

# Dark matter freeze-in from semi-production

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Based on [JHEP 06 \(2021\) 026 \[2104.05684\]](#) (with A. Hryczuk)

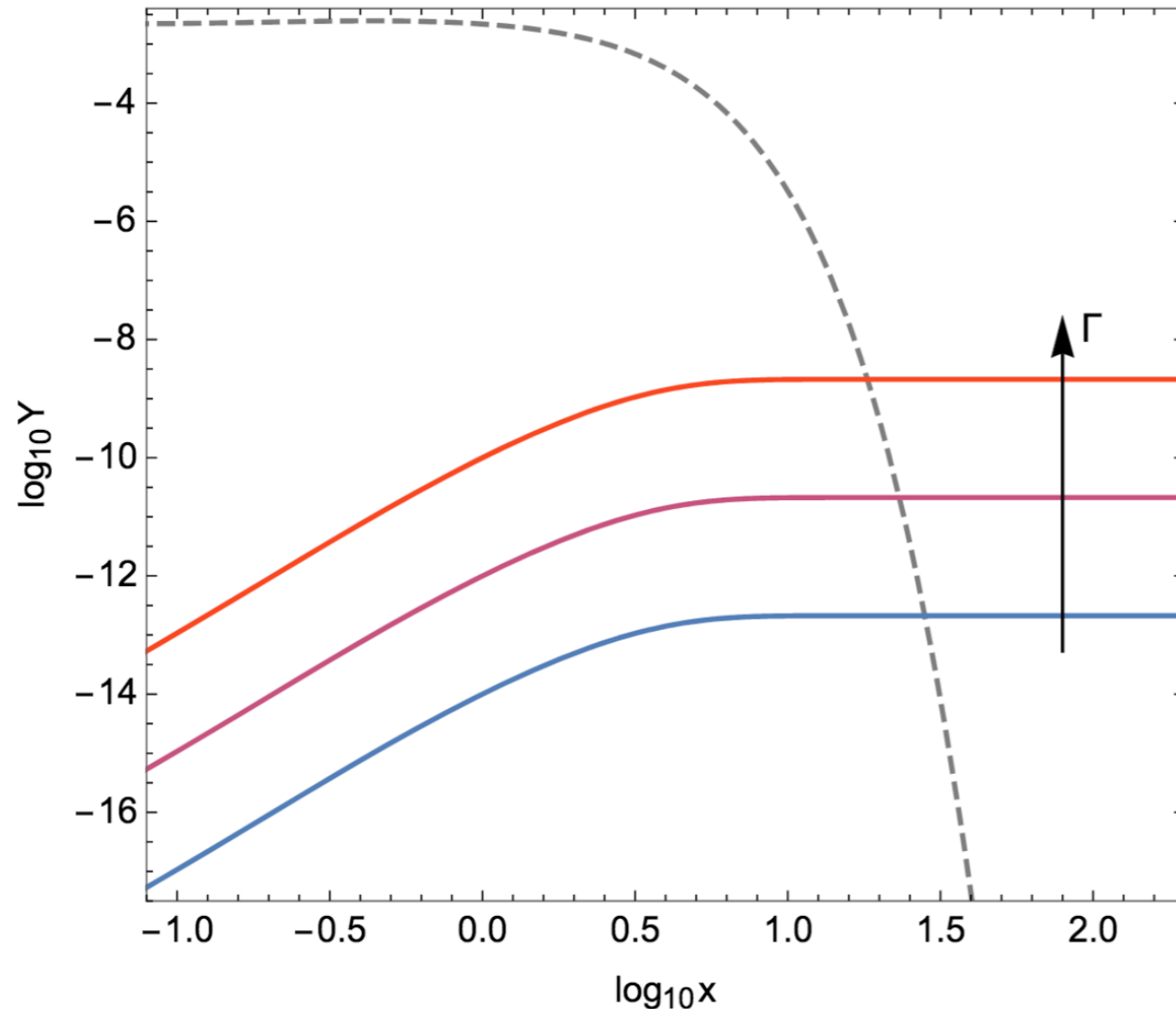
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# Freeze-in production of DM

From 1706.07442



$$x = \frac{m}{T}$$

- Dark matter interacts feebly with SM (**not in equilibrium**) and its initial concentration is tiny
- DM density freezes in when the density of SM is suppressed and the **production is inefficient**
- The production process can be a decay or annihilation

