



**COMPUTING FOR GRAVITATIONAL-
WAVE RESEARCH TOWARDS O4**

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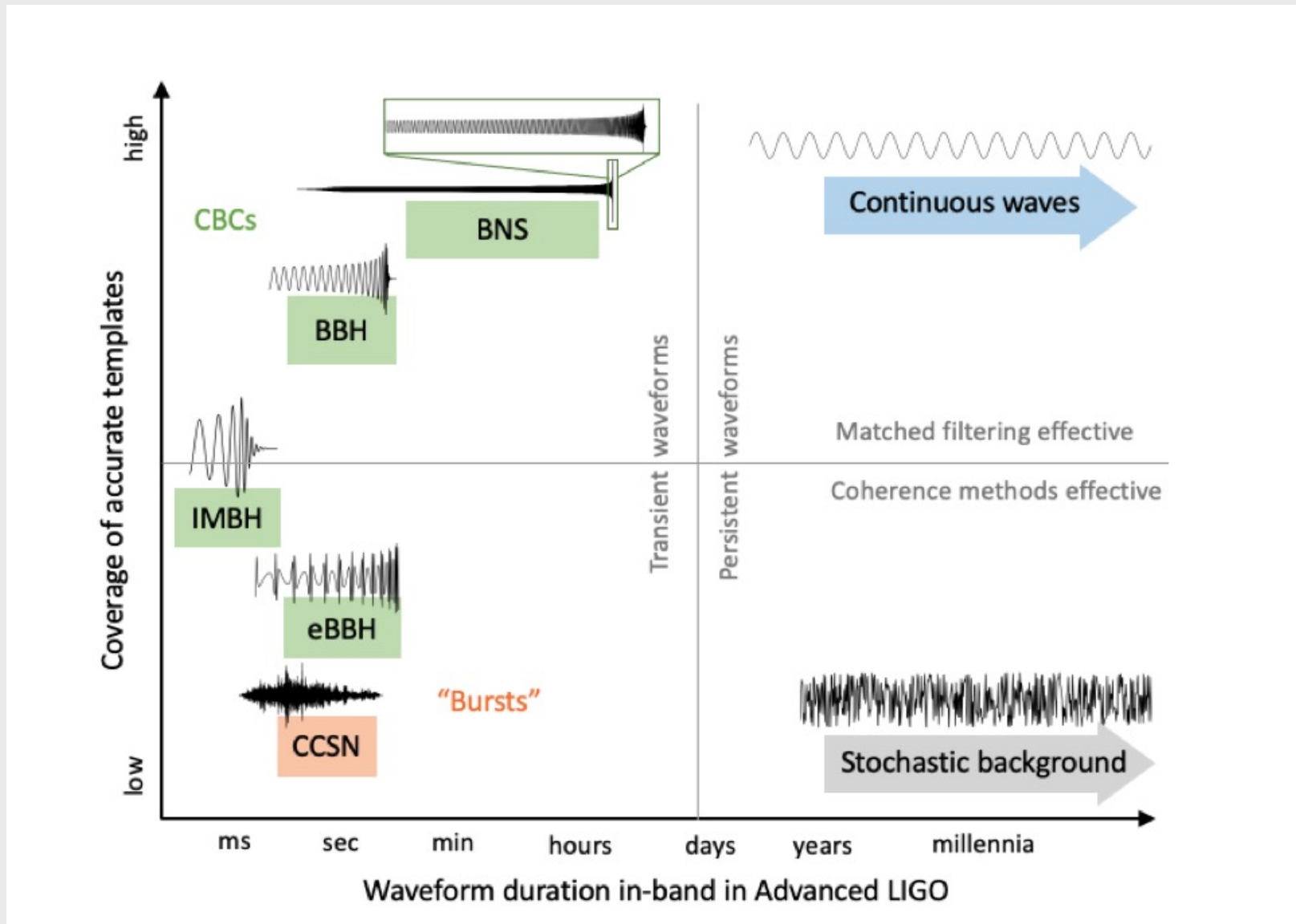
On behalf of the LVK Collaboration

Burst sources:

- **CBC: Compact Binary Coalescence**
 - Coalescing Compact Binary Systems (Neutron Star-NS, Black Hole-NS, BH-BH): Strong emitters, well modelled for much of the parameter space
- **Burst: Unmodeled transient bursts**
 - Asymmetric Core Collapse Supernovae: weak emitters, not well-modelled (“bursts”), transient
 - Cosmic strings, soft gamma repeaters, pulsar glitches,...
 - Who knows?

