

## *FASER2* **FPF Workshop #4** 31/1/2022





FASER

- If The existing FASER experiment is already set to probe new phase space.
- ▶ But FASER's size is heavily constrained by the available space underground
- ▶ The potential reach with an enlarged detector "FASER2" is under study
	- ! Decay Volume: Length=5m, Diameter=2m
- ▶ 4 orders of magnitude improvement in Reach
	- ! Angular acceptance of of all neutral pions:
		- $\triangleright$  0.6% in FASER
		- $\triangleright$  10% in FASER2
	- Improves sensitivities to LLPs produced in decays of heavy mesons
	- **Improves sensitivity to larger LLP masses**
- ▶ FASER already starting to give prood-of-principle of detector design and philosophy!

# FASER FASER **FASER Upgrade for HL-LHC**







# FASER FASER **FASER2 Reach**





[\[arXiv:1811.12522\]](https://arxiv.org/abs/1811.12522)



# FASER FASER **FASER2 Reach**





[\[arXiv:1811.12522\]](https://arxiv.org/abs/1811.12522)



# FASER FASER **FASER2 Reach**





[\[arXiv:1811.12522\]](https://arxiv.org/abs/1811.12522)



# FASER FASER **FASER2 Design**

- **I** Design considerations for FASER2
	-
	- $\triangleright$  More decay channels  $\rightarrow$  Need for particle ID
	-
	-
- ! Larger radius → Being on-axis less important
	-
- ! Larger detector → Larger background rate
	- → Different/cheaper technology
- $\triangleright$  Link to FASERv2  $\rightarrow$  Measure µ charge (and momentum) from (τ and μ) neutrino interactions
- ! Will be similar in philosophy to FASER
	- ▶ Can be optimised for different FPF scenarios
	- **Still much to be studied in** terms of possible detector configurations and







! Available space for FASER2 based on different facility scenarios:

# FASER FASER **FASER2 Design** | Parameters

### ! Possible detector configurations:



### **Scenario 2 (S2)**

- New dedicated cavern









# FASER FASER **FASER2 Design** | Veto

## ! The veto system will be similar to FASER.

- ▶ Reasonably simple to extend scintillator-based technology to cope with higher muons rates at HL-LHC.
	- ▶ Planning to use Fluka to get concrete estimate for muon rate.





 $\triangleright$  Factor 10 increase in radius  $\rightarrow$  100x increase in area to be instrumented.

# FASER FASER **FASER2 Design** | Tracking

- 
- ▶ Significant increase in detector length of FASER2 could allow larger decay product separations with different and possibly cheaper technology.

! Much more challenging to accommodate extended version of ATLAS SCT tracker module configuration, due to cost and services considerations.

Particles from IP





# FASER FASER **FASER2 Design** | Tracker technologies

- If There are two current main tracker technology candidates:
	- ! A SiPM and scintillating fibre tracker technology, such as LHCb's SciFi detector is a strong candidate
	- ▶ Monitored Drift Tube (MDT) technology, similar to ATLAS New Small Wheel also being considered
		- ! Requires the use of gases in the LHC tunnel that could be problematic for the Alcove scenario.





# FASER FASER **FASER2 Design** | Magnets

- $\triangleright$  To maintain sufficient field strength across the much larger aperture superconducting magnet technology is likely to be required
	- ! Suitable technology for this already exists and can be built for FASER2.
	- ! Cooling is one of the main obstacles but several possibilities:
		- **I** Use of cryocoolers
		- ! Share a single cryostat across several magnets





# FASER FASER **FASER2 Design** | Calorimeter

Particles from IP

! Sufficient spatial resolution to be able to identify particles ∼mm-cm separation;

- ! The calorimeter needs to have:
	-
	- ! Good energy resolution
	- ! Improved longitudinal separation with respect to FASER
	- $\blacktriangleright$  The capability to perform particle identification,
		- ▶ Separating e.g. electron and pions.
- ! Dual readout calorimetry is a good candidate to satisfy these requirements.





