

# Bremming Enhanced ALP Productions and FPF Sensitivity

4th Forward Physics Facility Meeting

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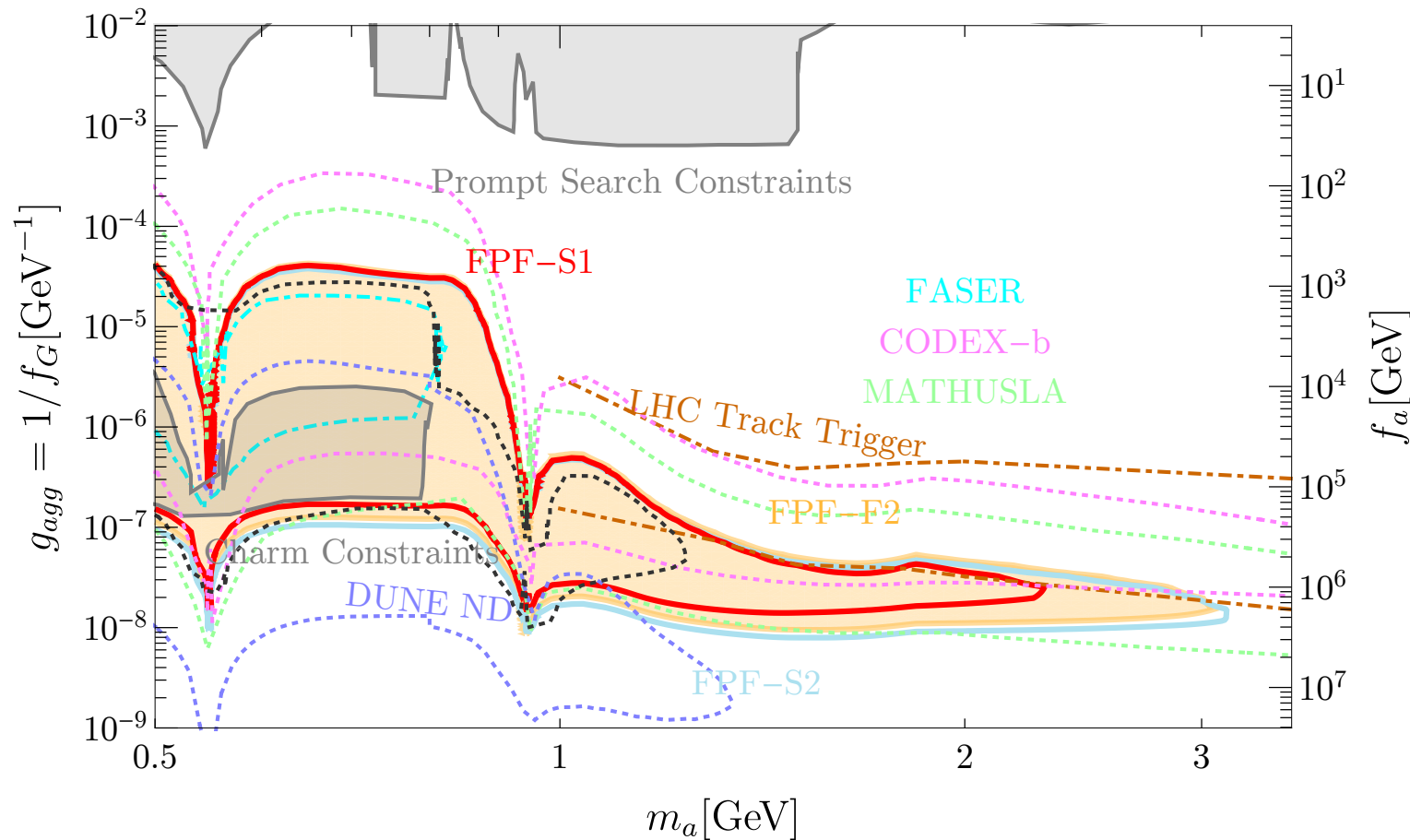
In collaboration with Zhen Liu

Paper in preparation  
Contribution to FPF whitepaper

# ALP Search at FPF

FPF provides a good platform for detecting the LLP in the very forward region.

Axion or Axion-like particle (ALP) solves the strong CP problem, good candidates for dark matter.



