

A Tale of Tails via Generalized Unitarity

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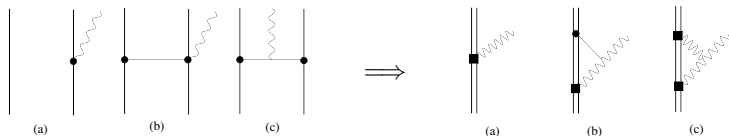


A Tale of Tails

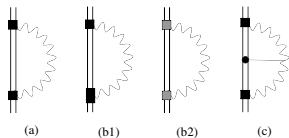
arXiv: [2202.04674](https://arxiv.org/abs/2202.04674)
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What We Are After...



- Hereditary effects \equiv Scattering of radiation from binary off its own gravitational background
- Radiation-reaction starts at 2.5PN [$n\text{PN} \equiv (v/c)^{2n}\text{PN}$]
- Scattering starts at 4PN – affecting BOTH **conservative** dynamics and **dissipative** observables [4PN = current complete precision frontier]



Radiation-Reaction POV

Traditional GR

Trail of Tails

- Studied for decades. 1st paper in PN theory – *Blanchet & Damour* 1988
- Steady progress along the years led by *Luc Blanchet*, e.g. LRR 2014
- Latest milestone: Tail-of-Tail-of-Tail ("TTT") *Blanchet+* 2016
arXiv:1607.07601 dissipation at $+4.5(=7\text{PN})$ though $+4\text{PN}$ incomplete
- Some overlap with simple Self-Force theory results

Challenges/Strengths

- Laborious, slow, derivation "very" gauge-dependent...
- CONSISTENT & SOLID
- Handles separately either dissipative or conservative sides
- Output of direct use to LIGO+ (even mostly exclusive)

QFT Toolbox to the Rescue!

Goldberger & Ross 2009

- Formulated EFT for radiative sector, from *Goldberger & Rothstein* 2004
- Clear separation of scales and standard Feynman toolbox
- Background field method to handle mixing of field modes
- Targets only dissipative flux – but reproduced then state-of-the-art (TT)
- Significant milestone: Classical RG flow of quadrupole moment!

Galley+ 2009–17

- Time reversal is broken – invoked Closed-Time-Path (CTP) formalism!
- Defined proper conservative AND dissipative contributions
- Higher loops \leftrightarrow Fewer graphs
- Reached only Tail level
- Several follow-ups by Foffa-Sturani+, Blumlein+

Our Mission

Goal: Set up Amplitudes-like framework to treat hereditary effects

Our Criteria:

- 1 Produce output for post-Newtonian binary inspirals – of use to LIGO+
- 2 Capture entirety of effects – both conservative + dissipative sides
- 3 Efficient & fast framework + machinery
- 4 Natural to generalize to all types of scattering + all PN orders

Other work in Amplitudes of different context and reach:

- Di-Vecchia+ arXiv:2101.05772, arXiv:2104.03256
- Bern+ arXiv:2112.10750
- Dlapa+ arXiv:2112.11296
- Jakobsen+ arXiv:2207.00569

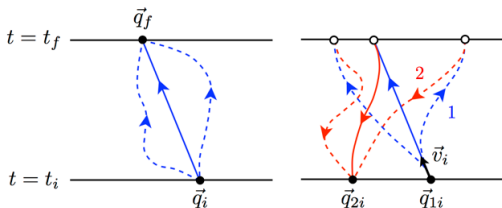
CTP 101

Closed-Time Path (CTP) Formalism

- Time reversal is broken
→ Initial-value problem ("in-in") instead of boundary-value problem (in-out)
- Non-equilibrium QFT – Schwinger/Keldysh formalism 1961/1964/5
- Degrees of freedom are formally doubled, and the action becomes:

$$S_{\text{CTP}}[\{\}_1, \{\}_2] \equiv S[\{\}_1] - S^*[\{\}_2]$$

- Worked out for classical context by Galley 2012, Galley+ arXiv:1412.3082



From Galley, arXiv:1210.2745, PRL 2013

CTP in Composite-Particle EFT

CTP in Worldline EFT

- Switch (from $\{1,2\}$) to $\{+,-\}$ basis
- For generic fields: $\phi_+ \equiv (\phi_1 + \phi_2)/2$, $\phi_- \equiv \phi_1 - \phi_2$
- For discrete DOFs: $q_+ \equiv (q_1 + q_2)/2$, $q_- \equiv q_1 - q_2$
- Classical propagator-matrix is off-diagonal

$$G_{+-} = \begin{pmatrix} 0 & G_{\text{adv}} \\ G_{\text{ret}} & 0 \end{pmatrix}$$

