# IceCube sterile neutrino searches

**Snowmass 2022 - Ben Smithers** 



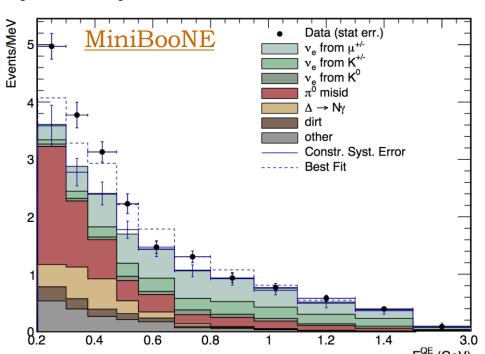


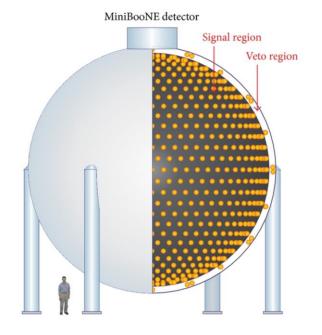


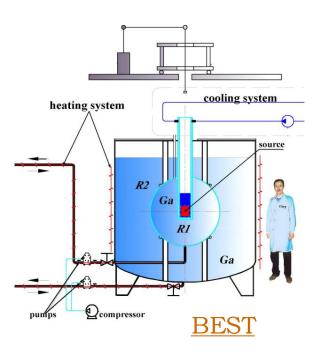


#### Motivations

- Anomalous MiniBooNE nu-e appearance results
- Could be addressed with 3+1 sterile neutrino model
  - Non-interacting flavor states,
  - "Light" mass-squared splitting ~1eV^2
- Many, many, more anomalous results since then







#### IceCube IceCube Lab **IceTop** 81 stations 324 optical sensors IceCube Array 86 strings including 8 DeepCore strings 5160 optical sensors 1450 m DeepCore 8 strings—spacing optimized for lower energies 480 optical sensors **Eiffel Tower** 324 m 2450 m 2820 m **Bedrock**

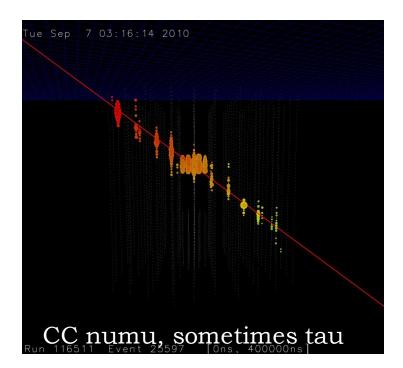


- Formally, IceCube Neutrino Observatory
- Array of 5160 light-sensing DOMs instrumented in south pole ice
- More densely instrumented region called DeepCore – sensitive to low-E oscillations
- Sparsely instrumented section sensitive to higher-E oscillations

#### **Event Morphologies**

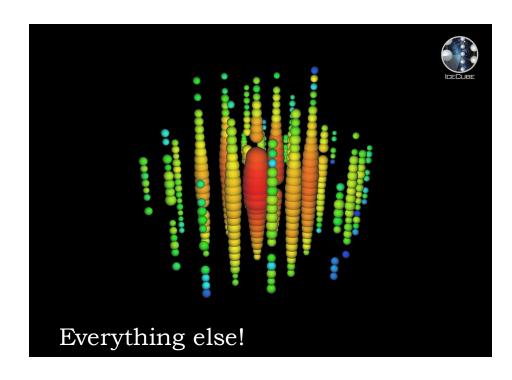
#### Tracks

- Poorly contained (lower energy resolution)
- Point with certainty (good angular resolution)
- Higher statistics



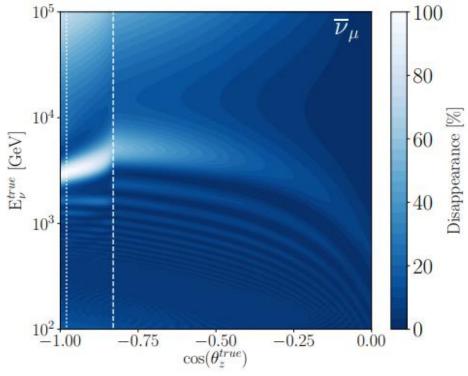
#### Cascades

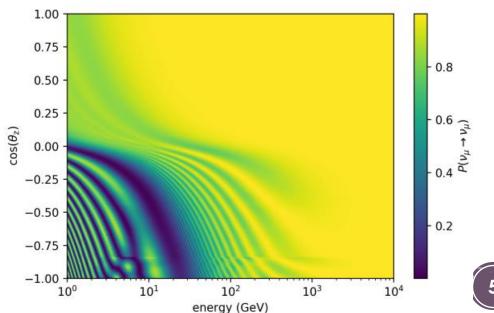
- Relatively easy to contain
- Blob-like (poor angular resolution)
- Low stats, often overlooked for BSM searches



