



Overview on RPV SUSY

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Roadmap

- The RPV MSSM:
Of Theorists' Dreams & Experimentalists' Nightmares
- Pheno 1: Taming the RPV Landscape
- Pheno 2: The Future is Bright
- Pheno 3: Let there be Light (SUSY)
- Outlook

The RPV MSSM

- Minimal SUSY: Poincare (+SUSY) invariance, SM gauge, MSSM content

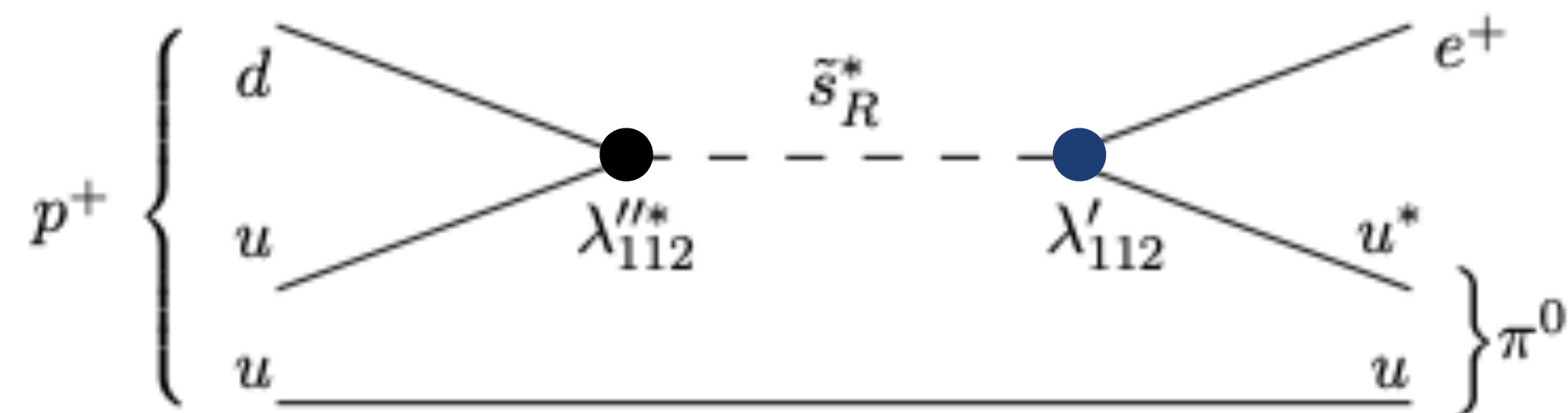
$$W = W_{\text{MSSM}} + \lambda_{ijk} L^i L^j E^k + \lambda'_{ijk} L^i Q^j D^k + \kappa_i L^i H^d \quad \text{LNV}$$

$$+ \lambda''_{ijk} U^i D^j D^k \quad \text{BNV}$$

- LNV + BNV

Proton decay

Bound $\sim 10^{34}$ yrs



The RPV MSSM

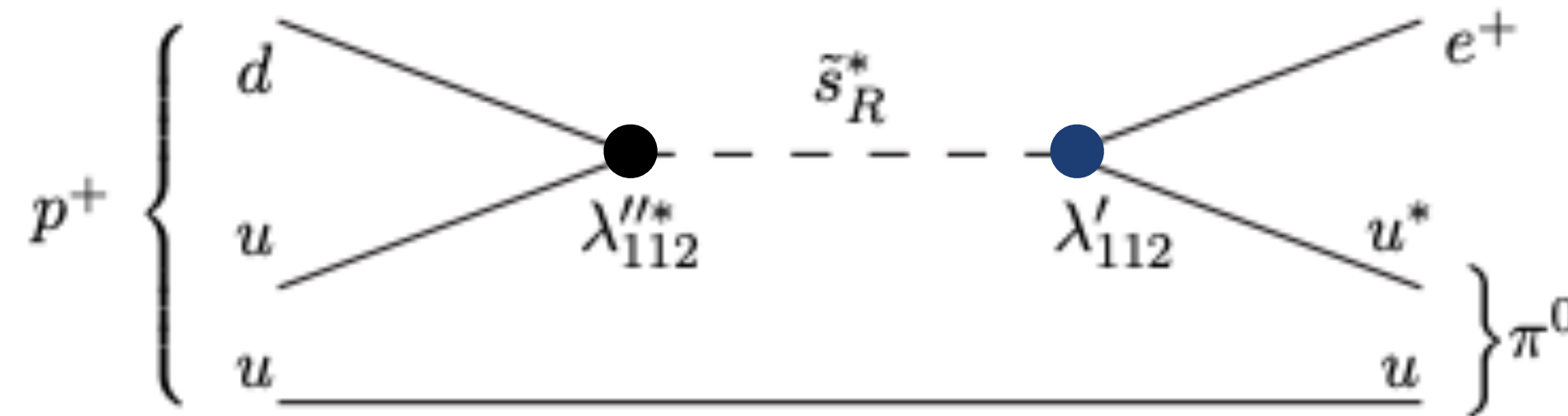
- MSSM: Impose R -parity

$$R = (-1)^{3(B-L)+2s} = \begin{cases} +1 & \text{for SM particles, higgses} \\ -1 & \text{for SUSY particles} \end{cases}$$

- $W = W_{\text{MSSM}} + W_{\text{LNV}} + W_{\text{BNV}}$ Proton stable! (@ tree-level)

The RPV MSSM

- But to protect the proton removing either ● or ● can be sufficient



- Indeed R is excessive! Baryon triality B_3 protects the proton (even better than R !)

$$W_{\text{RPV}} = W_{\text{MSSM}} + W_{\text{LNV}} + W_{\text{BNV}}$$

The RPV MSSM

Drastic changes RPV vs MSSM:

- Lightest SUSY particle (LSP) is stable -> **no longer!**
(odd in R)

- LSP is neutralino -> **no longer!**
(no longer DM candidate)

$$\text{LSP} \in \{\chi_1^0, \chi_1^\pm, \tilde{\nu}_L, \tilde{\ell}_{L,R}^\pm, \tilde{\tau}_1^\pm, \tilde{q}_{L,R}, \tilde{t}_1, \tilde{g}\}$$

- **Single SUSY production possible**

Colliders can
have 2x reach!