Heavy ALP  
Mass around GeV

Enhance FASER and  
FASER2 projection

Complementary to other  
experiments

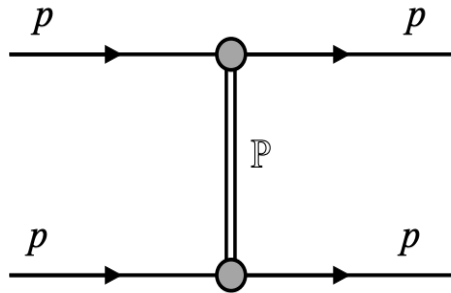
# ALP Production at LHC

Mixing with Mesons, Hard Scattering  
Bremsstrahlung

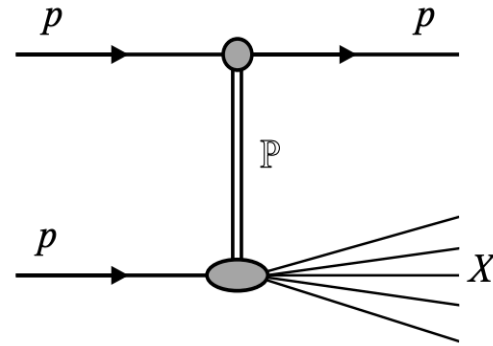
$$\frac{a}{8\pi f_a} \left( c_3 \alpha_3 G\tilde{G} + c_2 \alpha_2 W\tilde{W} + c_1 \alpha_1 B\tilde{B} \right)$$

**Gluon Dominated Region**

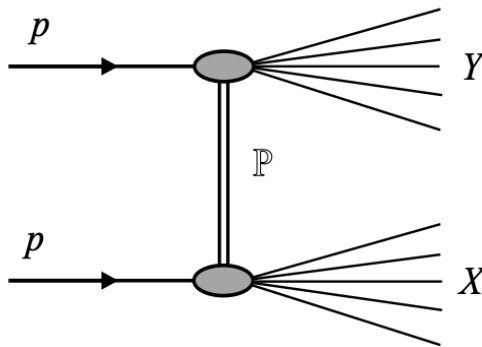
# Proton-Proton Scattering



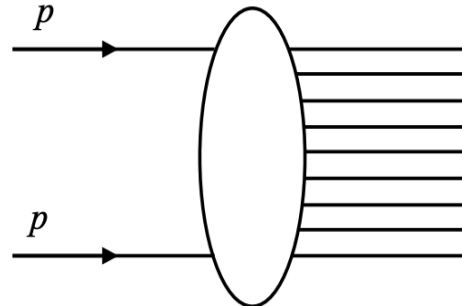
(a)



(b)



(c)



(d)

(a) Elastic Scattering

(b) Single Diffractive Scattering

(c) Double Diffractive Scattering

(d) Non-Diffractive Scattering

Particle Data Group (PDG) Prog. Theor. Exp. Phys. 2020, 083C01 (2020)  
Foroughi-Abari et al 2108.05900

We focus on the ALP as initial state radiation (ISR).

ISR from partons, proton?

Parton bremsstrahlung

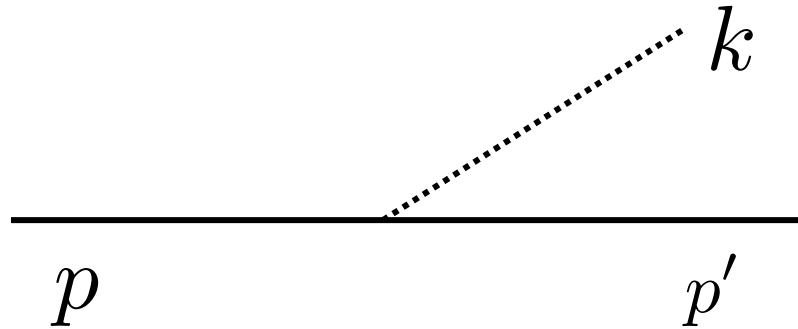
Our (preliminary) result shows that the ISR signal from gluon and quark bremsstrahlung is far smaller than the parton bremsstrahlung

Proton bremsstrahlung

Primary contribution.

