

# New developments in anomaly detection and model independent searches for new physics with machine learning

Benjamin Nachman

*Lawrence Berkeley National Laboratory*

[bpnachman.com](http://bpnachman.com)

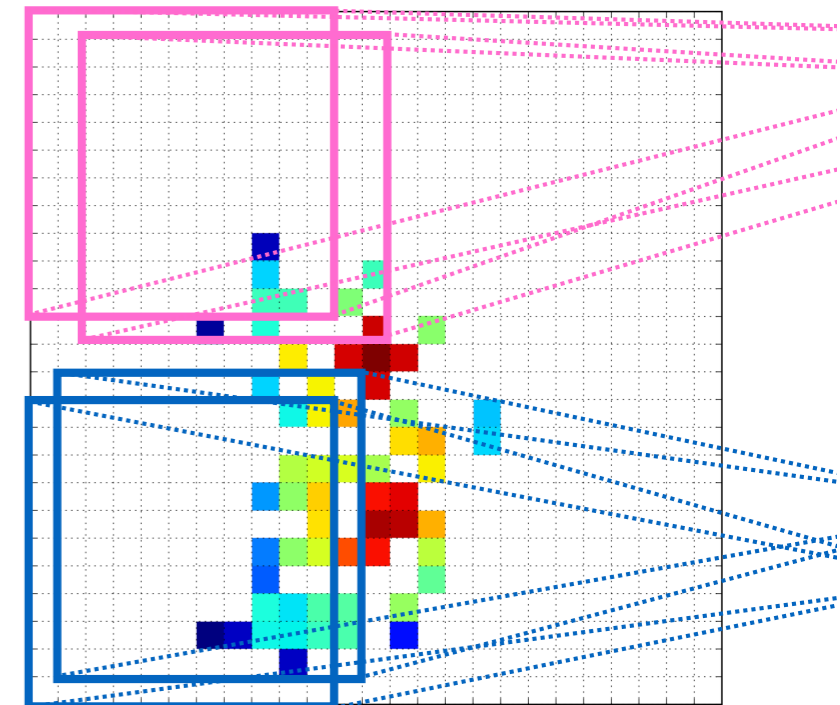
[bpnachman@lbl.gov](mailto:bpnachman@lbl.gov)



@bpnachman



bnachman



LHCP  
May 2022

# Questions in particle physics



**Theoretical** and **experimental** questions motivate a deep exploration **of the fundamental structure of nature**

Dark matter

Hierarchy problem

Strong CP

Flavor puzzles

Baryogenesis

Dark energy

We have performed thousands of hypothesis tests & have no significant evidence for physics beyond the Standard Model

**Three possibilities**



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**This is what keeps me up at night!**

**(3) We are not looking in the right place**

# The Path Forward



There are two complementary paths forward:

**(1) Identify new, specific, well-motivated places to look**

*This is still an incredibly important direction and has resulted in new directions like long-lived particle searches*

**(2) Look in many places all at once**

*Focus of today's talk!*

**This is what keeps me up at night!**

**(3) We are not looking in the right place**

# The Path Forward



There are two **complementary** paths forward:

(1) Identify new, specific, well-motivated places to look

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(2) Look in many places all at once

*Focus of today's talk!*

**There is no free lunch:** for any particular model, (2) will be less sensitive than (1). We need both search paradigms!

# Why not just look everywhere?



# Why not just look everywhere?

(a) There are a lot of places to look

$A \rightarrow BC$	$B = \text{SM}$									$B = \text{BSM}$	
	$e$	$\mu$	$\tau$	$q/g$	$b$	$t$	$\gamma$	$Z/W$	$H$		
$C = \text{SM}$	$e$	$Z'$	$R$	$R$	$LQ$	$LQ$	$LQ$	$L^*$	$L^*$	$L^*$	Many
	$\mu$		$Z'$	$R$	$LQ$	$LQ$	$LQ$	$L^*$	$L^*$	$L^*$	
	$\tau$			$Z'$	$LQ$	$LQ$	$LQ$	$L^*$	$L^*$	$L^*$	
	$q/g$				$Z'$	$W'$	$T'$	$Q^*$	$Q^*$	$Q'$	
	$b$					$Z'$	$W'$	$Q^*$	$Q^*$	$B'$	
	$t$						$Z'$	$Q^*$	$T'$	$T'$	
	$\gamma$							$H$	$H$	$Z_{KK}$	
	$Z/W$ $H$								$H$	$H^\pm/A$ $H$	
$C = \text{BSM}$	<p>Consider just the di-object search for resonant <math>A \rightarrow B C</math></p>									Many	

# Why not just look everywhere?



(a) There are a lot of places to look

	$e$	$\mu$	$\tau$	$q/g$	$b$	$t$	$\gamma$	$Z/W$	$H$	BSM $\rightarrow$ SM <sub>1</sub> $\times$ SM <sub>1</sub>				BSM $\rightarrow$ SM <sub>1</sub> $\times$ SM <sub>2</sub>			BSM $\rightarrow$ complex			
										$q/g$	$\gamma/\pi^0$ 's	$b$	...	$tZ/H$	$bH$	...	$\tau qq'$	$eqq'$	$\mu qq'$	...
$e$	[37,38]	[39,40]	[39]	$\emptyset$	$\emptyset$	$\emptyset$	[41]	[42]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	[43,44]	$\emptyset$	
$\mu$		[37,38]	[39]	$\emptyset$	$\emptyset$	$\emptyset$	[41]	[42]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	[43,44]	
$\tau$			[45,46]	$\emptyset$	[47]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	[48,49]	$\emptyset$	
$q/g$				[29,30,50,51]	[52]	$\emptyset$	[53,54]	[55]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
$b$					[29,52,56]	[57]	[54]	[58]	[59]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	[60]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
$t$						[61]	$\emptyset$	[62]	[63]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	[64]	[60]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
$\gamma$							[65,66]	[67-69]	[68,70]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
$Z/W$								[71]	[71]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
$H$									[72,73]	[74]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
BSM $\rightarrow$ SM <sub>1</sub> $\times$ SM <sub>1</sub>	$q/g$									$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
	$\gamma/\pi^0$ 's										[75]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
	$b$											[76,77]	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	$\emptyset$	
BSM $\rightarrow$ SM <sub>1</sub> $\times$ SM <sub>2</sub>	$tZ/H$																			
	$bH$																			
	...																			
BSM $\rightarrow$ complex	$\tau qq'$																			
	$eqq'$																			
	$\mu qq'$																			

Consider just the di-object search for resonant  $A \rightarrow B C$

...most cases are uncovered!



# Why not just look everywhere?



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- (a) There are a lot of places to look
- (b) You would find a lot of excesses

Best to cast a wide net in a smart way !

# Outline: Casting a Wide Net(work)



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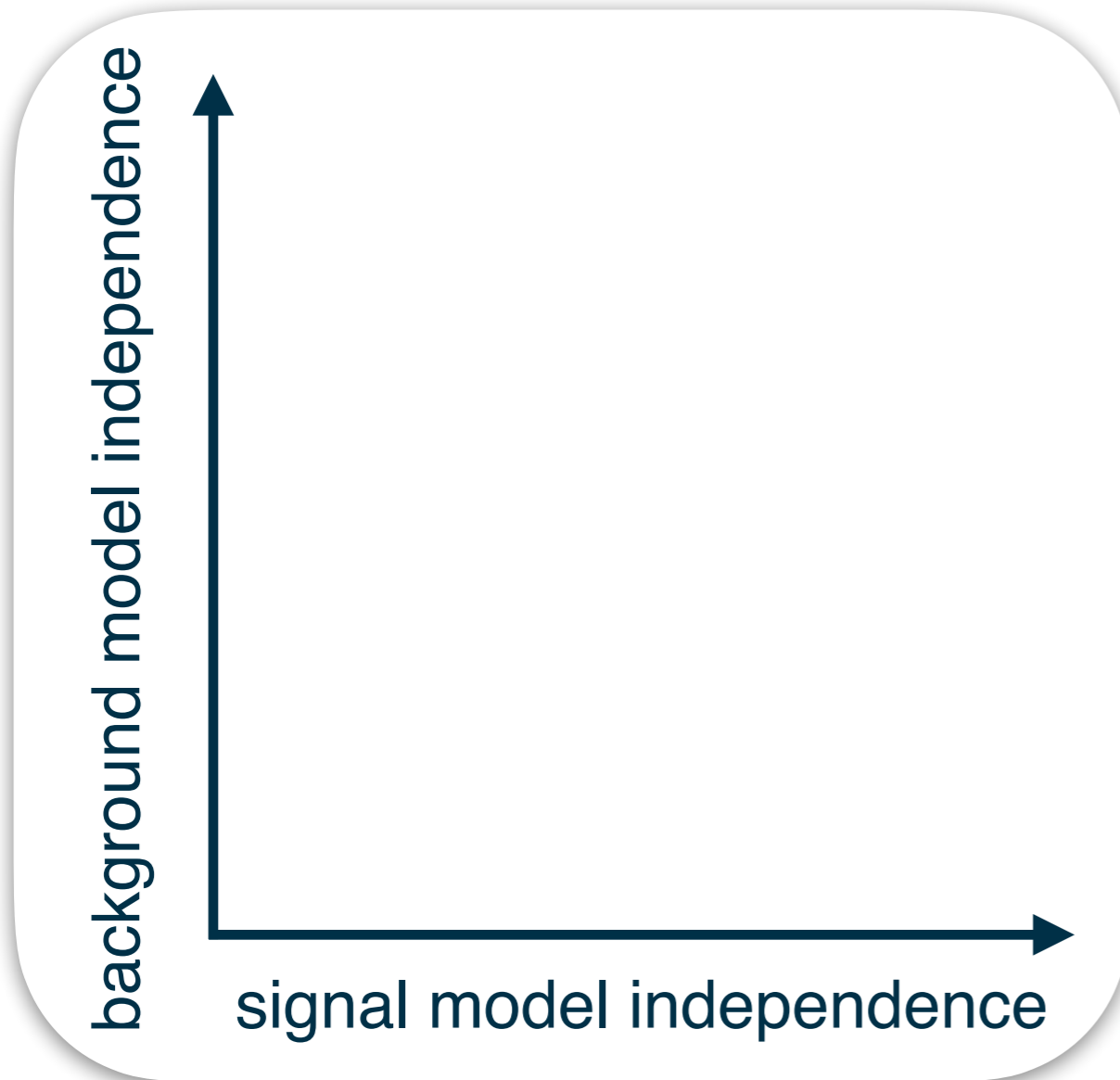
1. The landscape of model dependence
2. Overview of new ideas
3. Resonant anomaly detection
4. The future (and why you should be part of it!)