SUSY signature
no longer MET!

The RPV MSSM: A Theorist's Dream

RPV has many nice features; all following from minimal SUSY:

- Ingredients for baryogenesis
- Ingredients for neutrino masses
- Ingredients for LFUV, muon ($g-2$), leptoquarks, etc.
- Playground for many new opportunities: lepton PDFs, Long Lived Particles, etc.

The RPV MSSM: An Experimentalist's Nightmare

- Vanilla MSSM:

SUSY pair production \rightarrow Decay to Neutralino LSP \rightarrow MET signature

- RPV SUSY:

$$\text{sig.} = \left(\begin{array}{c} \tilde{q}\tilde{q} \\ \tilde{q}\tilde{g} \\ \tilde{g}\tilde{g} \\ \tilde{\ell}^+\tilde{\ell}^- \\ \tilde{\nu}\tilde{\nu} \\ \tilde{\chi}^0\tilde{\chi}^\pm \end{array} \right)_{\text{prod}} \otimes \left(\begin{array}{c} \tilde{\chi}_1^0 \\ \tilde{\chi}_1^\pm \\ \tilde{\nu}_i \\ \tilde{\ell}_i^\pm \\ \tilde{\tau} \\ \tilde{q} \\ \tilde{b} \\ \tilde{t} \\ \tilde{g} \end{array} \right)_{\text{possible LSP}} \otimes \left(\begin{array}{c} L_1 L_2 \bar{E}_1 \\ \dots \\ L_1 Q_1 \bar{D}_1 \\ \dots \\ \bar{U}_3 \bar{D}_2 \bar{D}_3 \end{array} \right)_{\text{LSP decay}}$$

45 couplings!

Pheno 1: Taming the RPV Landscape

How well can we exclude SUSY?



Vanilla MSSM

- > Missing energy searches
- > Coloured sector upto $\mathcal{O}(\text{TeV})$

RPV MSSM

- > ?
- > ?

Have we systematically investigated RPV MSSM?

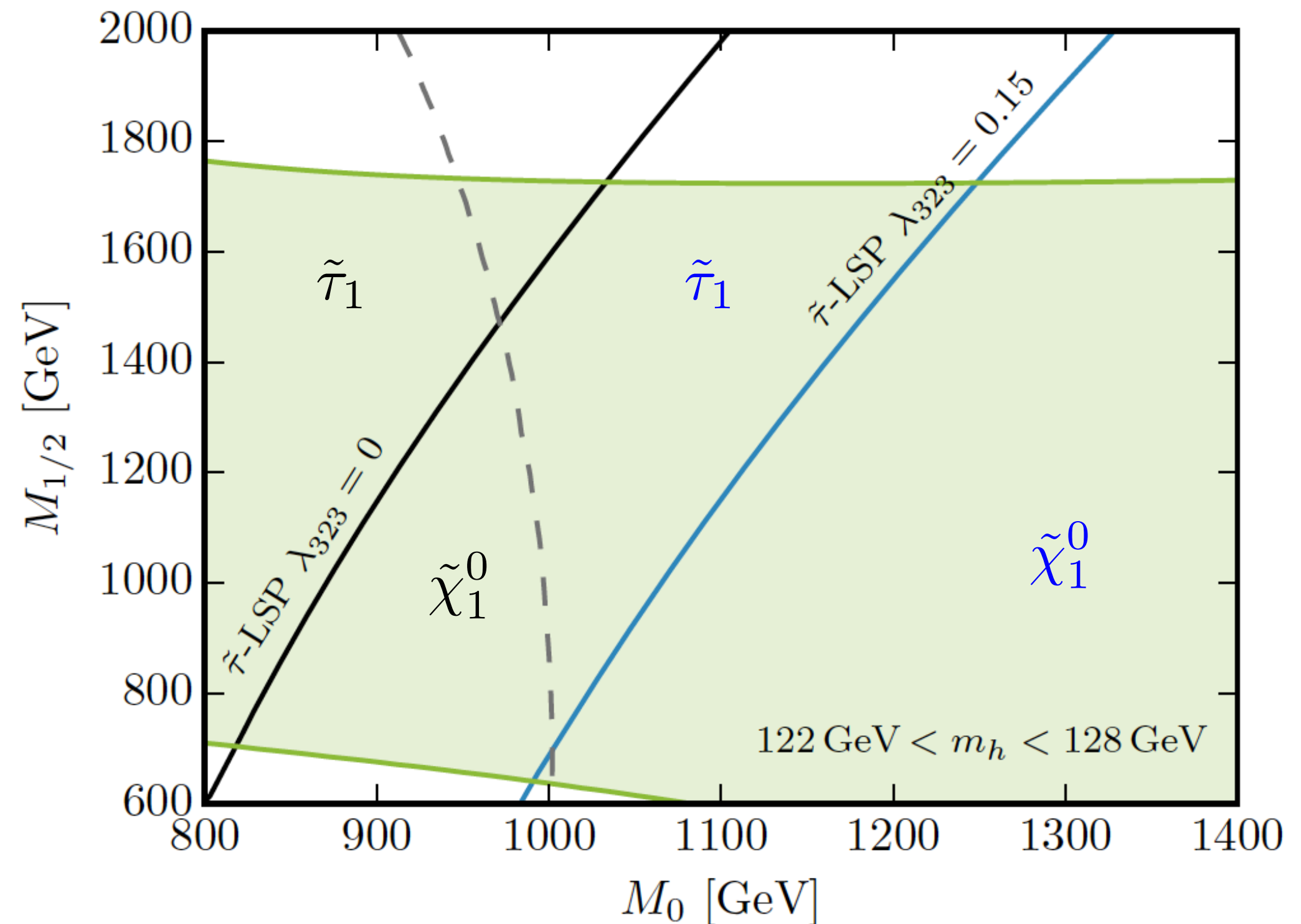
[arXiv: 1706.09418](https://arxiv.org/abs/1706.09418) [Dreiner et al.]

