

Dark matter at DUNE near detector

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In collaboration with:

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and Andrés. Gómez¹

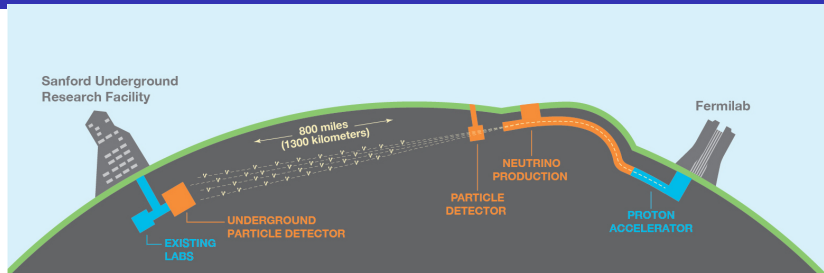
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Universidad EIA

Based on:

[arXiv:2109.11586](#)
[arXiv:22xx.xxxxx](#)



DUNE Near Detector

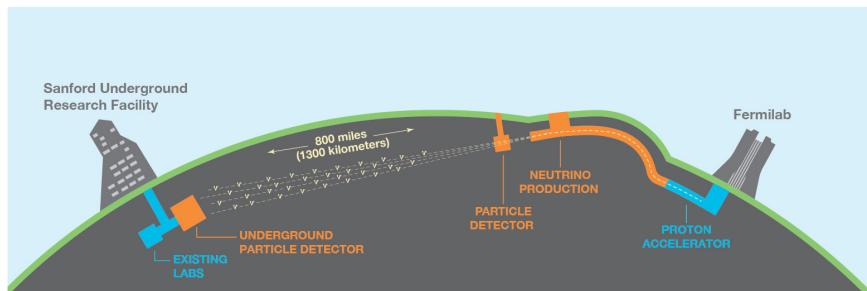


- Protons at 120 GeV collides against a fixed-target made of carbon, then a bunch of particles are created ($\pi^0, \eta, \eta', \omega, \pi^\pm, K^\pm, \dots$).

meson m	meson/POT	$\text{Br}(m^+ \rightarrow \mu^+ \nu_\mu)$	$\text{Br}(m^+ \rightarrow e^+ \nu_e)$
$\pi^+(\pi^-)$	4.3(4.0)	99.98 %	1.2×10^{-4} %
$K^+(K^-)$	0.39(0.27)	63.56 %	1.58×10^{-5} %

- Neutrinos are created after the decay of π^\pm, K^\pm, \dots
- Muon neutrinos are mostly created in the decays of π^\pm

Light dark matter at DUNE

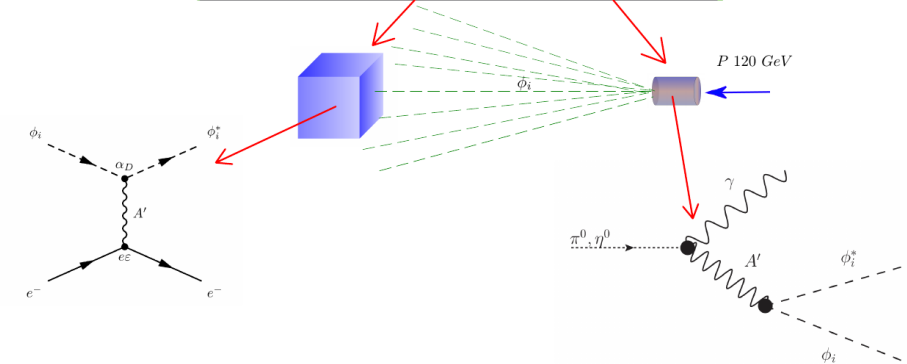
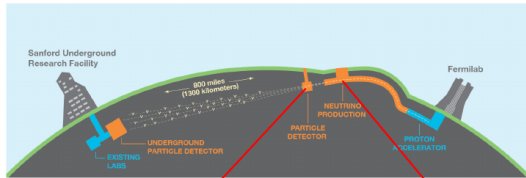