Typical annihilation cross section for freeze-in

$$\langle\sigma v\rangle \lesssim 10^{-40} \text{ cm}^3/s$$

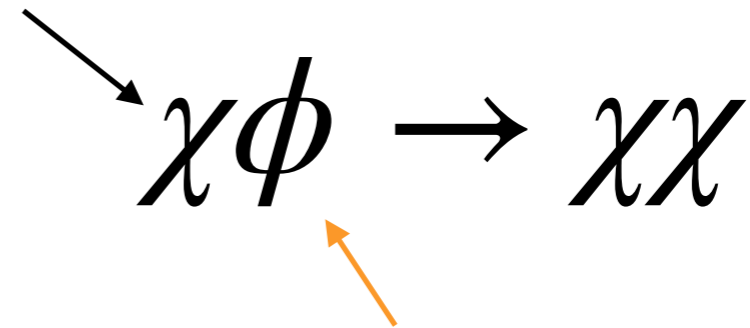
This is too weak for a near-future direct or indirect detection

IS IT POSSIBLE TO HAVE AN OBSERVABLE  
SIGNAL FROM FREEZE-IN?

# Semi-production

*Motivation for study*

Dark matter



New particle (mediator) or a SM state

What is different from the pair-annihilation freeze-in?

- The production rate is **proportional to the DM density**.  
Smaller initial abundance  $\rightarrow$  larger cross section
- Semi-production modifies the energy of DM particles in a non-trivial way, so the **temperature evolution can affect the relic density**

# Toy model

We start the investigation with a simple toy model

$$\mathcal{L}_{int} = \mathcal{L}_{SM} + \mathcal{L}_{\phi-SM} + \frac{\lambda}{2} \phi (\chi^3 + (\chi^*)^3) \quad \text{Z3 symmetry}$$

- Assume that  $\phi$  is **in equilibrium** with SM and  $\chi$  has some **tiny initial abundance** (e.g. from the gravitational production)
- The abundance of  $\chi$  is determined by the Boltzmann equation (we search for the model parameters that can reproduce the observed relic density of DM)

# Boltzmann equation (BE)

$$2E_i (\partial_t - H p \partial_p) f_i(p) = C[f_i]$$

Hubble rate  $\nearrow$   
 Distribution function  $\nearrow$   
 Collision term (interactions)  $\nwarrow$

- Integrating over the phase-space of the DM particle (and introducing comoving variables) one gets an equation for the **abundance**

