# The effects of magnetic fields on magnetic dipole moments

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Work done with Tuna Demircik, Gyurin Kim, Matti Järvinenar

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#### 1. Introduction

Motivation Measurement of Magnetic Moments The magnetic field effect

#### 2. Field-dependent magnetic moment

QED contributions QCD contributions

#### 3. Refraction

Refraction under magnetic field

#### 4. Conclusion

#### Anomalous magnetic moments of electrons and muons

▶ All massive particles with spin are known to have a magnetic dipole moment:

$$\vec{\mu} = \frac{g}{2} \frac{e}{m} \vec{S} = (a+1) \frac{e}{m} \vec{S}$$

- ▶ The magnetic moment is an intrinsic quantity of particles. Its agreement between theory and experiment has been a real triumph of QFT, since Schwinger calculated it in 1951.
- For electrons it is measured at ppb level (Haneke et. al 2008):

$$a_e^{\text{exp}} = 1159652180.73(28) \times 10^{-12}$$

The discrepancy is about  $2.4\sigma$  (Parker et. al 2018):

$$\Delta a_e \equiv a_e^{
m exp} - a_e^{SM} = (-87 \pm 36) \times 10^{-14}$$

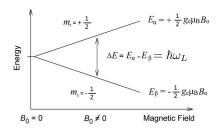
#### Anomalous magnetic moments of electrons and muons

- ▶ It served to test the consistency of GSW model of electroweak theory during its establishement. (Fujikawa-Lee-Sanda 1972)
- Currently it provides a test and also a hint for new physics beyond standard model.
- ► For muons it is measured at ppm level (FNAL 2021):

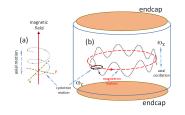
$$a_{\mu}^{\text{exp}} = 116592061(41) \times 10^{-11}$$

with 4.2
$$\sigma$$
 deviation,  $\Delta a_{\mu} = (251 \pm 59) \times 10^{-11}$  .

### Measurements of $a_e$ - Penning Trap (Old): U.W.



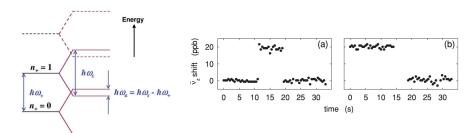
$$\Delta E = -\mu \cdot \vec{B}$$





$$\hbar\omega_c = 2\mu_B B_0$$

#### Measurements of $a_e$ - Improved Penning Trap: Harvard

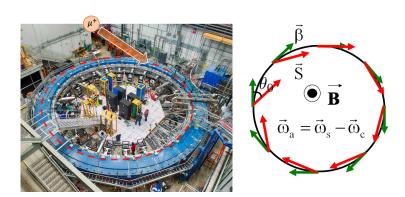


G. Gabrielse et al.(2004)

- Muons are unstable. It decays into electron 100% with lifetime of about  $2.2\mu \, \mathrm{s}$ .
- ► The key observation is that the difference between the spin precession frequency and the (relativistic) cyclotron frequency is given as

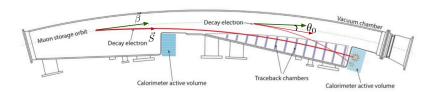
$$\omega_{\mathsf{a}} \equiv \omega_{\mathsf{s}} - \omega_{\mathsf{c}} = \mathsf{a}_{\mu} rac{\mathsf{e} \mathsf{B}}{\mathsf{m}} \,.$$

4. Conclusion

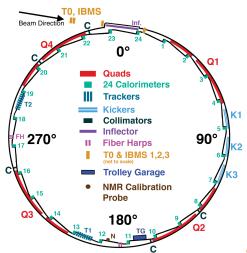


Fermi Lab Muon Storage Ring, slide from O. Kim (IBS)

▶ Because EW interaction is chiral, the electrons decay mostly along the muon spin direction with an angle  $\theta_0 = \omega_a t + \phi$ :

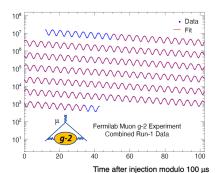


# Ring layout



► The count of positrons (electrons) at the detector as a function of time:

$$N(t) = N_0 e^{-t/\tau} \left[ 1 + A(E_e) \sin(\omega_a t + \phi) \right]$$
 (1)



▶ The operational definition of magnetic dipole moment is

$$\vec{\mu} \equiv g \frac{q \vec{S}}{2m} = - \left. \frac{\partial \mathcal{E}(\vec{B})}{\partial \vec{B}} \right|_{\vec{B}=0}$$
 (2)

- ▶ But in practice the magnetic fields never vanish. In fact one often needs strong magnetic fields, B > 10kG.
- In view of current accuracy, the effect of magnetic field may not be negligible.