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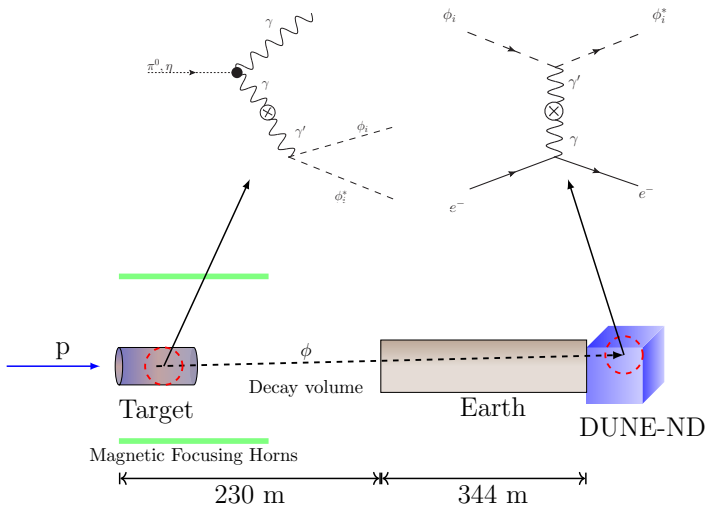
meson \mathbf{m}	meson/POT	$\text{Br}(\mathbf{m} \rightarrow \gamma\gamma)$
π^0	4.8	99.82 %
η	0.5	39.41 %

- LDM could be produced thanks to the mixing between $\gamma\text{-}\gamma'$

DUNE Near Detector



DUNE Near Detector



The Model

The SM is extended with:

- $G_{SM} \otimes U(1)_D \otimes Z_2 \otimes Z'_2$
- Two complex scalar singlets, $\phi_1 \sim (-1, +, -)$ and $\phi_2 \sim (-1, -, +)$
- One complex scalar singlet, $s \sim (2, -, -)$, which develops **VEV** and breaks $U(1)_D$

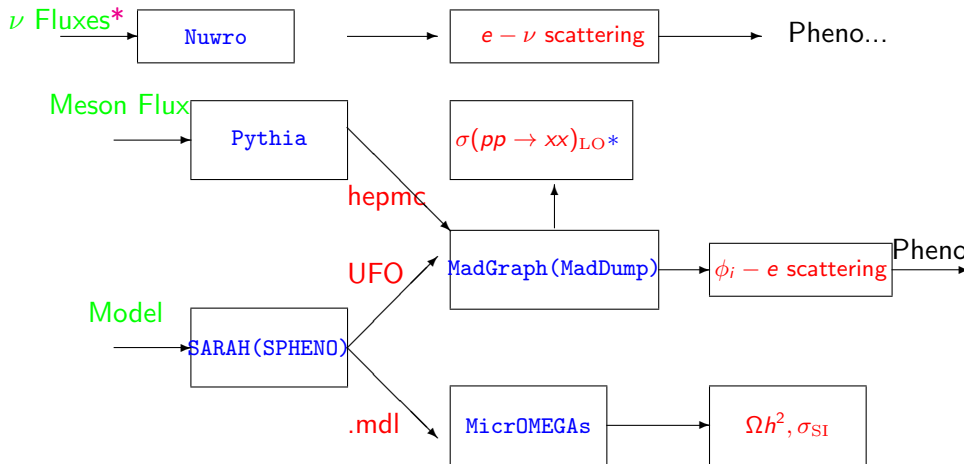
The Lagrangian reads:

$$\mathcal{L}_{\text{new}} \supset -\frac{\varepsilon}{2} F_{\mu\nu} F'^{\mu\nu} + \sum_{i=1}^2 |D_\mu \phi_i|^2 + V(H, s, \phi_1, \phi_2)$$

$$\mathcal{L}_\phi = \frac{ig_D}{2} \gamma'^\mu \sum_{i=1}^2 \left[\partial_\mu (\phi_i^\dagger) \phi_i - \phi_i^\dagger (\partial_\mu \phi_i) \right]$$

- After EWSB, $G_{SM} \otimes U(1)_D \otimes Z_2 \otimes Z'_2 \longrightarrow SU(3)_C \otimes U(1)_{\text{em}} \otimes Z_2 \otimes Z'_2$
the mixing of the fields (B_3^μ, A^μ, A'^μ) give rises to the boson Z , the photon γ and the dark photon $\gamma_D (\gamma')$.
- The dark photon mixes with the ordinary matter.
- ϕ_1 and ϕ_2 are stable and dark matter candidates.