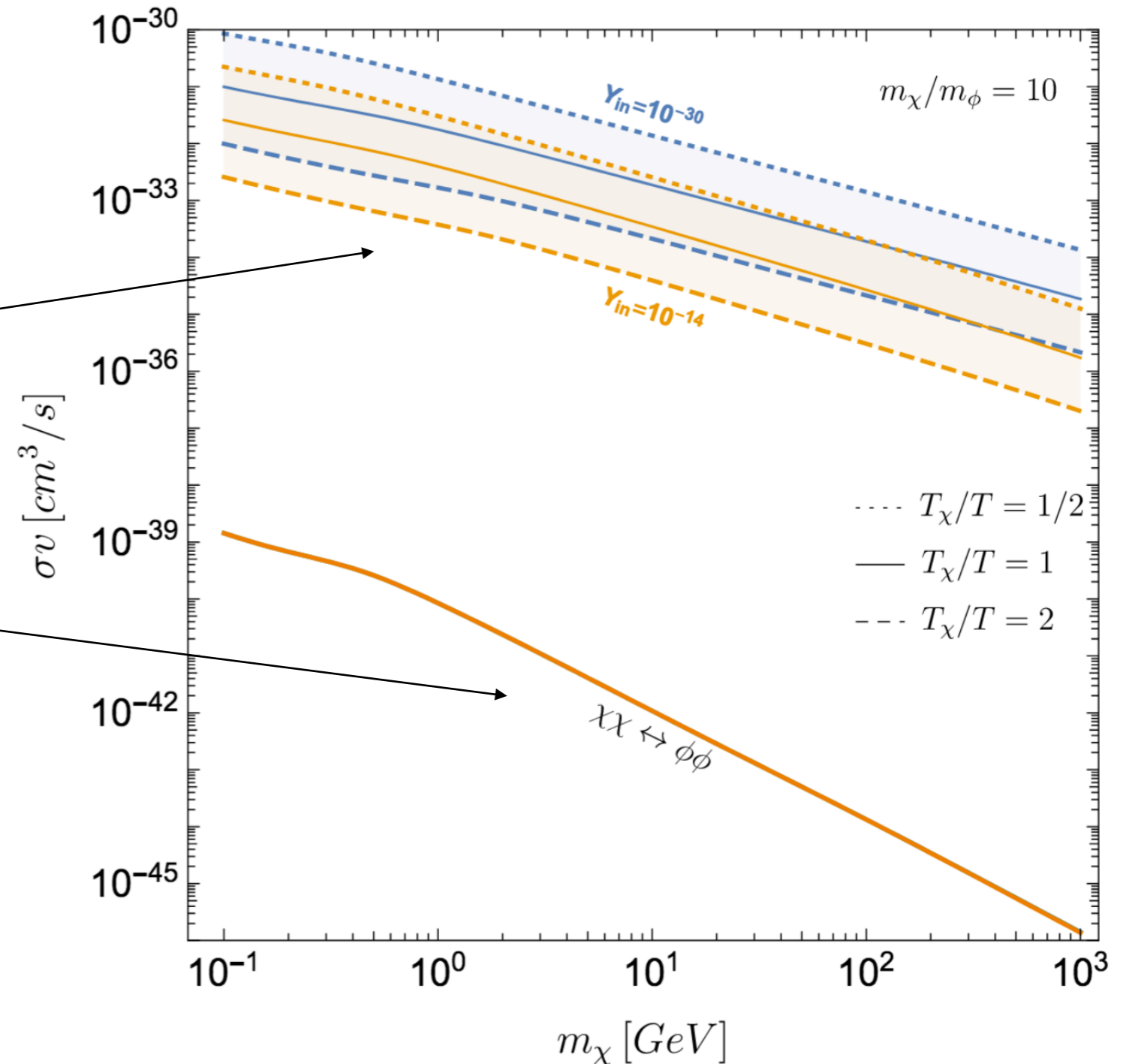
$$\frac{dY}{dx} = \frac{\langle \sigma v \rangle (T_\chi, T_\phi)}{xsH} Y$$

$$x = \frac{m}{T}$$

$$Y = \frac{n_{\text{DM}}}{s}$$

# Toy model

- Assume that the temperature of  $\chi$  is known
- Semi-production requires much **larger cross sections** than pair-production
- If  $\phi$  is out of equilibrium, even larger cross sections are possible



# Coupled system of BEs

- In reality we don't know the temperatures of  $\chi$  and  $\phi$  *a priori*, but they can be important!
- We assume that  $\phi$  and  $\chi$  are both self-thermalized (have an **equilibrium shape**) and solve for their density and temperature

$$\frac{Y'_i}{Y_i} = \frac{m_i}{x\tilde{H}} C_i^0,$$

$$\frac{y'_i}{y_i} = \frac{m_i}{x\tilde{H}} C_i^2 - \frac{Y'_i}{Y_i} + \frac{H}{x\tilde{H}} \frac{\langle p^4/E_i^3 \rangle}{3T_i}$$

$$C_i^0 \equiv \frac{g_i}{m_i n_i} \int \frac{d^3 p}{(2\pi)^3 E_i} C[f_i],$$

$$C_i^2 \equiv \frac{g_i}{3m_i n_i T_i} \int \frac{d^3 p}{(2\pi)^3 E_i} \frac{p^2}{E_i} C[f_i]$$

higher moment term

$$y_i \equiv m_i T_i / s^{2/3} \quad \text{Temperature parameter}$$



# Setting

- We now consider a more detailed example model, where  $\phi$  is a **scalar singlet** coupled to the Higgs doublet

Higgs portal interactions

$$\mathcal{L}_{\phi-SM} = A\phi H^\dagger H + \frac{\lambda_{h\phi}}{2}\phi^2 H^\dagger H - \mu_h^2 H^\dagger H + \frac{\lambda_h}{2}(H^\dagger H)^2$$

$$\mathcal{L}_{DS} = \frac{\mu_\phi^2}{2}\phi^2 + \frac{\mu_3^2}{3!}\phi^3 + \frac{\lambda_\phi}{4!}\phi^4 + \mu_\chi^2 \chi^* \chi + \frac{\lambda_\chi}{4}(\chi^* \chi)^2$$