***Anomaly* in this talk means unanticipated new physics (!)**

# Landscape of Model Dependence

15

**Suppose you want to search for a new signal process**

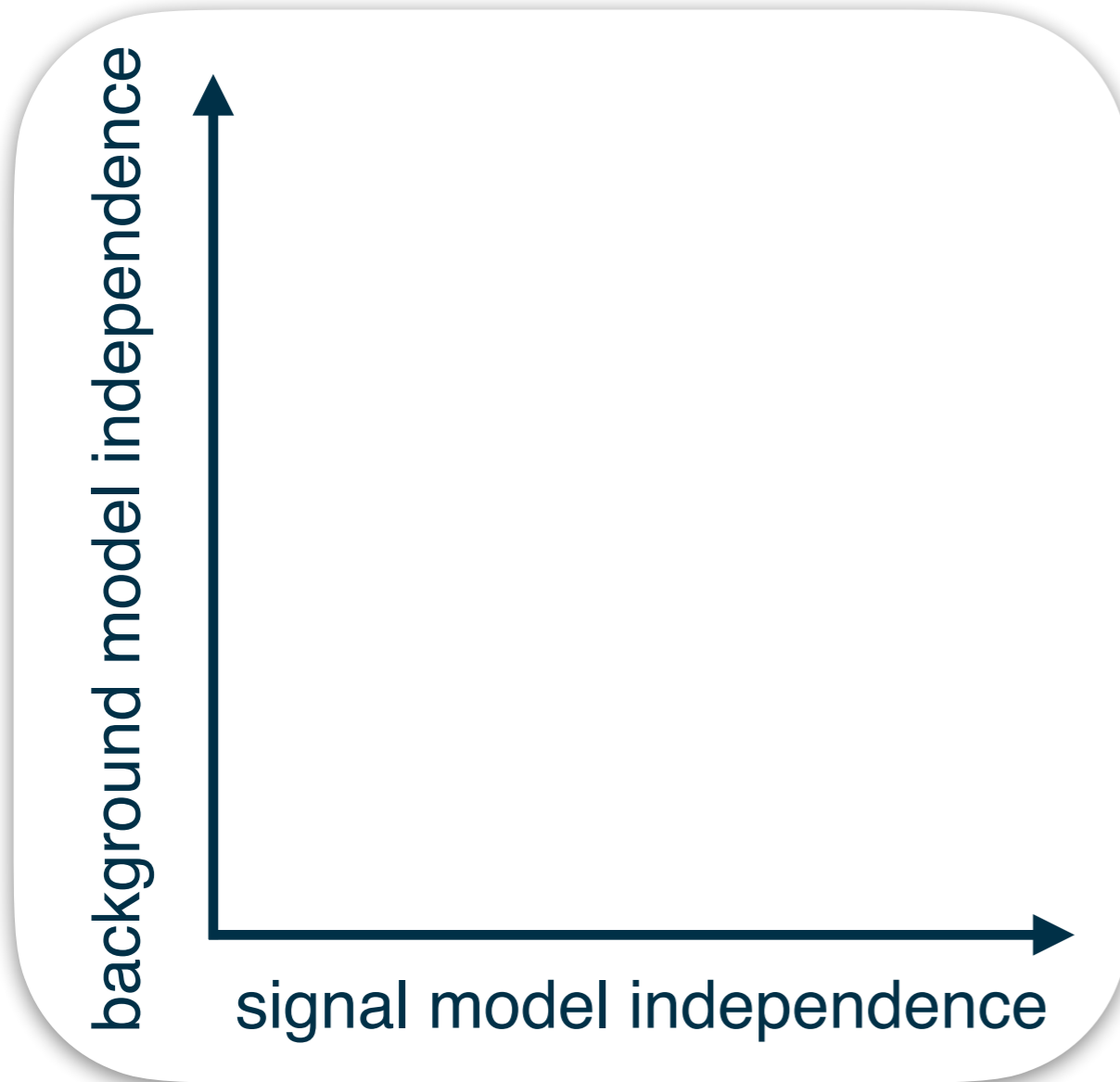


Signal sensitivity

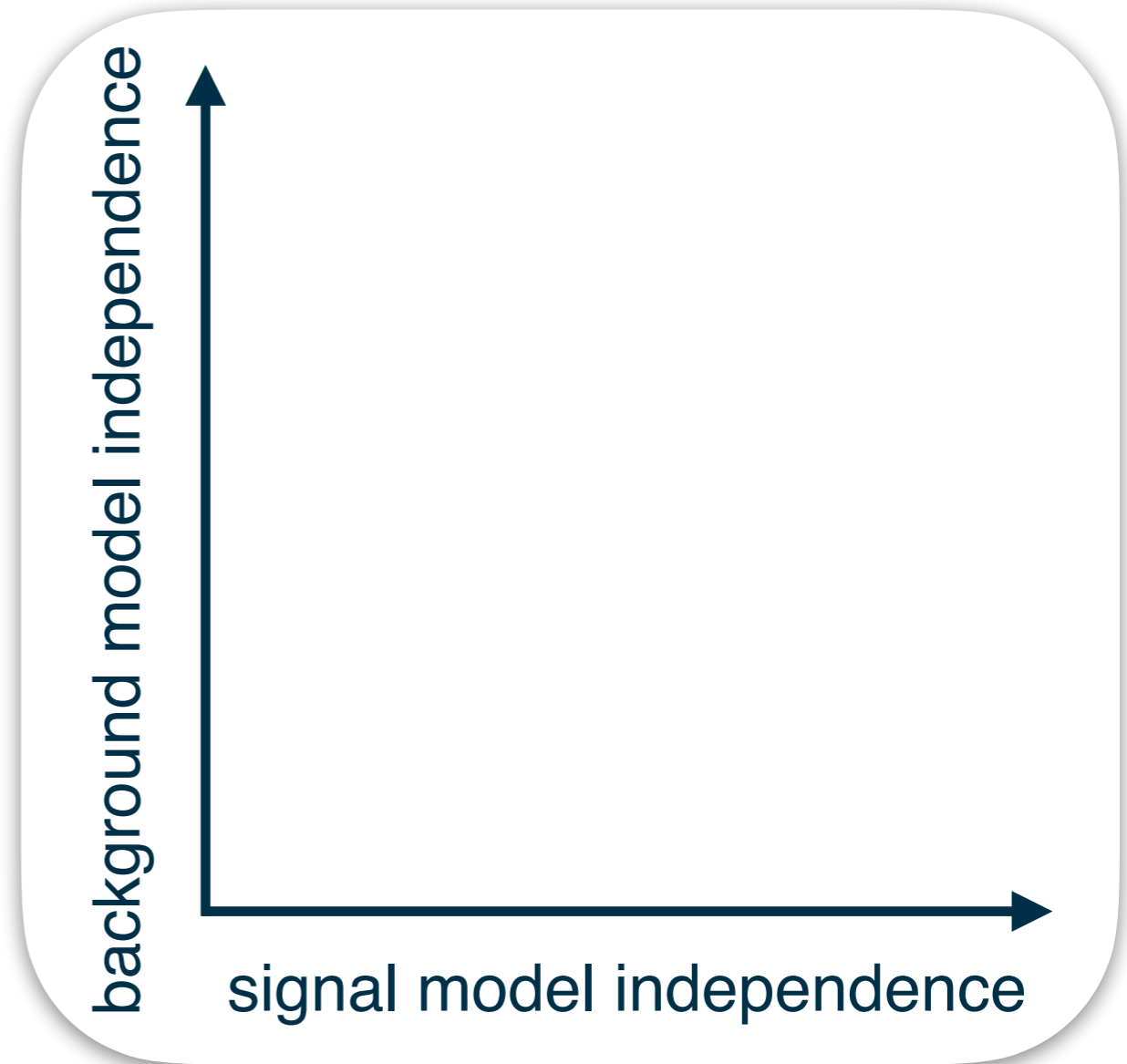
**Suppose you want to search for a new signal process**

# Landscape of Model Dependence

17



Signal sensitivity

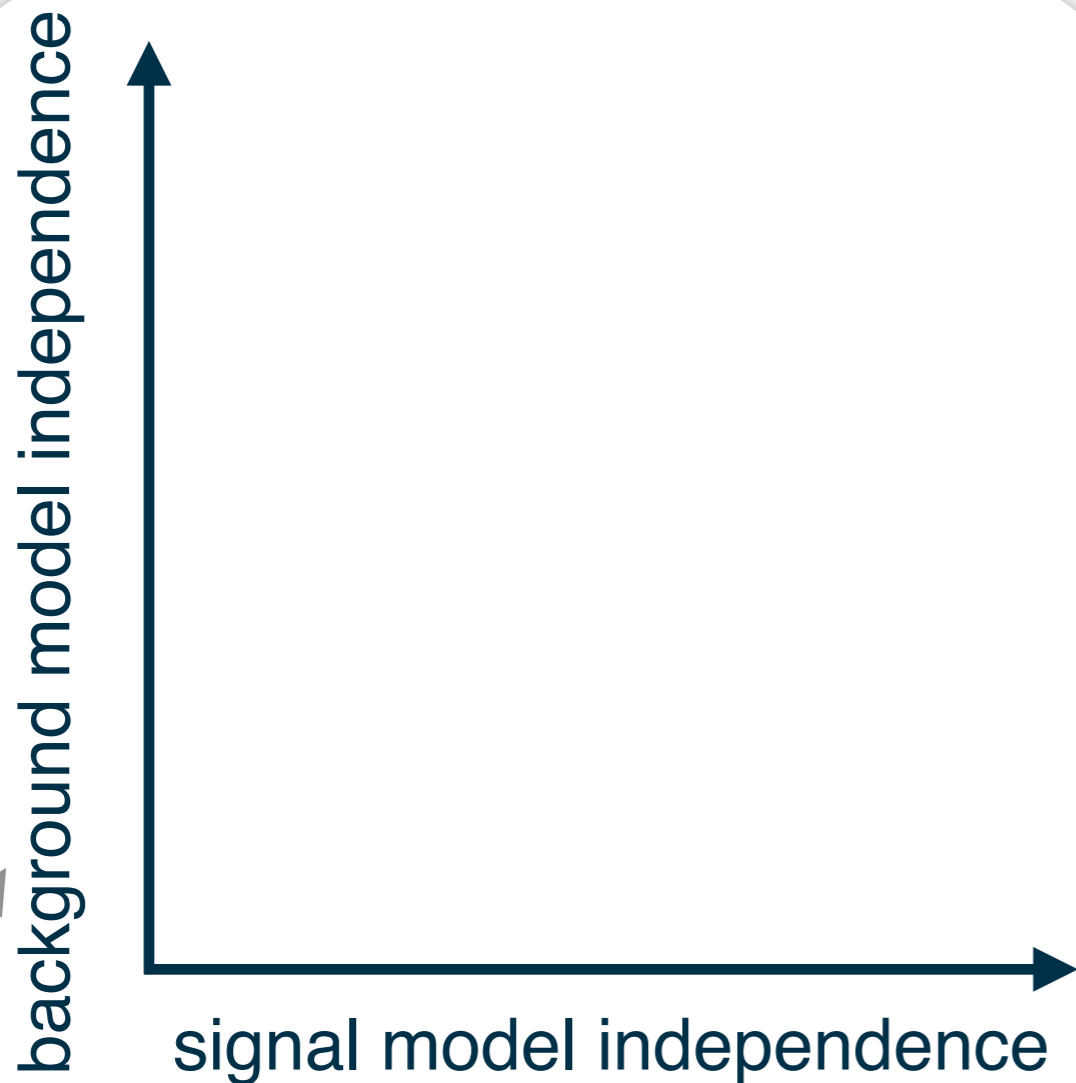


Background specificity

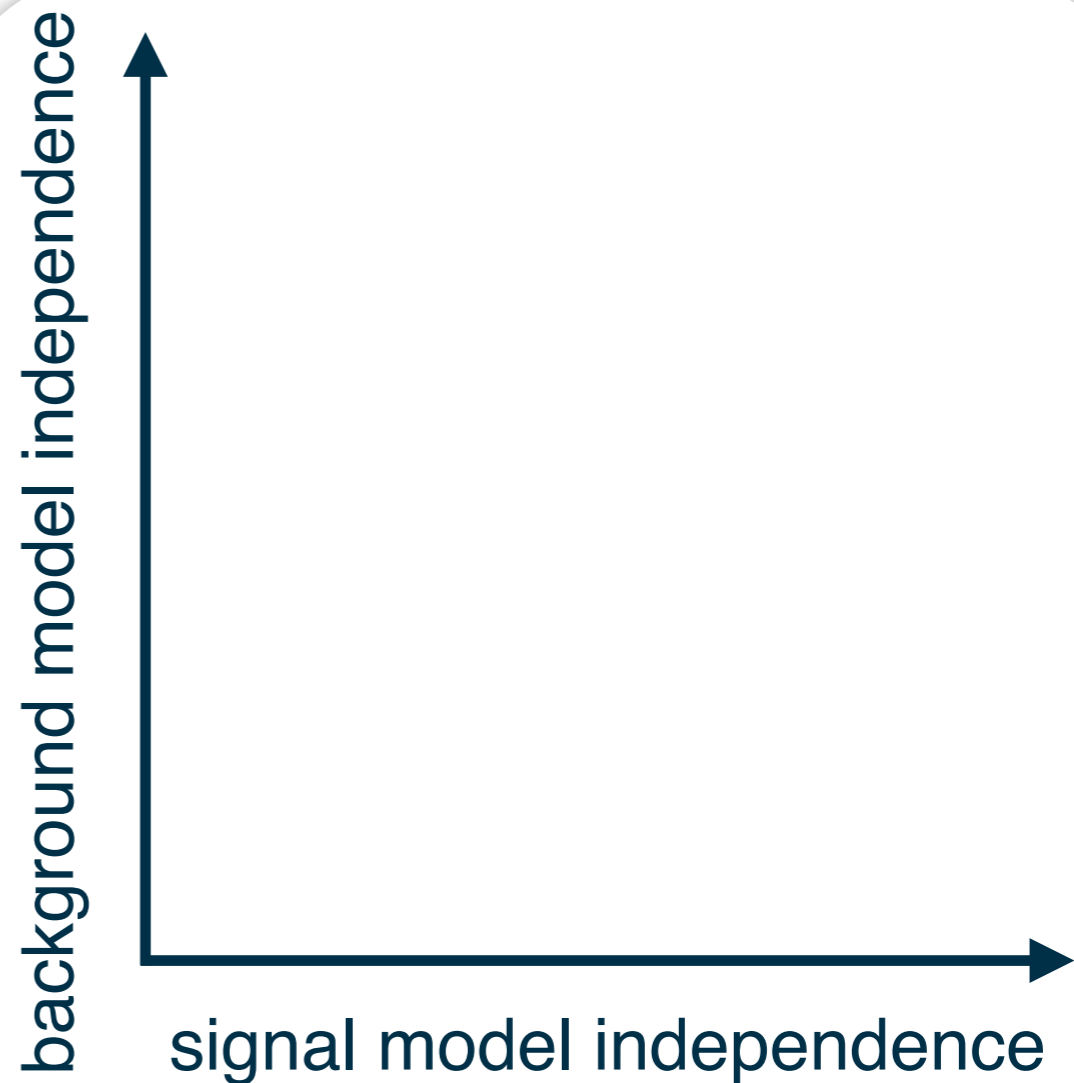
**Suppose you want to search for a new signal process**