Pheno 1: Taming the RPV Landscape

Assumptions:

- Work in CMSSM
Spectrum fixed by 5 parameters
LSPs: $\tilde{\chi}_1^0$, $\tilde{\tau}$ mainly

LSP	Required Couplings
$\tilde{\chi}_1^0$	$\Lambda_{\mathcal{R}_p} \ll 1$ or large M_0
$\tilde{\tau}_1$	$\Lambda_{\mathcal{R}_p} \ll 1$, small M_0 and large $M_{1/2}$
$\tilde{\tau}_1$	λ_{ij3} (dominantly $\tilde{\tau}_R$), λ'_{3jk} ($\tilde{\tau}_L$)
\tilde{e}_R	λ_{ij1}
$\tilde{\mu}_R$	λ_{ij2}
$\tilde{\nu}_e$	λ'_{1jk} , $\{j, k\} \neq \{1, 1\}^\ddagger$
$\tilde{\nu}_\mu$	λ'_{2jk}
\tilde{s}_R, \tilde{d}_R	λ''_{212} (degenerate LSPs)
\tilde{b}_1	$\lambda''_{123}, \lambda''_{213}, \lambda''_{223}^\ddagger$ (dominantly \tilde{b}_R)
\tilde{t}_1	λ''_{3jk} (dominantly \tilde{t}_R)



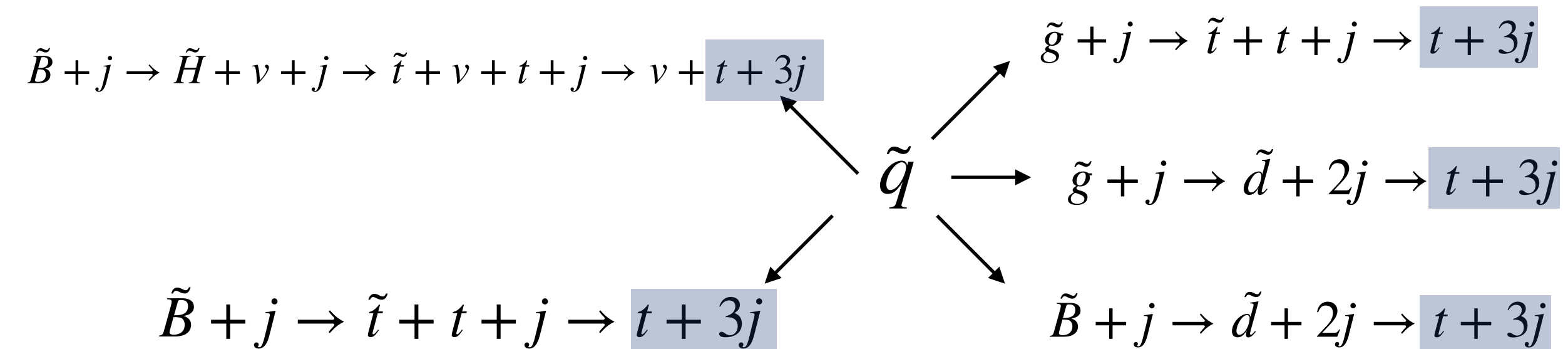
$$A_0 = -3 \text{ TeV}, \tan \beta = 30$$

Pheno 1: Taming the RPV Landscape

Main take away message:

- Can reduce RPV scenarios to set of few signatures

$$\{\text{LSP, coupling}\} = \{\tilde{q}, \lambda''_{312}\}$$



- Used this to identify gaps: Few in $\tilde{\chi}_0^1$ case; more in $\tilde{\tau}$
- Recast existing searches to cover gaps: RPV coverage is comparable to RPC for CMSSM!

Pheno 1: Taming the RPV Landscape

LSP	LLE (I)	LL ₃ E (II)	LLE ₃ (III)	LL ₃ E ₃ (IV)
$\bar{l}(\bar{\nu})$	$(3l + E_T^{\text{miss}}) / 4l$	$(2l + 1\tau + E_T^{\text{miss}}) / (2l + 2\tau)$	$(1l + 2\tau + E_T^{\text{miss}}) / (2l + 2\tau)$	$(3\tau + E_T^{\text{miss}}) / 4\tau$
\bar{e}	$(2l + E_T^{\text{miss}})$	$(2l + E_T^{\text{miss}}) / (1l + 1\tau + E_T^{\text{miss}})$	$4l + 2\tau + E_T^{\text{miss}}$	$(4l + 2\tau + E_T^{\text{miss}}) / (3l + 3\tau + E_T^{\text{miss}})$
$\bar{\tau}(\bar{\nu}_\tau)$	$(4l + 2\tau + E_T^{\text{miss}}) / (4l + 1\tau + E_T^{\text{miss}})$	$(3l + E_T^{\text{miss}}) / 4l$	$(2l + 4\tau + E_T^{\text{miss}}) / (2l + 3\tau + E_T^{\text{miss}})$	$(2l + 2\tau) / (1l + 2\tau + E_T^{\text{miss}})$
$\bar{\tau}_R$	$4l + 2\tau + E_T^{\text{miss}}$	$(4l + 2\tau + E_T^{\text{miss}}) / 3l + 3\tau + E_T^{\text{miss}}$	$2l + E_T^{\text{miss}}$	$(2l + E_T^{\text{miss}}) / (1l + 1\tau + E_T^{\text{miss}})$
\bar{g}	$4l + 4J + E_T^{\text{miss}}$	$(4l + 4J + E_T^{\text{miss}}) / (3l + 1\tau + 4J + E_T^{\text{miss}})$	$2l + 2\tau + 4J + E_T^{\text{miss}}$	$(2l + 2\tau + 4J + E_T^{\text{miss}}) / (1l + 3\tau + 4J + E_T^{\text{miss}})$
$\bar{q}/\bar{u}/\bar{d}$	$4l + 2j_1 + E_T^{\text{miss}}$	$(4l + 2j_1 + E_T^{\text{miss}}) / (3l + 1\tau + 2j_1 + E_T^{\text{miss}})$	$2l + 2\tau + 2j_1 + E_T^{\text{miss}}$	$(2l + 2\tau + 2j_1 + E_T^{\text{miss}}) / (1l + 3\tau + 2j_1 + E_T^{\text{miss}})$
$\bar{t}_L(\bar{b}_L)/\bar{t}_R$	$(4l + 2j_3 + E_T^{\text{miss}})$	$(4l + 2j_3 + E_T^{\text{miss}}) / (3l + 1\tau + 2j_3 + E_T^{\text{miss}})$	$(2l + 2\tau + 2j_3 + E_T^{\text{miss}})$	$(2l + 2\tau + 2j_3 + E_T^{\text{miss}}) / (1l + 3\tau + 2j_3 + E_T^{\text{miss}})$
\bar{b}_R	$(4l + 2b + E_T^{\text{miss}})$	$(4l + 2b + E_T^{\text{miss}}) / (3l + 1\tau + 2b + E_T^{\text{miss}})$	$2l + 2\tau + 2b + E_T^{\text{miss}}$	$(2l + 2\tau + 2b + E_T^{\text{miss}}) / (1l + 3\tau + 2b + E_T^{\text{miss}})$
$\bar{B}/\bar{W}/\bar{H}$	$4l + E_T^{\text{miss}}$	$(4l + E_T^{\text{miss}}) / (3l + 1\tau + E_T^{\text{miss}})$	$2l + 2\tau + E_T^{\text{miss}}$	$(2l + 2\tau + E_T^{\text{miss}}) / (1l + 3\tau + E_T^{\text{miss}})$