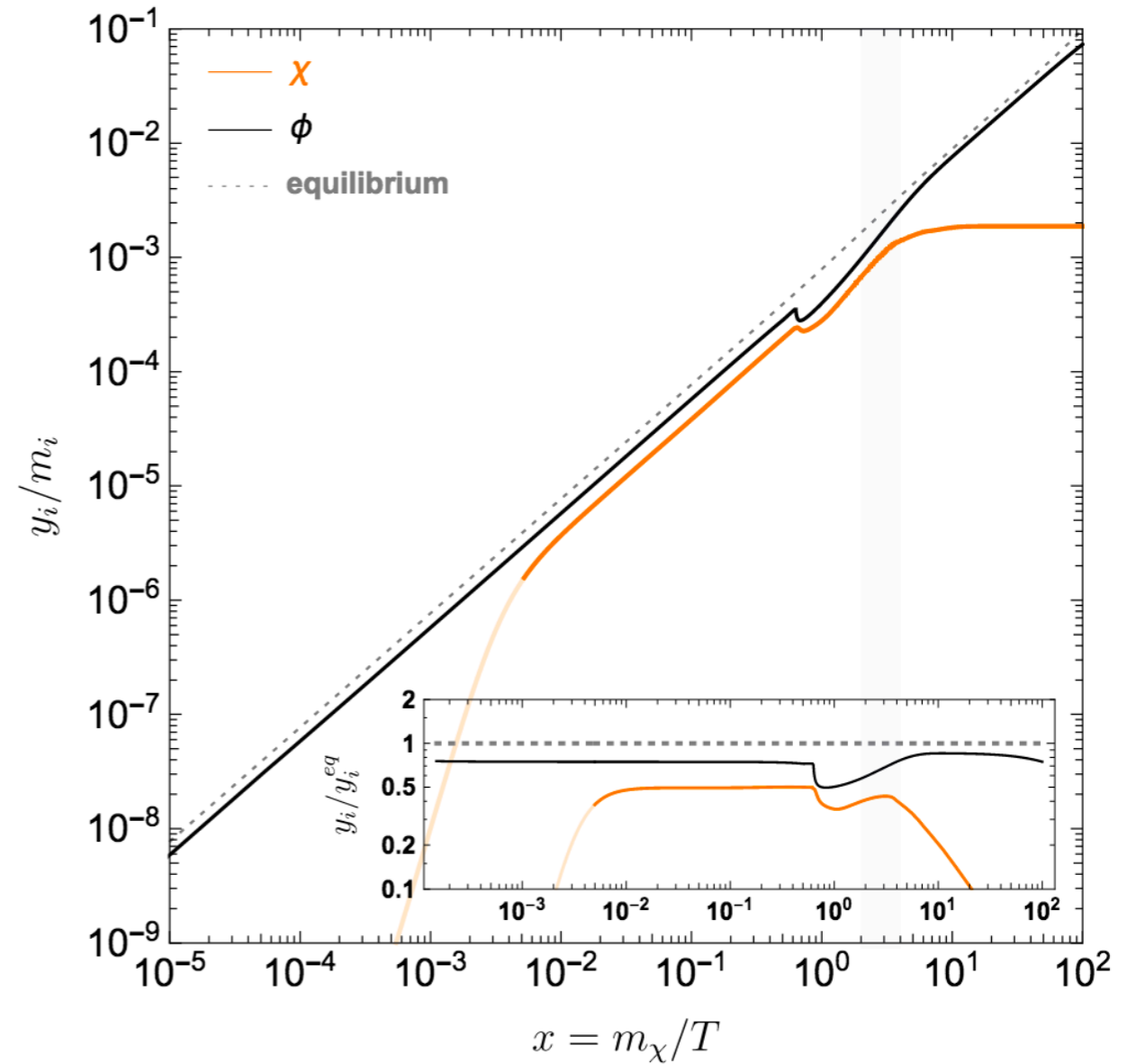
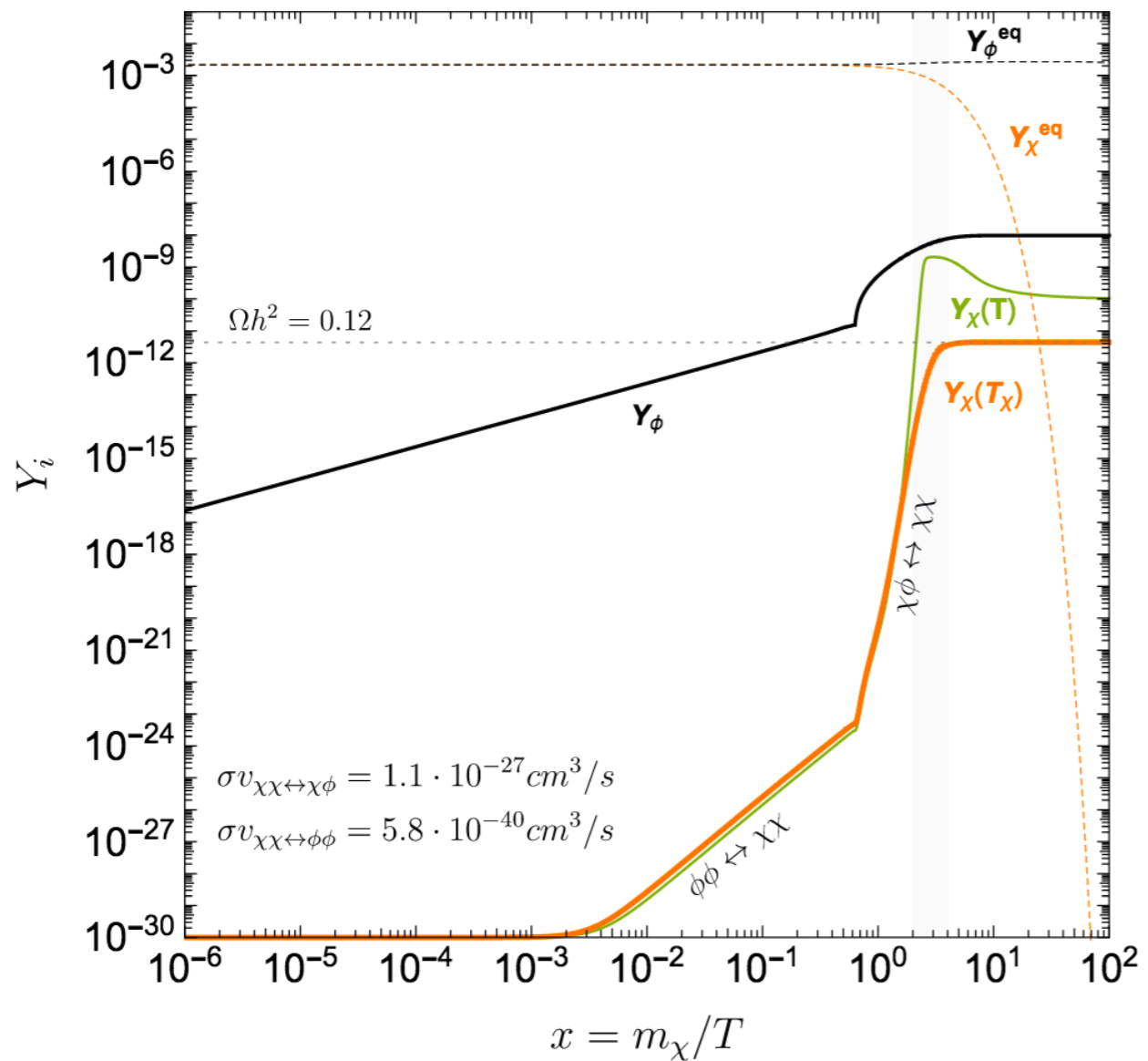
$$+ \frac{\lambda_1}{3!}\phi (\chi^3 + (\chi^*)^3) + \frac{\lambda_2}{2}\phi^2 (\chi^* \chi),$$