Continuous sources:

- **CW: Continuous waves**
 - Spinning neutron stars (known waveform, long/continuous duration)
 - All-sky and targeted searches
- **SGWB: Continuous stochastic background**
 - Cosmological stochastic background (residue of the Big Bang, cosmic GW background, long duration)
 - Astrophysical stochastic background

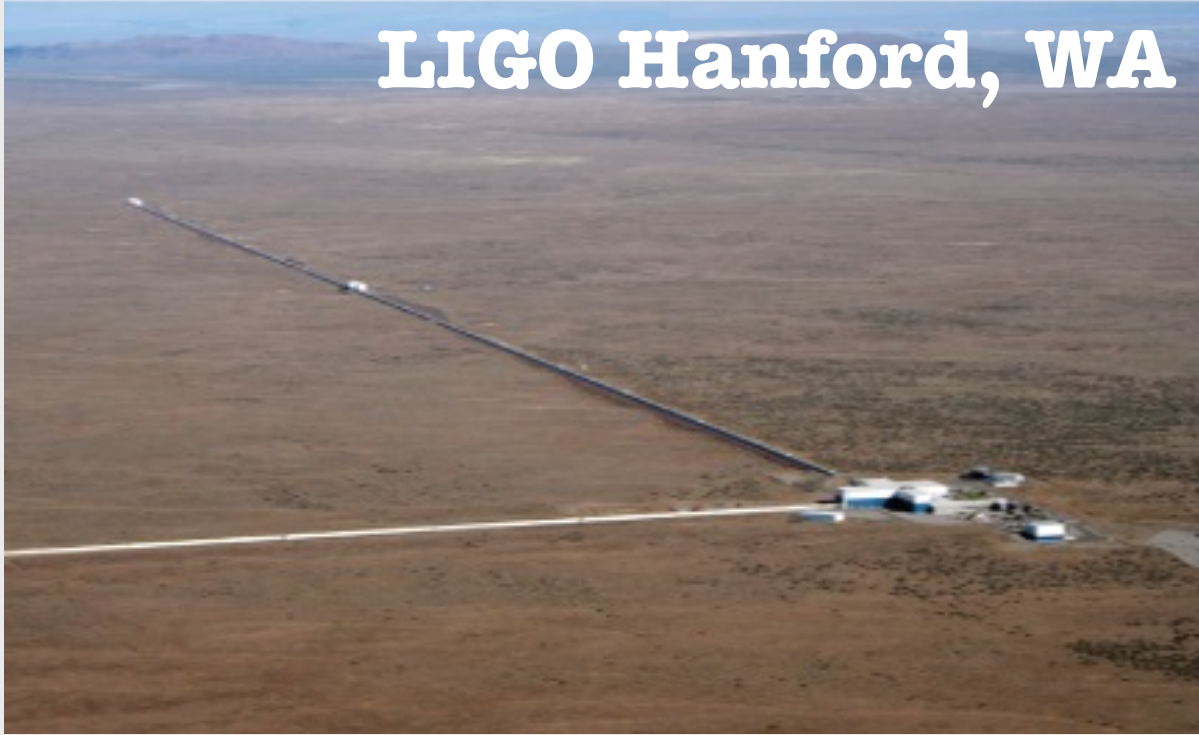


Jess McIver, D. H. Shoemaker, *Discovering Gravitational Waves with Advanced Ligo*, LIGO Document P2000530-v1 (2000)



European Gravitational
Observatory, Cascina (Italy)