#### Dominant Oscillations

- High Energies
  - ~500GeV to 10 TeV
  - Whole detector
  - BSM oscillations dominate
  - Both atmospheric and astrophysical
- Low Energies
  - ~5-50 GeV
  - DeepCore
  - Both BSM and regular effects intermingle
  - Atmospheric nu



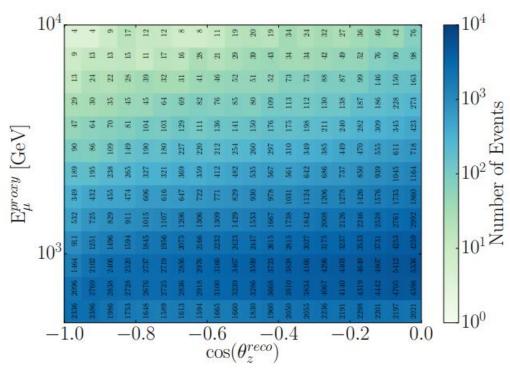


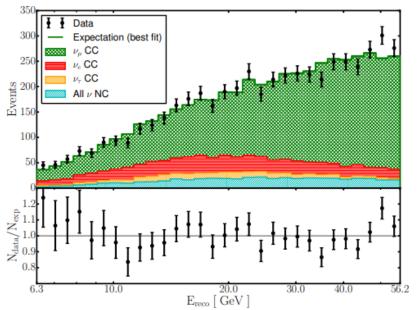


# © Past Analyses

#### Two Regimes

- High-energy, ~500GeV to 10 TeV, with
  - 8 years of IceCube
  - 305,735 up-going muon neutrino events
  - High-energy cascades analysis on the way
- Low-energy, ~5-50 GeV, with
  - 3 years of DeepCore
  - Approx 5118 events, assorted
  - OscNext, full 8-year analysis with
     ~260k events, coming soon
  - IceCube Upgrade will improve low-E sensitivity with a dense infill



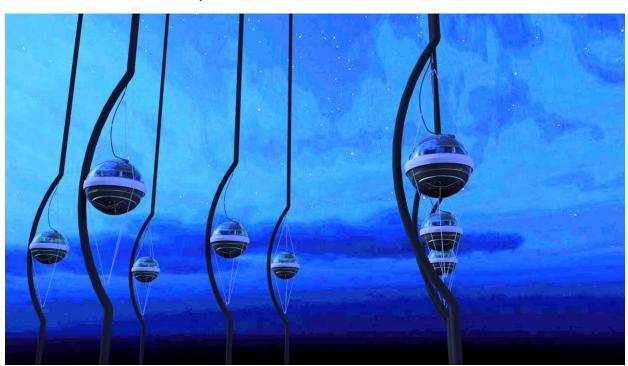




#### Systematic uncertainties

- High-energy, ~500GeV to 10 TeV
  - Hole Ice, Absorption/Scattering
  - DOM efficiency
  - Barr parameters, atmospheric density
  - Flux normalizations, slope
  - Cross section
  - Kaon energy loss rates

- Low-energy, ~5-50 GeV
  - Hole ice effects
  - DOM efficiency
  - Cross sections
  - Flux normalization, slope
  - nu/anu ratios

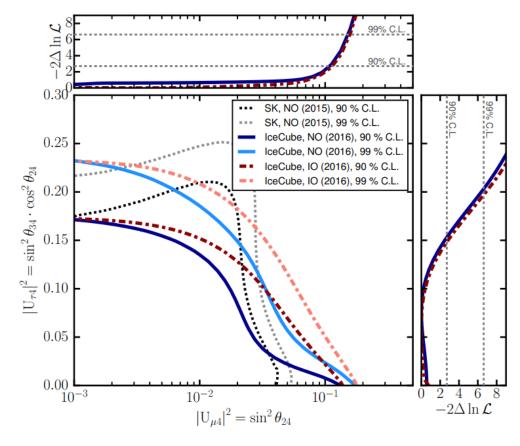


## Low-E Results (~5-50 GeV)

- Low-Energy DeepCore analysis
- All-flavor, all-interaction, up-going
- Fit to standard nu mixing parameters,

$$\Delta m_{32}^2 = 2.52 \cdot 10^{-3} \text{ eV}^2, \sin^2 \theta_{23} = 0.541$$

- First results consistent with 3neutrino model
- Nuisance parameters fit near nominal values