semi-production

pair-production

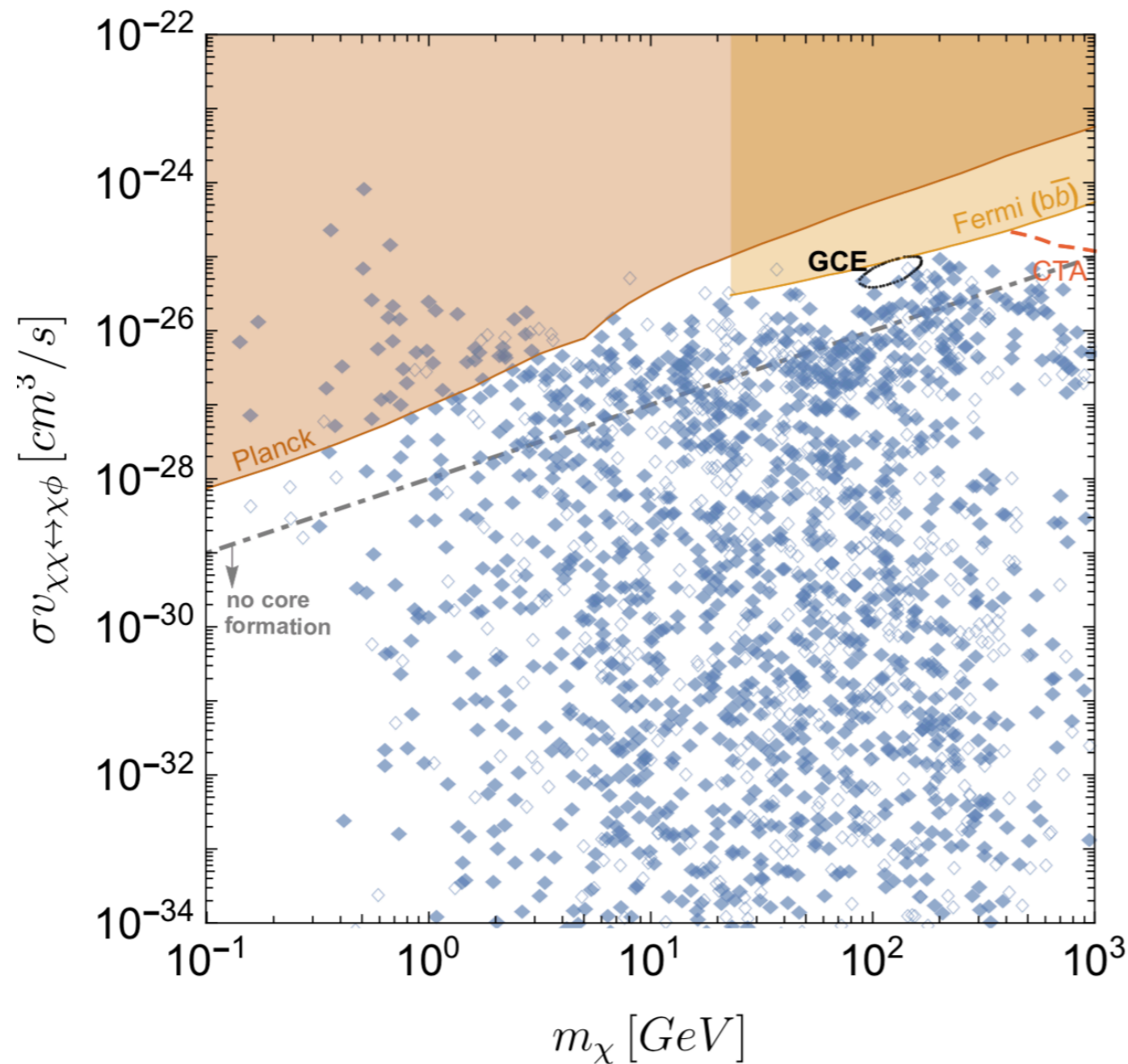
- $\phi$  gets a VEV, but  $\chi$  doesn't
- $m_\phi < 3m_\chi \rightarrow$  no decays

