

CRESST – direct dark matter experiment



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Laboratori Nazionali del Gran Sasso



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Valentyna Mokina for the CRESST collaboration
HEPHY OEAW

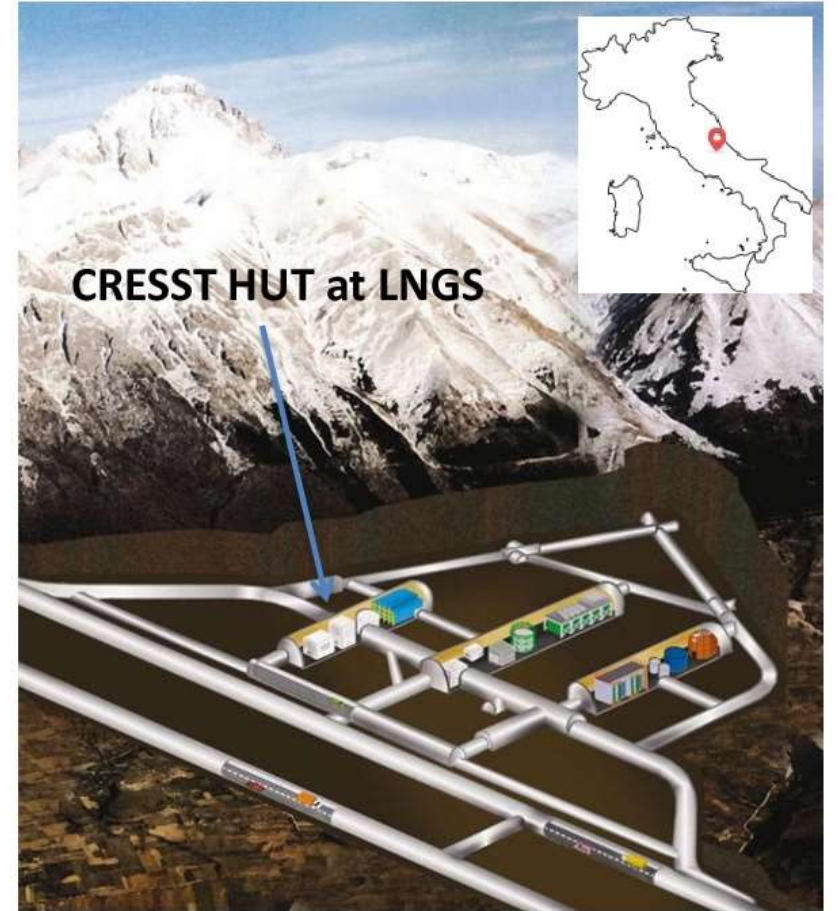
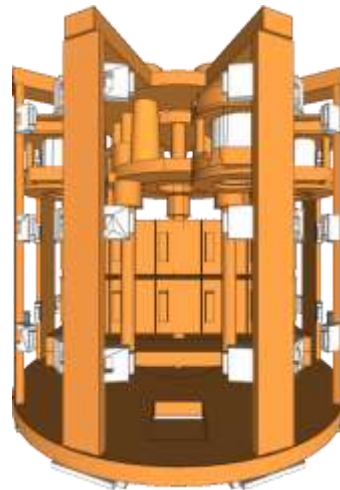
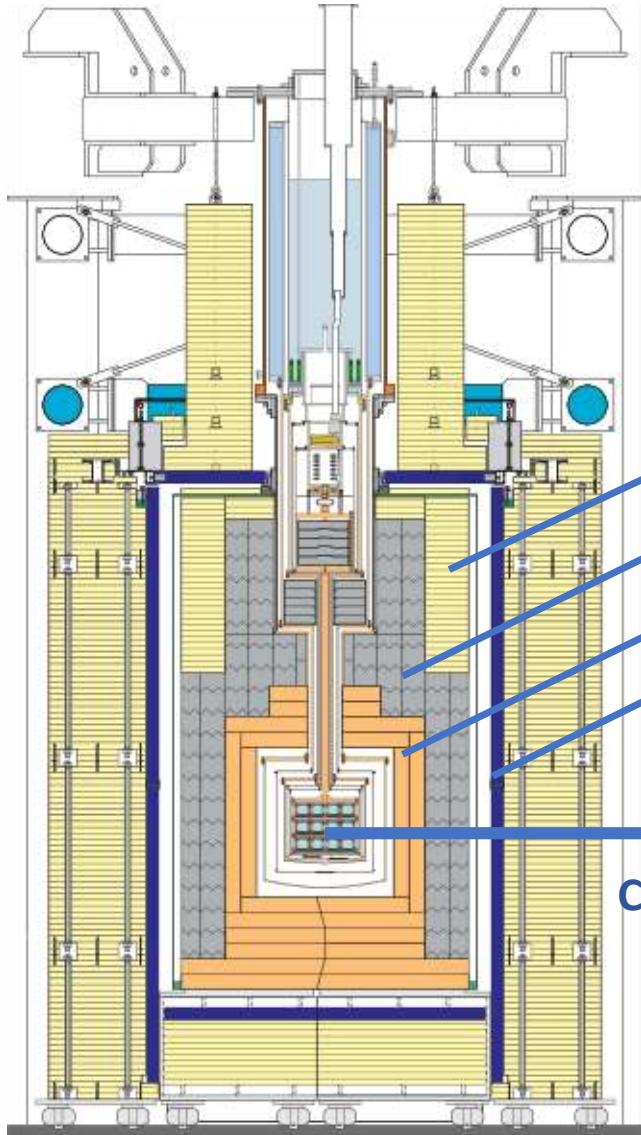
CRESST is located at LNGS (Laboratori Nazionali del Gran Sasso) in Italy

- Cryogenic scintillating calorimeter
- Target materials: CaWO_4 , Si, LiAlO_3 , Al_2O_3
- Read out channels: phonon and scintillation light

Shielding:

- polyethylene;
- lead;
- copper;
- muon veto system.