# Landscape of Model Dependence

18

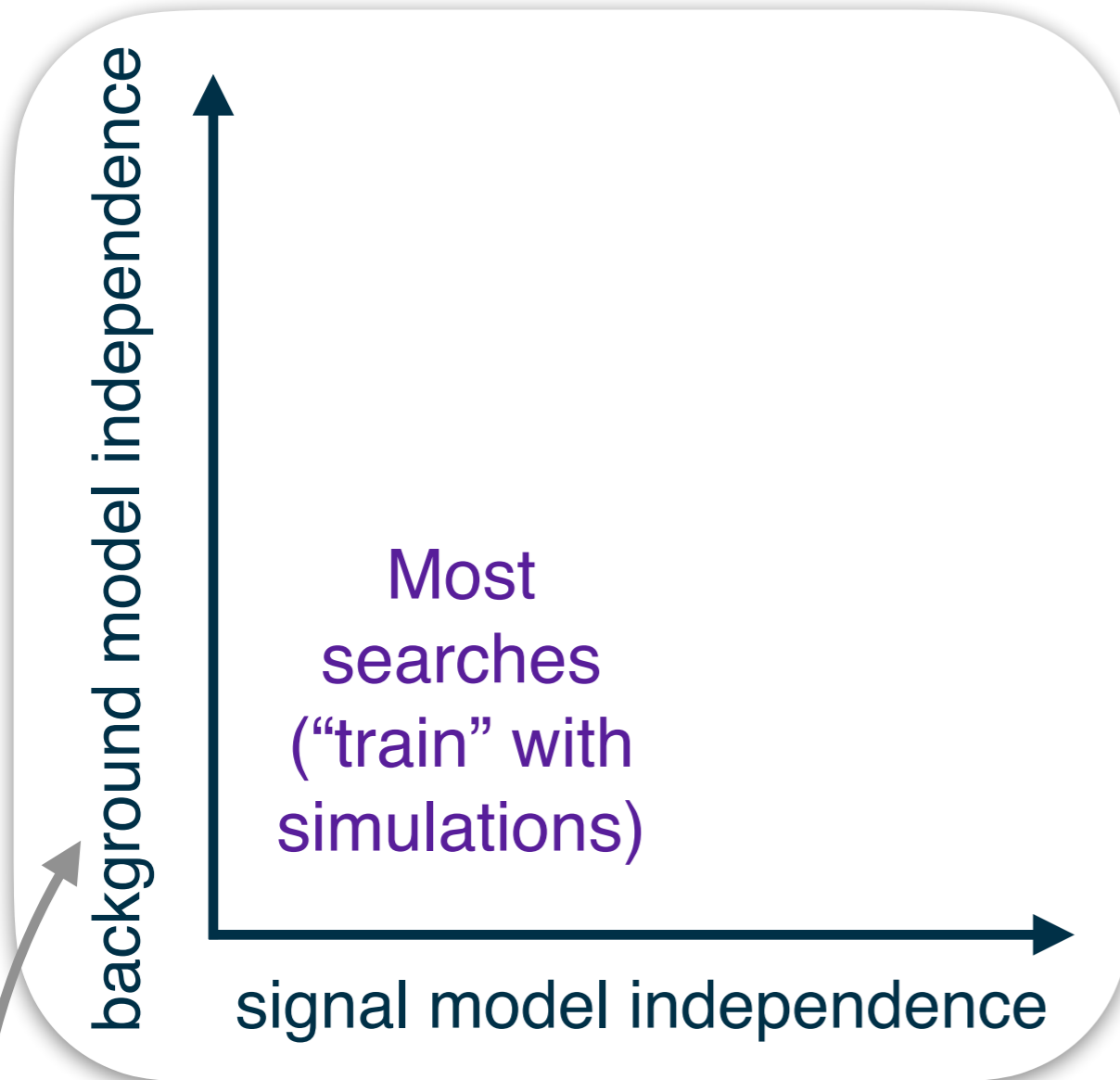


Signal sensitivity



Background specificity

*Standard Model*



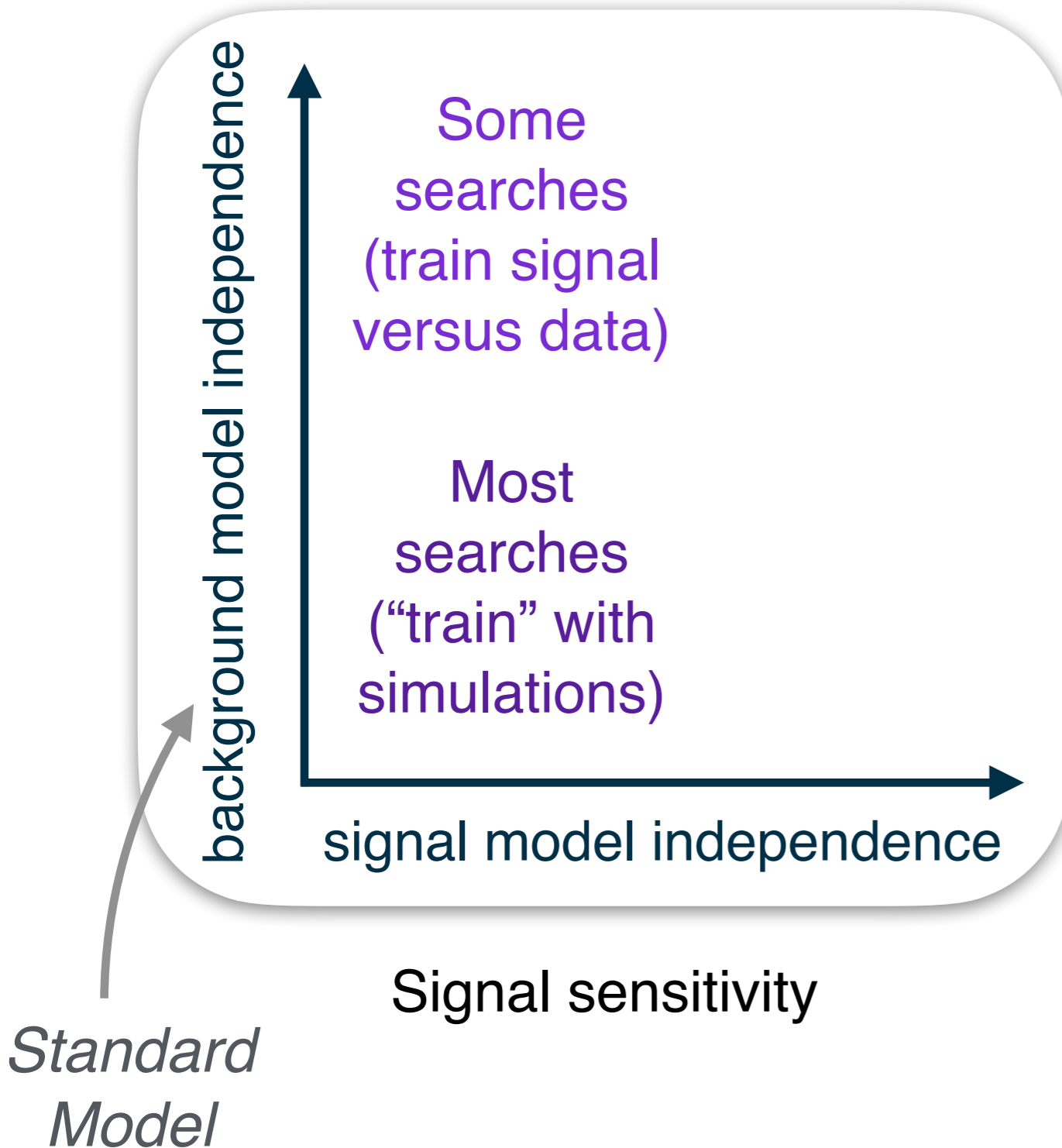
> 99% of searches at the LHC and elsewhere are of this type

"train" is in quotes because such searches may or may not use machine learning

Signal sensitivity

# Landscape of Model Dependence

20

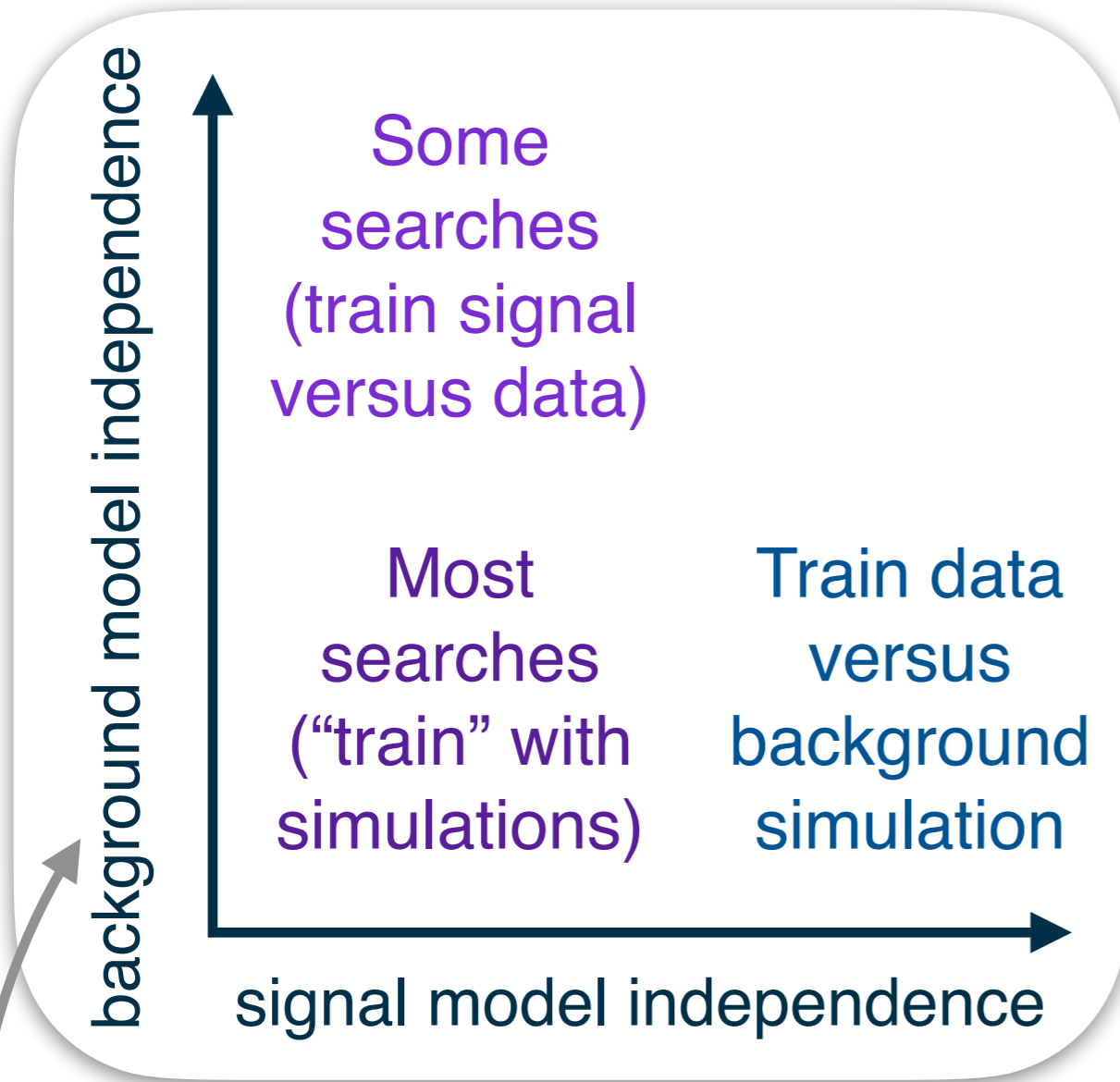


*e.g. signal  
simulation versus  
calibration data*

standard approach  
when signal is clean  
and well-understood,  
but background is  
not, e.g.  $h \rightarrow \gamma\gamma$