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▶ On dimensional grounds one expects the magnetic effect to be

$$\delta(g-2) \sim \frac{\alpha}{2\pi} \cdot \frac{eB}{m^2} = 3 \times 10^{-13} \left(\frac{B}{10 \, \mathrm{kG}}\right) \left(\frac{0.51 \, \mathrm{MeV}}{m}\right)^2 \, .$$

- The magnetic effect may be compatible with experimental uncertainty for electrons,  $\delta a_{\rm e}^{\rm exp} \simeq 2.8 \times 10^{-13}$  while way small for muons  $\delta a_u^{\rm exp} \simeq 54 \times 10^{-11}$
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#### Field-dependent magnetic moment

▶ If  $B \ll m^2/e$ , one can expand the magnetic moment g factor in powers of magnetic field:

$$g_e(B) \equiv rac{|\Delta \mathcal{E}|}{\mu_B B} = g_e(0) + \sum_{n=1}^{\infty} g_e^{(n)} \left(rac{eB}{m^2}
ight)^n \,.$$

From the exact electron propagator to the order  $\alpha$  under a constant magnetic the shift in the ground state energy is at one-loop :

$$\Delta \mathcal{E}^{\mathrm{QED}}(B) = \frac{\alpha}{2\pi} m \left[ -\frac{eB}{2m^2} + a_2 \left( \frac{eB}{m^2} \right)^2 + a_3 \left( \frac{eB}{m^2} \right)^3 + \cdots \right],$$

where 
$$a_2 = \left(\frac{4}{3} \ln \frac{m^2}{2eB} - \frac{13}{18}\right)$$
,  $a_3 = \left(\frac{14}{3} \ln \frac{m^2}{2eB} - \frac{32}{5} \ln 2 + \frac{83}{90}\right)$ .

► The anomalous magnetic moment at one-loop in QED is therefore given as

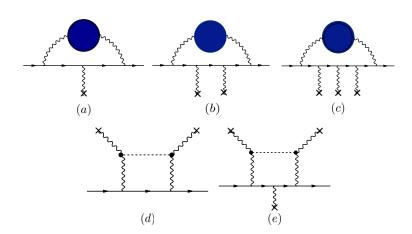
$$rac{1}{2}(g-2)_{
m QED}^{
m 1-loop} = rac{lpha}{\pi} \left[ rac{1}{2} \mp a_2 \left( rac{eB}{m^2} 
ight) - a_3 \left( rac{eB}{m^2} 
ight)^2 + \cdots 
ight] \, .$$

- ▶ The leading correction, linear in B, depends on the spin direction, which is coming from the  $(eB)^2$  contribution to the ground state energy.
- Since the  $(eB)^2$  in the energy is independent of B field direction, it contributes equally to both spin up and down
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## QCD contributions up to $B^2$



► The QCD corrections to the ground state energy are :

$$\Delta \mathcal{E}^{
m QCD}(B) = -\left(rac{lpha}{\pi}
ight)^2 rac{m}{2} \left[c_1\left(rac{eB}{m^2}
ight) + ar{c}_2\left(rac{eB}{m^2}
ight)^2 + ar{c}_3\left(rac{eB}{m^2}
ight)^3 + \cdots
ight]\,.$$

▶ The QCD corrections to magnetic moments are then

$$rac{1}{2}(g-2)_{
m QCD}^{
m LO} = \left(rac{lpha}{\pi}
ight)^2 \left[c_1 \pm ar{c}_2 \left(rac{eB}{m^2}
ight) + ar{c}_3 \left(rac{eB}{m^2}
ight)^2 \cdots
ight] \,.$$

From fig. (d) and (e)

$$c_2 = rac{1}{72\pi^2} \cdot rac{m^4}{f_\pi^2 m_\pi^2} \left( 30 \ln \left( rac{m^2}{m_\pi^2} 
ight) + 19 
ight) \left[ 1 + \mathcal{O} \left( rac{m^2}{m_\pi^2} 
ight) 
ight] \,.$$

$$c_3 = -\frac{1}{72\pi^2} \cdot \frac{m^4}{f_\pi^2 m_\pi^2} \left( 12 \ln \left( \frac{m^2}{m_\pi^2} \right) + 17 \right) \left[ 1 + \mathcal{O} \left( \frac{m^2}{m_\pi^2} \right) \right] .$$

### Direct measurements of magnetic moments

- ▶ In the current experiments of magnetic moments, the contribution linear in *B* is not directly measurable.
- ▶ The measurable effect is hence at  $B^2$  order:

$$\delta g_l(B) \simeq -\frac{2\alpha}{\pi} \left[ a_3 - \frac{\alpha}{\pi} \bar{c}_3 \right] \cdot \left( \frac{eB}{m^2} \right)^2 \simeq 1.3 \times 10^{-20}$$

$$(B=10\mathrm{kG} \text{ and } m=0.51 \mathrm{\ MeV})$$

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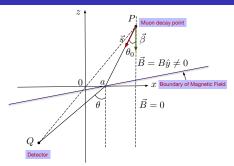
▶ The spin precession and cyclotron frequencies are proportional to the anomalous magnetic moment (Bargamann et al. 1959)

$$\omega_{\mathsf{a}} \equiv \omega_{\mathsf{s}} - \omega_{\mathsf{c}} = \mathsf{a}_{\mu} \frac{\mathsf{e} \mathsf{B}}{\mathsf{m}} \,.$$