Recently made more general

- All LSPs
- No assumptions about model

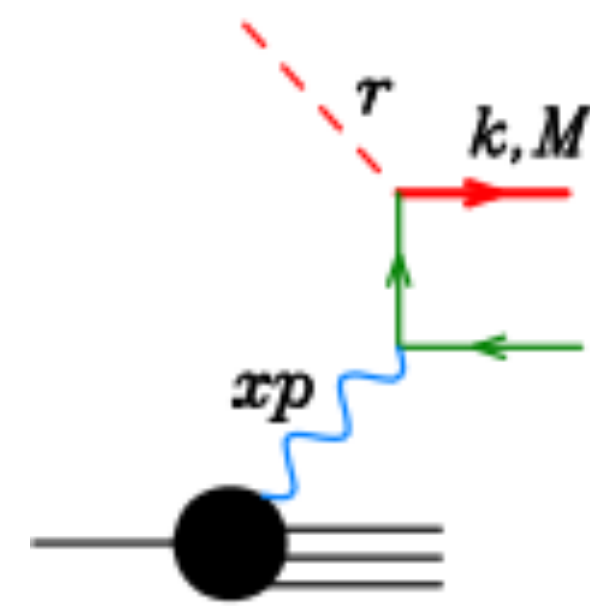
[work with H Dreiner, D Köhler, KY Sheng
in collab. with ATLAS, CMS]

- l : e/μ
- L : $e/\mu/\tau$
- j_i : light jet
- b : bottom jet
- t : top jet
- j_3 : b/t
- J : j_1/j_3

1. Four leptons 1: $4L + E_T^{\text{miss}}$ (- or + 2J or +4J)
2. Four leptons 2: $4L$
3. Two leptons: $2L + E_T^{\text{miss}}$
4. Three leptons: $3L + E_T^{\text{miss}}$
5. Five leptons: $5L + E_T^{\text{miss}}$
6. Six leptons: $6L + E_T^{\text{miss}}$

Pheno 2: The Future is Bright

- Lepton PDFs in proton (**suppressed** but...)

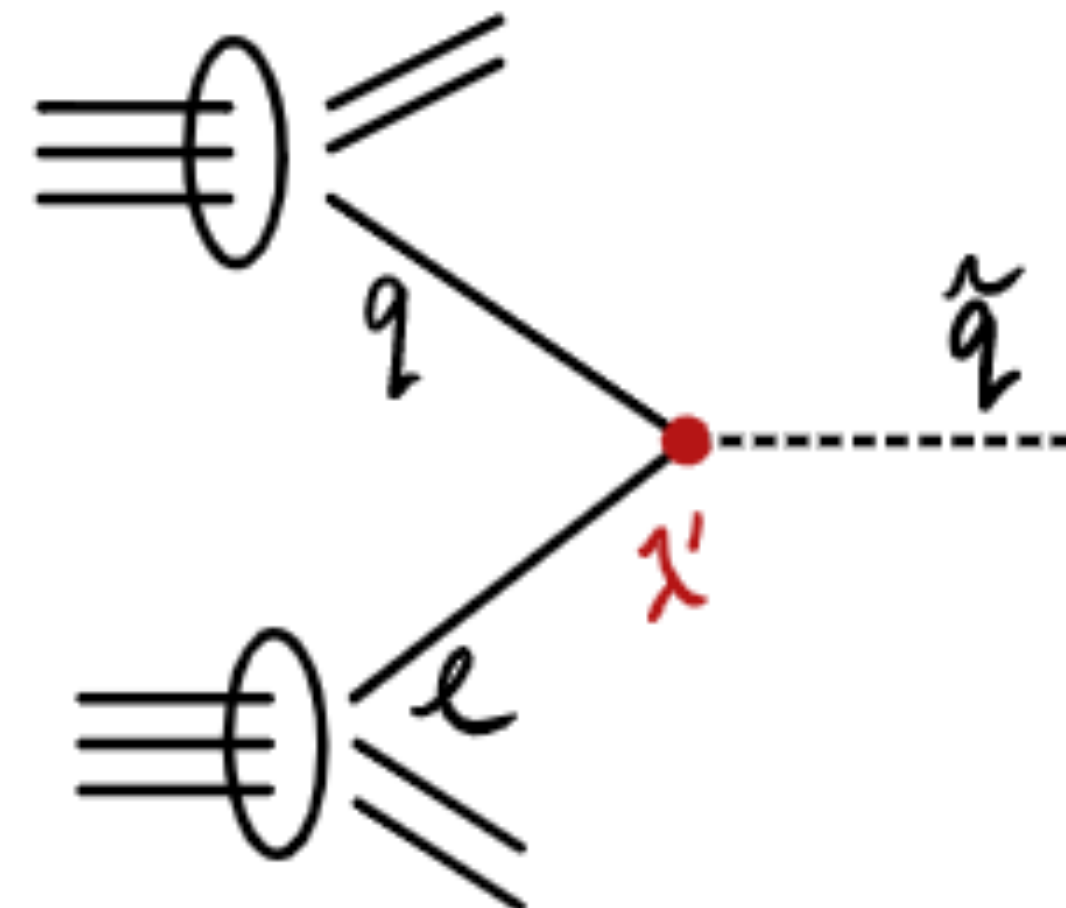


2 powers of α_{EM}

arXiv: 2005.06477 [Buonocore et al.]