CRESST-III Phase 1



Experimental location:

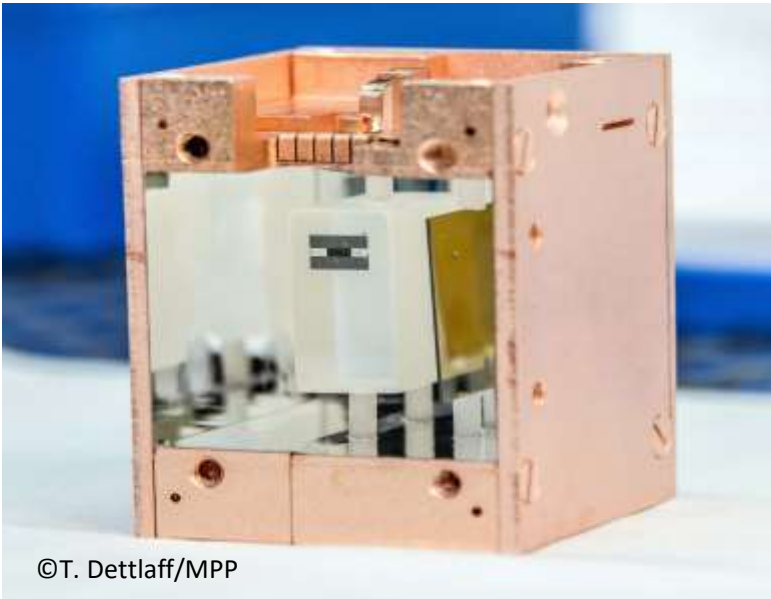
Average depth ~ 3600 m w.e.

Muon flux ~ $2.6 \times 10^{-8} \mu/\text{s}/\text{cm}^2$

Neutrons < 10 MeV: $< 10^{-6} \text{ n}/\text{s}/\text{cm}^2$

The CRESST experiment

Direct detection of dark matter particles via their scattering off target nuclei



©T. Dettlaff/MPP

- target material: CaWO_4 single crystals

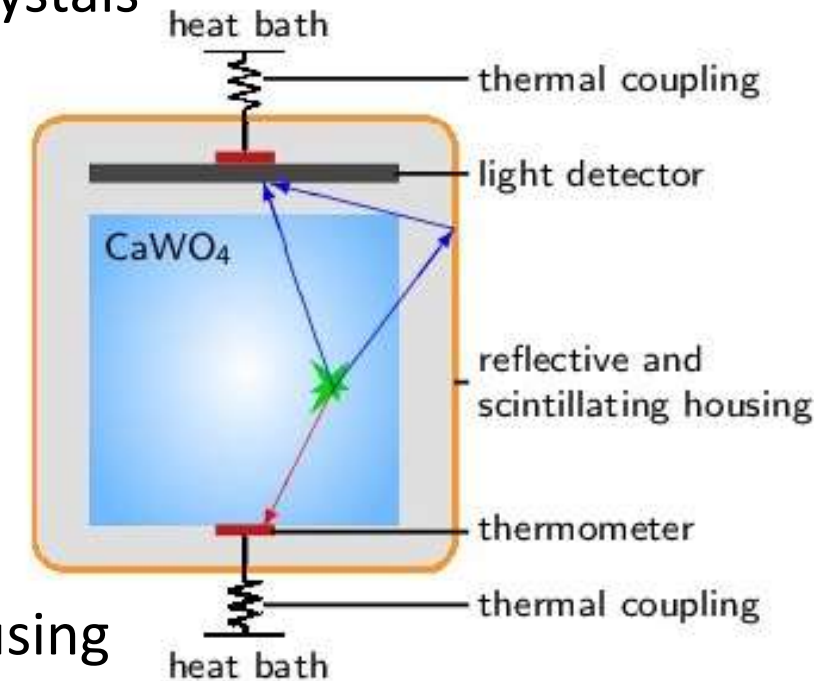
- particle interaction

→ heat (phonon) signal
read-out with thermometer

→ light signal
read-out with light detector

- reflective and scintillating housing

Target crystals operated as cryogenic calorimeters ($\sim 15\text{mK}$)