- Retarded and advanced propagators

$$G_{\text{ret}/\text{adv}}(x - x') = \int \frac{d^D p}{(2\pi)^D} \frac{e^{-ip_\mu(x-x')^\mu}}{(p_0 \pm i\epsilon)^2 - \vec{p}^2}$$

$+i\epsilon$ for retarded, $D \equiv d + 1$ with d for number of spatial dimensions

Integrating Out via Generalized Unitarity (GU)

Starting from

$$S_{\text{eff}(c)}[g_{\mu\nu}, y_c^\mu, e_{cA}^\mu] = -\frac{1}{16\pi G} \int d^4x \sqrt{g} R[g_{\mu\nu}] + S_{\text{pp}(c)}[g_{\mu\nu}(y_c), y_c^\mu, e_{cA}^\mu](\sigma_c)$$

with

$$S_{\text{pp}(c)}[h_{\mu\nu}, y_c^\mu, e_{cA}^\mu](t) = - \int dt \sqrt{g_{00}} \left[E(t) - \sum_{l=2}^{\infty} \frac{1}{l!} I^L(t) \nabla_{L-2} E_{i_{l-1}i_l} + \dots \right]$$

we integrate out all field modes to get an effective theory of dynamical multipoles.

Integrating out via GU (instead of Feynman)

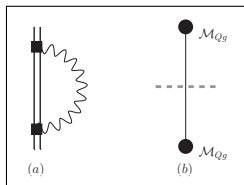
Multi-loop tools [Smirnov] + Generalized Unitarity:

Bern, Dixon, Dunabr, Kosower 1994, Britto, Cachazo, Feng 2004,...

$$S_{\text{eff}} = \int \frac{d\omega}{2\pi} \sum_{i \in \text{MI}} c_i(d, \omega) \mathcal{I}_i$$

$\{\mathcal{I}_i\}$ complete set of master integrals that span problem

Radiation Reaction



EFT and Amplitudes POVs

$$\begin{aligned}
 F^{(1)}(\lambda; \omega^2) &= \int \frac{d^d \ell_E}{(2\pi)^d} \frac{1}{(-\ell_E^2 + \omega^2)^\lambda} \\
 &= R(d, \lambda, \omega^{2n(\lambda)}) F^{(1)}(1; \omega^2)
 \end{aligned}$$

$$S_{RR} = \int \frac{d\omega}{2\pi} c_{RR}(\omega) F^{(1)}(1; \omega^2)$$

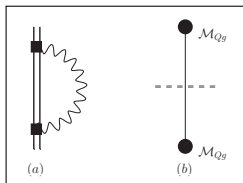
- Cuts assembled from tree amplitudes by public code `IncreasingTrees` [Edison+Teng 2020] via graviton state-sewing $\sum_{\text{states}} \varepsilon_k^{\mu\nu} \varepsilon_k^{\alpha\beta*} \equiv \mathcal{P}_k^{\mu\nu; \alpha\beta}$
- $\lambda_Q \equiv \sqrt{2\pi G_N}$

$$\begin{aligned}
 \mathcal{M}_{Qg} &\equiv \lambda_Q J^{\mu\nu} \varepsilon_{\mu\nu} \equiv \lambda_Q J^{\mu\nu} \varepsilon_\mu \varepsilon_\nu = -\lambda_Q I^{ab} \\
 &\times (k_0 k_a \varepsilon_0 \varepsilon_b + k_0 k_b \varepsilon_0 \varepsilon_a - k_a k_b \varepsilon_0 \varepsilon_0 - k_0 k_0 \varepsilon_a \varepsilon_b)
 \end{aligned}$$

■

$$\begin{aligned}
 c_{RR} &= \lambda_Q^2 J_1^{\mu\nu} \mathcal{P}^{\mu\nu; \alpha\beta} J_2^{\alpha\beta} \Big|_{P_\ell = \ell_E^2 - \omega^2 = 0} \\
 &= \delta(\ell_E^2 - \omega^2) \lambda_Q^2 \left(J_1^{\mu\nu} J_2^{\mu\nu} - \frac{J_1^{\mu\mu} J_2^{\nu\nu}}{d-1} \right)
 \end{aligned}$$