# Splitting Function

The signal rate for emitting an axion from the proton bremsstrahlung



$$\sigma = \int dz dp_T^2 \frac{d\mathcal{P}}{dz dp_T^2} \hat{\sigma}_{\text{NSD}}(\hat{s})$$

$$\frac{d\mathcal{P}}{dz dp_T^2} = \frac{1}{16\pi^2} \frac{z}{1-z} \frac{1}{(p'^2 - m_p^2)^2} \left[ \frac{1}{2} \sum |\mathcal{M}|^2 \right] F_S(k, p')^2$$

Hadronic Form Factor

$$F_S(k, p') = F_{1,p}^a(k^2) F_{pp^*a}(p'^2)$$

$$F_{pp^*a}(p'^2) = \frac{\Lambda^4}{\Lambda^4 + (p'^2 - m_p^2)^2} \quad \Lambda = 1 \sim 2 \text{ GeV}$$

# Structure Form Factor

Effective operator

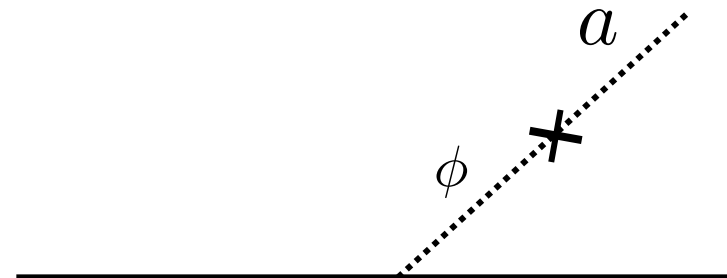
$$\frac{C_p(m_a)}{2f_a} \bar{p} \gamma^\mu \gamma_5 p \partial_\mu a$$

$$C_p(m_a) = C_p^0 + \sum_{\phi} \frac{C_{K,\phi} m_a^2}{m_a^2 - m_\phi^2} \frac{g_{\phi pp}}{m_p} - \sum_{\phi} \frac{C_{M,\phi} m_\phi^2}{m_a^2 - m_\phi^2} \frac{g_{\phi pp}}{m_p}$$

Kinetic mixing

Mass mixing

$C_{K,\phi}$   $C_{M,\phi}$  can be determined by the ChPT computation



$$m_a \ll m_{\pi^0} \quad C_p \rightarrow 0.47$$

Scenario 1:

$$|C_p^0| \approx 0.47 \quad \text{ignore the mixing}$$

Conservative!

Scenario 2:

$$|C_p^0| \approx 0 \quad g_{\pi^0 pp} = g_{\eta pp} = g_{\eta' pp}$$

Approximately show the mixing effects

# ALP Detection at FPF

After production at LHC, the ALP would propagate while decay

$$N_d = \int d \sin \theta_a dE_a \frac{dN_0}{dE_a d \sin \theta_a} \exp \left( - \frac{Dm_a}{\sqrt{E_a^2 - m_a^2} c \tau(m_a)} \right) \left[ 1 - \exp \left( - \frac{Lm_a}{\sqrt{E_a^2 - m_a^2} c \tau(m_a)} \right) \right]$$