# FASER FASER **FASER2 Design** | Muon ID

Particles from IP

- If The ability to identify electrons and muons would be important for:
	- ! Signal characterisation
	- ! Background suppression
	- **The interface with FASERv2**
- $\triangleright$  Iron will be placed after the calorimeter
	- ! Sufficient depth to absorb pions and other hadrons (~1m  $\approx$  50 X<sub>0</sub>, ~5  $\lambda$ )
- ! Followed by a detector for muon identification







# ! Projections created with the FORSEE tool:



! [Phys. Rev. D 104, 035012](https://arxiv.org/abs/2105.07077)

### ! <https://github.com/KlingFelix/FORESEE>

 $arXiv.org > hep-ph > arXiv:2105.07077$ 

### **High Energy Physics - Phenomenology**

[Submitted on 14 May 2021]

### FORESEE: FORward Experiment SEnsitivity Estimator for the LHC and future hadron colliders

### Felix Kling, Sebastian Trojanowski

We introduce a numerical package FORward Experiment SEnsitivity Estimator, or FORESEE, that can be used to simulate the expected sensitivity reach of experiments placed in the far-forward direction from the proton-proton interaction point. The simulations can be performed for 14 TeV collision energy characteristic for the LHC, as well as for larger energies: 27 and 100 TeV. In the package, a comprehensive list of validated forward spectra of various SM species is also provided. The capabilities of FORESEE are illustrated for the popular dark photon and dark Higgs boson models, as well as for the search for light up-philic scalars. For the dark photon portal, we also comment on the complementarity between such searches and dark matter direct detection bounds. Additionally, for the first time, we discuss the prospects for the LLP searches in the proposed future hadron colliders: High-Energy LHC (HE-LHC), Super proton-proton Collider (SppC), and Future Circular Collider (FCC-hh).

Comments: 11 pages, 3 figures, FORESEE code available at this https URL High Energy Physics - Phenomenology (hep-ph) Subjects: arXiv:2105.07077 [hep-ph] Cite as: (or arXiv:2105.07077v1 [hep-ph] for this version)

Search. Help | Adv

### $\equiv$  README.md

### **FORESEE: FORward Experiment SEnsitivity Estimator**

By Felix Kling and Sebastian Trojanowski

arXiv 2105.07077

### **Introduction**

We present the numerical package FORward Experiment SEnsitivity Estimator, or FORESEE, that can be used to simulate the expected sensitivity reach of experiments placed in the far-forward direction from the proton-proton interaction point. We also provide a comprehensive list of validated forward spectra of various SM species.

### Paper

Our main publication FORESEE: FORward Experiment SEnsitivity Estimator for the LHC and future hadron colliders provides an overview over this package. We recommend reading it first before jumping into the code.

### **Tutorials**

In the main folder in this repository, we provide tutorials for different LLP models: the dark photon, the dark Higgs, the ALP with W couplings and the up-philic scalar.

## ! **FASER2-default**

## ! **Scenario 1:**

! Significantly degraded sensitivity due to reduced decay volume length

## ! **Scenario 2:**



- ! Comparable sensitivity to FASER2-default, but somewhat improved due to larger decay volume length.
- ! Very small degradation in diagonal due to increased distance from IP.

**Dark Higgs** 

# FASER FASER **FPF Scenarios** | Dark Higgs

## ! **FASER2-default**

## ! **Scenario 1:**

! Significantly degraded sensitivity due to reduced decay volume length

EXTERNATION:<br>
Diameter of detector much  $\sum_{i=1}^{\infty} 10^{-4}$ more important here. Due to larger angle emission from B-hadrons of LLP.

 $10^{-1}$ 





## ! **Scenario 2:**

### ! **Scenario 2**: ▶ Check effect of different production and decay modes



- ! **Production modes** rather different than for FASER
	- ▶ Pion decay at low mass
	- $\triangleright$  Then eta decay
	- ▶ Then Dark Bremsstrahlung

## ! **Scenario 2**:



- ! **Decay modes** also very different to FASER
	- **Electron decay at low mass**
	- **Muon decay**
	- ▶ Hadrons



## ! **Scenario 2**:



# $\zeta$  Simulation studies

- ▶ Started Geant4 simulations of possible FASER2 designs
	- ▶ Focussing on magnets and particle separations
	- Impacts tracker and calorimeter design considerations
- ! Using events generated with FORESEE as input to G4
	- ! LLP spectra and decays handled by FORESEE
	- ▶ Currently only looking at Dark Photon decay to e+ e-.