HEP Tools implementation for Signal and Background



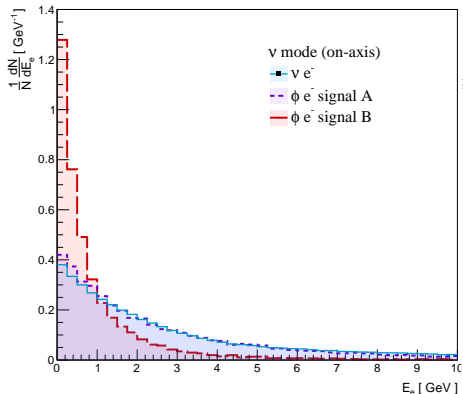
<https://home.fnal.gov/ljf26/DUNEFluxes/>

Signal and Background on-axis

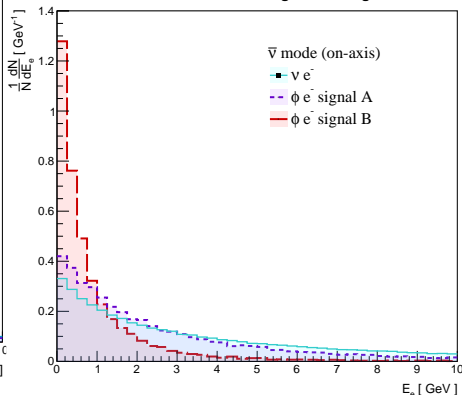
Benchmark Points

- **BPA** $\rightarrow m_{\gamma'} = 30$ MeV, $m_{\phi_1} = 20$ MeV, $m_{\phi_2} = 1.7$ GeV, $\alpha_D \varepsilon^4 = 10^{-13}$
- **BPB** $\rightarrow m_{\gamma'} = 90$ MeV, $m_{\phi_1} = 30$ MeV, $m_{\phi_2} = 1.7$ GeV, $\alpha_D \varepsilon^4 = 10^{-13}$

Signal and Bkg at DUNE ND



Signal and Bkg at DUNE ND



Benchmark Point

$$m_{\gamma'} = M_{A'} = 3m_{\phi_1}, m_{\phi_2} = 1.7 \text{ GeV}, \alpha_D = 0.1,$$