LIGO Hanford, WA

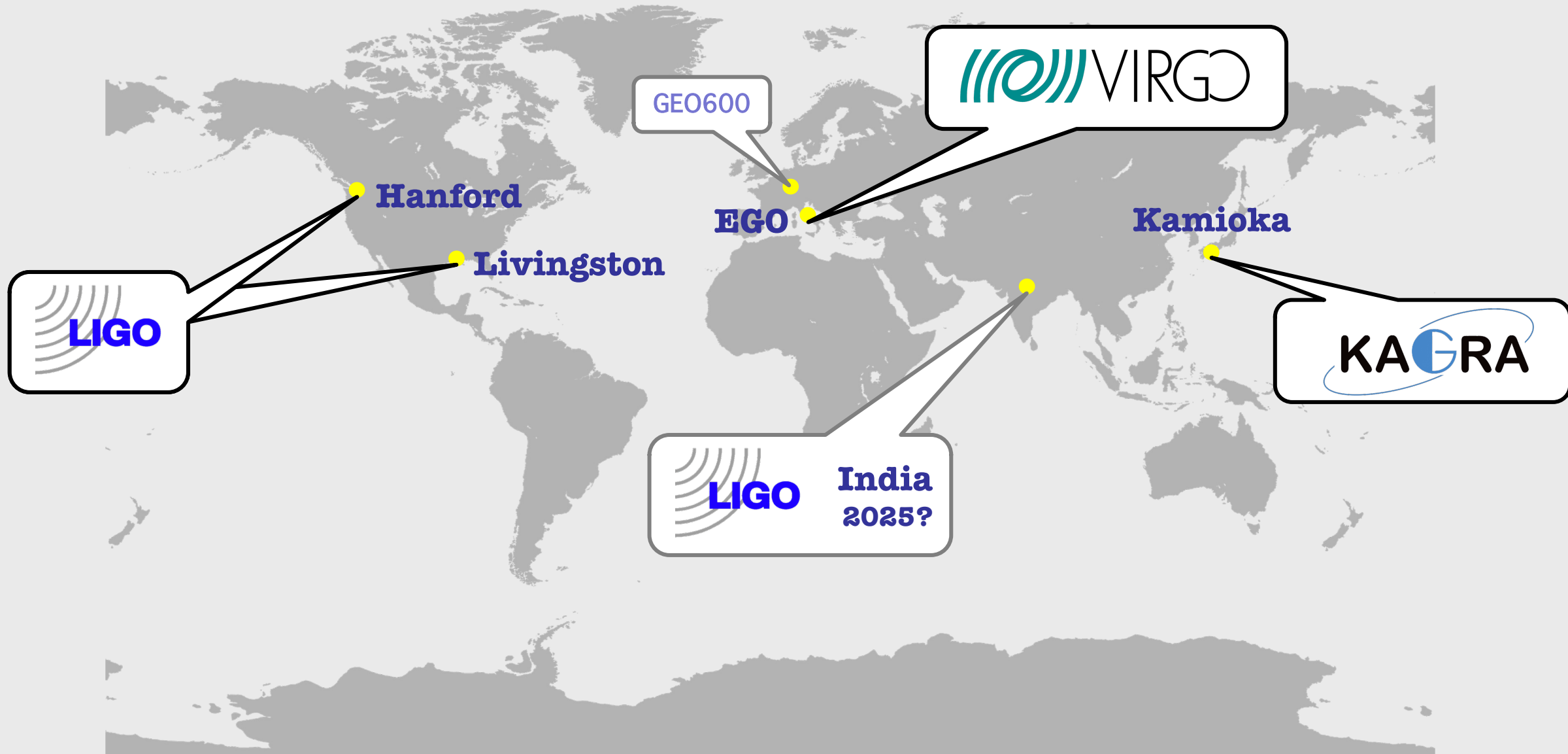


LIGO Livingston, LA



Caltech MIT/Ligo Laboratory

A WORLDWIDE NETWORK



THREE COMPUTING DOMAINS

**On-site
infrastructure**

Online

- Data acquisition and pre-processing
- Instrument control
- Environmental monitoring
- ...

**Plain old HTC
(and some HPC)**

Offline

- Deep searches
- Offline parameter estimation
- Detector Characterization (DetChar)
- (Template bank generation)