# Landscape of Model Dependence



Signal sensitivity

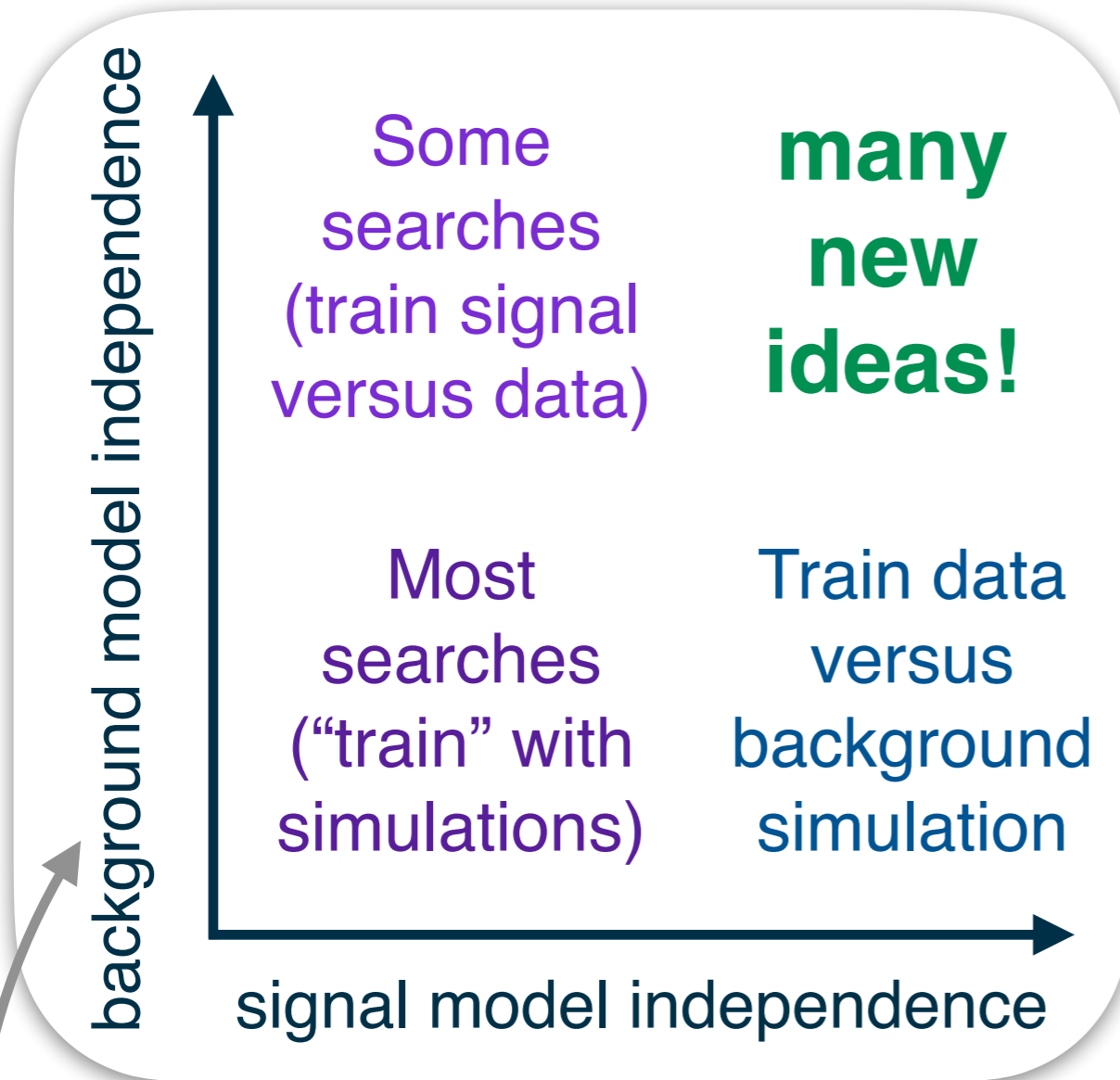
*Standard  
Model*

signal model  
*independent*  
background model  
*dependent*

There is a history of these searches at the LHC, Tevatron, HERA, LEP

# Landscape of Model Dependence

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Signal sensitivity

*Standard Model*

There are many new ideas that make use of modern machine learning

The goal is to learn **directly from data**, injecting as little bias as possible

For a recent review, see G. Karagiorgi, G. Kasieczka, S. Kravitz, BN, D. Shih, Nature Reviews Physics (2022), 2112.03769


I like to categorize new ideas based on the core assumption about the BSM, which is intimately related to the technique ***supervision***

**Unsupervised** = no labels

**Weakly-supervised** = noisy labels

**Semi-supervised** = partial labels

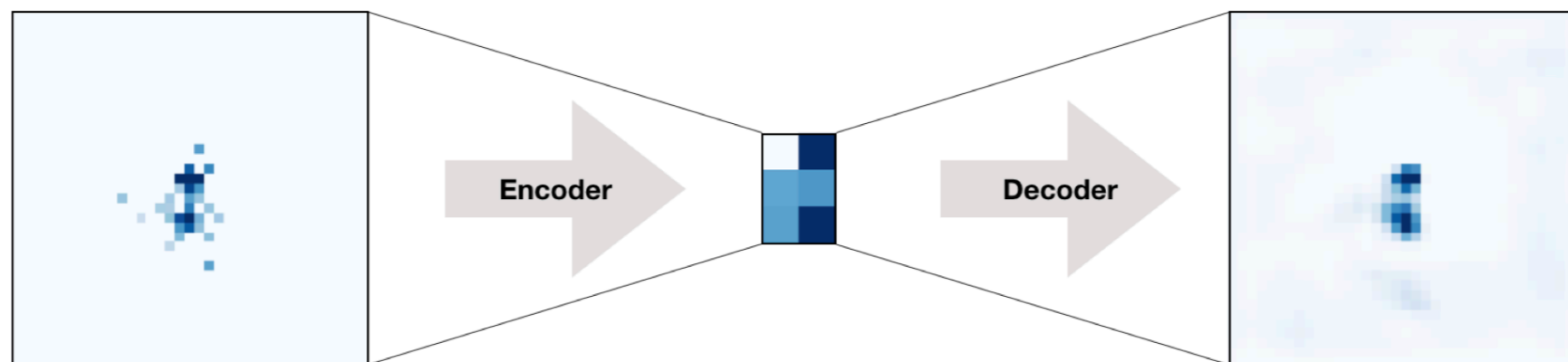
**Supervised** = full label information



*This is most searches. You simulate the signal (label = 1), simulate the background (label = 0) and “train” a classifier to distinguish the 1’s from the 0’s.*

**Unsupervised** = no labels

Typically, the goal of these methods is to look for events with **low  $p(\text{background})$**

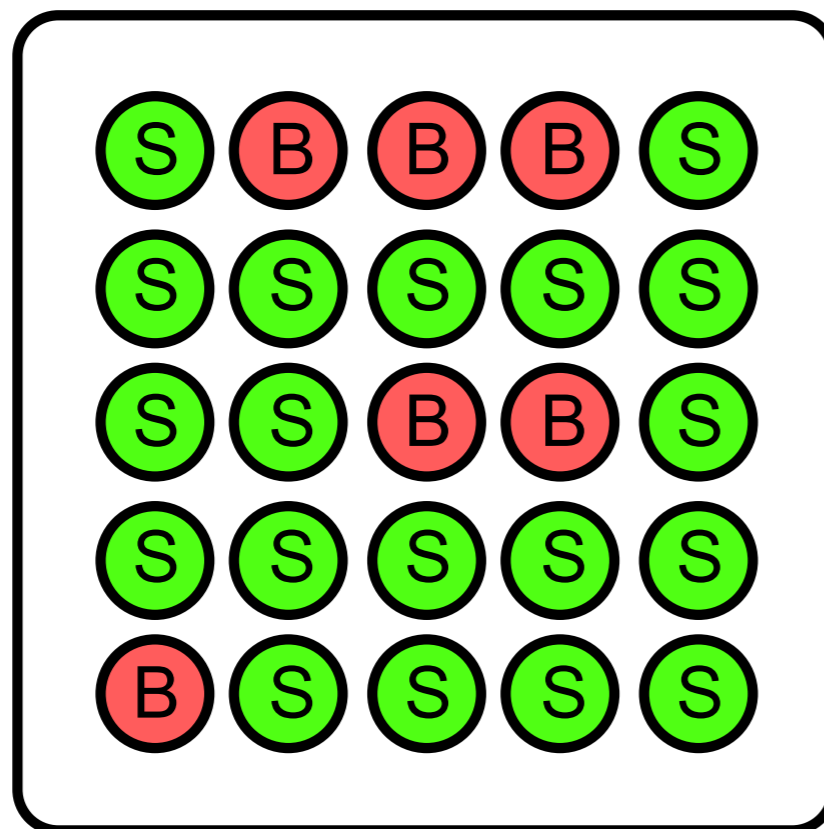


One strategy (autoencoders) is to try to compress events and then uncompress them. When  $x = \text{uncompress}(\text{compress}(x))$ , then  $x$  probably has low  $p(x)$ .

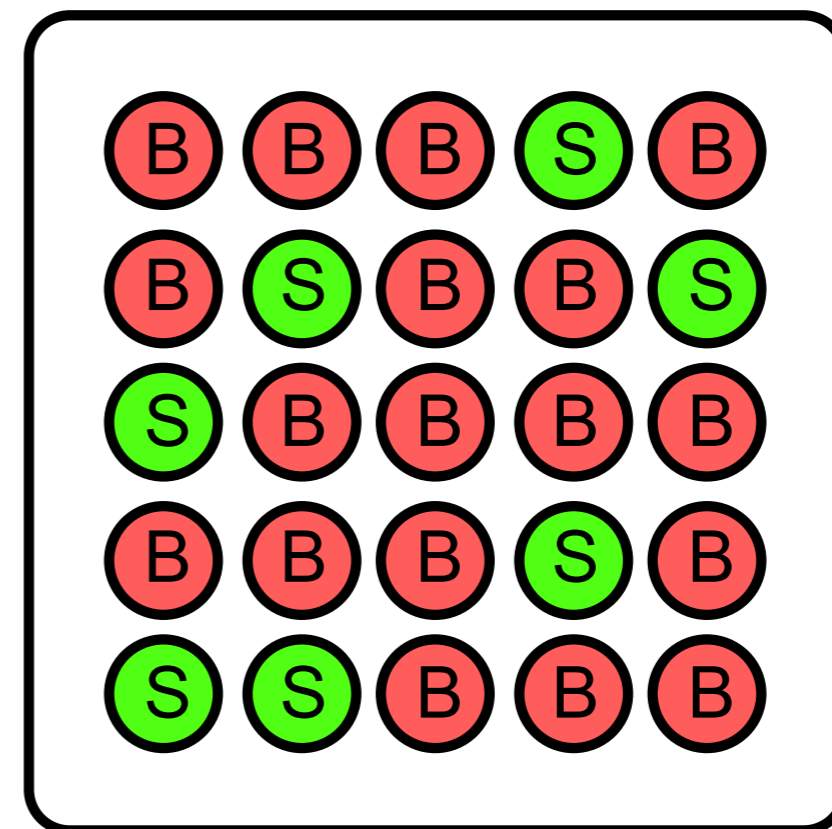
**Weakly-supervised** = noisy labels

Typically, the goal of these methods is to look for events with high  $p(\text{possibly signal-enriched})/p(\text{possibly signal-depleted})$

Signal enriched



Signal depleted



# Semi-supervised

**Semi-supervised** = partial labels

Typically, these methods use some signal simulations to build signal sensitivity



vs



e.g. SM background versus many signals



Approach:	<b>Unsupervised</b>	<b>Weakly supervised</b>
<b>BSM assumption</b>	Signal is rare (low $p$ )	Signal is an over density (high $p$ ratio)
<b>Main drawback</b>	rare is not invariant* under coordinate transformations!	need two samples

\*for a detailed discussion about this, see K. Desai, BN, J. Thaler, 2112.05722

Approach:

**Unsupervised**

**Weakly supervised**

**BSM assumption**

Signal is rare  
(low  $p$ )

Signal is an  
over density  
(high  $p$  ratio)

**Main drawback**

rare is not invariant\*  
under coordinate  
transformations!