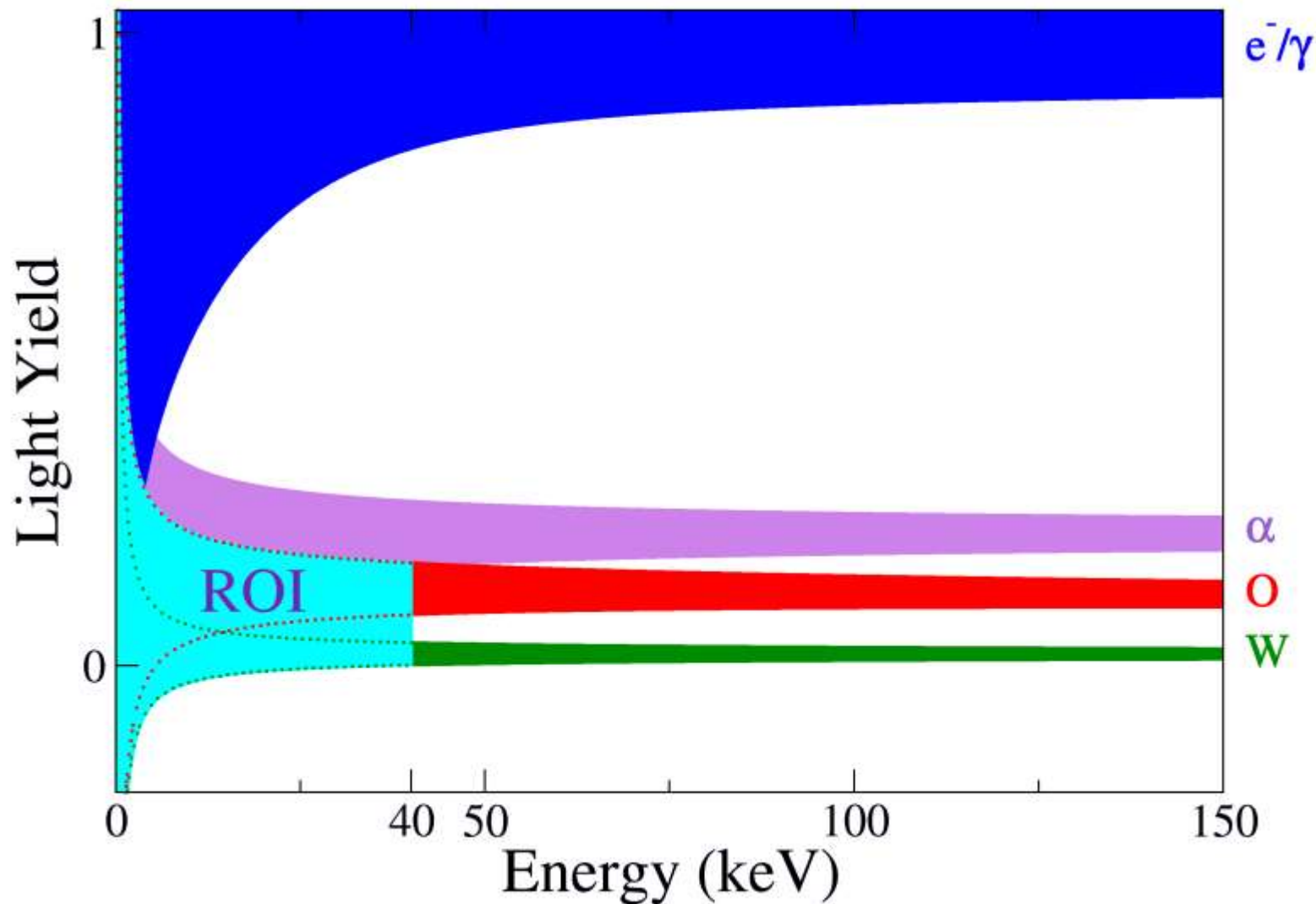
Event discrimination

$$\text{Light Yield} = \frac{\text{Light signal}}{\text{Phonon signal}}$$

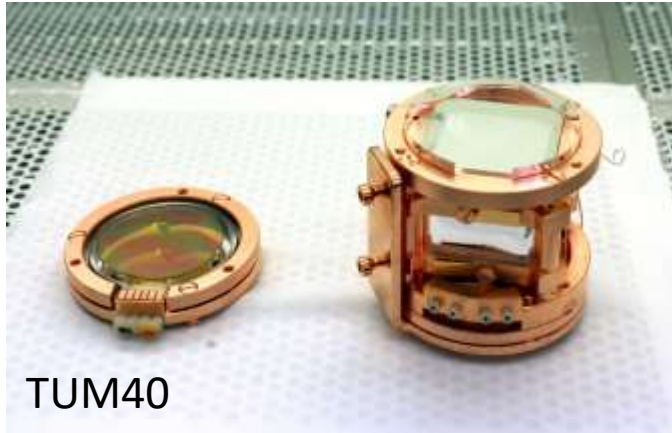
Characteristic of the event type

Excellent discrimination between potential signal events (**nuclear recoils**) and dominant radioactive background (**electron recoils**)

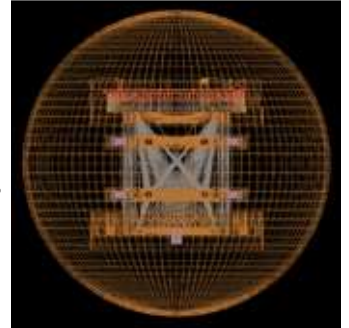
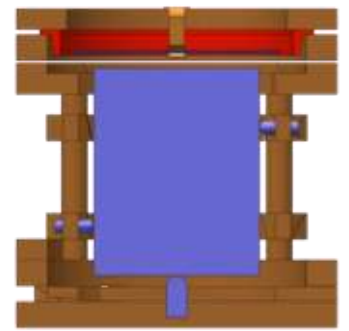
ROI: region of interest for dark matter search