# Evolution of density and temperature



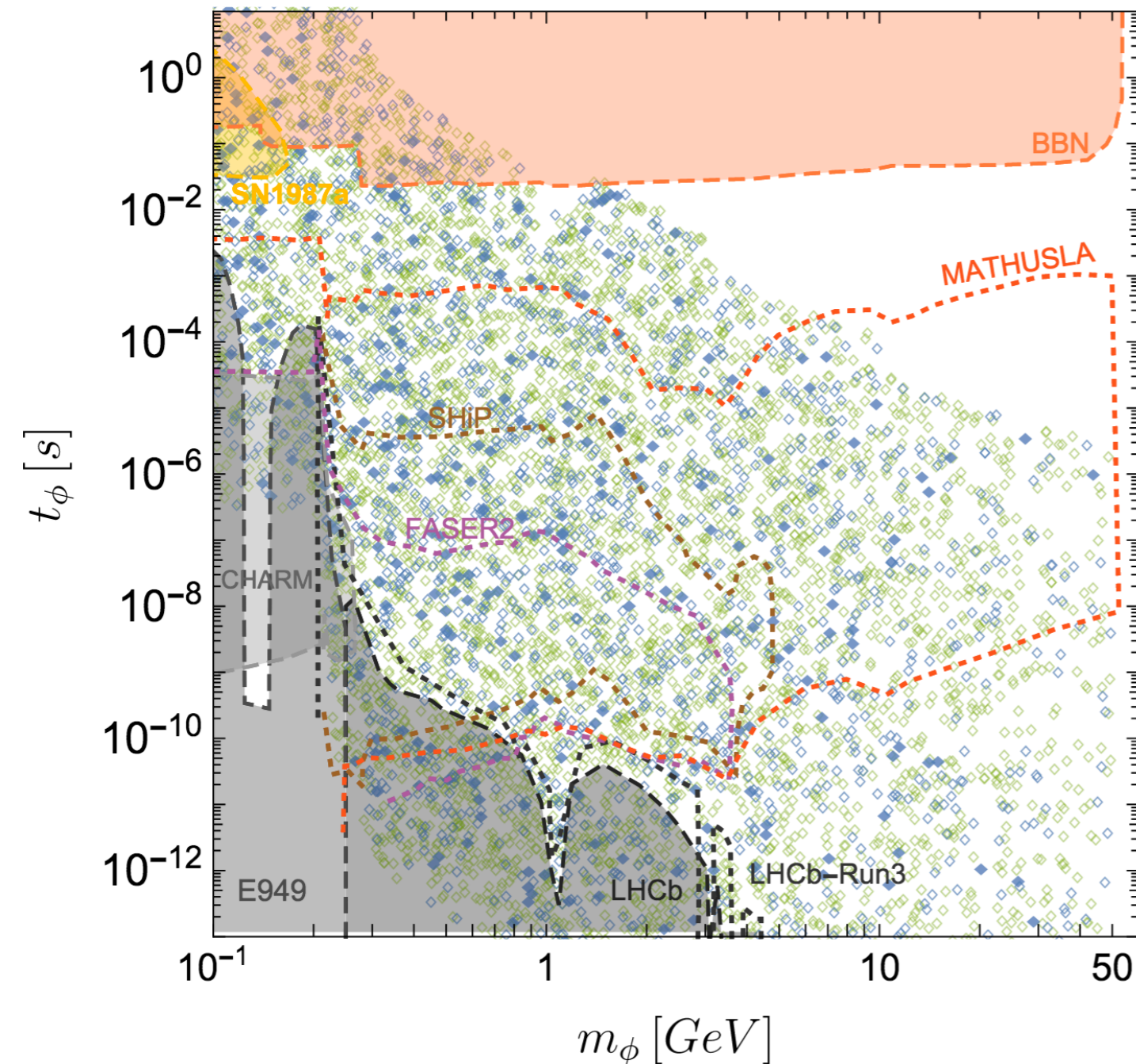
$$m_\chi = 100 \text{ GeV}, \mu_\phi = 1 \text{ GeV}, \lambda_1 = 1.1 \times 10^{-2}, \lambda_2 = 10^{-8}, \lambda_{h\phi} = 6 \times 10^{-11}$$

# Indirect detection constraints and predictions



- The results of the scan in the parameter space for the DM production dominated by the **semi-annihilation** processes.
- The **coloured** squares indicate the points, which are **within the reach of the future searches** for the mediator  $\phi$  and the empty ones are beyond these prospects.
- The points above the grey dot-dashed line can potentially **explain the core formation** in dSph [1803.09762]

# Long-lived particle searches



- The constraints on the properties of the mediator  $\phi$  and the prospects for its detection.
- The blue points correspond to the DM production dominated by the semi-annihilation, while the green ones – by the pair-annihilation.

# Conclusion

- We have studied the freeze-in mechanism based on the semi-annihilation process
- Semi-production freeze-in requires **larger cross sections** than the pair-production freeze-in
- This mechanism can be incorporated in various models and promises an interesting phenomenology that is within **the reach of near-future experiments:**
  - Indirect detection
  - Collider searches (FPF)

**Thank you!**