need two samples

**Canonical example:  
resonances!**

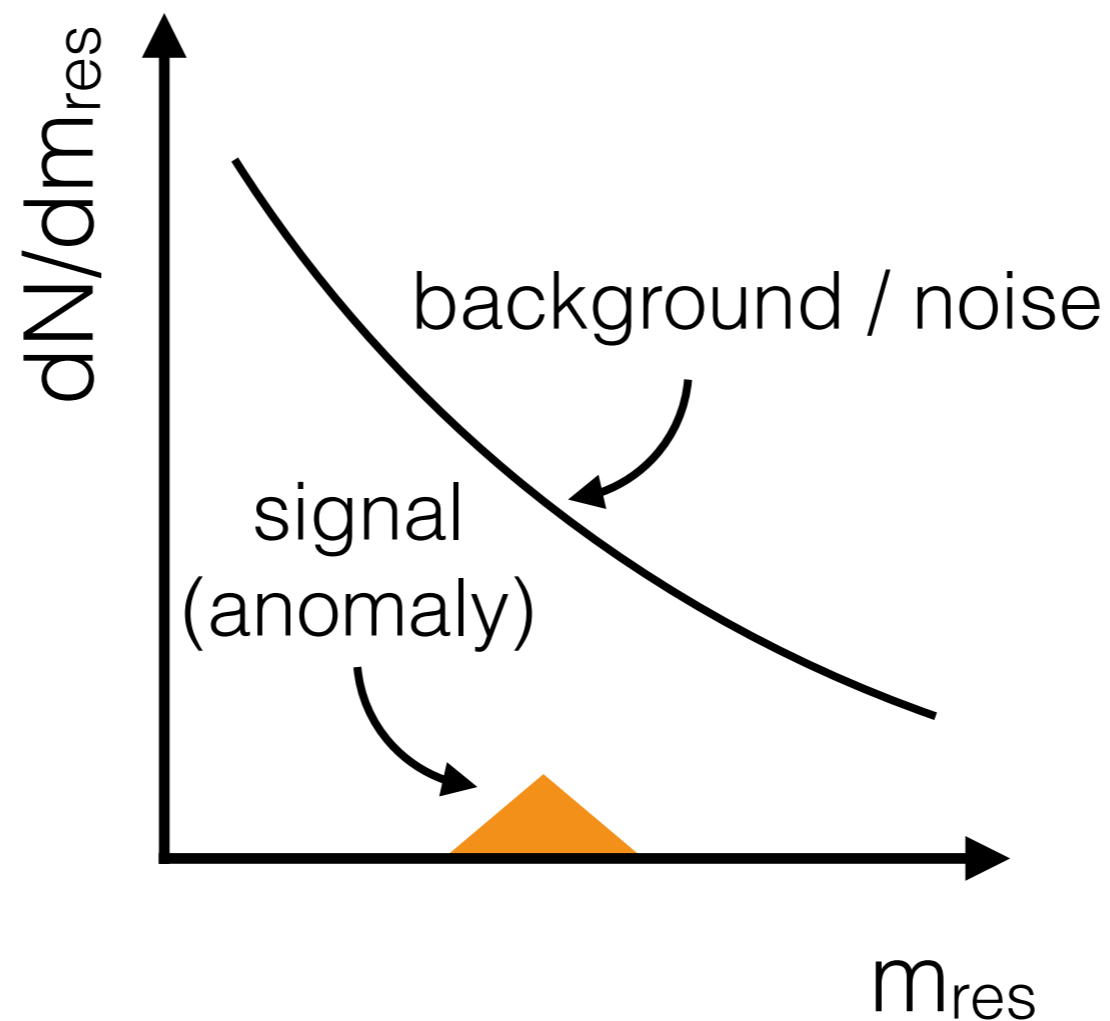
\*for a detailed discussion about this, see K. Desai, BN, J. Thaler, 2112.05722



# Resonant Anomalies

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A relatively general, but powerful assumption is that the anomaly is localized somewhere in phase space.

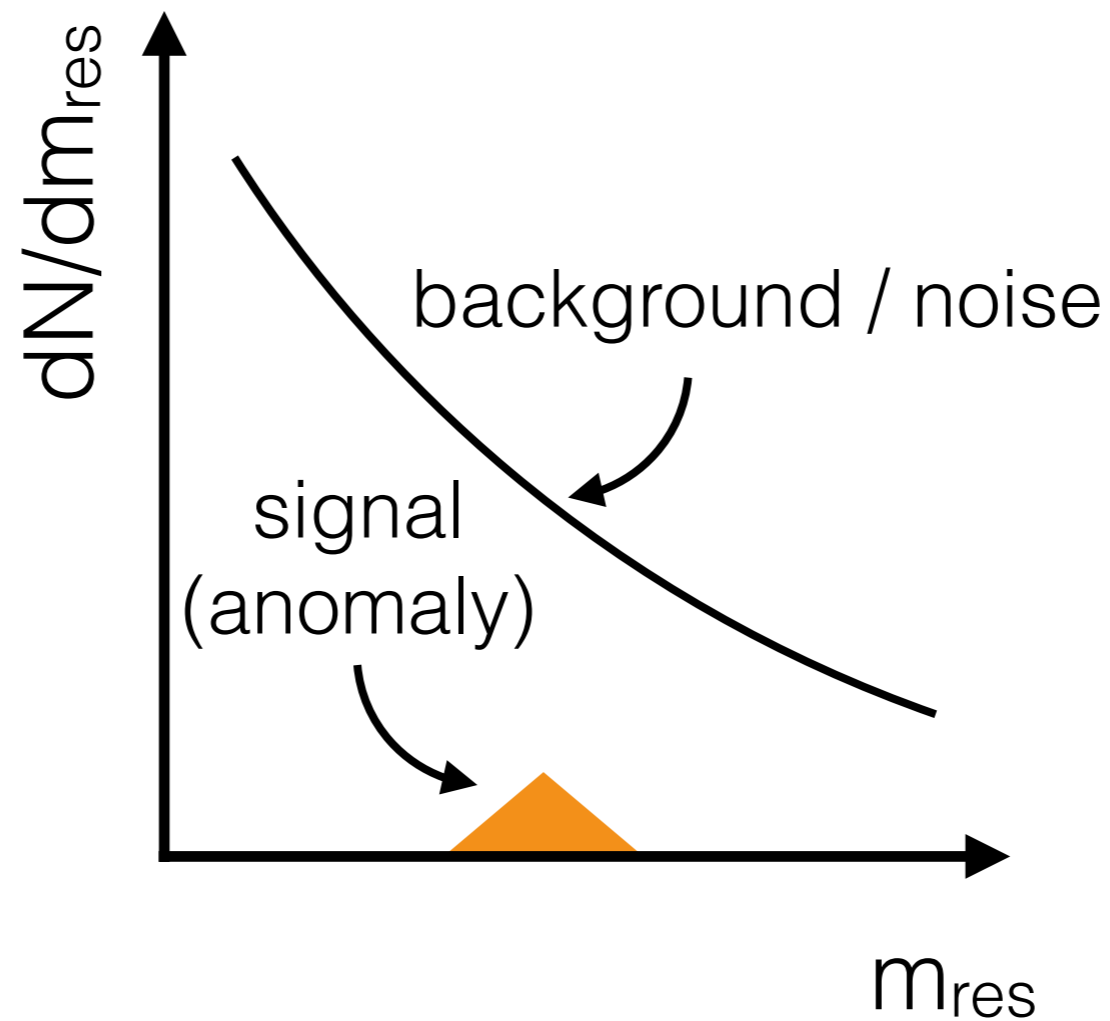


Generically true when there are on-shell new particles.

# A Worked Example: dijets @ LHC

30

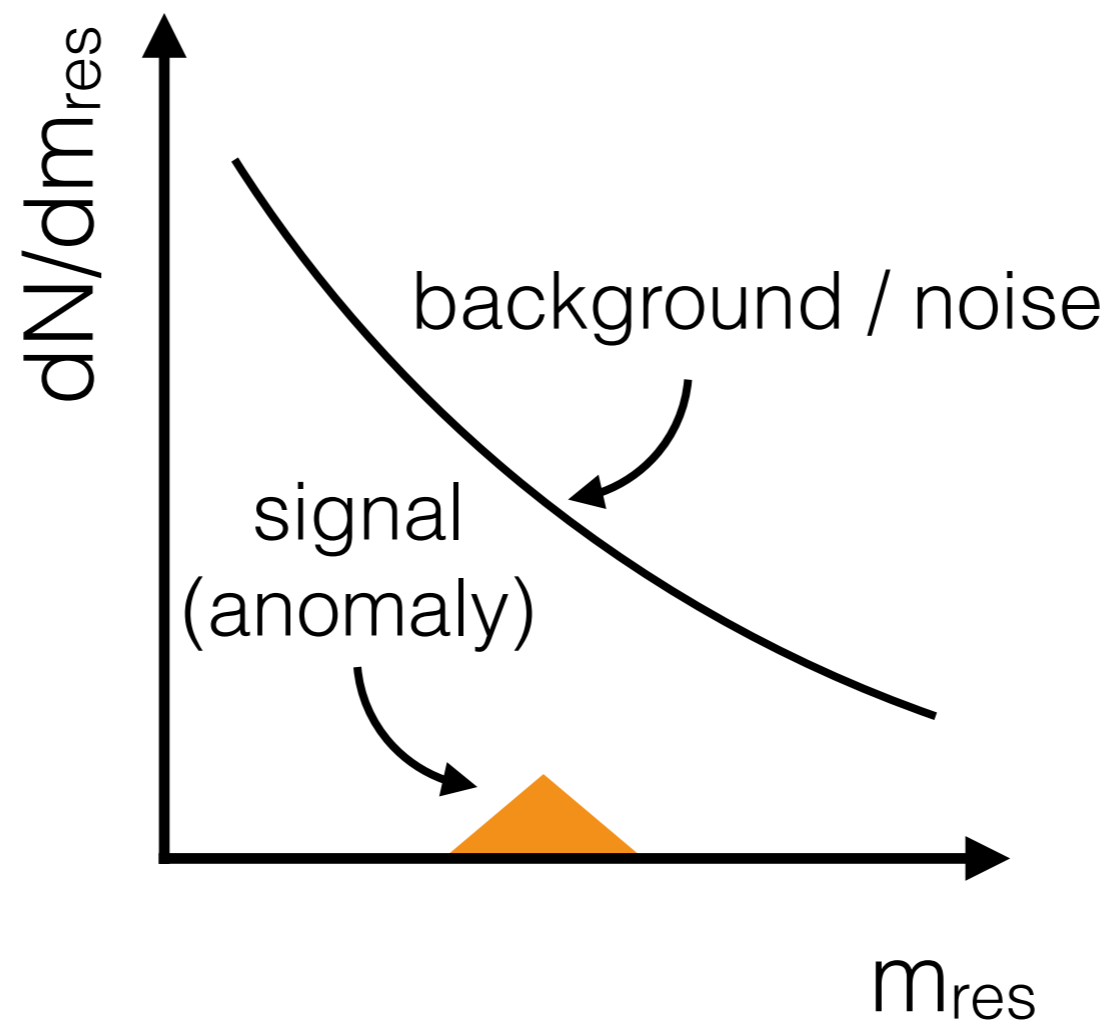
I'll walk you through a weakly-supervised approach.



# A Worked Example: dijets @ LHC

31

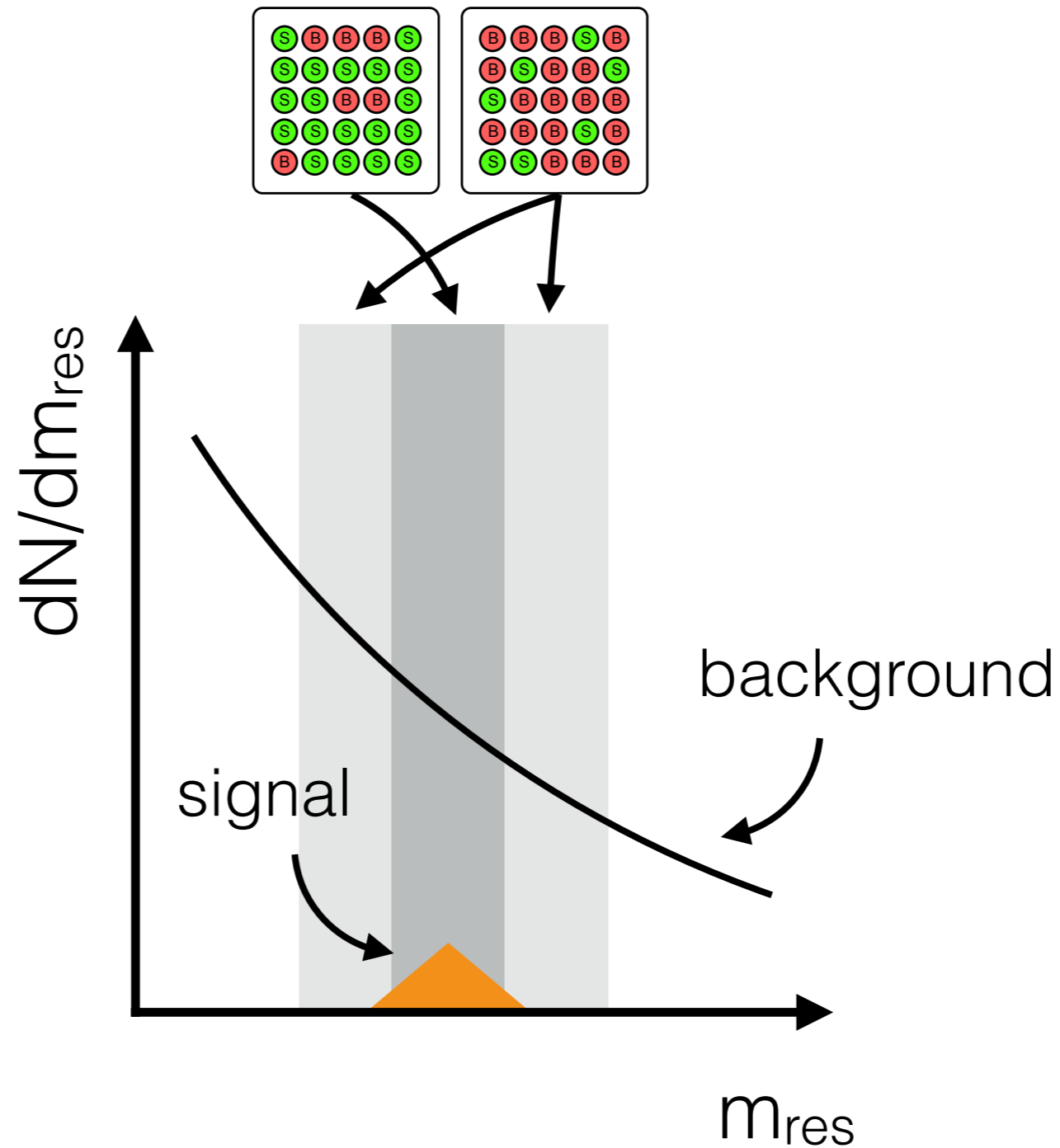
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First: we will need to generate noisy labels.