- HL-LHC can overcome rare
Currently 200 fb^{-1} , HL-LHC: 250 fb^{-1} per yr

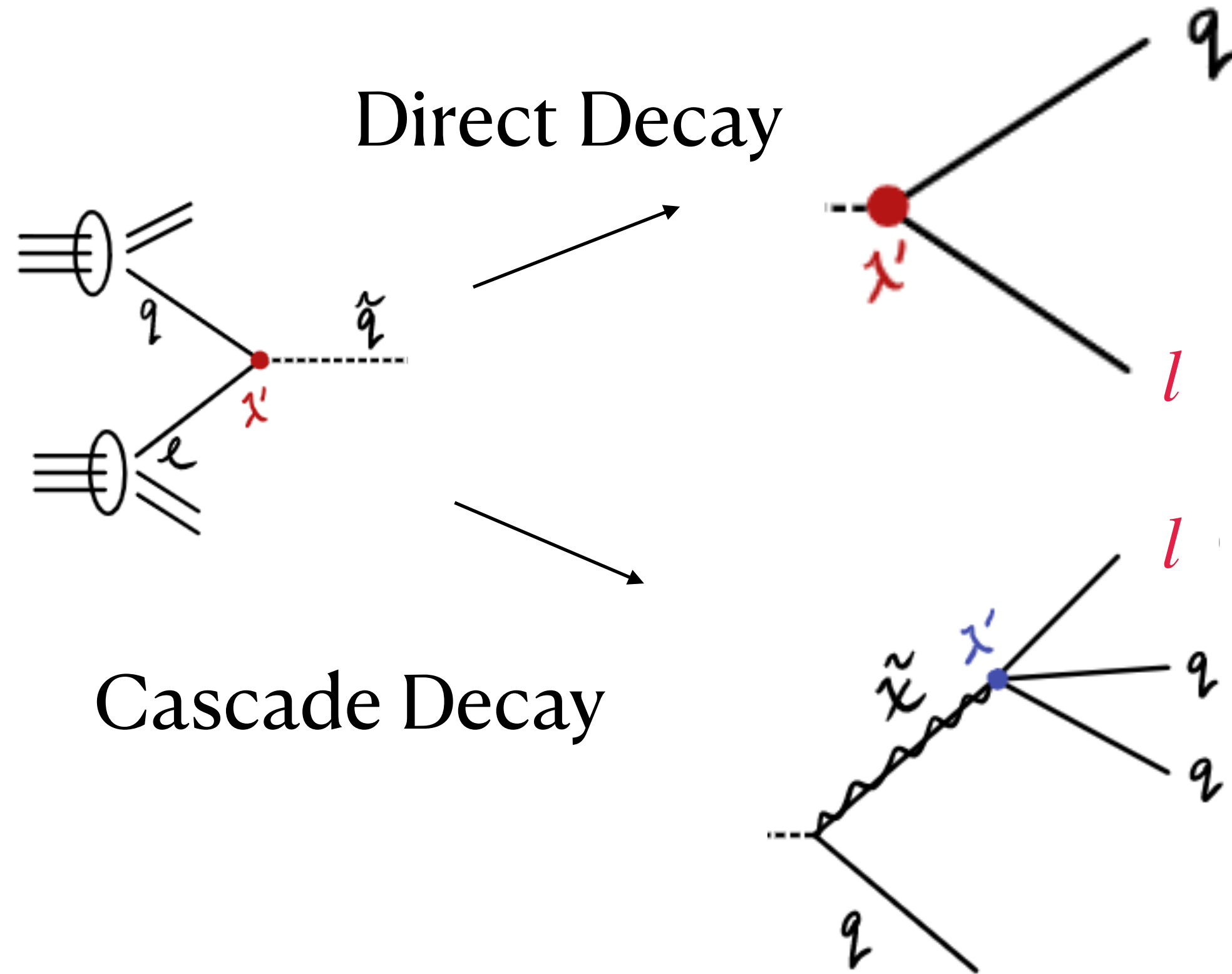
- Single production possible in RPV



2x Kinematic reach!

arXiv: 2112.12755 [work with H Dreiner, T Opferkuch, VM Lozano]