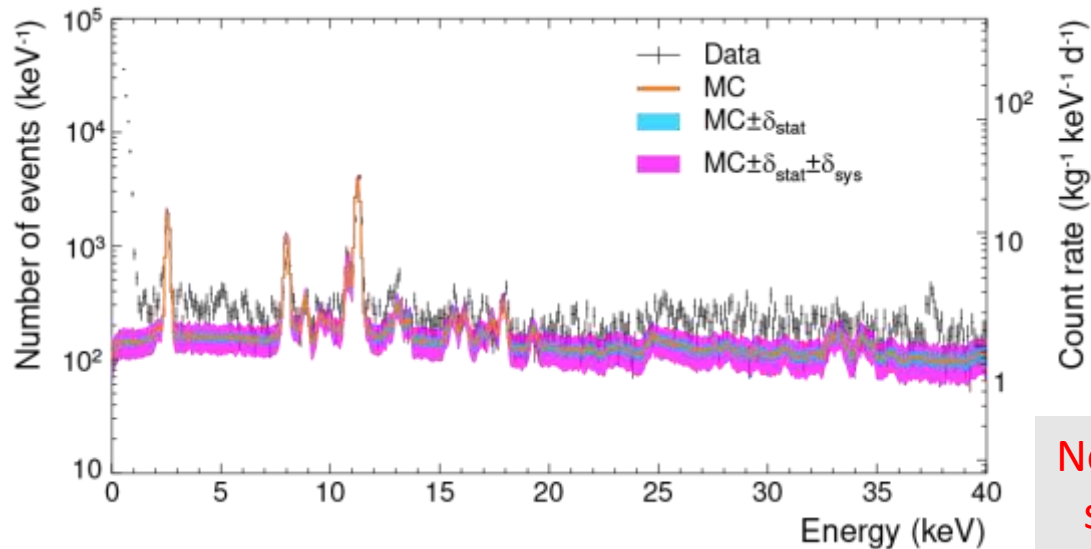
One module background simulation



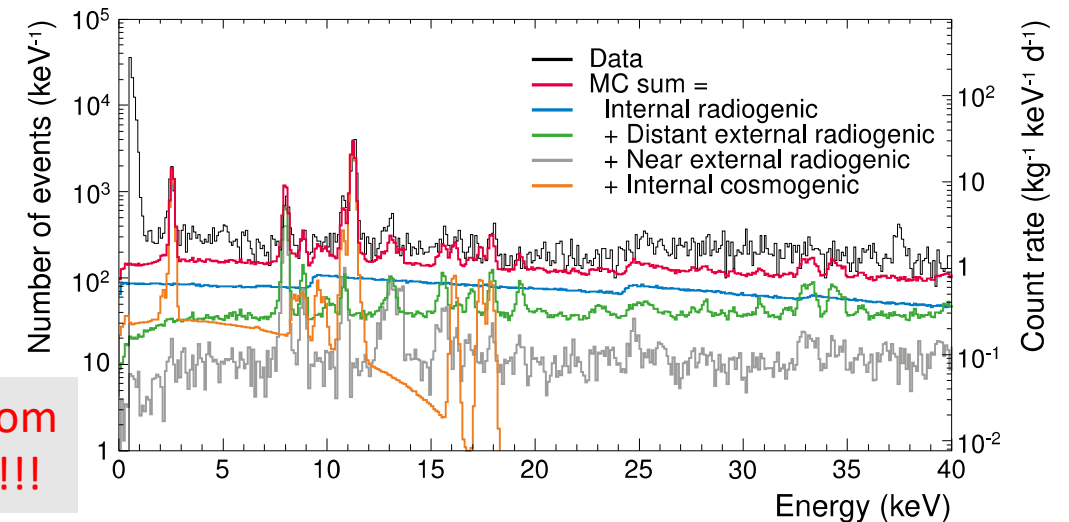
- Geant4 based electromagnetic background model for the CRESST experiment;
- Study of cosmogenic activation of CaWO_4 crystal scintillator;
- Simplified geometry reproduced already up to 68% background in ROI;
- Foundation for more detailed models of the actual CRESST detector modules.



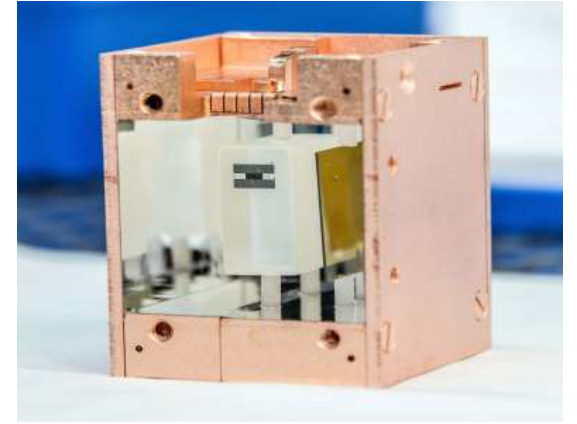
Up to $68 \pm 16\%$ of background can be reproduced with simulations



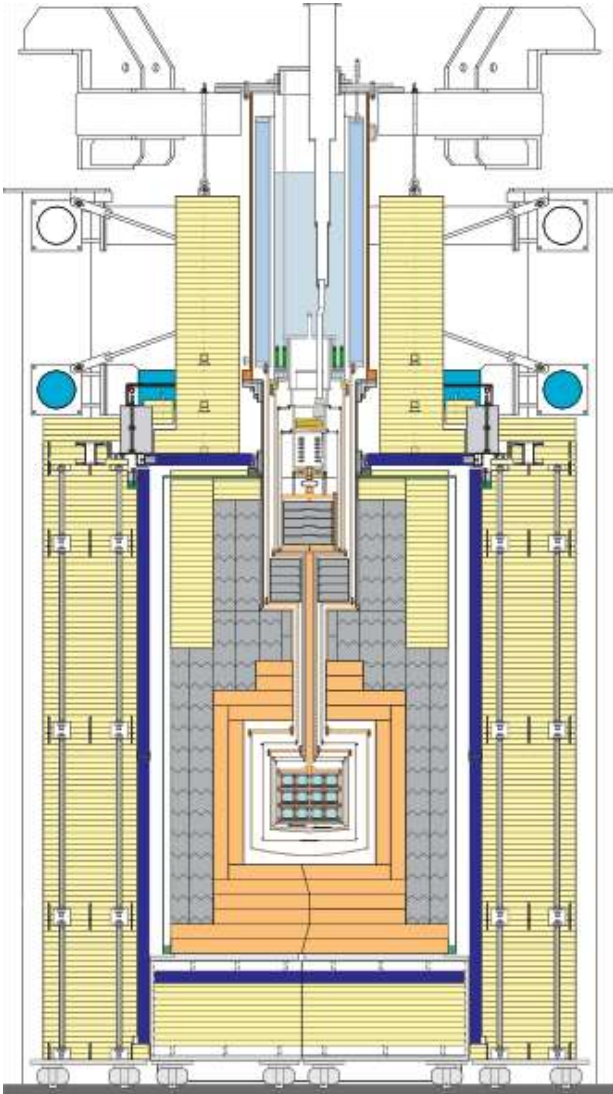
No input from screening!!!



Screening campaign



CRESST-III detector



CRESST set-up

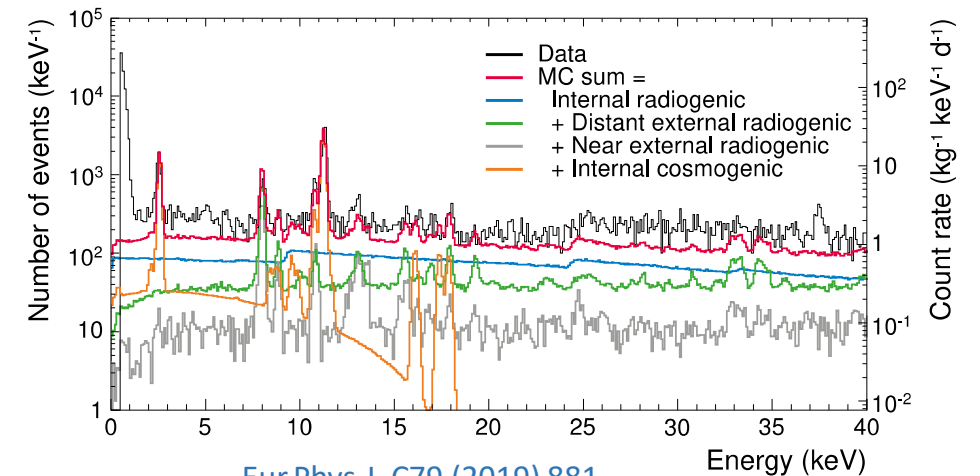
Materials:

- copper -> HPGe ✓ ICP-MS ✓ bulk ^{210}Po ✓ NAA ↻
- crystals (CaWO_4) -> bolometric meas. ↻
- reflective foil -> ICP-MS ✓
- bronze clams -> ICP-MS ✓
- polyethylene -> HPGe ✓ ICP-MS ✓
- lead -> ICP-MS ✓
- connectors -> HPGe ✓
- brass -> ICP-MS ✓
- pins -> ICP-MS ✓



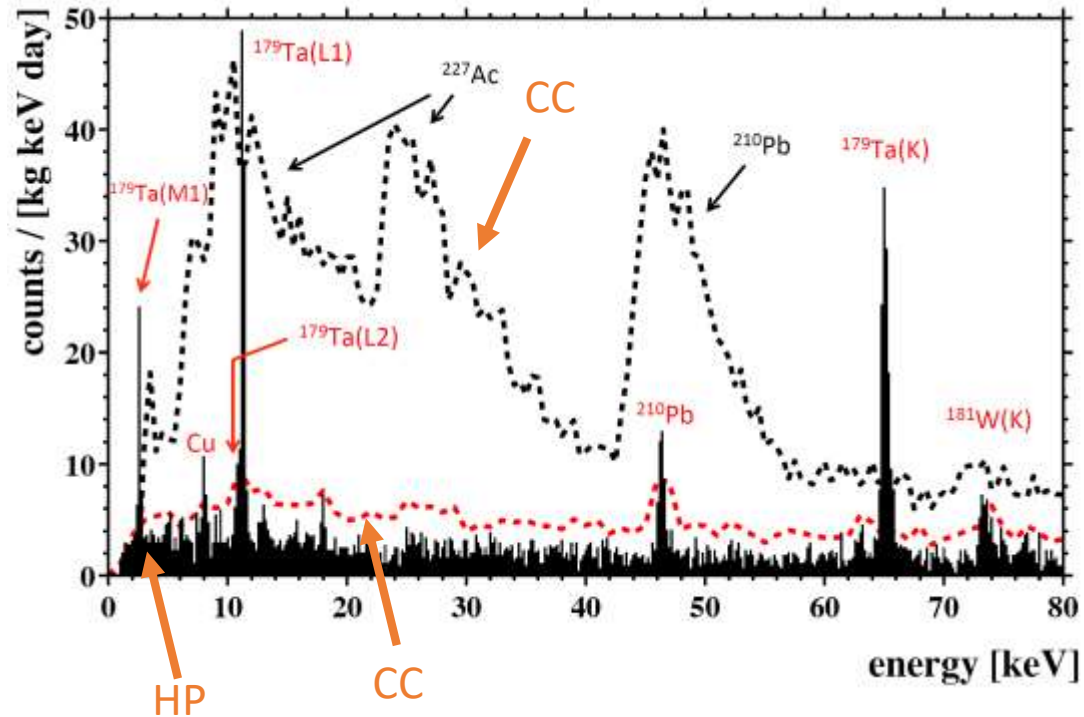
Simulation

Up to $68 \pm 16\%$ of background can be reproduced with simulations



Eur.Phys.J. C79 (2019) 881
Eur.Phys.J. C79 (2019) 987

High-purity scintillating CaWO_4 crystals



Commercial crystals (CC) \longrightarrow home production (HP)

A factor of 2-10 decrease in the background

Radiochemical Purification of CaCO_3

1. Transform CaCO_3 powder into aqueous solution of $\text{Ca}(\text{NO}_3)_2$
2. Mix solution with an extractor (TOPO) dissolved in n-Dodecan
 \rightarrow Impurities move from the $\text{Ca}(\text{NO}_3)_2$ solution to the extractor solution
3. Extraction of $\text{Ca}(\text{NO}_3)_2$ solution
4. Remove precipitated CaWO_4
5. Washing with alkaline solution and water

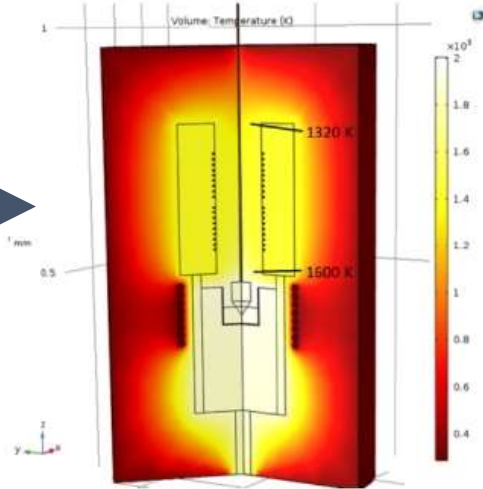
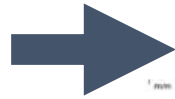


High-purity scintillating CaWO_4 crystals



CaWO_4 powder

1.5 kg of purified CaWO_4 powder produced via co-precipitation using $(\text{NH}_4)_2\text{WO}_4$ solution (NEW!)



temperature gradients during the growth to study the reduction of internal stresses

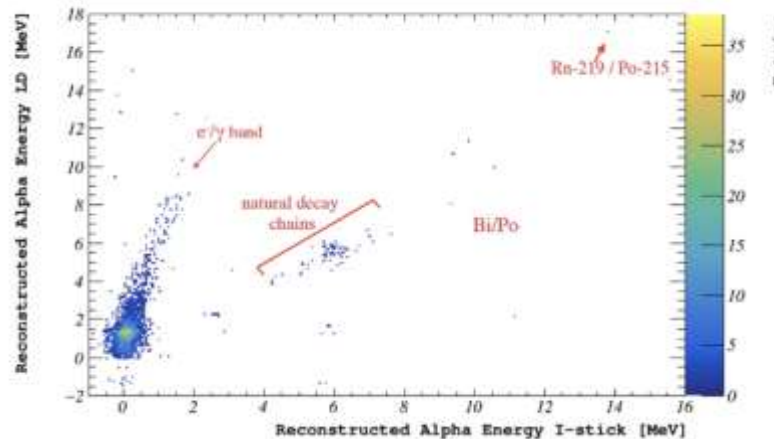


New crystal
New level of radiopurity



oxygen atmosphere at 1400°C for 20 h

	Activity, (mBq/kg)
Total alpha activity	0.535 ± 0.055
^{147}Sm and ^{180}W	0.046 ± 0.016
Single alpha lines	0.454 ± 0.051
Bi-Po cascades	0.029 ± 0.013
$^{219}\text{Rn} - ^{215}\text{Po}$ decay	0.006 (1 event)
Total alpha activity nat. decay chains	0.489 ± 0.053



Valentyna Mokina - HEPHY OEAW

Total α activity (3.08 ± 0.04) mBq/kg for TUM40 [1] and for TUM93 [2] crystal is (0.489 ± 0.05) mBq/kg.