Josh McFayden | FPF | 31/1/2022









# $\zeta$  Simulation studies

- ▶ Started Geant4 simulations of possible FASER2 designs
	- ▶ Focussing on magnets and particle separations
	- Impacts tracker and calorimeter design considerations
- ! Using events generated with FORESEE as input to G4
	- ! LLP spectra and decays handled by FORESEE
	- ▶ Currently only looking at Dark Photon decay to e+ e-.

### Josh McFayden | FPF | 31/1/2022















4000

5000

# FASER FASER **Simulation studies** | Scenario 1 ! **Scenario 1:**

6000

7000

8000

- ! 1T field gives ~5mm separation at Station 1 with 1.5m DV.
- ! At Station 3/Calo need ~20 mm resolution for good separation.







0

1000

2000

4000

5000

6000

7000

8000F

# FASER FASER **Simulation studies** | Scenario 1 ! **Scenario 1:**



- ! 1T field gives ~5mm separation at Station 1 with 1.5m DV.
- ! At Station 3/Calo need ~20 mm resolution for good separation.
- ! Larger radius = softer LLPs







# FASER FASER **Simulation studies** | Scenario 1 ! **Scenario 1:**

- ! 1T field gives ~5mm separation at Station 1 with 1.5m DV.
- ! At Station 3/Calo need ~20 mm resolution for good separation.
- ! Larger radius = softer LLPs













# FASER FASER **Simulation studies** | Scenario 2 ! **Scenario 2:**



! Significantly increased detector/ magnet lengths results in large separations even at first station.





0

1000

2000

3000

4000

5000

7000

# FASER FASER **Simulation studies** | Scenario 2 no B-field ! **Scenario 2:**



- ! Significantly increased detector/ magnet lengths results in large separations even at first station.
- ! Even without any magnetic field separations are comparable to Scenario 1.



# **E** Simulation studies | Scenario 2 E = 1 TeV ! **Scenario 2:**



- ! Significantly increased detector/ magnet lengths results in large separations even at first station.
- ! Even without any magnetic field separations are comparable to Scenario 1.
- ▶ But need to consider the effect on the reach of high momentum LLPs
	- ▶ E.g. much smaller separations at high energy





28

0

50

100

150

200

250

300

# FASER FASER **Simulation studies** | Scenario 2 coupling ! **Scenario 2:**



- ! Significantly increased detector/ magnet lengths results in large separations even at first station.
- ! Even without any magnetic field separations are comparable to Scenario 1.
- ! But need to consider the effect on the reach of high momentum LLPs
	- ! E.g. much smaller separations at high energy
	- **EXT** Lower couplings require larger boost to reach decay volume
	- ! **Need to see full effect on reach!**













# $\zeta$  Simulation studies

- ! Possible detector technologies
	- **Tracker** 
		- ! SCT:
			- $\triangleright$  Pitch = 80 um, resolution = 17 um
			- Naive hit cluster ~300 um
		- ▶ SciFi:
			- $\triangleright$  Pitch = 250 um, Resolution = 100 um
			- Naive hit cluster: ~1 mm?
		- $\blacktriangleright$  Reduced resolution of SciFi detector seems acceptable even in 1 $\Box$
	- ! Calorimeter
		- ▶ Dual readout
		- ! Resolution: <10 mm? (Tested with fibre diameter of 1mm)
		- ! Dual readout calorimeter resolution sufficient at 3rd tracker station

### ! Also need to extend studies to include LLP mass reconstruction

- I Not possible without magnet
- ▶ Want to understand what is possible with a magnet.







 $\triangleright$  Could actually be too much - particles may be bent out of the detector and miss the last

# FASER FASER **Other considerations**

 $\triangleright$  There are models where  $\pi_0$ s are important - may want to reconstruct these from

 $\triangleright$  As the  $\pi_0$  will be high energy this can be challenging as 2 photons will be very collimated and

## ! Separations of up to >1m

- tracking station and also the calorimeter
- $\triangleright$  For high energy particle momentum measurements alignment of the tracker planes becomes very important
	- ! (and will likely dominate the momentum resolution over the hit resolution).
	- ▶ With an electromagnet alignment can be constrained with field off data
		- ! Need to know detector movements when magnet is ramped up may require (optical?) alignment system.
- calorimeter
- the decay position of the  $\pi_0$  will be unknown.