Pheno 2: The Future is Bright

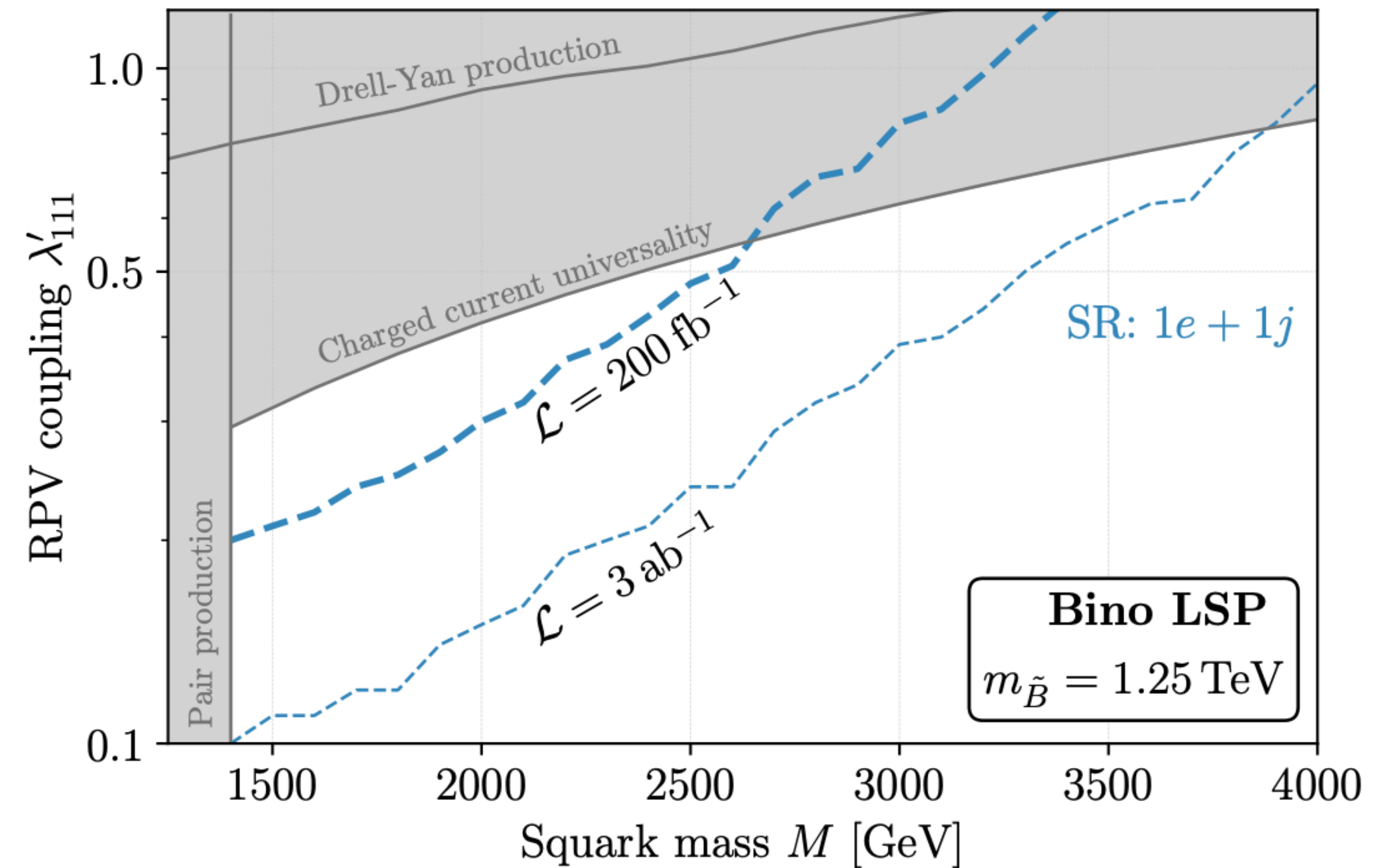
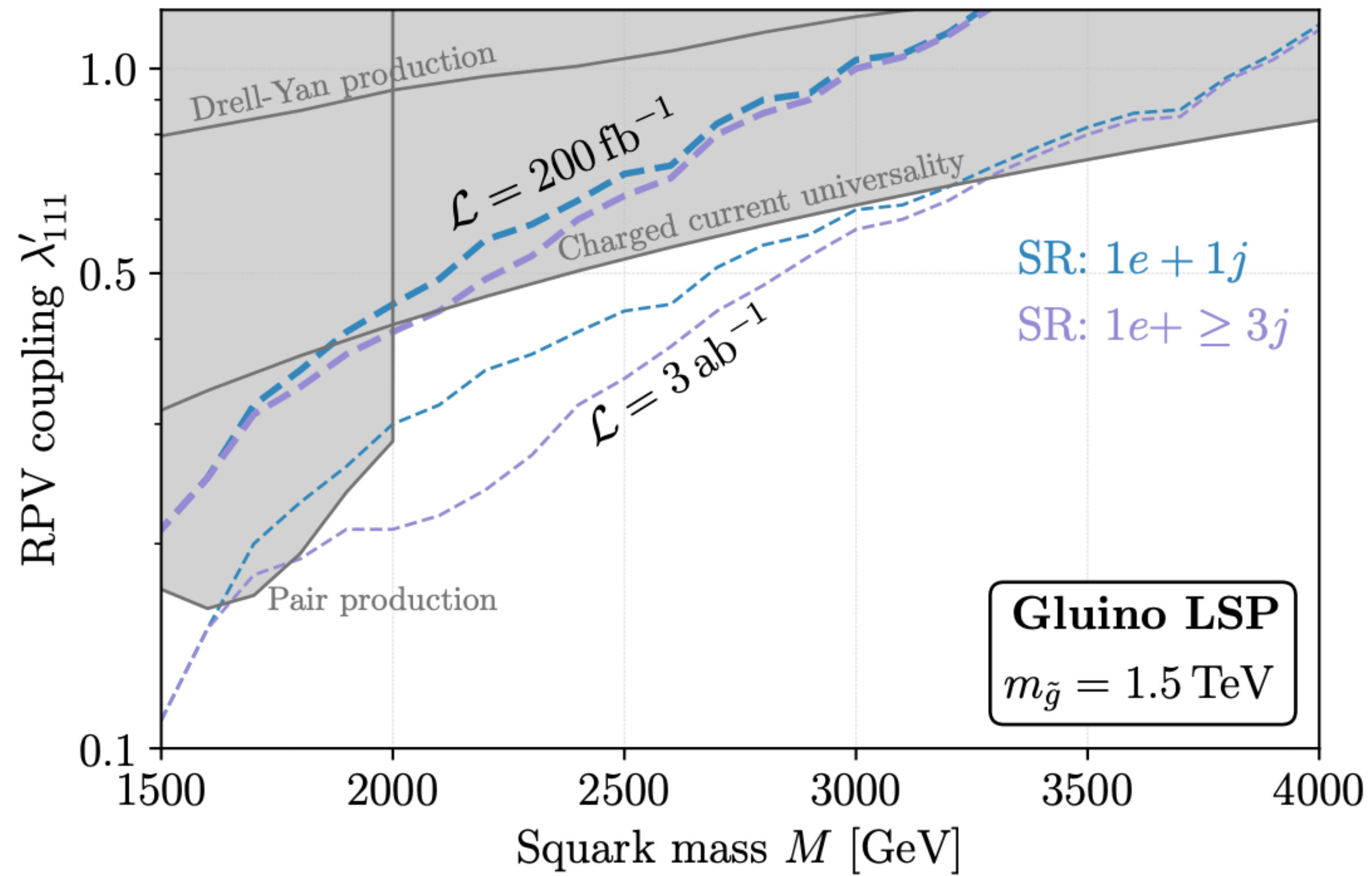


Both Single
Lepton Channels!

Very model independent!