Increase by a factor of 6.3 ± 0.7

[1] R. Strauss et al., JCAP 2015 06, 030 (2015)

[2] A. Kinast et al., J LTPHys. (2022) [10.1007/s10909-022-02743-7](https://doi.org/10.1007/s10909-022-02743-7)

Screening campaign

Sample	Th [pg/g]	U [pg/g]
1. brass screw	70 ± 21	14 ± 4
2. superconducting cable with copper matrix	77 ± 23	84 ± 25
3. bronze clamp	162 ± 49	690 ± 207
4. pins	20 000 ± 6 000	176 000 ± 53 000
5. copper circuit	734 ± 220	283 ± 85

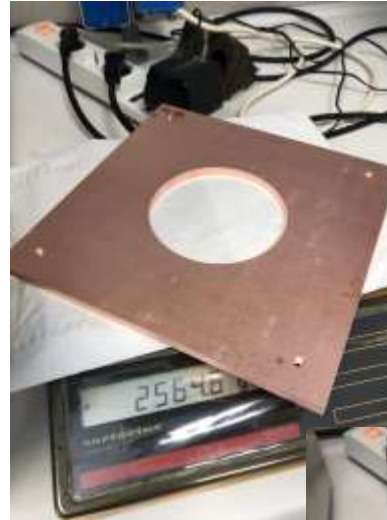


ICP-MS at LNGS

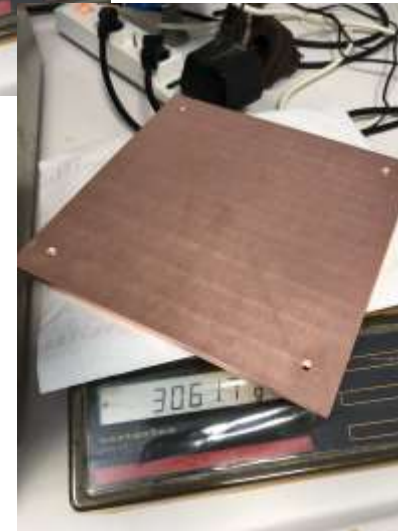


Sample	²³² Th [μBq/kg]	²³⁵ U [μBq/kg]	²³⁸ U [μBq/kg]
Copper *	< 2	< 0.43	< 6.2
Copper CUORE	< 2	< 0.46	< 65

* more sensitive result



16 plates



18 plates

Chain	Nuclide	Activity, (mBq/kg)
²³² Th	²²⁸ Ra	<0.024
	²²⁸ Th	<0.021
²³⁵ U		<0.05
²³⁸ U	²³⁴ Th	<3.5
	^{234m} Pa	<0.76
	²²⁶ Ra	<0.02
	⁴⁰ K	<0.19
	¹³⁷ Cs	<0.0056
	⁴⁶ Sc(83.8d)	0.029 ± 0.006
	⁴⁸ V(15.97d)	<0.04
	⁵⁴ Mn(312d)	0.051 ± 0.009
	⁵⁹ Fe(44.5d)	0.042 ± 0.011
	⁵⁶ Co(77.2d)	0.054 ± 0.008
	⁵⁷ Co(272d)	< 0.14
	⁵⁸ Co(70.9d)	0.5 ± 0.05
	⁶⁰ Co(5.28y)	0.046 ± 0.006

HPGe at LNGS

ICP-MS and HPGe measurements of shielding materials

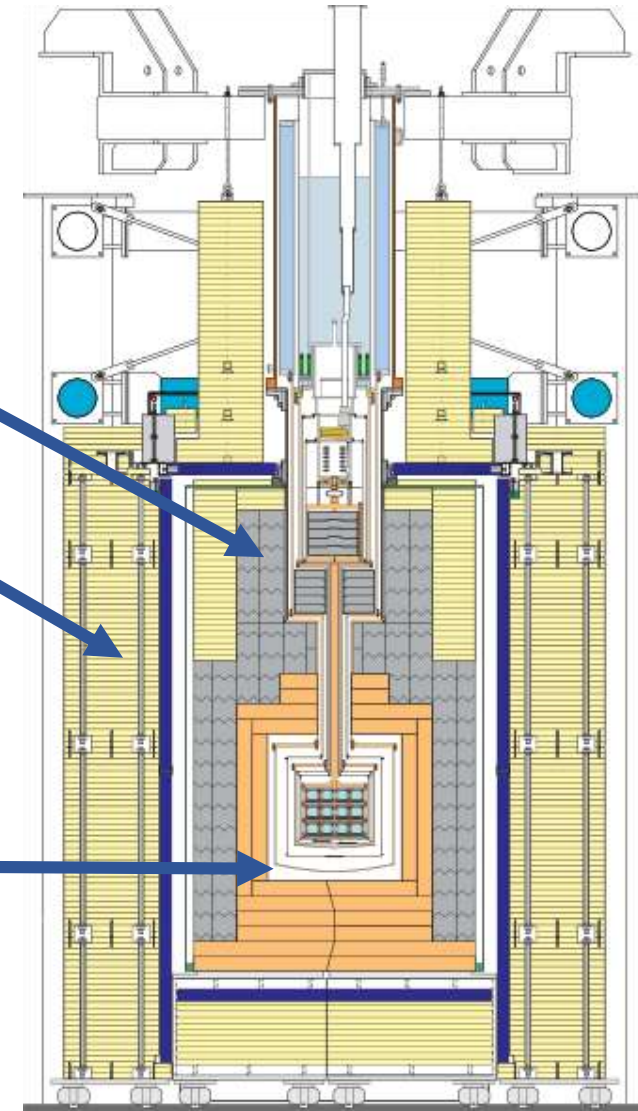
ICP-MS at LNGS

Needs an update!!