# A Worked Example: dijets @ LHC

32

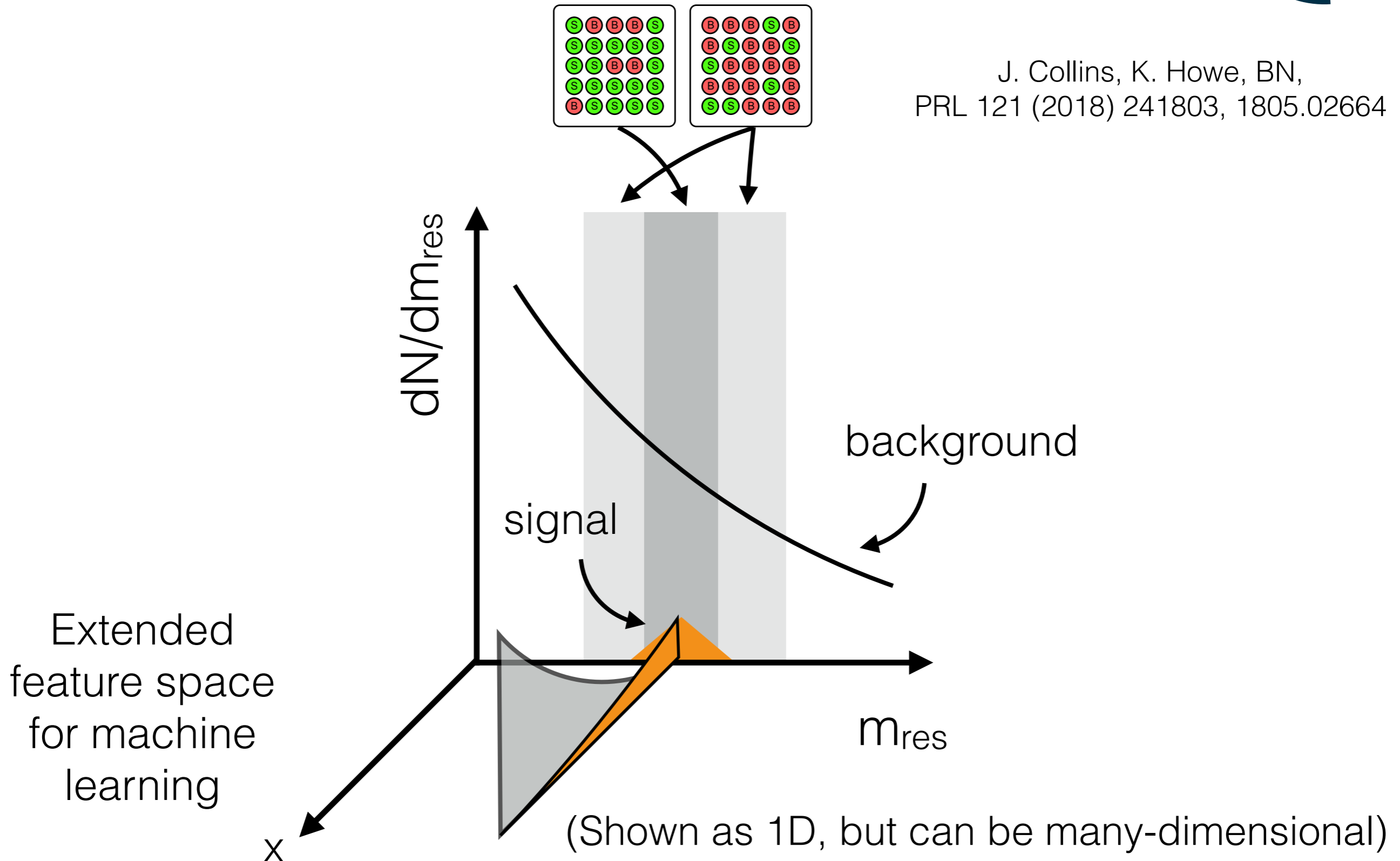


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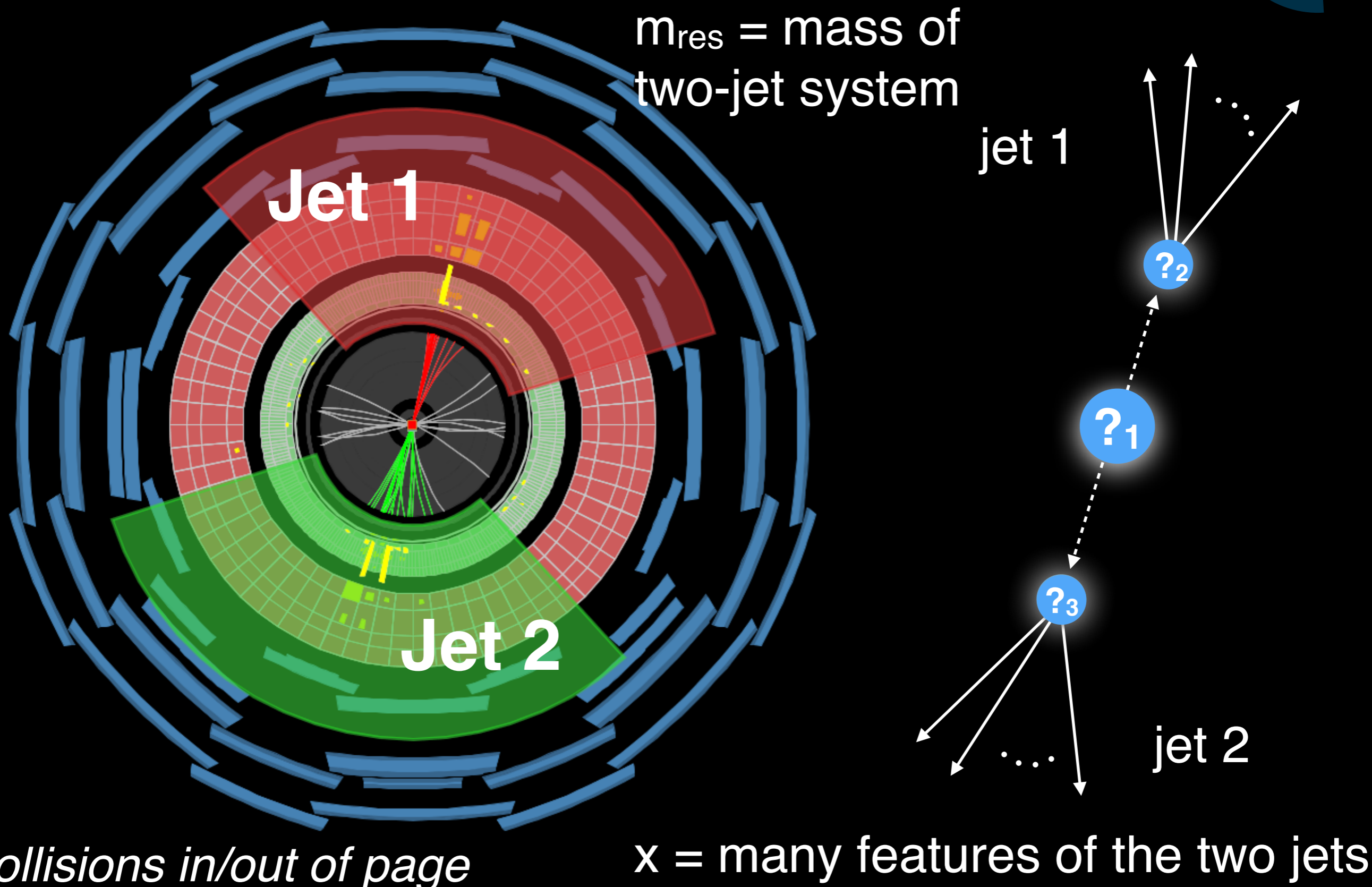
# A Worked Example: dijets @ LHC

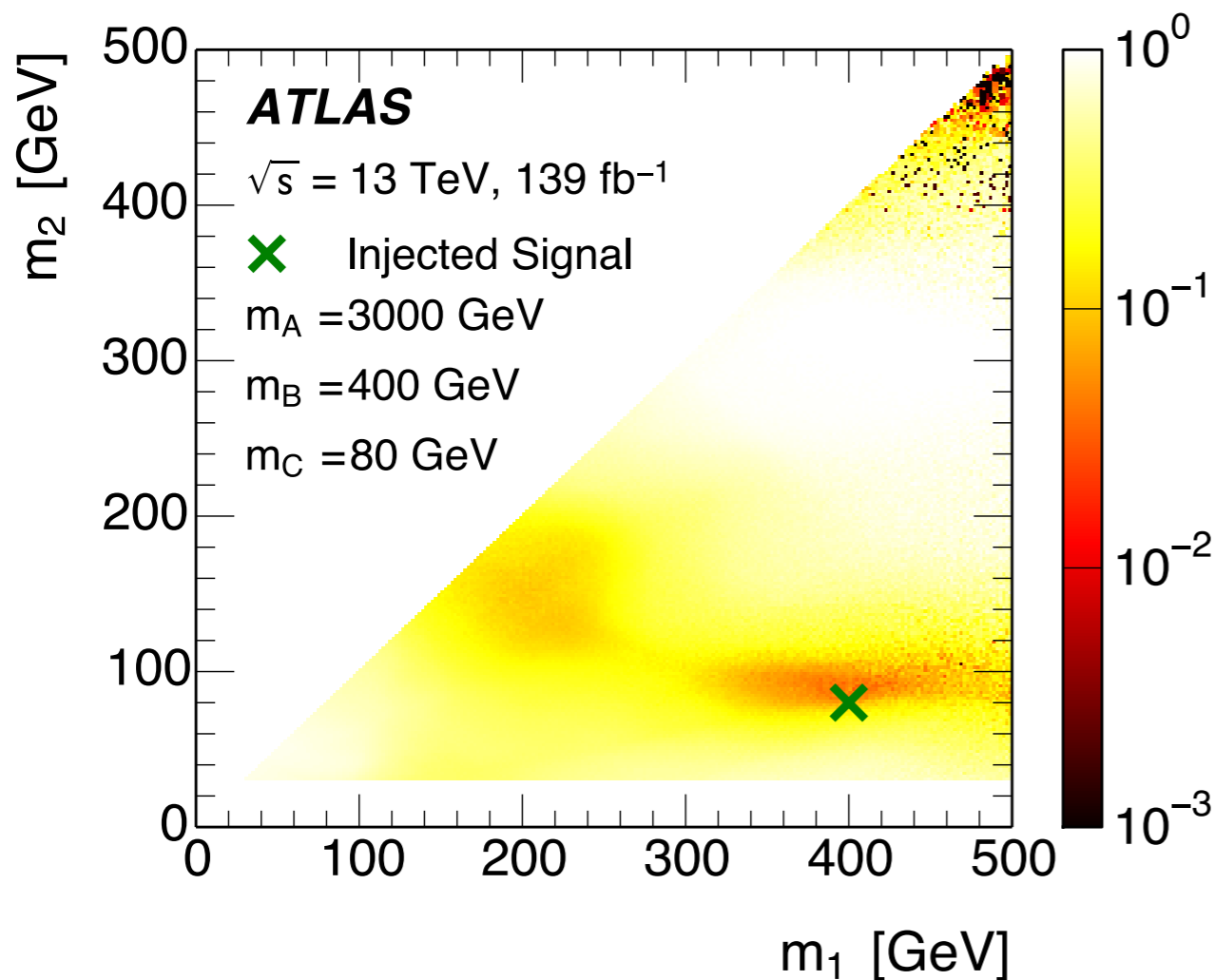


J. Collins, K. Howe, BN,  
PRL 121 (2018) 241803, 1805.02664



# Example: two-jet search





res = mass of

A first version of this search has been performed by **ATLAS!**

*Phys. Rev. Lett.* 125 (2020) 131801, 2005.02983



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Physics Briefing

## Machine learning qualitatively changes the search for new particles

Tags:  
[machine learning](#),  
[analysis](#)

13 May 2020 | By [ATLAS Collaboration](#)

While powerful, the approach I've just described has multiple challenges when scaling up.