# FASER FASER **Conclusions**

- If the physics potential of a larger scale successor to FASER is clear.
- ▶ Possible scenarios for this larger detector are being explored and initial reach studies strongly indicate a preference for a FPF with a dedicated new cavern.
- ! Possible detector technologies already being identified.
- **I** Design simulations in Geant4 started
	- ! Refining understanding of detector technology needs and optimal layouts.
	- ▶ Machinery used for simply signature so far
	- ! Several important studies still to perform:
		- ! Determine possible mass reconstruction performance (also including material in simulation) ▶ Extend reach studies to other Dark Higgs and other more complicated decay channels
		-





# **Back-ups**



# FASER FASER **Target scenarios** | Dark Photon







# FASER FASER **Target scenarios** | Dark Higgs







# FASER FASER **Target scenarios** | Dark Higgs






## FASER FASER **Target scenarios** | ALP









## FASER FASER **Target scenarios** | ALP









 $\frac{1}{3}$  0.16 0.14  $0.12$  $0.1$ 0.08 0.06 0.04  $0.02$  $\boldsymbol{0}$ 











## FASER FASER **Validation** | 1 TeV









# FASER FASER **Simulation studies** | FASER









# FASER FASER **Simulation studies** | Scenario 1







# FASER FASER **Simulation studies** | Default FASER2











# FASER FASER **Simulation studies** | Scenario 2





# $\zeta$  Simulation studies











1

### FASER FASER **Simulation studies** | Coupling ! **Scenario 2:**







# FASER FASER **Simulation studies** | Mass







# **E** Simulation studies | Scenario 2 E = 1 TeV



## $E$  Dual readout calo

Josh McFayden | FPF | 31/1/2022







- ! **FASER2** (incl default)
- ! **S1 (alcove)** 
	- ! Worse than very narrow F2

# FASER FASER **FPF Scenarios** | Dark photons



#### FASER FASER **FPF Scenarios** | Dark photons ! **FASER2** (incl default) **NA48**  $10^{-3}$ ! **S1 (alcove)**  ! Worse than very narrow F2 **HPS** ! **S2 (enlarged UJ12)**  $10^{-4}$ relic target  $\boldsymbol{\omega}$ ! Worse than very narrow F2 Kinetic Mixing

 $10^{-5}$ 

 $10^{-6}$ 

 $10^{-7}$   $\frac{1}{10^{-2}}$ 



#### FASER FASER **FPF Scenarios** | Dark photons ! **FASER2** (incl default) **NA48**  $10^{-3}$ ! **S1 (alcove)**  ! Worse than very narrow F2 **HPS** ! **S2 (enlarged UJ12)**  $10^{-4}$  $-$ targe ic to GMA ! Worse than very narrow F2 Mixing  $\eta$ Nucal ! **S3 (new cavern)** Kinetic I  $10^{-5}$ ! Better than but comparable to F2 ! Slight shift in diagonal due to  $10^{-6}$ increased distance from IP. E137  $10^{-7}$   $\frac{1}{10^{-2}}$



#### FASER FASER **FPF Scenarios** | Dark photons ! **FASER2** (incl default) **NA48**  $10^{-3}$ ! **S1 (alcove)**  ! Worse than very narrow F2 **HPS** ! **S2 (enlarged UJ12)**  $10^{-4}$  $-$ targe ic to GMA ! Worse than very narrow F2 Mixing  $\eta$ Nucal ! **S3 (new cavern)** Kinetic I  $10^{-5}$ ! Better than but comparable to F2 ! Slight shift in diagonal due to  $10^{-6}$ increased distance from IP.  $10^{-7}$   $\frac{1}{10^{-2}}$





- ! **FASER2** (incl default)
- ! **S1 (alcove)** 
	- ! Widest S1 design comparable to  $D=1m$  F2.

# FASER FASER **FPF Scenarios** | Dark Higgs



- ! **FASER2** (incl default)
- ! **S1 (alcove)** 
	- ! Widest S1 design comparable to  $D=1m$  F2.
- ! **S2 (enlarged UJ12)**
	- ! Widest S2 design comparable to D=1m F2.