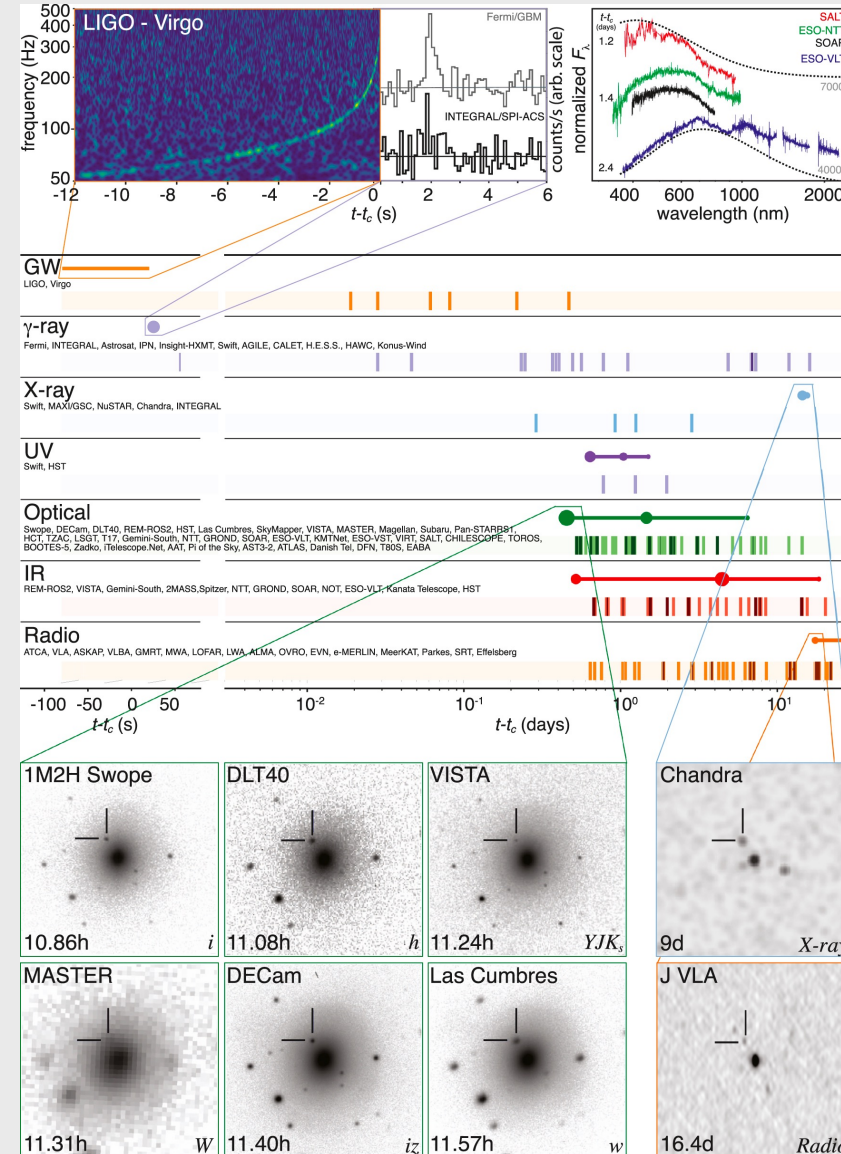
Here's the fun

Low-latency

- Candidate search
- Sky localization
- LL parameter estimation
- Alert generation and distribution

MULTIMESSENGER ASTRONOMY

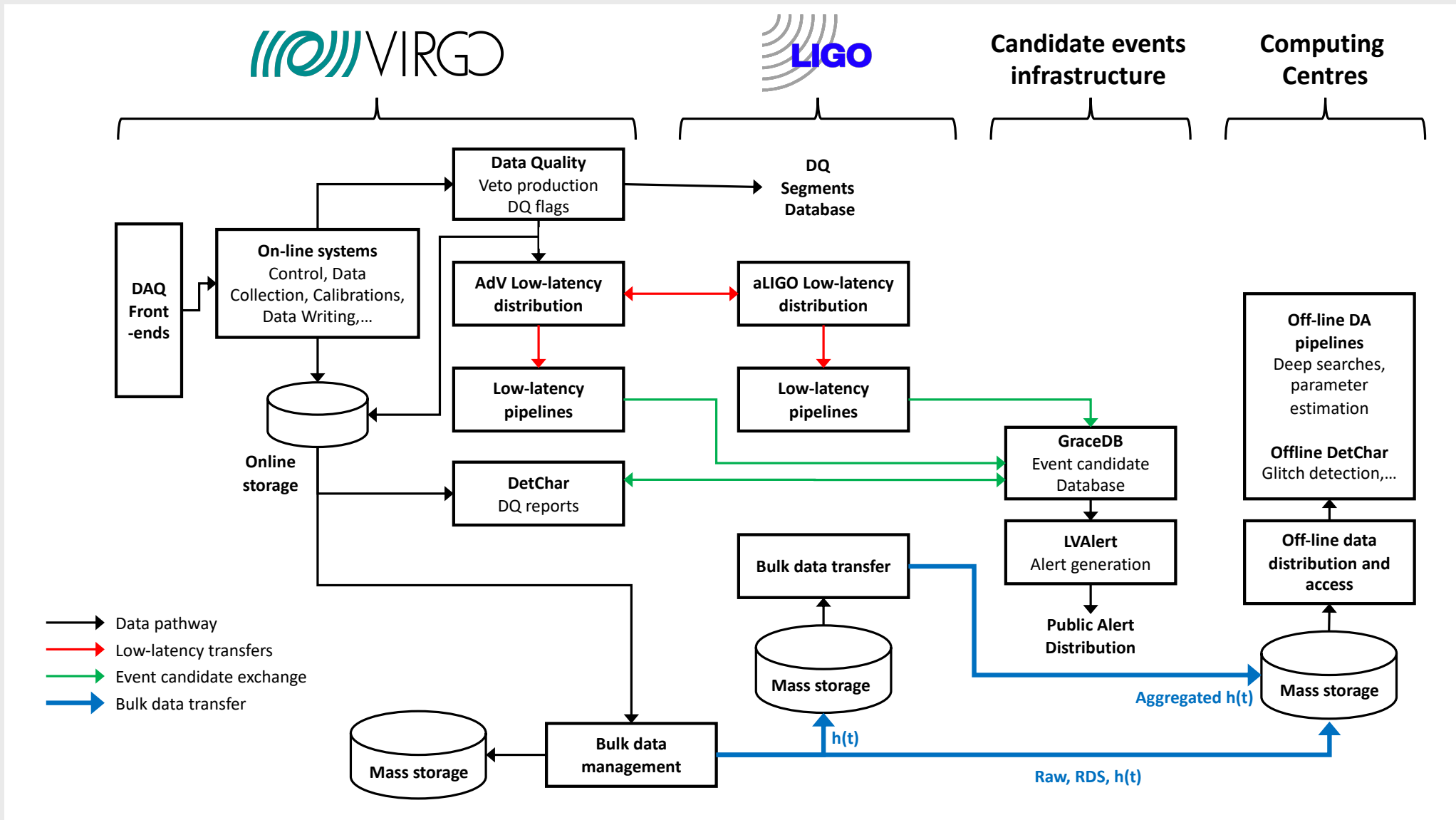
- Only GW170817 so far...
- Public Alerts are “Triggers” for ground- and space-based EM observatories
- In O3 average latency was $\sim 1/2$ hour
- Target latency for O4 is \sim minute
- However, for some events “early warning” alerts with negative latency are possible