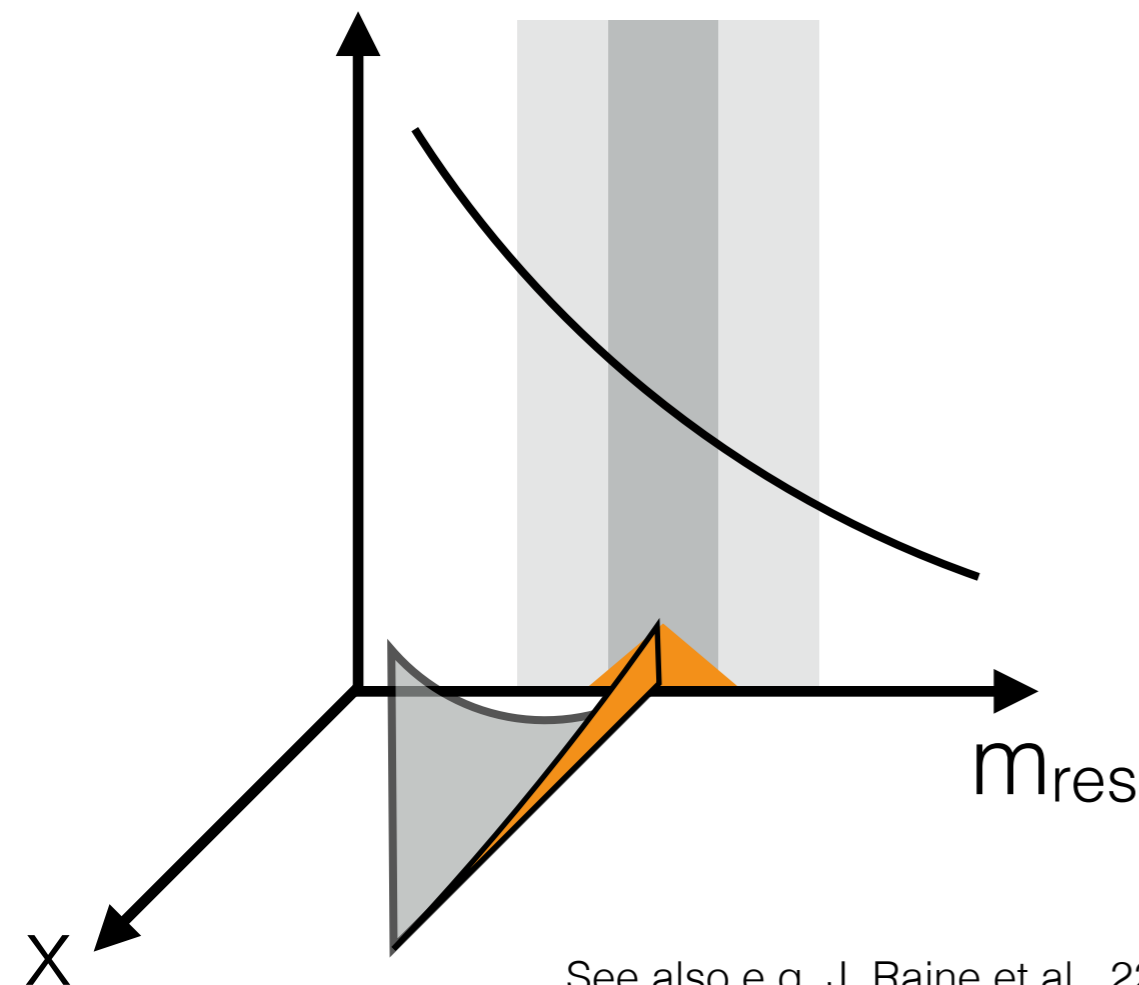


While powerful, the approach I've just described has multiple challenges when scaling up.

## Example Challenge: Decorrelation

The approach doesn't work if  $m_{\text{res}}$  and  $x$  are strongly related.

For instance, consider the extreme case where  $m_{\text{res}}$  is part of  $x$ .



See also e.g. J. Raine et al., 2203.09470

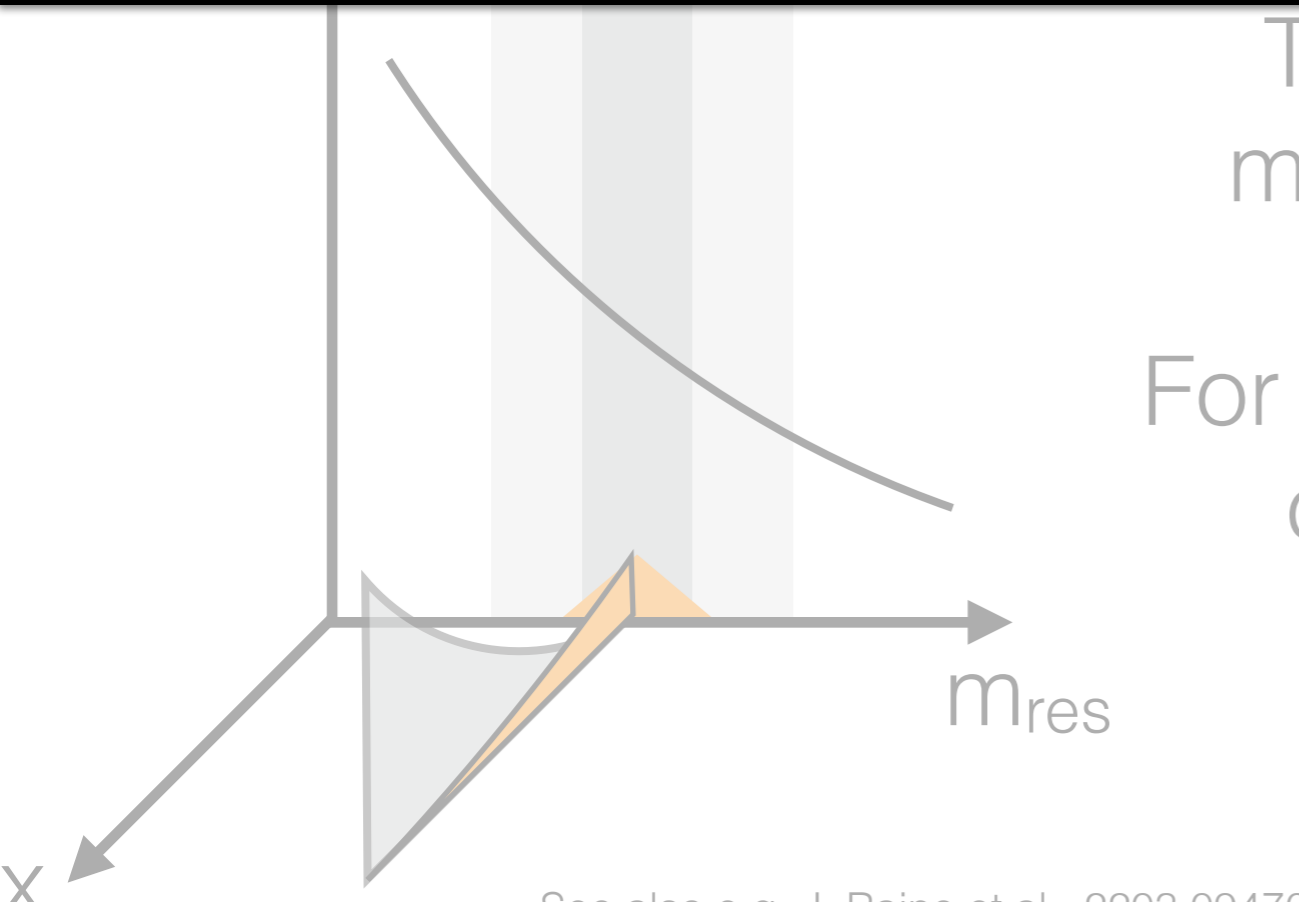
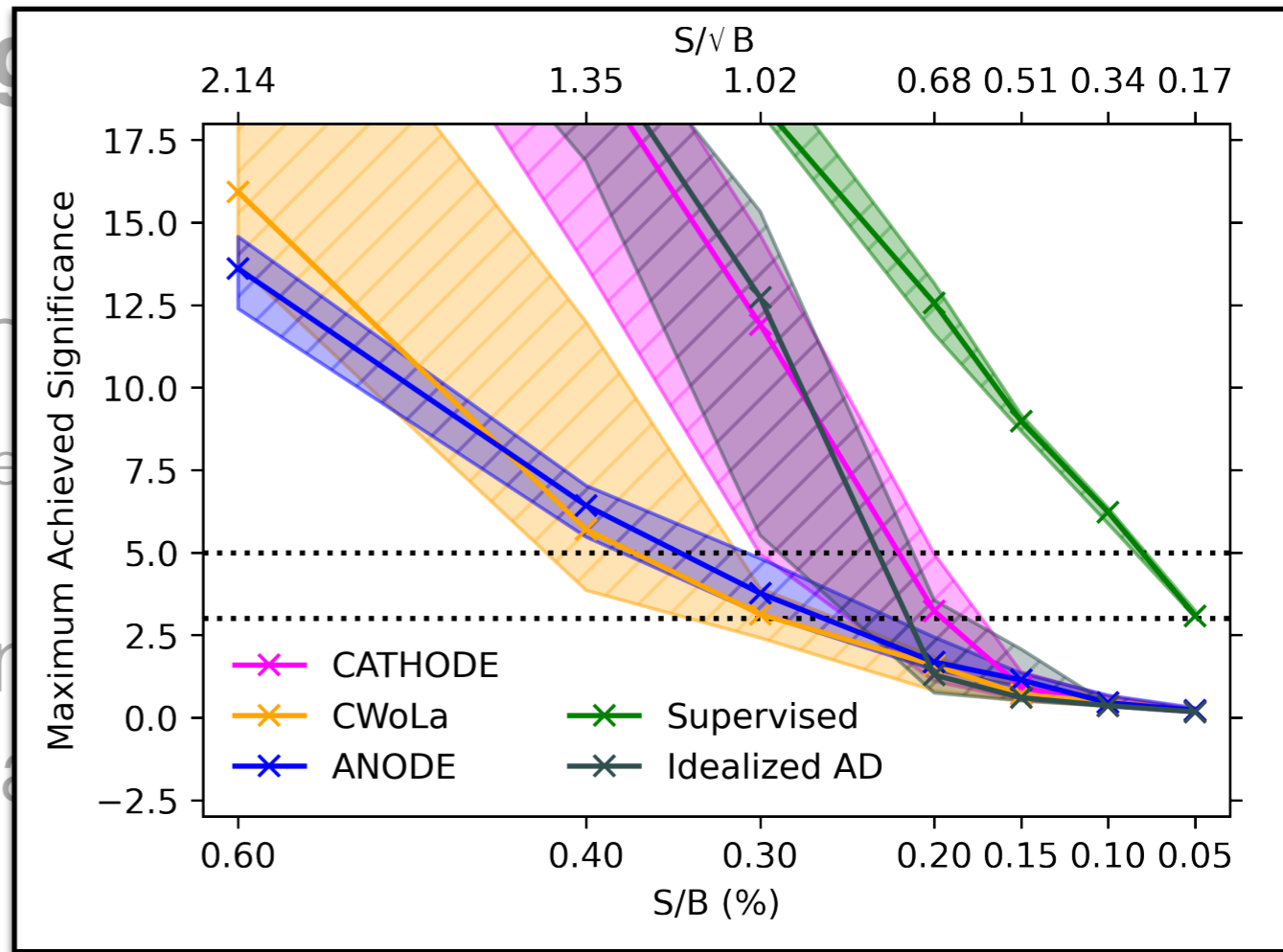
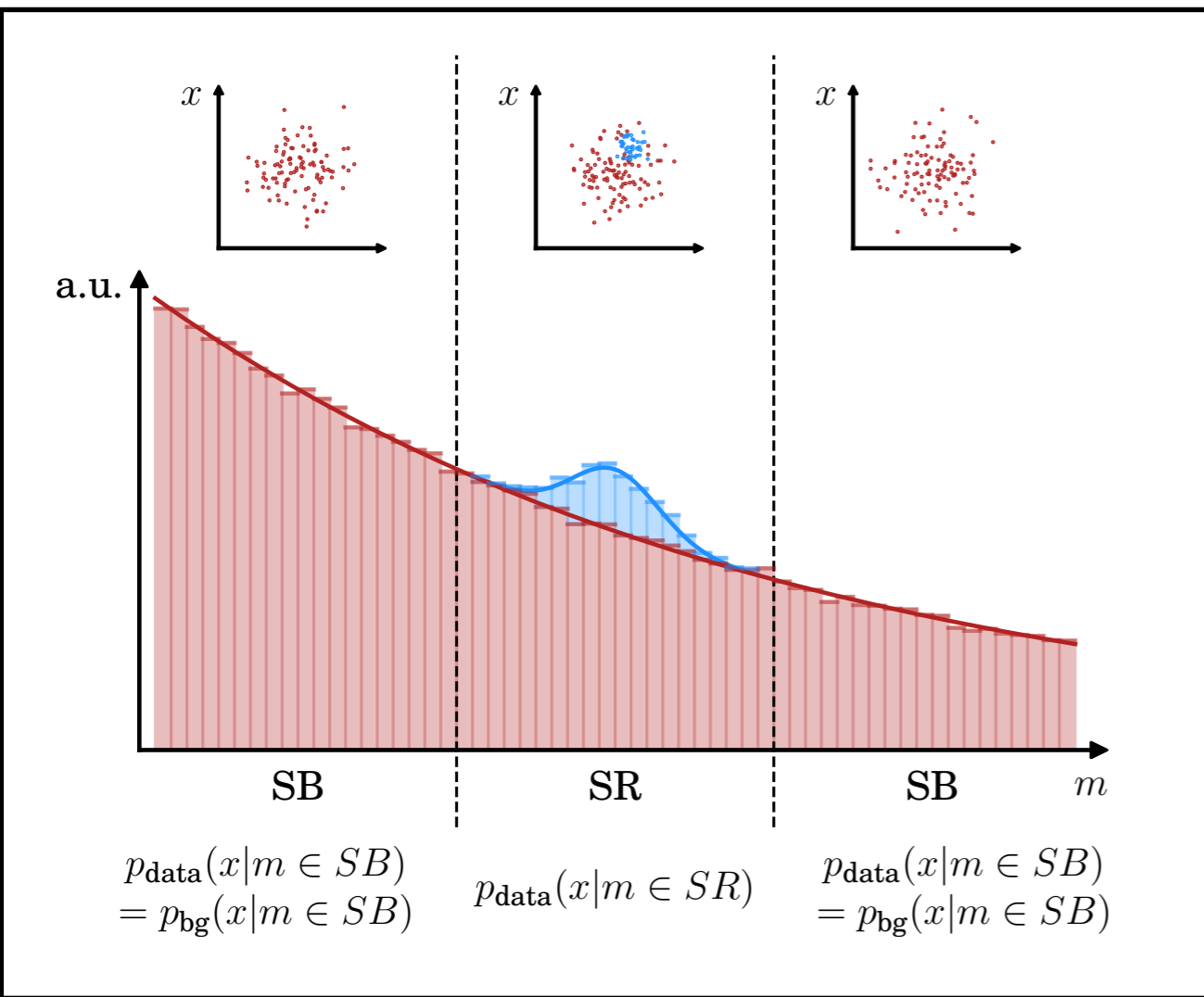
K. Benkendorfer, L. Le Pottier, BN, 2009.02205