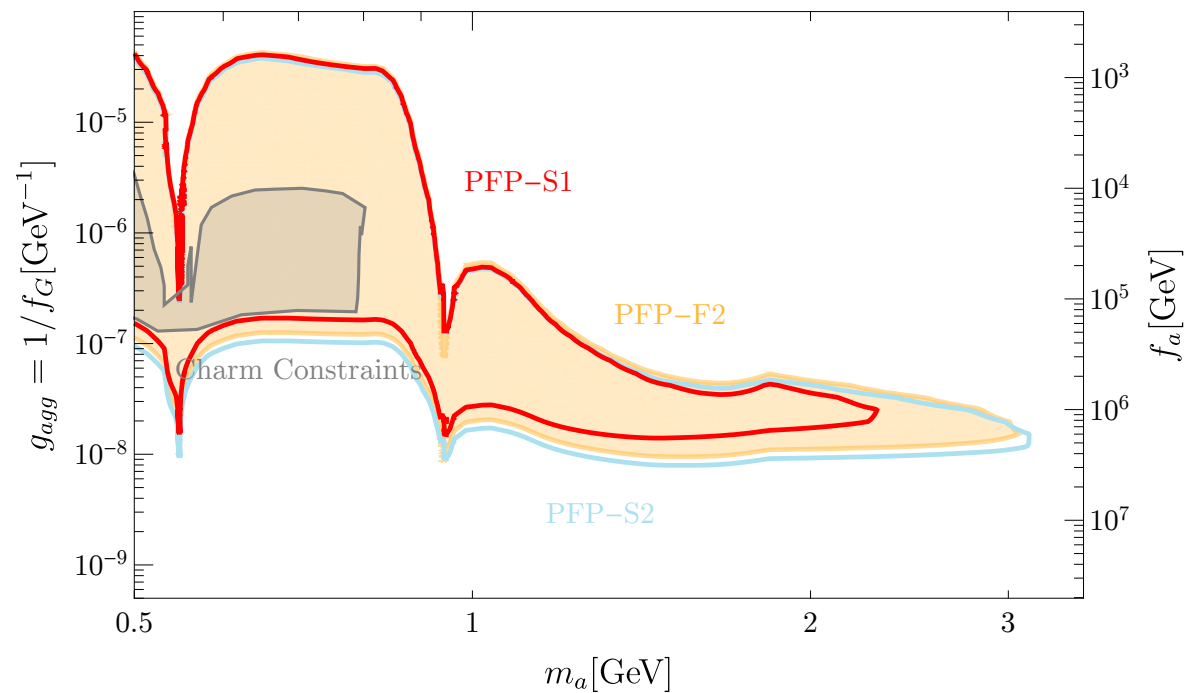
Production distribution

Decay behavior

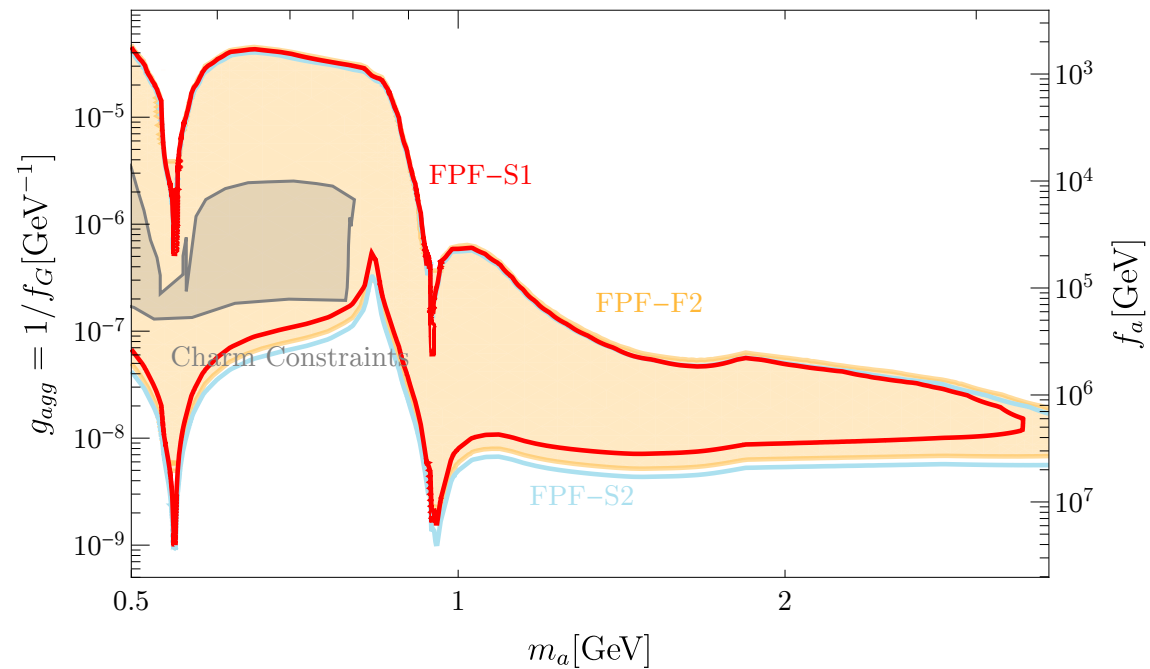
FPF Scenario	Distance to IP [m]	Available Length [m]	Decay Volume Length [m]	Available Diameter [m]	Decay Volume Diameter [m]
F2: Original FASER2	480	15	5	2	2 (/ 1 / 0.5)
S1: UJ12 Alcoves	500	5	1.5 (/ 2)	1.52	2 / 1 (/ 0.5)
S2: Purpose-Built Facility	620	25	10 (/ 15 / 20)	2	2 / 1 (/ 0.5)