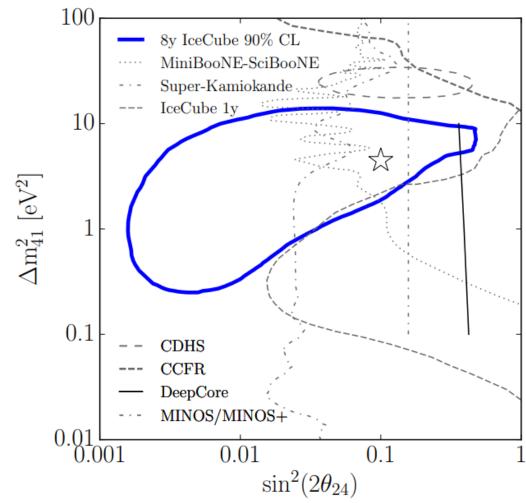
10.1103/PhysRevD.95.112002



#### High-E Results (~500 GeV - 10TeV)

- High Energy, matter effect
- Fits to all nuisance parameters
- Closed contour, best fit
  - $\sin^2(2\theta_{24})=0.10$ ,  $\Delta m_{41}^2=4.5$ eV<sup>2</sup>
- Exclusion contour at 99% CL
- Potentially statistically weak signal hint at 90%CL
- Motivates cross-checking in other channels

- Published in
  - PRD 10.1103/PhysRevD.102.052009
  - PRL 10.1103/PhysRevLett.125.141801



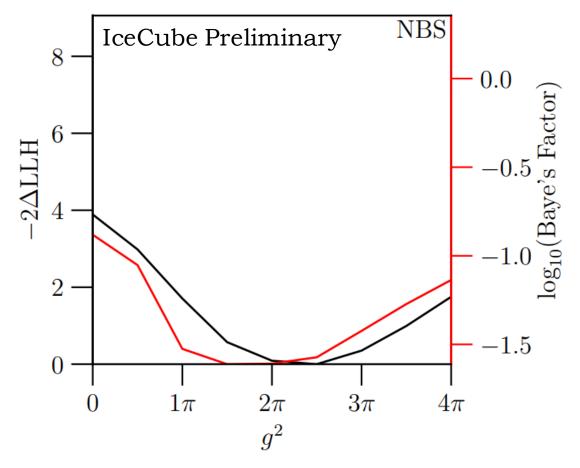


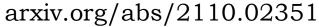
## Sterile Decay Sensitivity

- An additional mass, flavor state with decay
- Same 8-year through-going muon sample
- Sterile state with lifetime

$$\frac{1}{\tau} = \Gamma = \frac{g^2 m_4}{16\pi}$$

• Analysis fits to,  $\Delta m_{41}^2, \sin^2(\theta_{24}), g^2$  frequentist and Bayesian model comparison



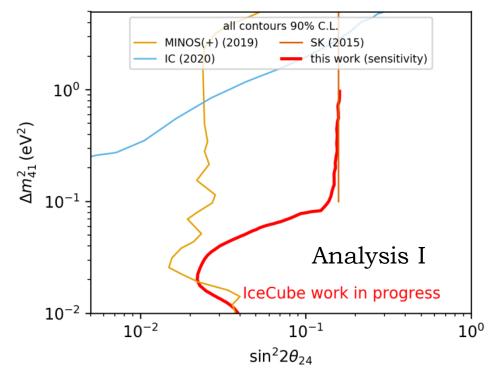


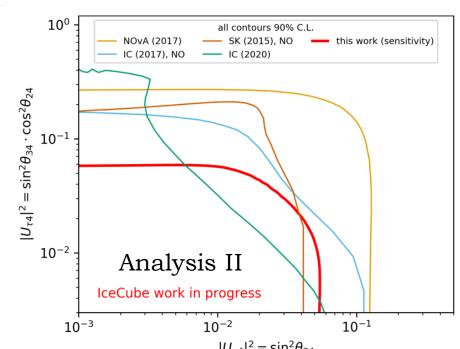


# (12) Upcoming analyses

## OscNext Analysis

- Full 8 years of DeepCore data
- 5-300 GeV analysis
- 260k events in total
- Multiple sub-analyses
  - In both,  $\Delta m_{32}^2, \theta_{23}$  free
  - Analysis II  $\delta_{24}$  free
- Improved systematic uncertainties
  - Interpolation between GENIE and CSMS DIS cross-sections
  - DOM eff, hole/bulk ice