“Multi-messenger Observations of a Binary Neutron Star Merger”

B. P. Abbott *et al.* 2017 *ApJL* 848 L12
doi:10.3847/2041-8213/aa91c9

COMPLEX OVERALL DATA FLOWS

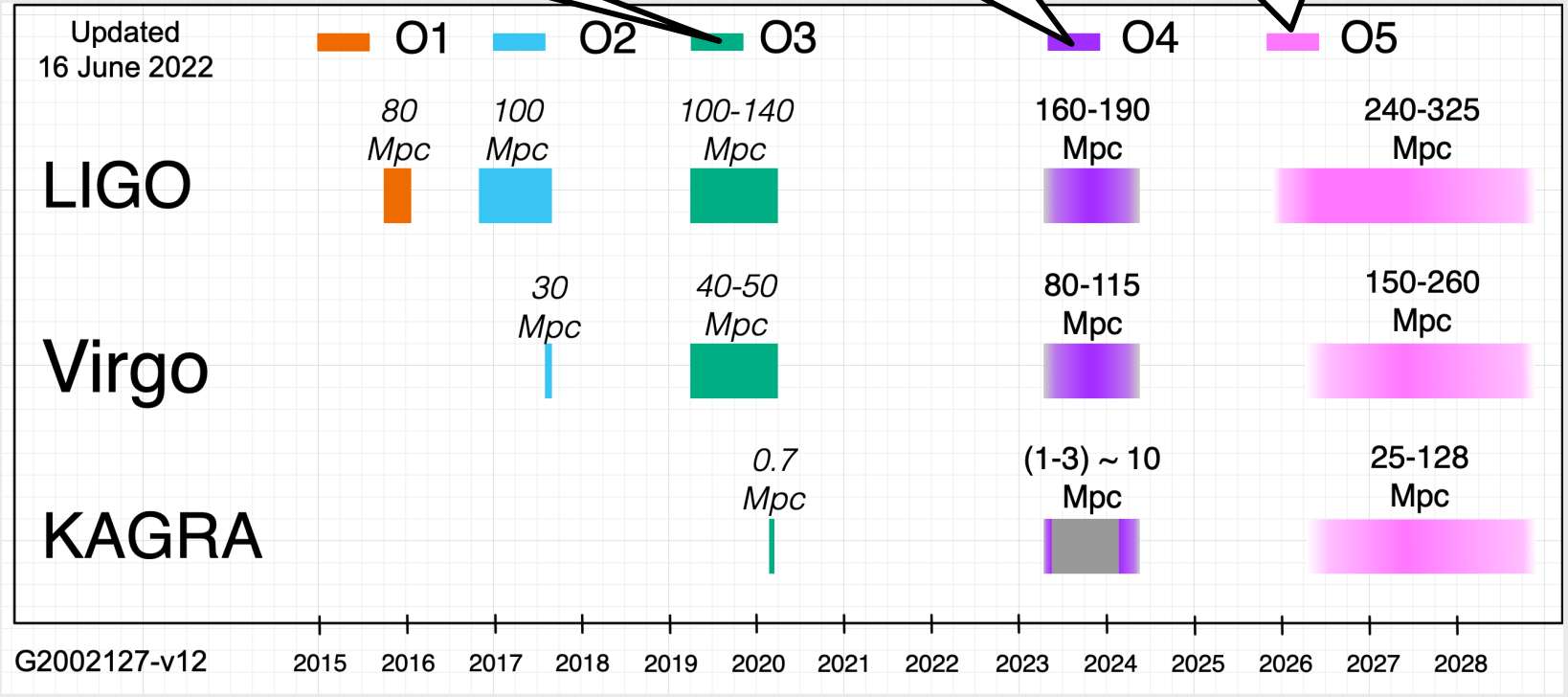


The screenshot shows the GraceDB Public Alerts page for LIGO/Virgo O3. The page title is "LIGO/Virgo O3 Public Alerts" with "Detection candidates: 56". The table below lists several events with their details.

Event ID	Possible Source (Probability)	UTC	GCN	Location	FAR	Comments
S200316bj	MassGap (>99%)	March 16, 2020 21:57:56 UTC	GCN Circulars Notices VOE		1 per 446.44 years	
S200311bg	BBH (>99%)	March 11, 2020 11:58:53 UTC	GCN Circulars Notices VOE		1 per 3.5448e+17 years	
S200308e	NSBH (83%), Terrestrial (17%)	March 8, 2020 01:19:27 UTC	GCN Circulars Notices VOE		1 per 8.757 years	RETRACTED