Radiation Reaction



$$C_{RR} = \delta(P_\ell) \lambda_Q^2 \frac{(d+1)(d-2)}{(d+2)(d-1)} \omega^4 \kappa_{ab}(\omega)$$

with trace of CTP quadrupole DOFs

$$\kappa_{ab}(\omega) = I_a^{ij}(-\omega) I_{ij,b}(\omega)$$

with $a, b \in \{+, -\}$

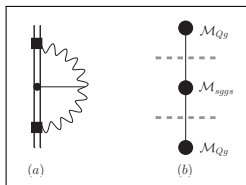
$$S_{RR} = \frac{2\pi G_N}{5} \int \frac{d\omega}{2\pi} \omega^4 \sum_{a,b \in \{+,-\}} \kappa_{ab}(\omega) F^{(1)}(1_{ab})$$

with retarded and advanced propagators $F^{(1)}(1_{-+}/1_{+-}) \equiv F^{(1)}(1; (\omega \pm i\epsilon)^2)$

$$S_{RR} = -i \frac{G_N}{5} \int_{-\infty}^{\infty} \frac{d\omega}{2\pi} \omega^5 I_-^{ij}(-\omega) I_{+,ij}(\omega)$$

■ ϵ_d contributions will matter for higher-order renormalization

Tail



3 momenta invariants but plugging into FIRE:

$$\begin{aligned}
 F^{(2)}(1_X, 1_Y, 0) &= \int \frac{d^d \ell_1 d^d \ell_2}{(2\pi)^{2d}} \frac{1}{(-\ell_1^2 + \omega_X^2)(-\ell_2^2 + \omega_Y^2)} \\
 &= F^{(1)}(1; \omega_X^2) F^{(1)}(1; \omega_Y^2)
 \end{aligned}$$

■

$$S_T = \int \frac{d\omega}{2\pi} c_T(\omega) F^{(2)}(1_X, 1_Y, 0)$$

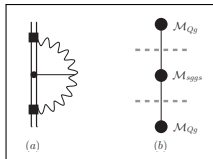
■ Now we need an amplitude of 2 massive scalars and 2 gravitons.

It is extracted from IncreasingTrees, and expanded in the large-mass limit:

$$\begin{aligned}
 \mathcal{M}_{sggs}(m_s \rightarrow \infty) &= \frac{\lambda_g \lambda_E}{\omega_{k_2}^2} \frac{1}{2(k_2^\mu k_{3,\mu})} \left[(k_2^\mu k_{3,\mu}) \varepsilon_2^0 \varepsilon_3^0 \right. \\
 &\quad + \omega_{k_2} ((\varepsilon_3^\mu k_{2,\mu}) \varepsilon_2^0 - (\varepsilon_2^\mu k_{3,\mu}) \varepsilon_3^0) \\
 &\quad \left. - \omega_{k_2}^2 (\varepsilon_2^\mu \varepsilon_{3,\mu}) \right]^2 + \mathcal{O}(m_s^{-1})
 \end{aligned}$$

■ $\mathcal{M}_{sgs} \equiv \lambda_E (p^\mu p^\nu / m_s^2) \varepsilon^\mu \varepsilon^\nu$ fixes $\lambda_E \equiv -\sqrt{8\pi G_N} E$, similarly $\lambda_g \equiv -\sqrt{32\pi G_N} E$

Tail



$$\begin{aligned}
 C_{1,1,0}^{(2)} &= \sum_{\text{states}} \mathcal{M}_{Qg}(-\omega) \mathcal{M}_{sggs} \mathcal{M}_{Qg}(\omega) \Big|_{P_{\ell_1}=0, P_{\ell_2}=0} \\
 &= \lambda_Q^2 \delta(P_{\ell_1}) \delta(P_{\ell_2}) \\
 &\quad \times J_{I(-\omega)}^{\mu\nu} P^{\mu\nu;\alpha\beta} \mathcal{M}_{sggs}^{\alpha\beta;\gamma\sigma} P^{\gamma\sigma;\rho\tau} J_{I(\omega)}^{\rho\tau} \Big|_{m_s \rightarrow \infty}
 \end{aligned}$$

- Reducing again and evaluating integrals with appropriate CTP prescriptions:

$$S_T = \frac{2}{5} G_N^2 E \int \frac{d\omega}{2\pi} \omega^6 \kappa_{-+}(\omega) \left[\frac{1}{\epsilon_d} + \log\left(\frac{\omega^2}{\mu^2}\right) - i\pi \operatorname{sgn}(\omega) \right]$$

- In general, poles will cancel out from renormalization and lower-order zeros

Derivation of Physics

Conservative Dynamics

From the causal effective actions radiation-reaction forces are obtained by varying wrt CTP DOFs, $\{q_{\pm}^i\}$, then taking the physical limit, $q_+^i \rightarrow q^i$ and $q_-^i \rightarrow 0$.