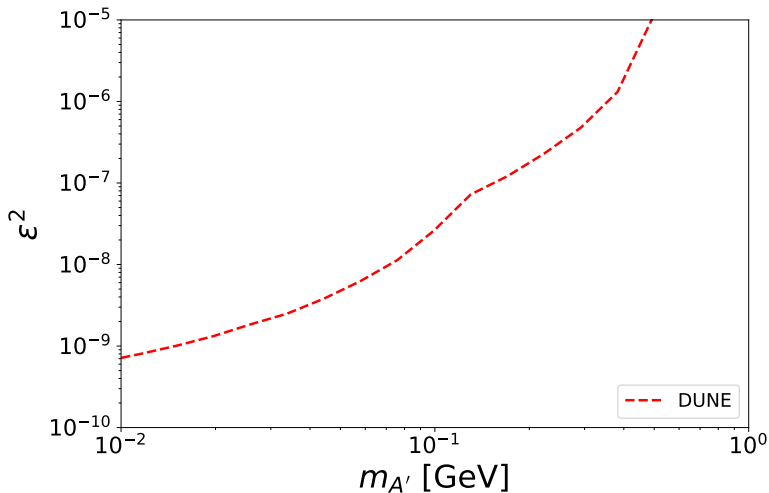


Figure 1: Expected sensitivity at 90% CL

Benchmark Point

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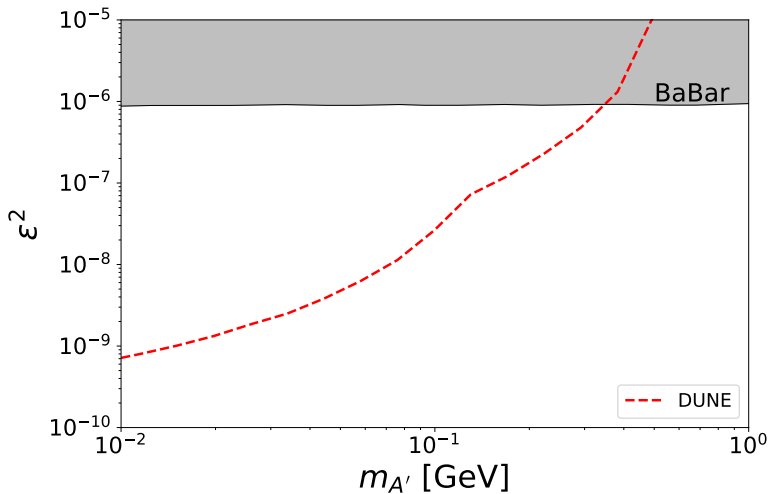


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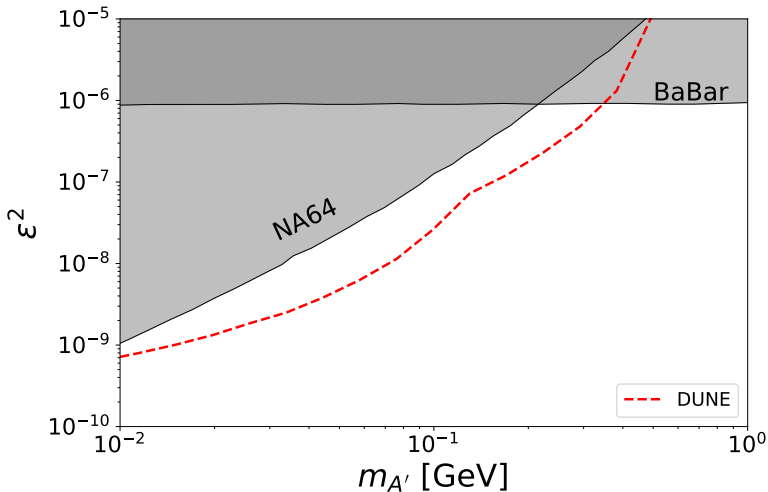


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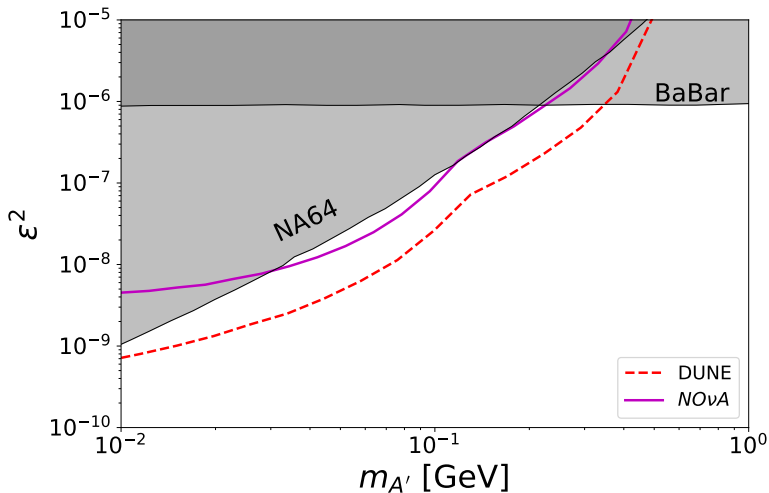


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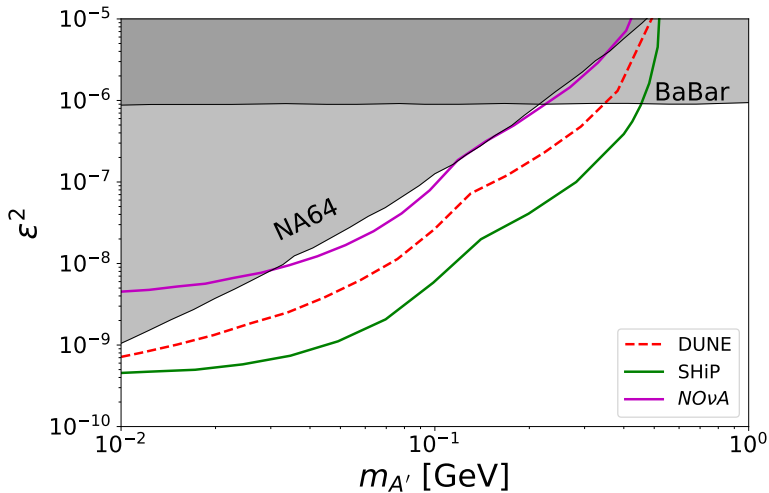