S200303ba	BBH (86%), Terrestrial (14%)	March 3, 2020 12:15:48 UTC	GCN Circulars Notices VOE		1 per 2.4086 years	RETRACTED
	BBH (89%),	March 2, 2020	GCN Circulars		1 per 3.3894	

- Centralized aggregation point for information about candidate gravitational-wave events
- Provides web interface and programmatic API
- Works together with a notification service (XMPP-based LVAAlert, being replaced by HOPSKOTCH/Kafka-based igwn-alert for O4)
- GWCelery: a package for annotating and orchestrating LIGO/Virgo alerts, built on the Celery distributed task queue.

10² CBC det./yr
10³ CBC det./yr
10⁴ CBC det./yr



Then:

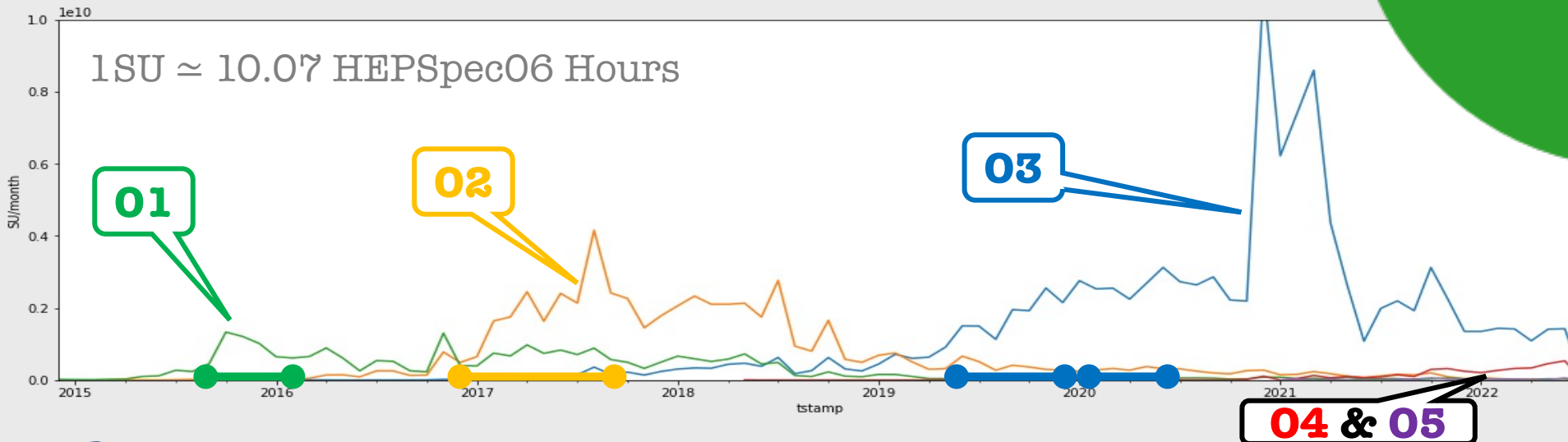
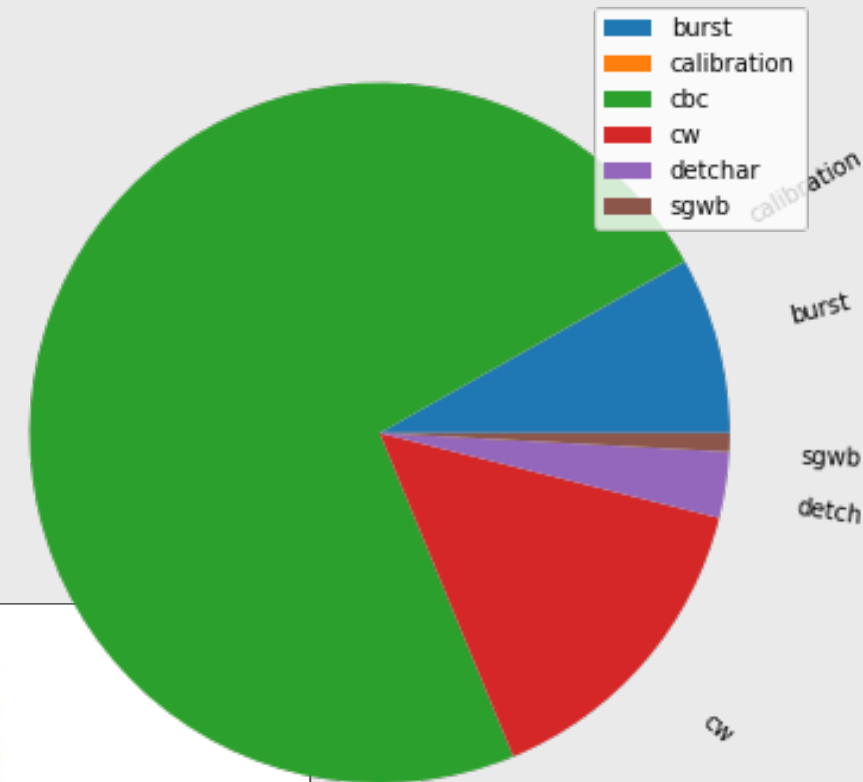
- “Post-O5” plans being prepared
- Projects for Third-generation interferometers being proposed (Einstein Telescope, Cosmic Explorer)
- Growth in “computing size” (relatively) gradual



