



## Taming hadronisation corrections for collider observables

#### Silvia Ferrario Ravasio

University of Oxford

#### A life in phenomenology: a conference in honour of Paolo Nason, University of Milano – Bicocca, 15<sup>th</sup> September 2022

"All-orders behaviour and renormalons in top-mass observables", SFR, Nason, Oleari "Infrared renormalons in kinematic distributions for hadron collider processes," SFR, Limatola, Nason "On linear power corrections in certain collider observables", Caola, SFR, Limatola, Melnikov, Nason "Linear power corrections to  $e^+e^-$  shape variables in the 3-jet region", Caola, SFR, Limatola, Melnikov, Nason, Ozcelik

## A typical collider event



Ingredients to describe a lepton collision

- Hard process ( $Q\sim 100~{\rm GeV}$ ): fixed order expansion in the strong coupling  $\alpha_s(Q)$
- multiple soft and/or collinear emissions, with  $Q > k_{\perp} > \Lambda$ , with  $\Lambda \sim 1$  GeV. Tools: analytic resummation (more accurate) or parton shower algorithms (more flexible)
- <u>Hadronization corrections</u>: phenomenological models (Lund or cluster) from MC event generators, or analytic models

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## Hadronization models for shape observables



• Event shapes to perform precise measurements of  $\alpha_s$ .

• Non-perturbative linear power corrections  $\propto 1/Q$  must be provided in order to fit the data!

 $etty (S_a \rightarrow 0)$ 

Isotropic  $(S_0 \rightarrow 1)$ 

 $\bullet$  Analytic models: shift the peturbative prediction by a constant amount  $\propto 1/Q$ 

$$\Sigma(O) \to \Sigma(O - \mathcal{N} \bigtriangleup O)$$

universal Independent of  $O(\Phi)$ 

We need to control linear NP corrections if we want percent or permille precision at  $Q \approx 100$  GeV!



#### Linear power corrections in collider observables

• We need to know when to expect linear power corrections!



If present in the **trans**verse momentum of gauge bosons, we will never match the experimental precision!

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## Estimating power corrections

 $\bullet$  Several sources of non-perturbative corrections, e.g. the Landau pole  $\Lambda$  in the QCD coupling constant



$$\alpha_{\rm s}(\boldsymbol{Q}) = \frac{1}{2b_0 \log\left(\frac{\boldsymbol{Q}}{\Lambda}\right)}; \quad b_0 = \frac{11 \mathcal{C}_{\rm A}}{12\pi} - \frac{n_l \mathcal{T}_{\rm R}}{3\pi} > 0$$

which leads to an **instrinsic ambiguity** when integrating over the soft momenta.

$$\int_0^Q \mathrm{d}k \, k^{p-1} \alpha_{\mathrm{s}}(\boldsymbol{k}) = \left[ Q^p \times \frac{p}{2 \, b_0} \sum_{n=0}^\infty \left( \frac{2 \, b_0}{p} \, \alpha_{\mathrm{s}}(Q) \right)^{n+1} \, \boldsymbol{n}! \right]$$

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• The ambiguity has to cancel with contributions arising from physics beyond perturbation theory: estimate of non-perturbative effects. The smallest term in the series is

$$Q^p \sqrt{\frac{\alpha_s(Q)p\pi}{b_0}} e^{-\frac{p}{2b_0\alpha_s}} = \sqrt{\frac{\alpha_s(Q)p\pi}{b_0}} \Lambda^p$$

## Large- $n_f$ limit

• Ambiguity related to the appearance of the Landau pole can be studied in the large number of flavour  $n_f$  limit, which allows to perform all-orders computations exactly.

$$\begin{array}{c} \overbrace{000} & \overbrace{000} & = & \overbrace{0000} & + & \overbrace{000} & \overbrace{1} \\ \hline{\frac{-ig^{\mu\nu}}{k^2 + i\eta}} & \rightarrow & \frac{-ig^{\mu\nu}}{k^2 + i\eta} \times \frac{1}{1 + \Pi(k^2 + i\eta, \mu^2, \epsilon) - \Pi_{ct}} \\ \Pi(k^2 + i\eta, \mu^2) - \Pi_{ct} & = \alpha_s(\mu) \left( -\frac{n_f T_R}{3\pi} \right) \left[ \log \left( \frac{|k^2|}{\mu^2} \right) - i\pi\theta(k^2) - \frac{5}{3} \right] + \mathcal{O}(\epsilon) \end{array}$$