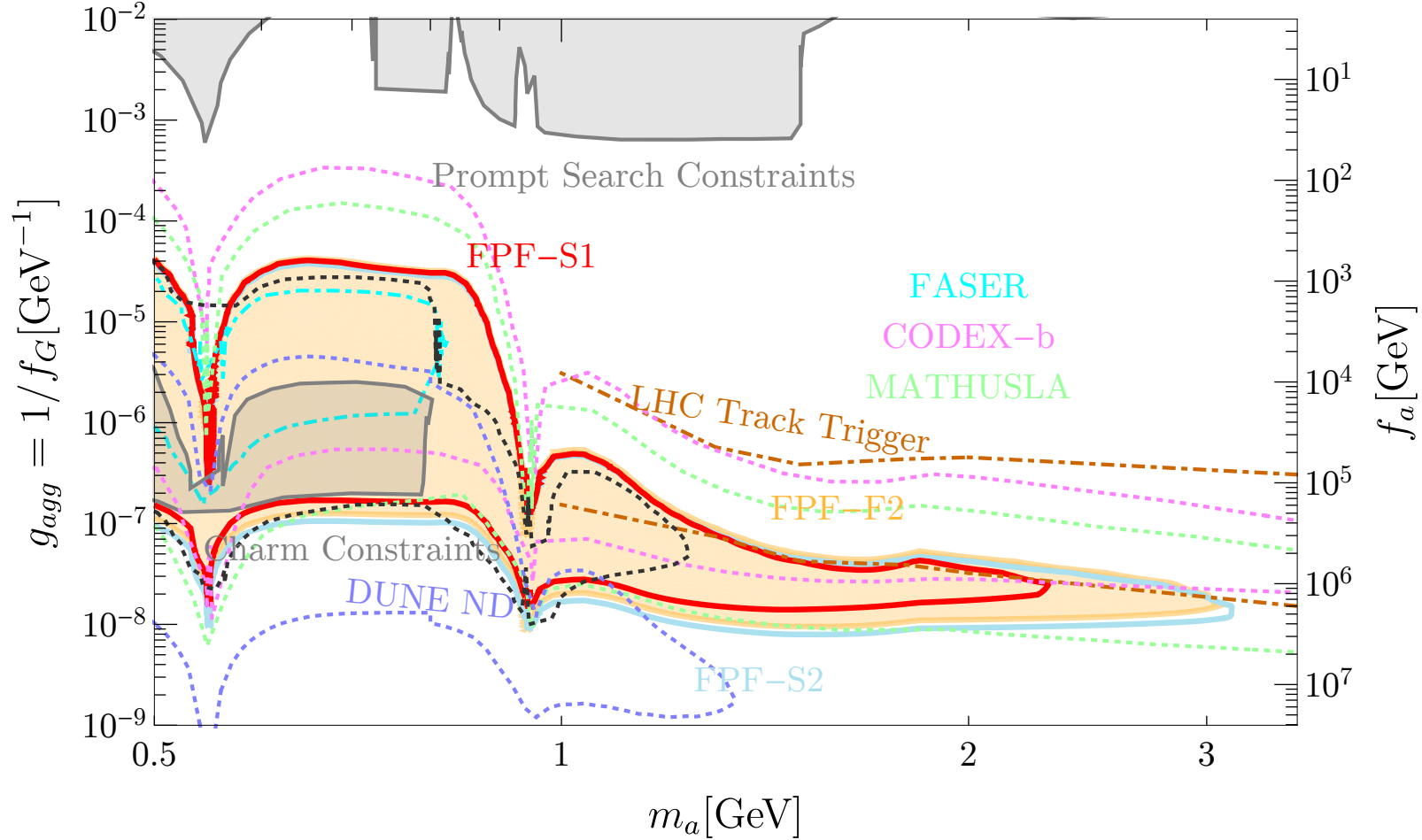
Scenario 1: Conservative



Scenario 2: Mixing effects



Scenario 1:  $|C_p^0| \approx 0.47$  ignore the mixing Conservative!