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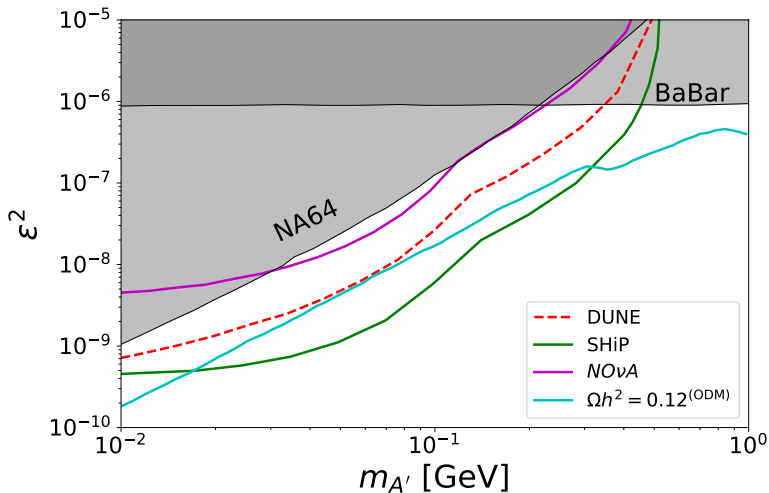


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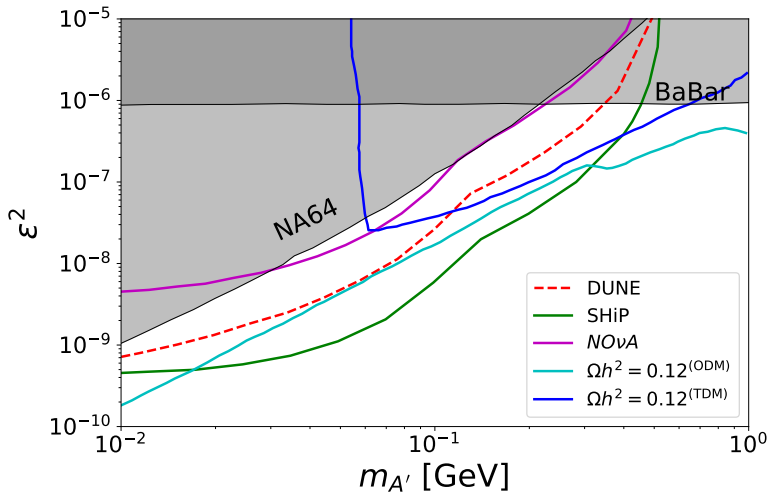


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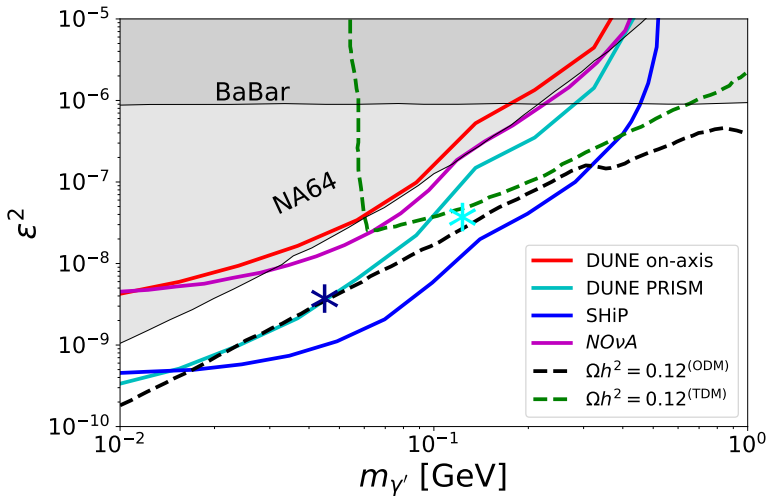


Figure 1: Expected sensitivity at 90% CL

How to distinguish Signal from Background?