Sample	Th [pg/g]	U [pg/g]
lead	< 200	< 200
polyethylene outer	< 1300	< 500

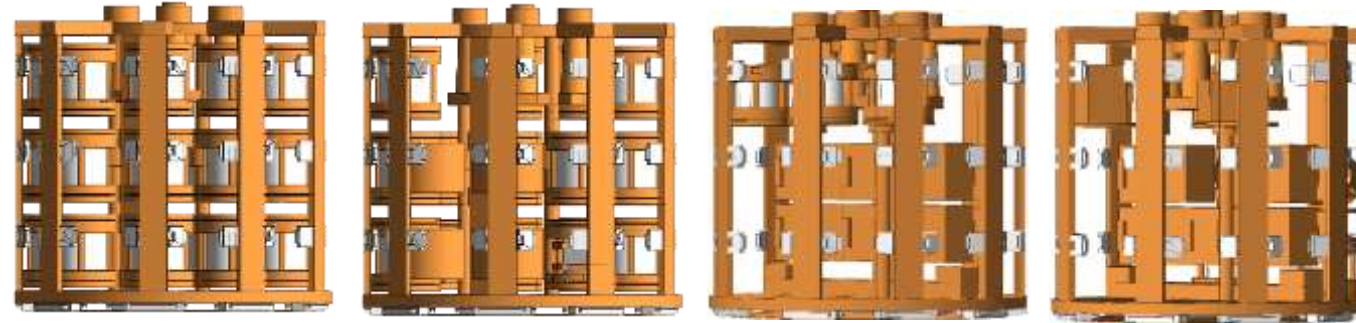
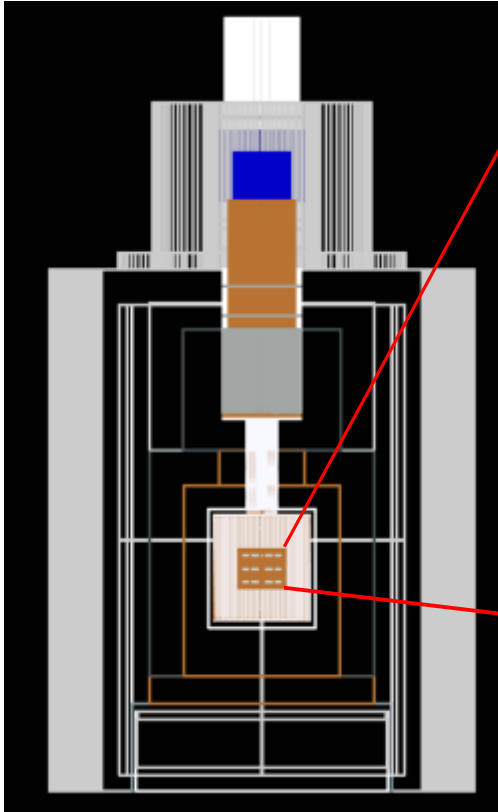
HPGe at LNGS

Sample	Th [pg/g]	U [pg/g]
polyethylene inner	100 ± 30	82 ± 9



From one module to full setup simulation

DAWN visualization of the Carousel with detectors as implemented in Geant4

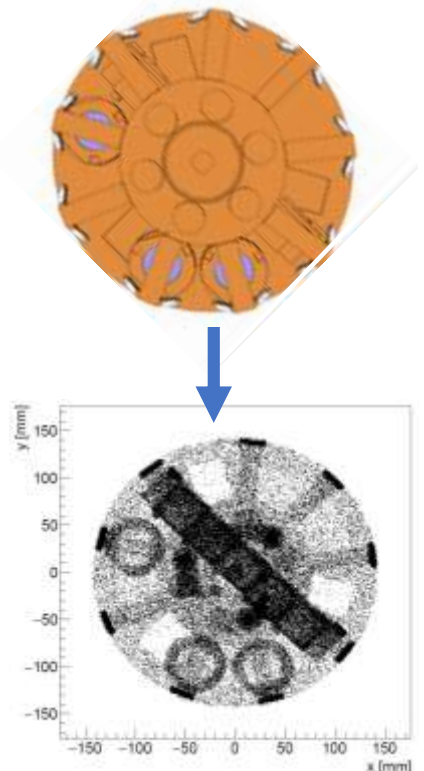


CRESST-II
phase 1
730kg result
[Eur.Phys.J.
C72\(2012\)1971](#)

CRESST-II
phase 2
TUM40, Lise
[Eur.Phys.J.
C74\(2014\)3184,
C76\(2016\)25](#)

CRESST-III
phase 1
Detector A
[Phys.Rev.
D100\(2019\)102002](#)

CRESST-III
phase 1 Run 2



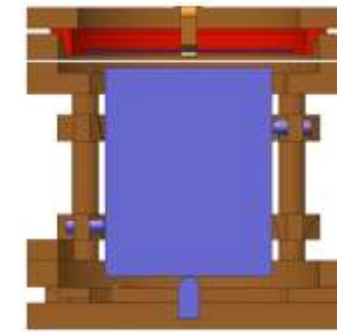
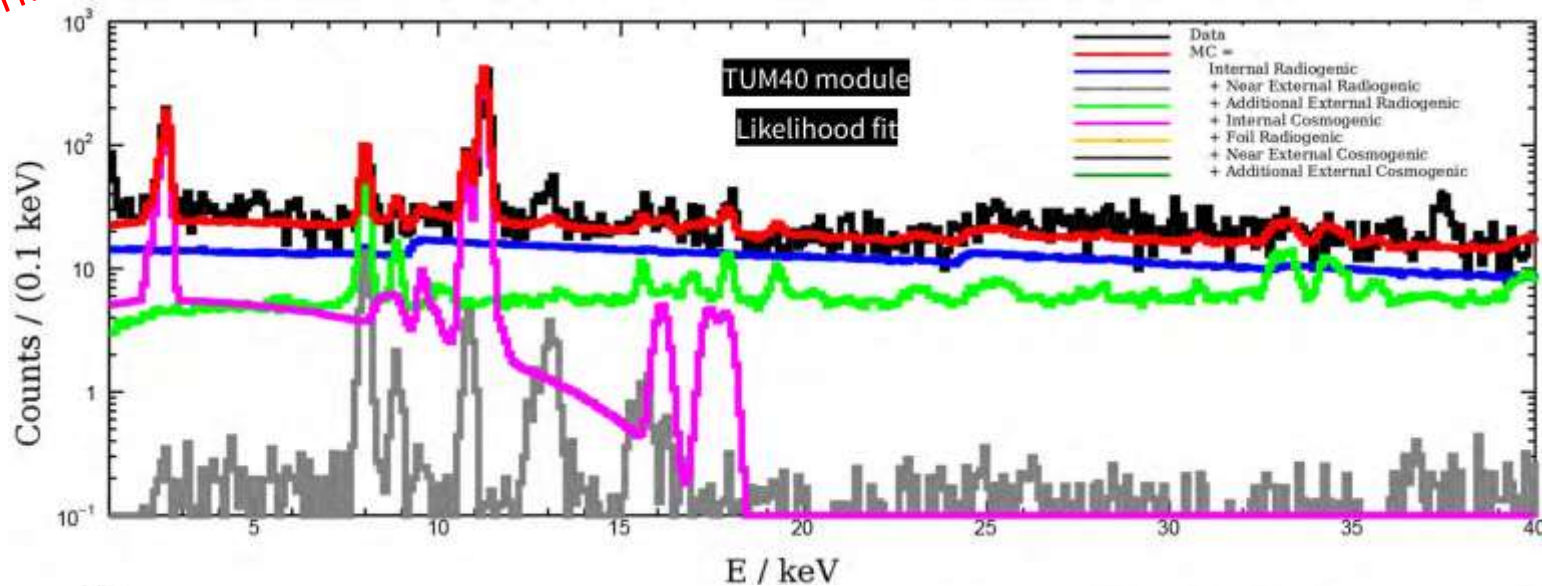
CRESST set-up