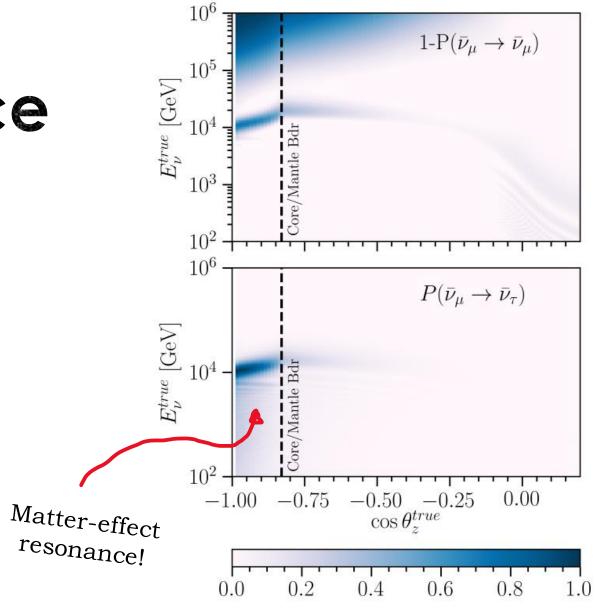






#### Tau Appearance

- $\nu_{\mu} \rightarrow \nu_{s} \rightarrow \nu_{\tau}$  resonance expected for non-zero  $\theta_{24}, \theta_{34}$
- Up-going antineutrinos, passing through the Earth's core
- Leads to muon disappearance, tau appearance
- Potential for cascade appearance, direct tau appearance

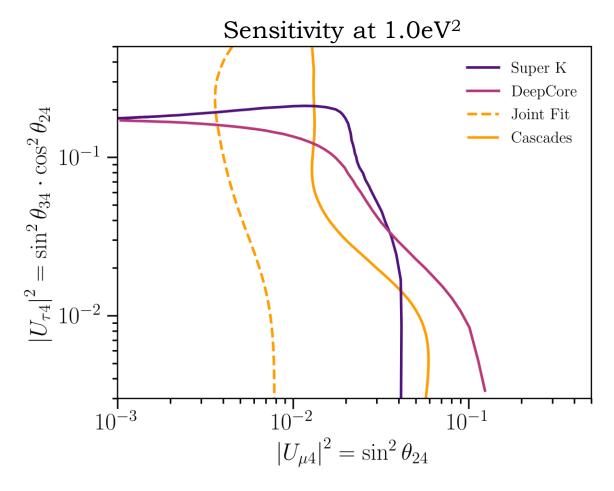


Accepted by PRD: arxiv.org/abs/2111.08722



## Calculated Sensitivity

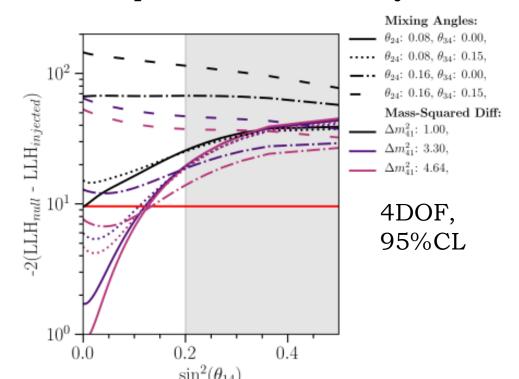
- Calculated using public IC effective areas, reconstruction efficiency
- Could discover signatures in  $v_{\tau}$  appearance
- Simplified systematic uncertainties
  - Flux normalization, shape
  - Ice Absorption/scattering
- Considering both
  - Cascade-only sensitivity
  - Joint track-cascade sensitivity



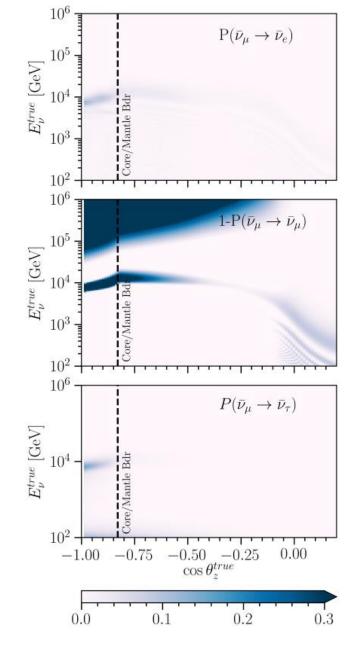


## Probing $\theta_{14}$

- Recent BEST results further support gallium anomaly
- Non-zero  $\theta_{14}, \theta_{24}, \theta_{34}$  could lead to similar resonant  $v_e, v_\tau$  appearance
- Will be able to probe BEST anomaly







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#### Outlook

- Upcoming IceCube OscNext will improve upon previous 3yr DeepCore analysis
- High-Energy analyses incorporating cascade events
- IceCube poised to make direct tauappearance measurement
- IceCube will be able to probe the BEST anomaly



# Thank you for your time! Questions?