- In the previous result no cuts were implemented at all, and the test-statistic used for the sensitivity was based in the whole number of events for signal and background energy histograms. The sensitivity at 90% was computing using

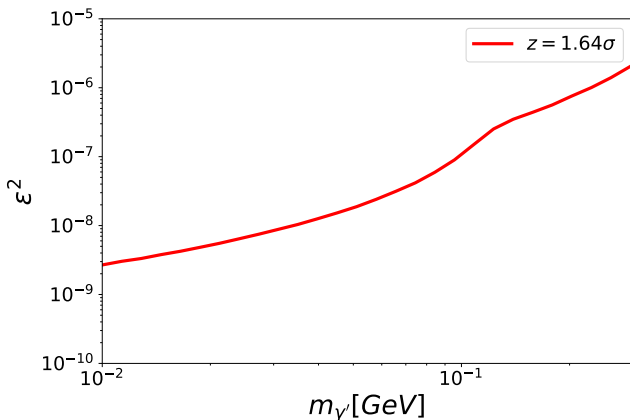
$$Z = \frac{N^\phi}{\sqrt{N^\nu + \sum_{i \in (\nu_e, \nu_\mu, \bar{\nu}_e, \bar{\nu}_\mu)} (k_i N_i^\nu)^2}} \geq 1.64 ,$$

- **Our new study** is based on the exploration of **ML techniques** to distinguish signal from background and improve the potential exclusion reach of DUNE.
- Machine Learning (ML) can be used as supervised learning to perform classification.

Basic idea and preliminary results: Sensitivity for DUNE-ND: One year running

Benchmark Points

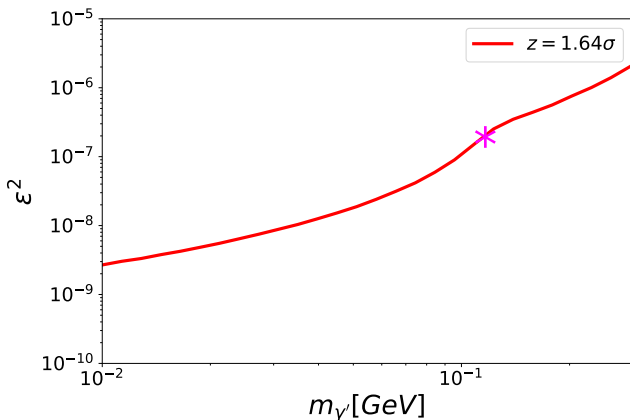
$$m_{\gamma'} = 0.11 \text{ GeV}, \varepsilon^2 = 1.5 \times 10^{-7} \rightarrow \varepsilon^2 = 4.5 \times 10^{-8}$$