- Implementation of full geometry (each Run, all detectors);
- Adaption of the e.m. background model to the actual CRESST detector modules;
- Contamination levels from material assays conducted within;
- Simulation of neutron background;
- Study of cosmogenic activation of CaWO_4 crystal scintillator;
- Surface contaminations studies.

Simulating a homogeneous contamination in all parts made of Cu inside the Carousel.

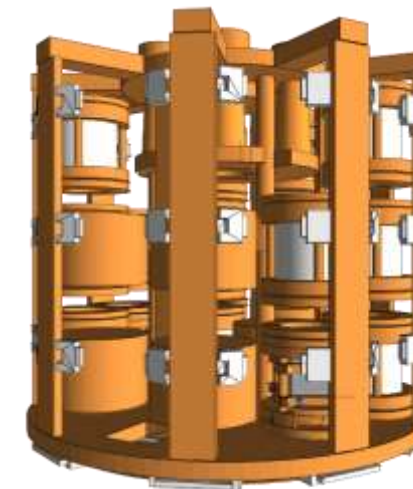
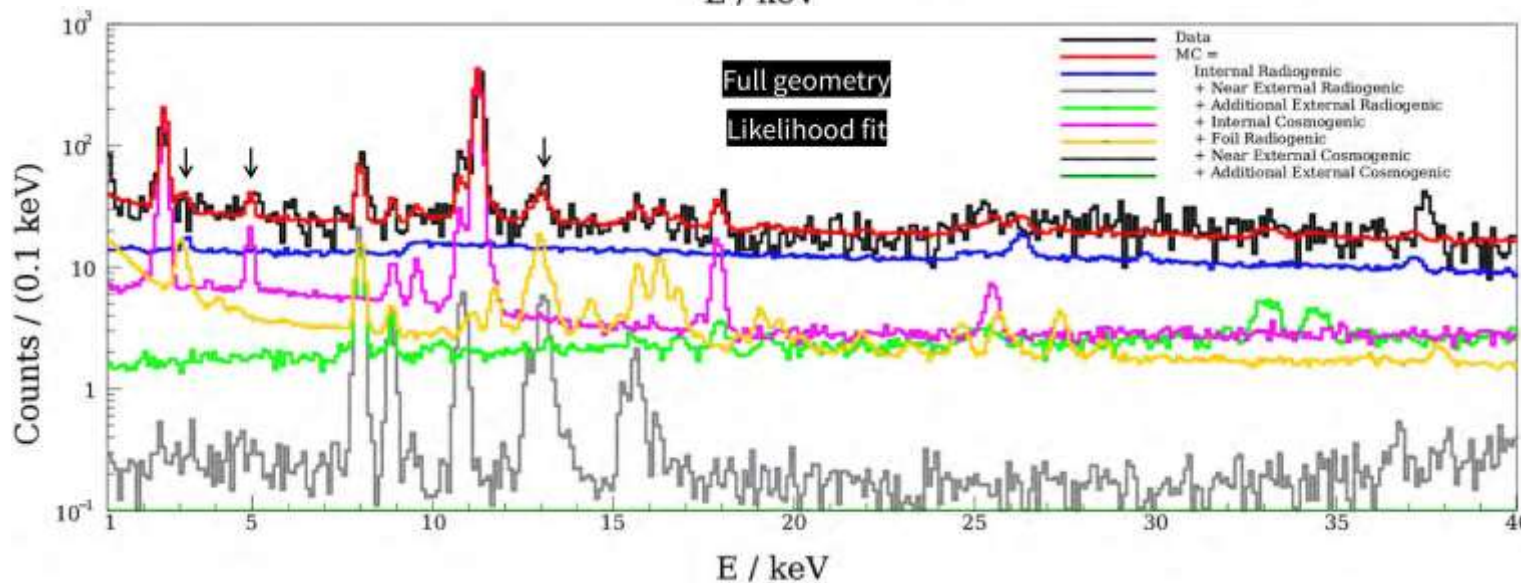
Work in progress

Likelihood model of background



- Bronze
- Copper
- BC400
- CaWO₄
- Sapphire

Up to $68 \pm 16\%$ of background can be reproduced with simulations



$97 \pm 0.3\%$

Publication in preparation

Conclusions

- CRESST operates a new generation of TUM-grown crystals with improved radiopurity due to chemical purification of their raw materials;
- The screening campaign is ongoing to understand the activity concentration of different isotopes in materials used in the experiment;
- The results of these studies are used as an input for simulation of the background of the CRESST experiment (development of sub-keV Monte Carlo model).
- Likelihood model of the background is developed and allows to reproduce up to 97%.

Waiting for dark matter