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• naive non-abelianization at the end of the computation (large  $b_0$ )

$$\Pi(k^2 + i\eta, \mu^2) - \Pi_{\rm ct} \rightarrow \alpha_{\rm s}(\mu) \underbrace{\left(\frac{11C_{\rm A}}{12\pi} - \frac{n_l T_{\rm R}}{3\pi}\right)}_{b_0} \left[\log\left(\frac{|k^2|}{\mu^2}\right) - i\pi\theta(k^2) - C\right]$$

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### Large- $b_0$ approximation for realistic collider processes

$$O = \int d\Phi \frac{d\sigma(\Phi)}{d\Phi} O(\Phi) = O_{\rm LO} - \frac{1}{\pi b_0} \int_0^\infty d\lambda \frac{d}{d\lambda} \left[ \frac{T(\lambda)}{\alpha_{\rm S}(\mu)} \right] \arctan\left[\pi b_0 \alpha_{\rm S}(\lambda e^{-C/2})\right]$$

•  $\lambda$  can be thought as a gluon mass / virtuality

• 
$$T(\lambda) = \int d\Phi_b V_\lambda(\Phi_b) O(\Phi_b) + \frac{\lambda^2}{\pi b_0} \int d\Phi_{q\bar{q}} R_{q\bar{q}}(\Phi_{q\bar{q}}) O(\Phi_{q\bar{q}}) \delta(m_{q\bar{q}}^2 - \lambda^2)$$
  
•  $T(\lambda) \xrightarrow{\lambda \to 0} O_{\text{NLO}}$ 

[S.F.R, Nason, Oleari '19]

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#### Large- $b_0$ approximation for realistic collider processes



## Assessing power corrections from first principles

• Linear power corrections in transverse momentum of the Z boson [Salam and Slade, JHEP 11 (2021), 220] can limit the ultimate theoretical precision achievable

$$\frac{\Lambda}{p_{\perp}} = \frac{1 {\rm GeV}}{30 {\rm GeV}} = {\rm 3\%}$$

Current theoretical err  $\approx 3\%$ Experimental error  $\approx 0.3\%$ 

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Within our framework, we found no evidence of linear **non-perturbative** power corrections [SFR, Limatola, Nason, JHEP **06** (2021), 018]

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When the ultimate theoretical precision is not spoilt by linear power corrections?

When the observable is **inclusive** with respect to QCD radiation

[Caola, <u>SFR</u>, Limatola, Melnikov, Nason, JHEP **01** (2022), 093]



• Linear power corrections are present in event shapes (thrust, C-parameter...).

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- Linear power corrections are present in event shapes (thrust, C-parameter...).
- Strong coupling constant determinations lead
  - $\alpha_s = \mathbf{0.1179(10)}$  world average
  - $\alpha_s=\!\mathbf{0.1135(10)}$  from Thrust

[Abbate et al., Phys. Rev. D 86 (2012), 094002]



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• Linear power corrections for V=0 known for a long time ...

[Dokshitzer, Webber, Phys. Lett. B 404 (1997), 321-327], [Dokshitzer et al., JHEP 05 (1998), 003]

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- ullet ... and assumed to be valid also for V>0
- But for the C-parameter [Luisoni, Monni, Salam, Eur. Phys. J. C 81 (2021) no.2, 158]

 $\frac{\rm Linear \ power \ correction \ at \ }{C=0.75} \approx 0.48$  Linear power correction at C=0



## Revisiting NP corrections for event shapes

photon

• Since we cannot deal with gluons, to study NP corrections away from the two-jet limit in the large- $n_f$  limit we consider the process  $\mathbf{Z} \to q\bar{\mathbf{q}}\gamma$ .