OVERALL COMPUTING RESOURCES USAGE

- Including offline, low-latency and DetChar
- Overall CPU for O3 was ≈ 7000 MHS06 Hours
- Large peaks after end of observation periods (frequency-domain analyses)
- As expected, interest for older data wanes as new data become available
- Required computing power for O4 about $1.5 \div 2 \times O3$
- Overall: about 10% of an LHC experiment

O3 by AG



- Raw interferometer data don't grow much with increasing instrument sensitivity
 - In O3 we were writing $o(1\text{PB})$ per year of raw data per detector
 - $h(t)$ (or “strain”, the physics channel) + $o(10^5)$ control channels
 - Pre-processed data for final user analysis is more than 1 order of magnitude smaller
 - $1.5 \times$ expected in O4 (more control channels!)
- What grows is the amount of useful scientific information embedded in the data
 - And the computing power needed to wring it out

Need to define and deploy a common and sustainable GW computing environment

- Provide a **uniform runtime environment** for offline pipelines
- **Full interoperability** between Virgo, LIGO (and KAGRA)
- Provide scalability and the **opportunity to exploit heterogeneous resources**
- Adopt **mainstream, widely used tools**, leveraging upon HEP experience

Enter **IGWN** – the International Gravitational Waves observatories Network

- A coordination effort aimed at jointly discussing the computing policy, management, and architecture issues of LIGO, Virgo, and KAGRA.

So, for example...