Dissipative Observables

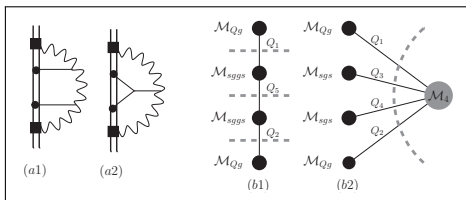
- No proper derivation of dissipative observables from CTP effective actions existed in the literature.
- We worked out generic derivation starting from generalized Noether theorem:

$$\frac{dE}{dt} = -\frac{\partial L}{\partial t} + \dot{q}^i \left[\frac{\partial K}{\partial q_-^i} \right]_{\text{PL}} + \ddot{q}^i \left[\frac{\partial K}{\partial \dot{q}_-^i} \right]_{\text{PL}} + \dots$$

- L conservative potential = one of time histories
- K non-conservative potential = 2 time histories CANNOT be separated
- CTP quadrupoles as generalized DOFs

$$\int dt \frac{dE}{dt} = \Delta E = \int d\omega \frac{dE}{d\omega}$$

Tail-of-Tail (TT)



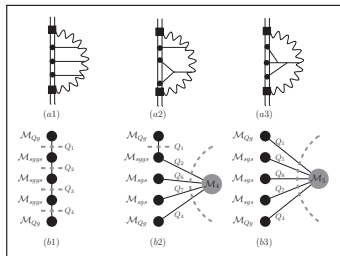
- After reduction of the integral family:

$$S_{\text{TT}} = \int \frac{d\omega}{2\pi} \left[c_1(\omega) F^{(3)}(1, 1, 0, 0, 1, 0) + c_2(\omega) F^{(3)}(1, 1, 1, 1, 0, 0) \right]$$

- We need the 4-graviton amplitude but no special kinematics needed there
- No effective action previously derived

$$S_{\text{TT}} = \frac{107}{175} G_N^3 E^2 \int \frac{d\omega}{2\pi} \omega^7 \kappa_{-+}(\omega) \left[\pi \operatorname{sgn}(\omega) + i \left[\frac{2}{3\epsilon_d} + \log \left(\frac{\omega^2}{\mu_1^2} \right) \right] \right]$$

Tail-of-Tail-of-Tail (TTT)



$$S_{\text{TTT}} = \int \frac{d\omega}{2\pi} \left[c_1 F_{1,1,1,1,0,0,0,0,0,0}^{(4)} + c_2 F_{1,0,0,1,1,1,1,0,0,0}^{(4)} \right. \\ \left. + (c_3 F_{1,0,1,1,1,1,0,0,0,0}^{(4)} + c_4 F_{1,1,0,1,0,1,1,0,0,0}^{(4)}) \right]$$

- No effective action previously derived
- Result is quite lengthy with ζ_2 and $\zeta_3 \equiv \text{Apéry's constant}$

Dissipative Observables

Power Spectra

With $\kappa(\omega) \equiv I^{ij}(-\omega)I_{ij}(\omega)$

$$\frac{dE_{RR}}{d\omega} = -\frac{G_N}{5\pi} \omega^6 \kappa(\omega)$$

$$\frac{dE_T}{d\omega} = -\frac{2}{5} G_N^2 E \omega^7 \kappa(\omega)$$

$$\frac{dE_{TT}}{d\omega} = \frac{428}{525\pi} G_N^3 E^2 \omega^8 \log(\omega/\mu_1) \kappa(\omega)$$

$$\frac{dE_{TTT}}{d\omega} = \frac{856}{525} G_N^4 E^3 \omega^9 \log(\omega/\mu_2) \kappa(\omega)$$

For T and TT – agreement w Bini+Geralico 2021, TTT – new

Flux for Circular Orbits

We can specialize to circular orbit with orbital frequency Ω .

Perfect agreement of our results with traditional Self-Force/PN theory.

Summary & Outlook

- Effective actions that encapsulate all conservative and dissipative physics, including TT, TTT – never previously derived
- Derivation of dissipative observables matches with results where available
- Set up Amplitudes-like framework + machinery for hereditary effects: Directly applicable to PN binary inspirals, proven to 4-loop, efficient & fast
- Competitive with traditional GR and standard EFT methods

Prospects:

- Scattering from generic multipole source off any multipole source background
- Extension to all PN orders

Both entail tree amplitudes with massive particles of any spin!