Cascade End	Example	Signal
\tilde{B}	$\tilde{d} \rightarrow \tilde{B} + 1j$	$1l + 3j$
\tilde{W}	$\tilde{d} \rightarrow \tilde{g} + 1j \rightarrow \tilde{q} + 2j \rightarrow \tilde{W} + 3j$	$1l + 5j$
\tilde{g}	$\tilde{d} \rightarrow \tilde{g} + 1j$	$1l + 3j$
\tilde{q}	$\tilde{d} \rightarrow \tilde{g} + 1j \rightarrow \tilde{q} + 2j$	$1l + 3j$
\tilde{d}	—	$1l + 1j$
\tilde{u}	$\tilde{d} \rightarrow \tilde{g} + 1j \rightarrow \tilde{u} + 2j$	$1l + 5j$
\tilde{l}	$\tilde{d} \rightarrow \tilde{g} + 1j \rightarrow \tilde{q} + 2j$ $\rightarrow \tilde{W}^0 + 3j \rightarrow \tilde{l} + 1l + 3j$	$1l + 5j$
$\tilde{\nu}$	$\tilde{d} \rightarrow \tilde{g} + 1j \rightarrow \tilde{q} + 2j$ $\rightarrow \tilde{W}^\pm + 3j \rightarrow \tilde{\nu} + 1l + 3j$	$1l + 5j$
\tilde{e}	$\tilde{d} \rightarrow \tilde{B} + 1j \rightarrow \tilde{e} + 1l + 1j$	$3l + 2j$

Pheno 2: The Future is Bright

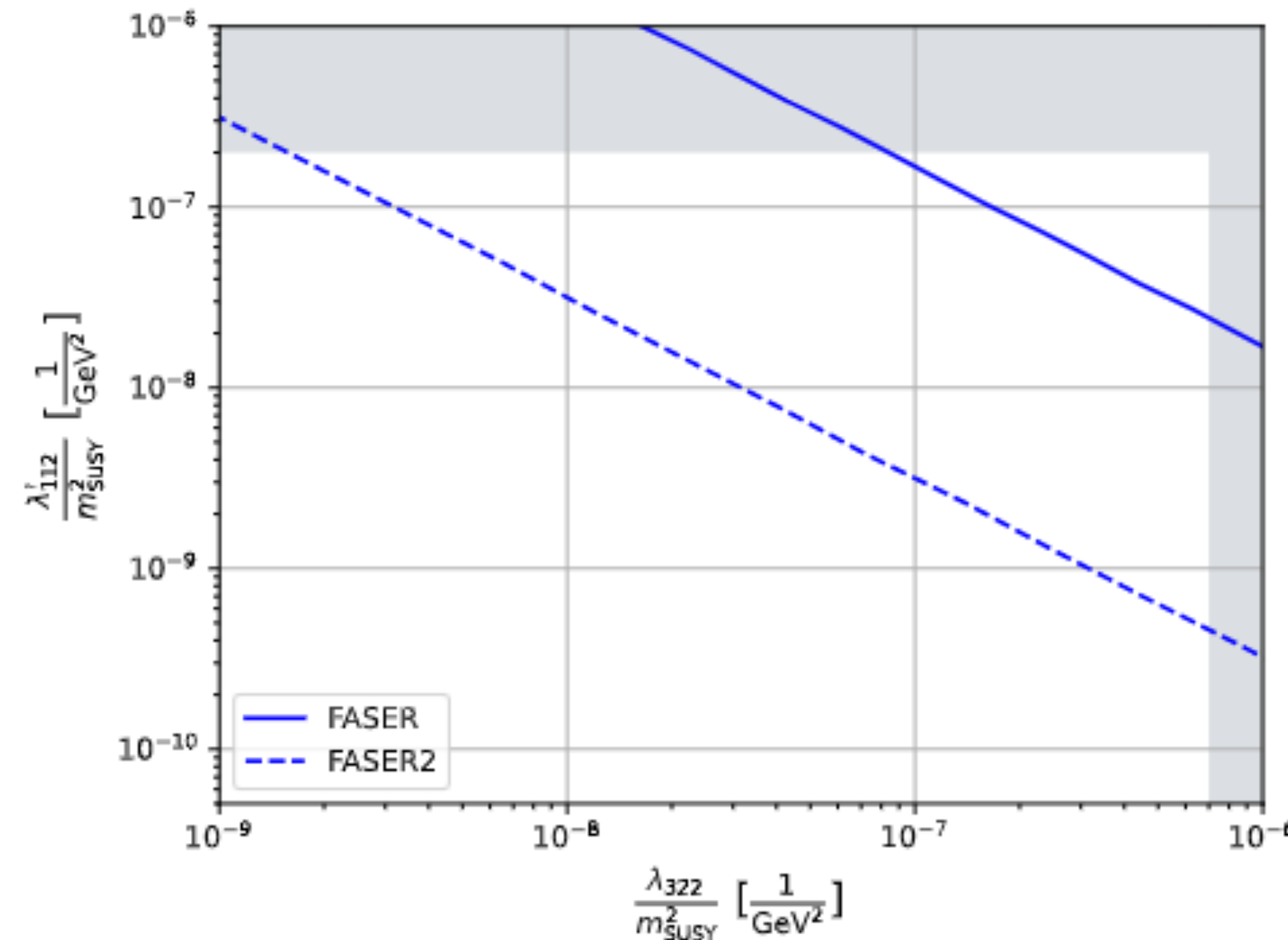
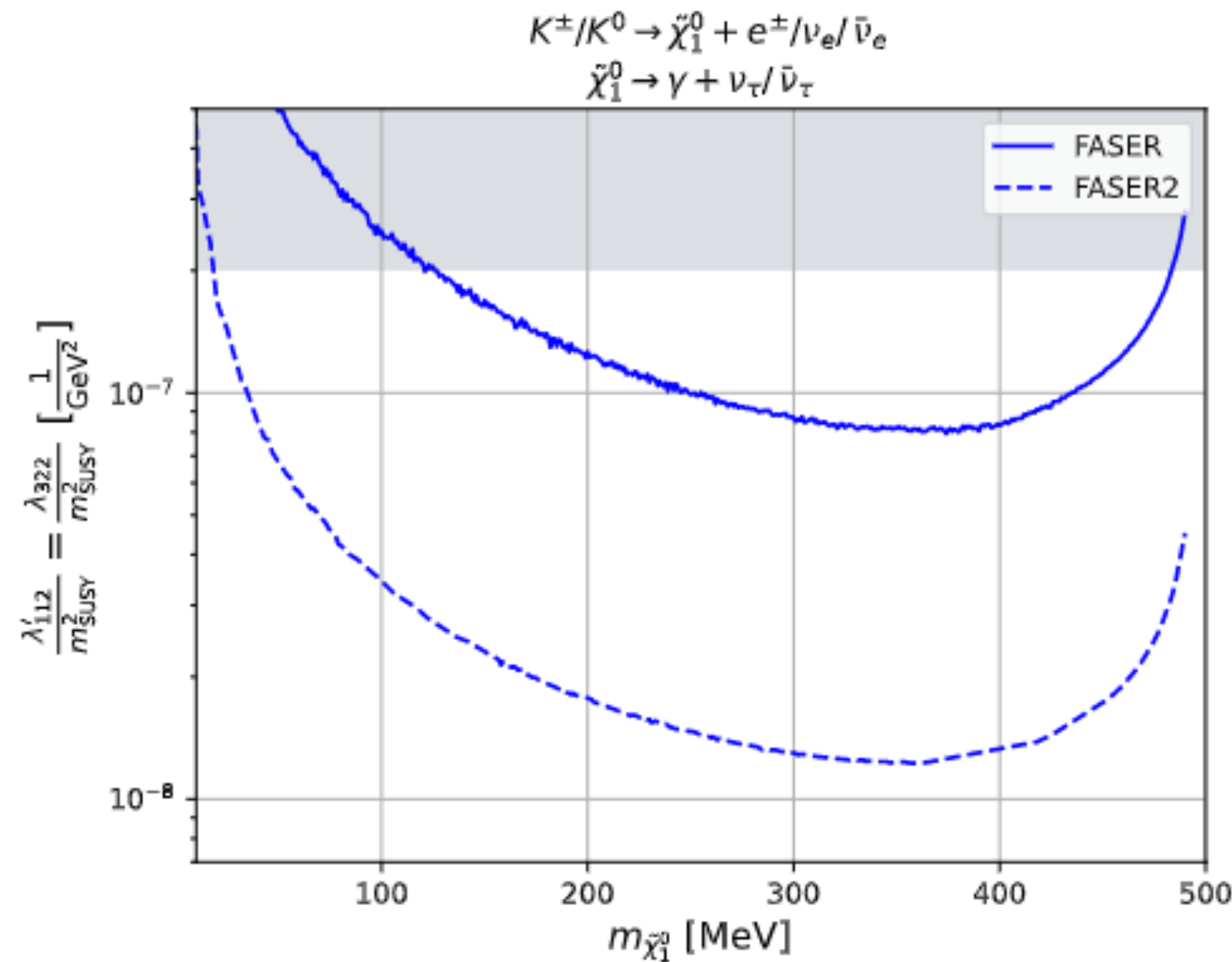