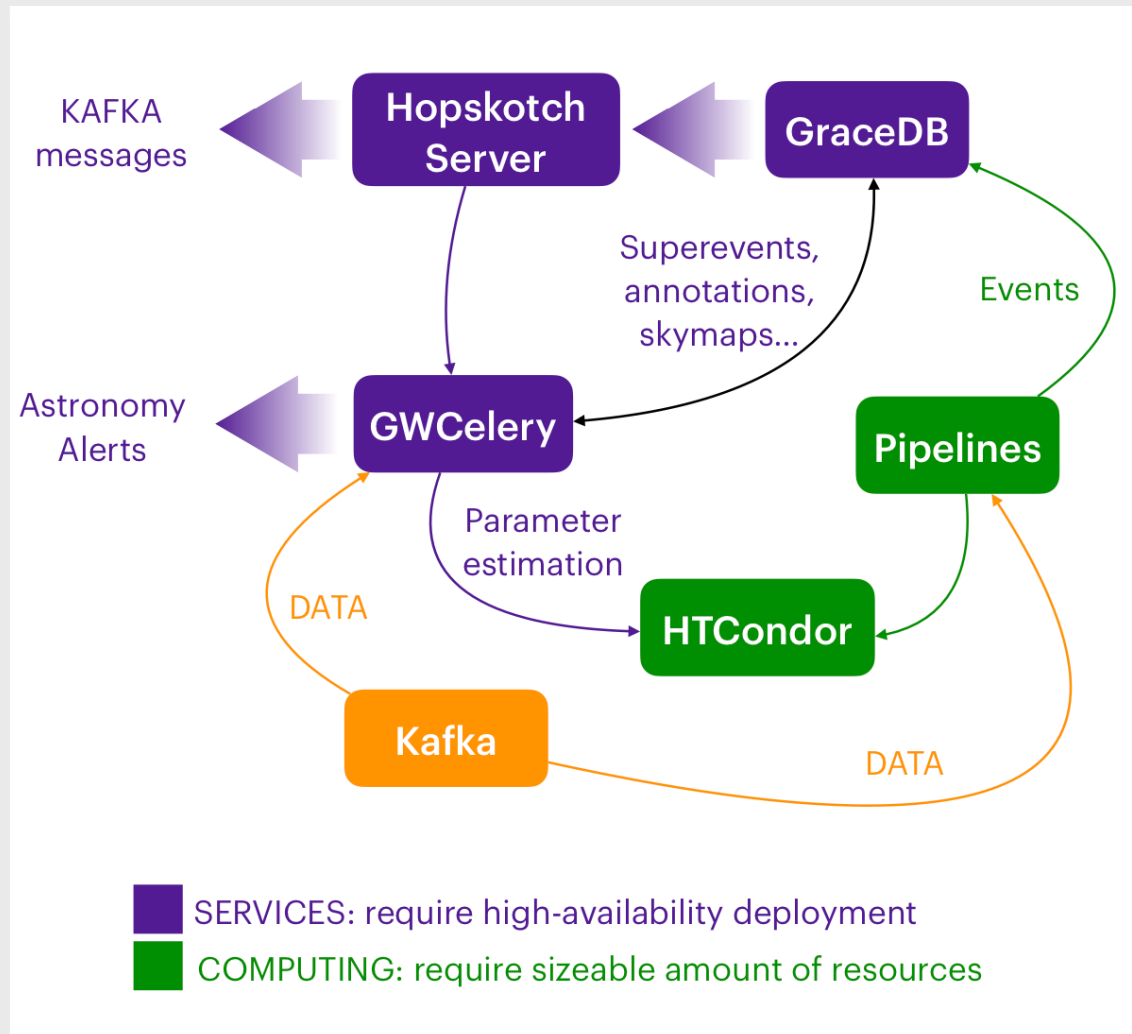
- Use Apache Kafka for low-latency data transfers
- Move towards a fully distributed offline computing model
 - Based on HTCondor for workload management and CVMFS+StashCache for data distribution
 - Strong support from OSG and HTCondor community
- Gradually adopt Rucio for data management tasks
 - E.g. transfer of $h(t)$ data to StashCache Origin servers for distribution
- Streamline code management
 - GitLab, CI, Conda Environments, Containers,...
- ...

THE EU SIDE OF THE IGWN



MODERN LOW-LATENCY SERVICES DEPLOYMENT

- Upgrade of main services (GWCelery and GraceDB)
 - And exploring a complete redesign
- Deployment on Kubernetes in INFN-Cloud at CNAF, INFN's Tier-1 centre in Bologna
 - With the idea of also having on-demand CPU resources for running jobs (not yet there)
 - Production and Development instances running at CNAF
 - Sandboxed configuration option for self-contained installation on Minikube
 - GitLab CI for deployment and configuration
 - Monitoring via Prometheus/Grafana
- Get ready for exploitation of upcoming “Scientific Cloud” resources



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- Still a very large number of computing tools and services to maintain
 - Even adopting mainstream technologies, some high-level services are unavoidably proprietary
- Pipelines are very heterogeneous and difficult to port to distributed environment
 - Python, C++, Matlab,...
 - Essentially no common framework
 - Many have been developed to run in specific environments (data access technology, runtime environment,...)
- Difficult to recruit (and keep) the right computing skills
 - ...which translates into a large personpower gap for computing

- GW computing started small but is coming of age
- Increased interferometer sensitivity does not imply proportional growth in data size...
- ...but computing power requirements do grow.
- No quantum leaps O4 → O5 → “post-O5” → 3G
- A large effort is being made to adopt modern computing technologies and strategies, and to integrate into the wider physics computing community
- A large personpower gap is slowing down the transition, but most important targets for O4 are within reach