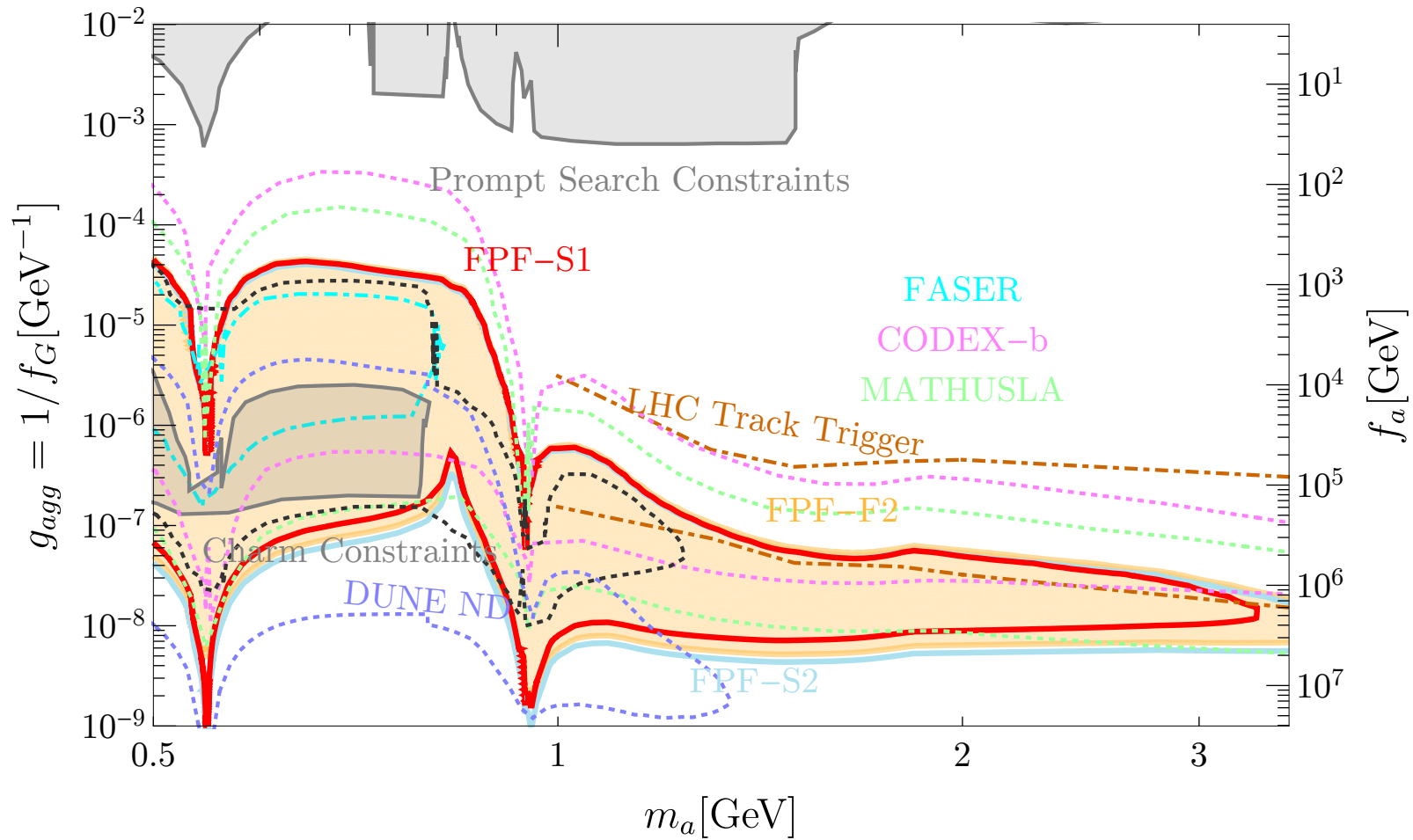


DUNE ND: 2011.05995  
 FASER: 1806.02348

LHC Track: 1911.12364  
 FASER2: 1811.12522

CODEX-b: 1911.00481  
 MATHUSLA: 1606.06298

Scenario 2:  $g_{\pi^0 pp} = g_{\eta pp} = g_{\eta' pp}$       Approximately show the mixing effects.



# Conclusion

The ALP production from proton bremsstrahlung process can enhance the previous FASER and FASER2 projected sensitivity and complementary for other experiments.

*Thank you!*