▶ The number of electrons N(t) detected at the electromagnetic calorimeter is then

$$N(t) = N_0 e^{-t/\tau} \left[ 1 + A \sin \left( \omega_a t + \phi \right) \right].$$

### Refraction under magnetic field



► When electrons exit the magnetic field, they are refracted, following Maupertuis' principle:

$$\delta \int_{P}^{x=a} \left( p_1 dl + \vec{A} \cdot d\vec{r} \right) + \delta \int_{x=a}^{Q} p_2 dl = 0.$$

▶ The momentum satisfies

$$p^2 + m^2 = (\mathcal{E} - \Delta \mathcal{E}(B))^2.$$

The refracted angle  $\theta$  is related to the incident angle  $\theta_0 = \omega_a t + \phi$  as

$$\tan \theta - \tan \theta_0 = \gamma_* \tan \theta_0 \cdot \sec^2 \theta_0 \cdot \frac{\delta m}{m},$$

where  $\gamma_* = m\mathcal{E}/p_2^2 \approx m/p_2$  and  $\delta m = \Delta \mathcal{E}(B)$ .

 $\blacktriangleright$  When  $\tan \theta_0$  is small, the increment in the refraction angle

4. Conclusion

$$\delta\theta = -\gamma_* \frac{\delta \textit{m}}{\textit{m}} \tan\theta_0 \sim 10^{-16} \left(\frac{1~{\rm GeV}}{\textit{p}_2}\right) \left(\frac{\textit{B}}{10\,{\rm kG}}\right) \tan\theta_0 \,.$$

► For a generic angle  $\theta_0$ , we find

$$\sin\theta = \frac{\sin\theta_0 \left(\cos^2\theta_0 + \gamma_*\delta m/m\right)}{\sqrt{\cos^6\theta_0 + \sin^2\theta_0 \left(\cos^2\theta_0 + \gamma_*\delta m/m\right)^2}}$$

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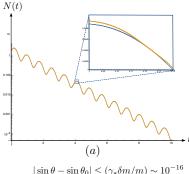
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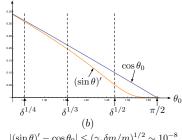
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The shape of the electron distribution changes due to the refraction by the magnetic field,  $\delta^n = \pi/2 - (\gamma_* \delta m/m)^n$ :

4. Conclusion



 $\sin \theta - \sin \theta_0 | \le (\gamma_* \delta m/m) \sim 10^{-16}$ 



 $|(\sin \theta)' - \cos \theta_0| \le (\gamma_* \delta m/m)^{1/2} \sim 10^{-8}$ 

▶ Due to the refraction, the oscillatory part of the shape of positron count changes:

$$0 \lesssim |\sin \theta - \sin \theta_0| \lesssim \gamma_* \frac{\delta m}{m} \sim 10^{-16}$$
.

- Near the peak (valley) the new shape is very flat for the range of angles  $\Delta\theta_0 \simeq (3\gamma_*\delta m/m)^{1/4} \sim 10^{-4}$ . Two slopes differ by  $(\gamma_*\delta m/m)^{1/2} \sim 10^{-8}$  in this region.
- Since each run subset  $10^9$  positrons are detected at Fermi lab experiment, one should detect  $10^5$  positrons in  $\Delta\theta_0$ , which is sufficient to distinguish two distributions.

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- ▶ We have calculated the effect of magnetic fields on the anomalous magnetic moment up to B<sup>2</sup>, both QED and QCD.
- The accuracy of experiments for magnetic moments is now high enough, ppb level, to be sensitive to magnetic fields.
- Currently only the quadratic effect is measurable:

$$\delta g_{\rm e}(B) \simeq -\frac{2\alpha}{\pi} \left[ a_3 - \frac{\alpha}{\pi} \bar{c}_3 \right] \cdot \left( \frac{eB}{m_{\rm e}^2} \right)^2 \simeq 1.3 \times 10^{-20}$$

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- The slope of the shape changes significantly near the peak (valley),  $|\pi/2 \theta| < (3\gamma_* \delta m/m)^{1/4} \sim 10^{-4}$ .
- By measuring this change, especially the slope difference,  $(\gamma_* \delta m/m)^{1/2} \sim 10^{-8}$ , one could measure the refraction due to the magnetic field.
- Furthermore to improve the experimental uncertainties, one should consider this effect, since  $\omega_a = a_\mu eB/m$  is determined by fitting the distribution shape of detected positrons.

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