Basic idea and preliminary results: Sensitivity for DUNE-ND: One year running

Benchmark Points

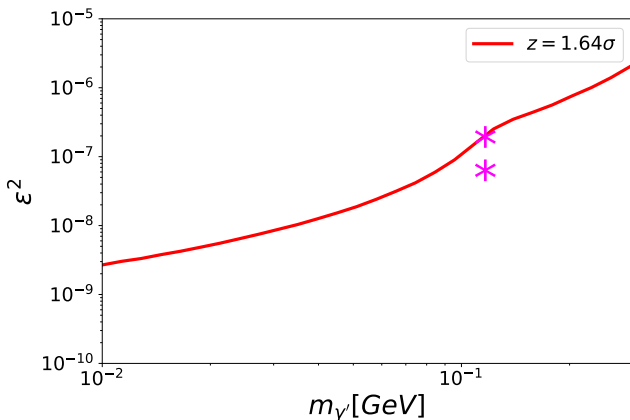
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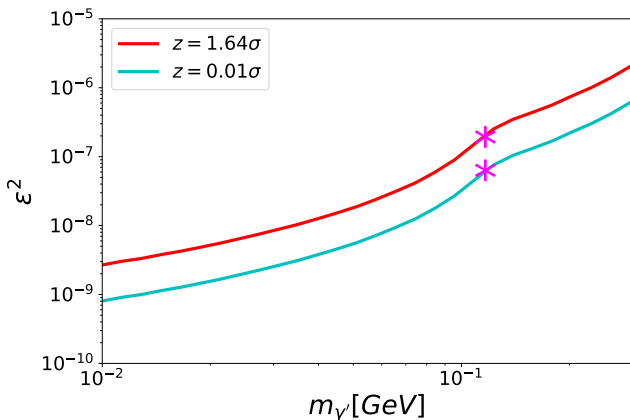
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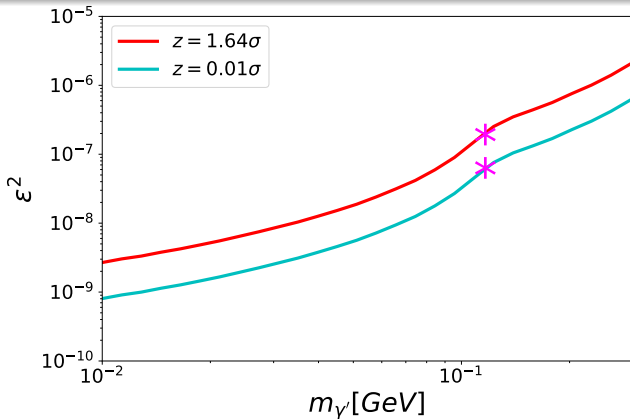
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For the new point (in cyan line) the expected sensitivity using ML (xgboost) is of 1.68σ

Log-Likelihood ratio for two hypothesis

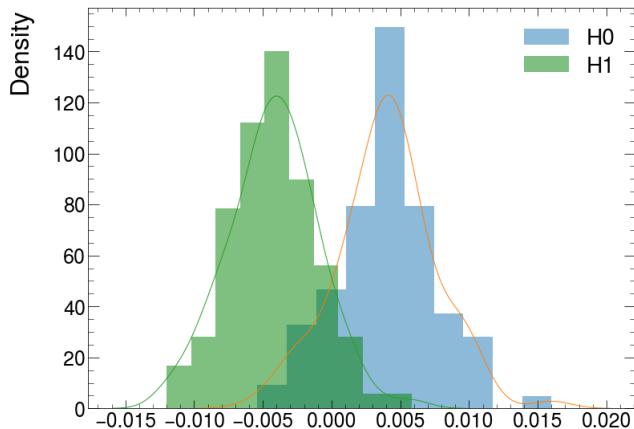


Figure 2: LLR for two hypothesis

Benchmark Points

$M_{\gamma'} = 0.11 \text{ GeV}, \varepsilon^2 = 1.5 \times 10^{-7}$ [Left] $\varepsilon^2 = 4.5 \times 10^{-8}$ [Right]

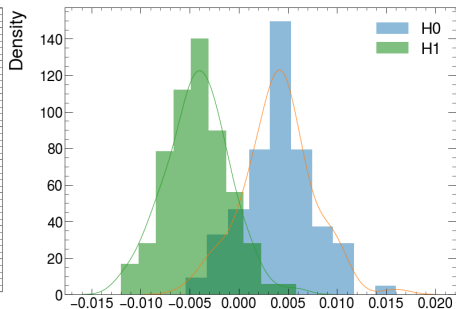
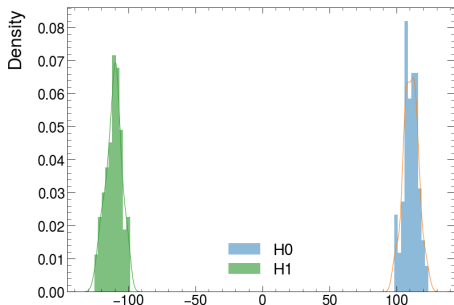


Figure 3: LLR Benchmark point 1 (up) Figure 4: LLR Benchmark point (down)

Conclusions

- Beam Dump experiments has the potential to explore a big portion of the parameter space of the model.
- The correct relic density can be explained in regions that Beam Dump experiments will soon probe.
- Off-axis analysis implemented.
- ML techniques has the potential to increase the exploratory reach of DUNE
- A full analysis is under way

Thanks for your attention