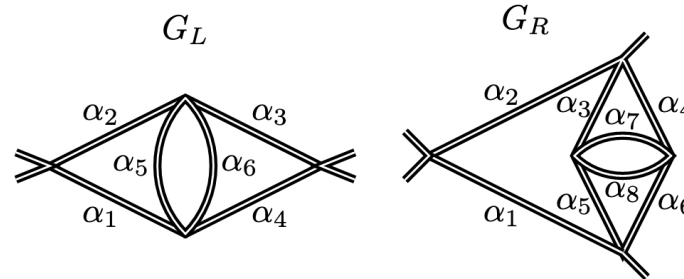
$$\partial \mathbf{F}_G \supset \begin{cases} \mathbf{U}_\gamma \times \mathbf{F}_{G/\gamma} & \text{on } a_\gamma = L_\gamma, \\ \mathbf{F}_\gamma \times \mathbf{U}_{G/\gamma} & \text{on } a_\gamma = L_\gamma + 1. \end{cases}$$

$$(1+c)(L_{\gamma_1} + L_{\gamma_2}) + (c)_{\gamma_1} + (c)_{\gamma_2} \geq (1+c)(L_{\gamma_1 \cup \gamma_2} + L_{\gamma_1 \cap \gamma_2}) + (c)_{\gamma_1 \cup \gamma_2} + (c)_{\gamma_1 \cap \gamma_2}.$$

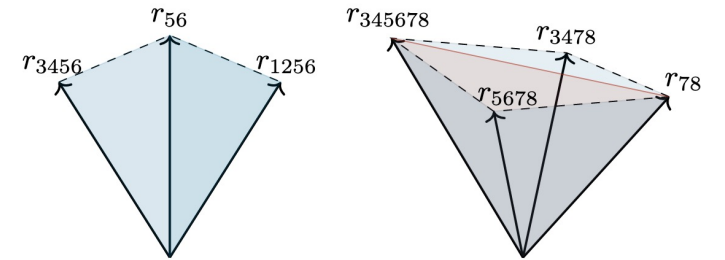


$$M \left\{ \begin{array}{c} N \\ \text{grid diagram} \end{array} \right\} = \frac{2^{MN}}{(MN)!} \frac{C_{M,N}}{\varepsilon^{MN} s^N t^M} + \dots,$$

$$M \left\{ \begin{array}{c} N \\ \text{grid diagram} \end{array} \right\} = \frac{F_{M,N}(s, t, p_i^2)}{(MN)!} \frac{C_{M,N}}{\varepsilon^{MN}} + \dots,$$



$$C_{M,N} = (MN)! \frac{\text{sf}(M-1)\text{sf}(N-1)}{\text{sf}(M+N-1)}$$



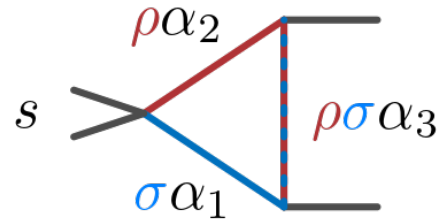
$$\text{sf}(N) = 0!1!2! \dots N!$$

If there's time:

**Glimpse at the Laurent expansion
in dimensional regularization, $D=4-2\epsilon$**

(inspired by [\[Brown, Kreimer '11\]](#) in BPHZ renormalization
and [\[Brown, Dupont '19\]](#) in string theory amplitudes)

Simplest example

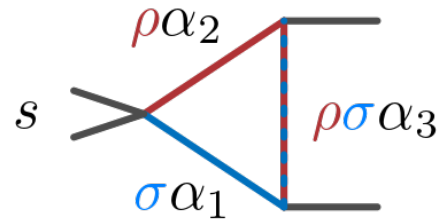


$$= \frac{\Gamma(1+\epsilon)}{(-s)^{1+\epsilon}} \left\{ \frac{1}{\epsilon^2} + \int \frac{d^3\alpha}{\text{GL}(1)} \left[\left(\frac{\alpha_1}{\alpha_1+\alpha_2+\alpha_3} \right)^{-\epsilon} d \log \left(\frac{\alpha_1}{\alpha_1+\alpha_2+\alpha_3} \right) - \left(\frac{\alpha_1}{\alpha_1+\alpha_3} \right)^{-\epsilon} d \log \left(\frac{\alpha_1}{\alpha_1+\alpha_3} \right) \right] \right. \\ \left. \wedge \left[\left(\frac{\alpha_2}{\alpha_1+\alpha_2+\alpha_3} \right)^{-\epsilon} d \log \left(\frac{\alpha_2}{\alpha_1+\alpha_2+\alpha_3} \right) - \left(\frac{\alpha_2}{\alpha_2+\alpha_3} \right)^{-\epsilon} d \log \left(\frac{\alpha_2}{\alpha_2+\alpha_3} \right) \right] \right\}$$



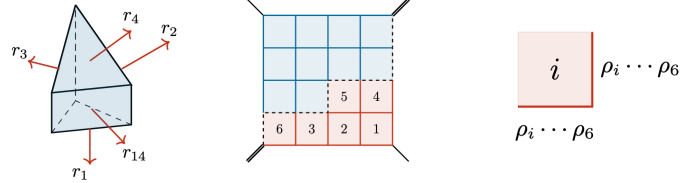
All integrals manifestly finite

Series[0, {ϵ, 0, 0}]



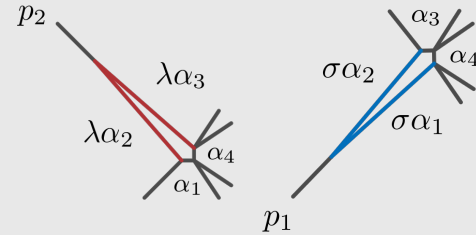
$$= \frac{\Gamma(1+\epsilon)}{(-s)^{1+\epsilon}} \left\{ \frac{1}{\epsilon^2} + \underbrace{\int_0^\infty d \log \left(\frac{\alpha_1+1}{\alpha_1+\alpha_2+1} \right) \wedge d \log \left(\frac{\alpha_2+1}{\alpha_1+\alpha_2+1} \right)}_{-\zeta_2} + \mathcal{O}(\epsilon) \right\}$$

3. Computational scheme

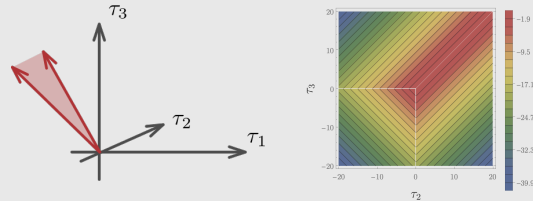


Combinatorics of UV/IR divergences is summarized by Feynman polytopes

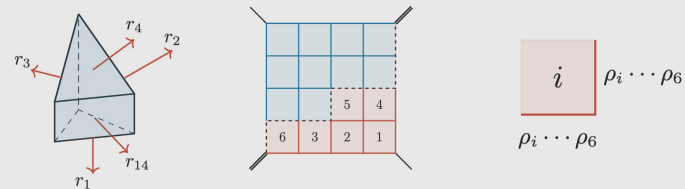
1. Worldline interpretation



2. Divergences as redundancies



3. Computational scheme



Thank you!