A. Hallin et al., 2109.00546

A. Andreassen, BN, D. Shih, PRD 101 (2020) 095004, 2001.05001

BN and D. Shih, PRD 101 (2020) 075042, 2001.04990

approach I've just described  
 scales well when scaling up.



K. Benkendorfer, L. Le Pottier, BN, 2009.02205

**A. Hallin et al., 2109.00546**

A. Andreassen, BN, D. Shih, PRD 101 (2020) 095004, 2001.05001

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See also e.g. J. Raine et al., 2203.09470

While powerful, the approach I've just described has multiple challenges when scaling up.

We also need to benchmark new approaches.

## The LHC Olympics 2020

A Community Challenge for Anomaly Detection in High Energy Physics



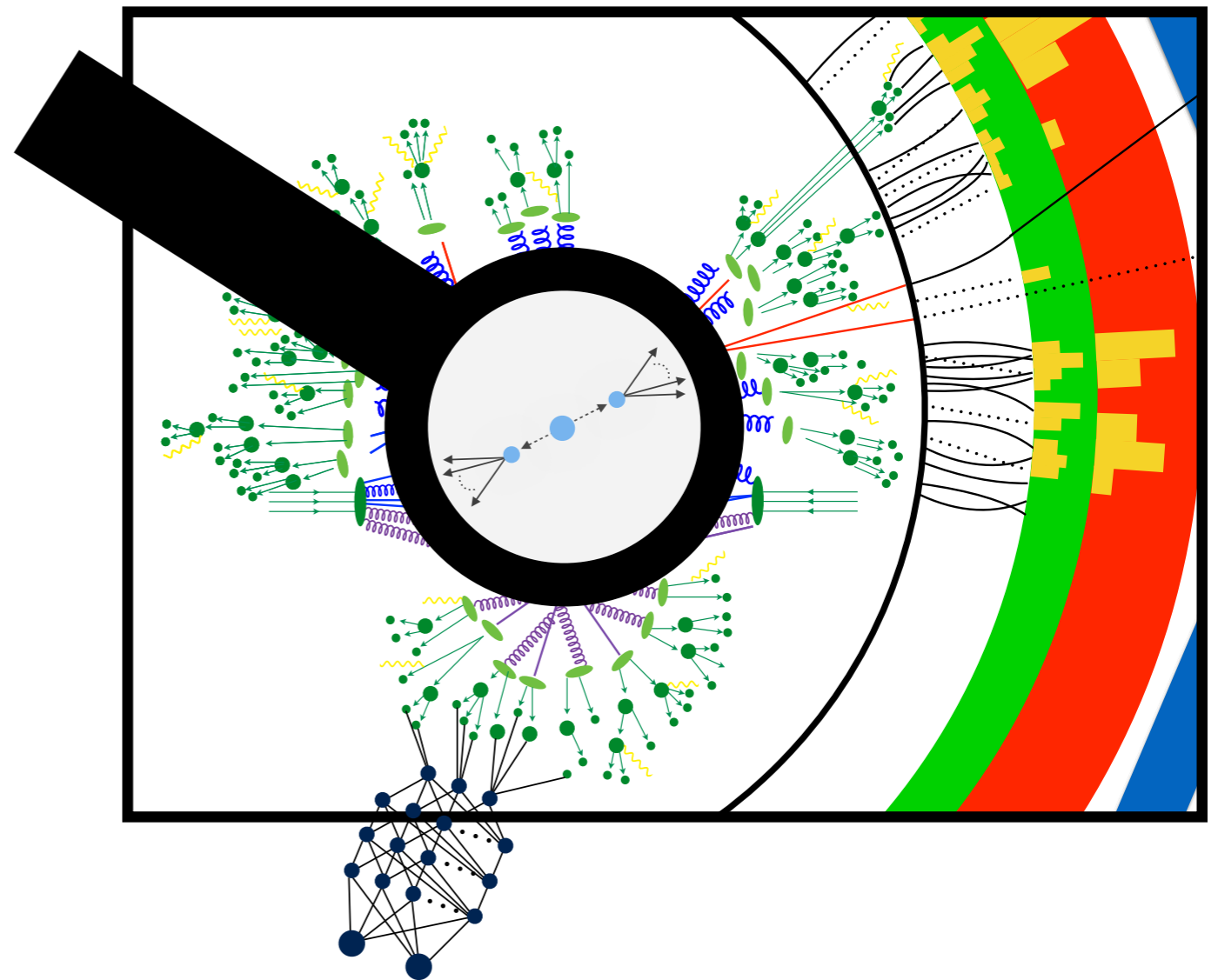
Comparing model independent approaches is difficult, which is why we put together the LHC Olympics datasets + challenge

(see also [Dark Machines](#) and [ADC2021](#))

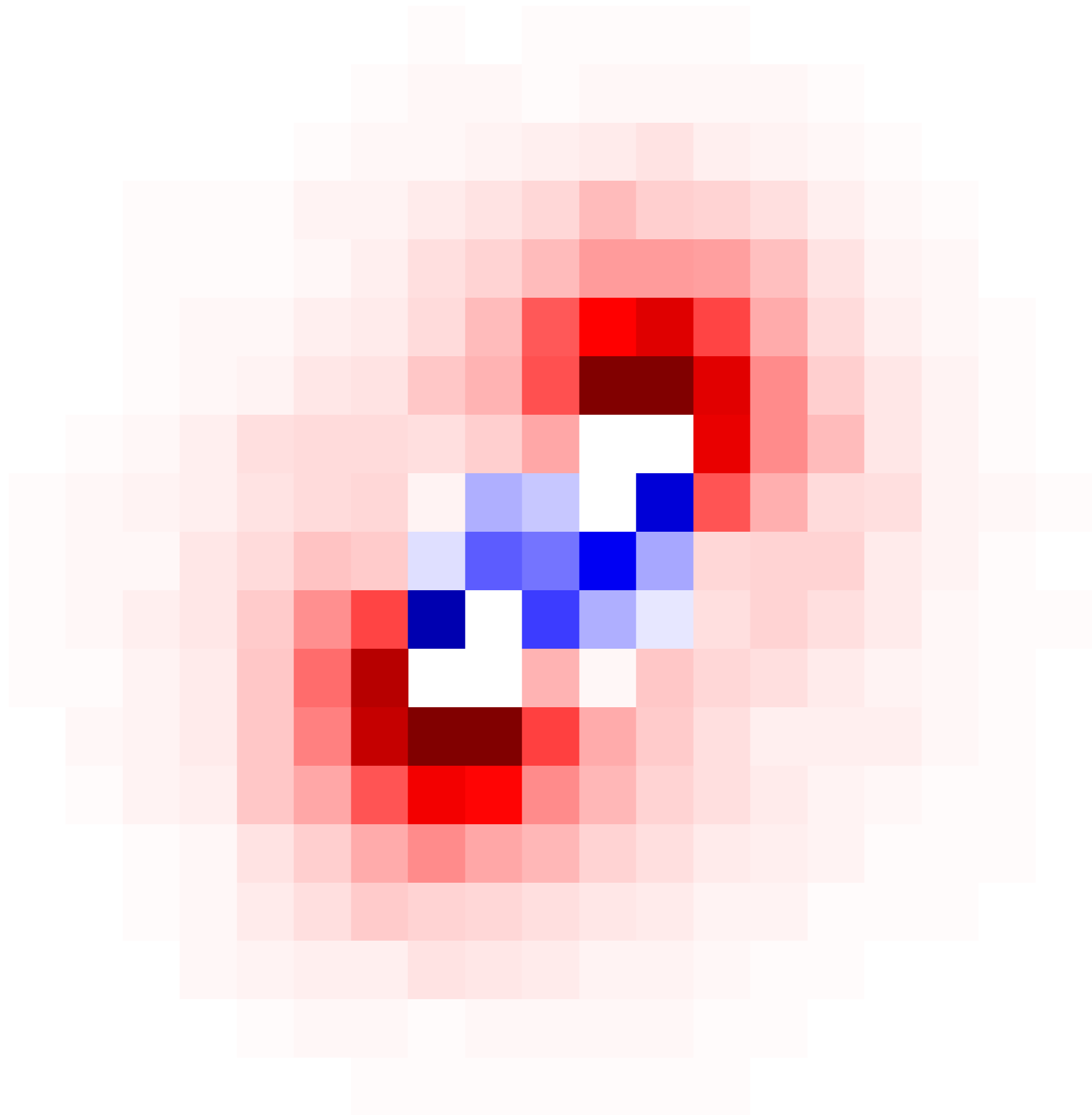
It is an exciting time to work on anomaly detection for the LHC and beyond!

This is a rapidly growing area with lots of room for innovation (and from physicists!)

We will need many approaches to achieve broad coverage



See the [Living Review](#) for more refs!



Fin.

# Backup



# Results with data

Phys. Rev. Lett. 125 (2020) 131801, 2005.02983