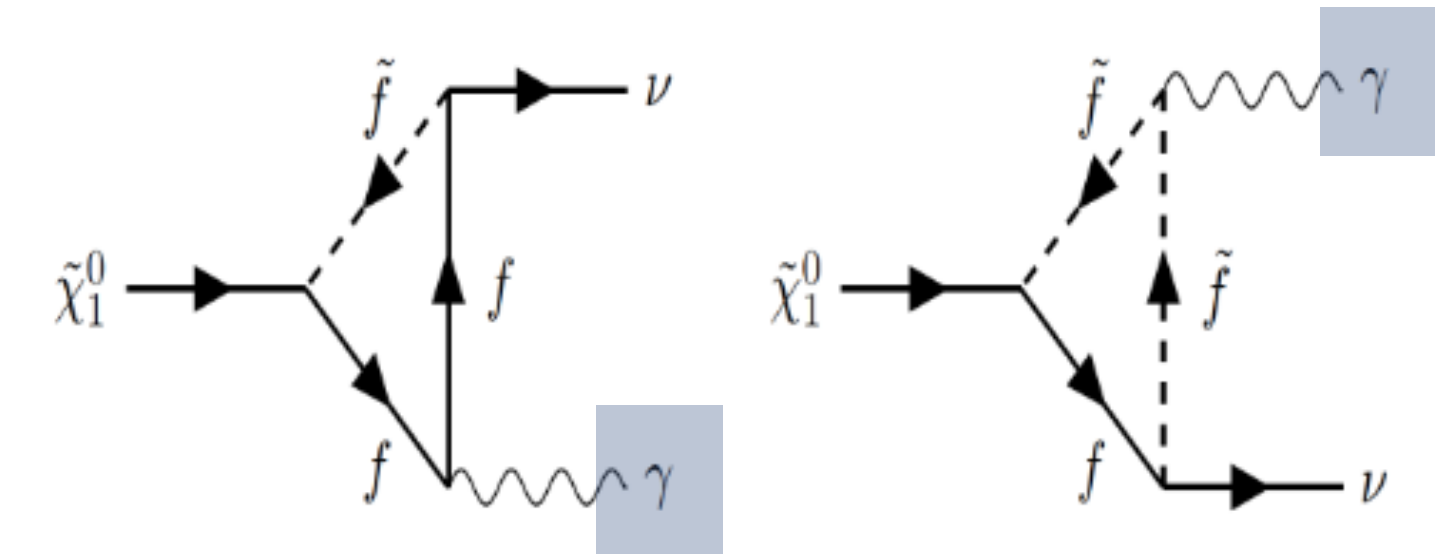


Pheno 3: Let There be Light (SUSY)

- Is there still scope for light SUSY?
- If $\tilde{\chi}_0^1$ bino-like, no mass bounds from colliders
- Cosmology requires $m_{\tilde{\chi}_0^1} \gtrsim 45$ GeV if dark matter (Lee, Weinberg)
- But unstable neutralino (as in RPV) can be massless! -> This is a real gap in our searches

Pheno 3: Let There be Light (SUSY)

- Long-lived particle program at LHC: **FASER, MATHUSLA, CODEX-b, ANUBIS, etc.**
- Very light (sub-GeV) $\tilde{\chi}_0^1$ produced in meson decays via RPV
- Long-lived, far detector decay



Sensitive to
all $\lambda_{ijj}, \lambda'_{ijj}$

Outlook

- RPV SUSY is just as well motivated as RPC SUSY
- Signatures are very different; many complex possibilities
- Can classify systematically in model-independent way -> towards more complete coverage
- Very recent developments: Lepton PDFs and LLPs
RPV SUSY shows up here too
Can implement model-independent analyses here too!

Thanks for your time!