 $10^{-5}$ 

 $10^{-1}$ 

 $10^{-3}$ 

## FASER FASER **FPF Scenarios** | Dark Higgs





#### **E** Forward Physics Facility | Dark Higgs ! **FASER2** (incl default) ! **S1 (alcove)**   $10^{-3}$ ! Widest S1 design comparable to D=1m F2. ! **S2 (enlarged UJ12)** ! Widest S2 design comparable to D=1m F2. ! **S3 (new cavern)** ! Better than but comparable to F2  $10^{-5}$ ! Slight shift in diagonal due to increased distance from IP.  $10^{-1}$



- ! **FASER2** (incl default)
- ! **S1 (alcove)** 
	- ! Widest S1 design comparable to  $D=1m$  F2.
- ! **S2 (enlarged UJ12)**
	- ! Widest S2 design comparable to D=1m F2.

# FASER FASER **FPF Scenarios**| Dark Higgs

 $10^{-3}$ 

 $10^{-5}$ 

 $10^{-1}$ 

#### ! **S3 (new cavern)**

- ! Better than but comparable to F2
- ! Slight shift in diagonal due to increased distance from IP.



 $\bar{E}$  Light Weak DM Motivation ! The **LHC experiments** are producing incredible results, extending reach to more extreme phase-spaces and performing increasingly precise measurements.

! But the lack of any observation of BSM physics motivates **looking elsewhere** too.





# $\bar{E}$  Light Weak DM Motivation

▶ Main region of interest is for new particles that satisfy DM relic density requirements.





- ! One of the defining characteristics of weakly interacting light particles is their **long lifetime.**
- ! Distinct signatures
- ! But could still be produced in large numbers in hadron decays at ATLAS!



### $\bar{E}$  Light Weak DM Motivation









#### ! FASER is a new experiment at CERN!

! Data-taking starts in Run 3

#### ! Detector is 480m from ATLAS IP1

- **If** Directly in line with beam collision axis.
- $\triangleright$  Transverse radius of only 10cm covering the mrad regime (η>9.1)
- 
- $\triangleright$  From only 10-8 of solid angle 1% of  $\pi_0$ s are in acceptance.

## FASER FASER **LDM** | FASER

# $\blacktriangleright$  Inelastic pp cross section is huge  $\rightarrow$  10<sup>16</sup> collisions in Run 3  $\rightarrow$  10<sup>17</sup> π, 10<sup>13</sup> B

#### ! Old SPS → LEP tunnel

- **I** On line-of-sight (with some digging)
- $\triangleright$  Shielded by FASER layout  $\equiv$  high energy neutrino beamline
- $\blacktriangleright$  Low beam  $k$ 
	-







## FASER FASER **FASER Location**

#### ▶ The TI12 service tunnel just happens to be in just the right place for FASER:



#### FASER FASER **FASER Location**! In relation to ATLAS at Point 1

# IP1<br>ATLAS

forward jets

#### FASER layout  $\equiv$  high energy neutrino beamline

charged particles (P<7 TeV)



neutrino, dark photon

LHC tunnel

TI22 tunney

FASER<sub>V</sub>









! From first scouting photos…

#### FASER layout  $\equiv$  high energy neutrino beamline

forward jets

 $1 - 1 - 1$ 

**IHC** magnets

# FASER FASER **FASER Now Installed!**





# FASER FASER **FASER Now Installed!**

FASER

根本

 $Decay$ 

forward jets

**IHC** magnets





# FASER FASER **FASER Now Installed!**

FASER Decay



27 K. 7

forward jets





#### The detector consists of:

- Scintillator veto
- 1.5m long decay volume

- 2m long spectrometer
- 



**Calorimeter**





## FASER FASER **Target scenarios** | Dark photon






- ! Expected sensitivity of FASER for **dark photons**
- ! Detector signature:
	- $A' \rightarrow e+e-$
	- ! Charged tracks appearing in decay volume
	- $\triangleright$  Opposite charges separate through detecto
	- ! Significant energy deposit in calorimeter
- **> Sensitivity** 
	- ! Considers all production channels
		- ! Assumes no background, requires N=3 events
	- $\triangleright$  Reach limited by decay length (high  $\varepsilon$ ) and production rate (low ε)

## FASER FASER **Target scenarios** | Dark photon

## ! **New parameter space probed with just 1 fb-1 in 2022**