• For many event shapes such as thrust and C-parameter, collinear contributions are exponentially suppressed, so the leading soft approximation is sufficient to compute  $T(\lambda)$ 

$$O_{n+1}(p_1,\ldots,p_n,\boldsymbol{k}) \approx O_n(\tilde{p}_1,\ldots,\tilde{p}_n) + \underbrace{\frac{\boldsymbol{k}_{\perp}}{Q} f_n(\varphi,\eta,\{\tilde{p}_i\})}_{\Delta O}, \quad \lim_{\eta \to \pm \infty} f_n(\varphi,\eta,\{\tilde{p}_i\}) \propto \exp^{-|\eta|} = 0$$

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 ${\ }$  With a smooth mapping  $\Phi_{n+1}\to \tilde{\Phi}_n$  the calculation largely simplifies

$$\begin{aligned} k &= z_1 \tilde{p}_1 + z_2 \tilde{p}_2 + k_\perp \\ p_{1,2} &\approx (1 - z_{1,2}) \tilde{p}_{1,2} \end{aligned} \Rightarrow \quad \frac{d\Sigma(O < o)}{d\lambda} = \int d\Phi \, \delta(O(\Phi) - o) \frac{d\sigma}{d\Phi} \left[ \mathcal{M} \frac{2C_F \alpha_s}{\pi} \int \frac{dk_\perp}{k_\perp} dy \frac{d\varphi}{2\pi} \Delta O(k_\perp, \varphi, \eta; \Phi) \, \delta(k_\perp - \lambda) \right] \end{aligned}$$

[Caola, S.F.R., Limatola, Melnikov, Nason, '21, +Ozcelik, '22]

where  $\mathcal{M}$  is the Milan factor [Dokshitzer, Lucenti, Marchesini, Salam, '98] computed in the 2-jet limit!

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• The Milan factor  $\mathcal{M}$  takes into account the difference between the emission of a **soft gluon** of  $k_{\perp} = \lambda$ , and the emission of an off-shell gluon decaying in a **pair of quarks** with  $m_{q\bar{q}} = \lambda$ .

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• We assume the same formulae for more complex final states, with  $C_F$  replaced by the proper color factor for each dipole

#### Results

Non-negligible kinematic dependence!

[Caola, S.F.R., Limatola, Melnikov, Nason, Ozcelik, '22]





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#### Results

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• Why people did not obtain these results before?

• Do we believe in these results?

#### Why the results are new



- Different kinematic mapping prescription to build the phase space for an additional soft parton, lead to different results away from the two jet limit.
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#### Why the results are new



- Different kinematic mapping prescription to build the phase space for an additional soft parton, lead to different results away from the two jet limit.
- We identified which ones are correct: those who are analytic in the soft limit.
- Our results coincide with the results from [Luisoni, Monni, Salam '20] obtained using the smooth PS mappings (Catani-Seymour/Antenna/PanScales)
- Is it a fundamental constraint we want for a **Parton Shower** generator?

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## Preliminary fits of $\alpha_s$



Taming hadronisation corrections for collider observables

### Not yet the end of the story



- Power corrections very quickly drops away from the strict two jet limit;
- Our estimate on non-power corrections assumes the **fixed-order calculation** as perturbative baseline
- For  $\tau \leq 0.05$  the cross section is overly dominated by singular terms / we need **resummation** also the evaluate non-perturbative corrections.



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#### How do we reconcile the two pictures?

The transition between the two-jet and three jet behaviour should be more smooth, as it happens when a fixed order calculation is combined with resummation.



### Conclusion and outlooks

- It is of utmost importance to tame hadronisation corrections
  - When do we expect linear power corrections?
  - e How do we calculate them?

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- With this framework we could investigate any infrared safe observable (for processes without gluons at LO ...)
- We showed inclusive observables are free from linear power corrections
- More insights on the calculation of hadronisation corrections for event shapes
- ... although some freedom is taken during the "non-abelianization" phase, and we do not have a definitive recipe!

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• Due to the confinement, asymptotic states are ill defined in QCD: the pole mass has a  $\mathcal{O}(\Lambda)$  ambiguity.

$$m_p = \overline{m}(\overline{m}) \sum_{i=0} c_i \alpha^i \text{ till } \mathcal{O}(\alpha_s^4) \quad \text{[Marquard, Smirnov}^2, \text{ Steinhauser '15]}$$
$$m_r = 1.270 \pm 0.212 \pm 0.205 \pm 0.289 \pm 0.529 \pm 0.68V$$

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$$m_b = 4.180 + 0.398 + 0.198 + 0.144 + 0.135 + \dots \text{ GeV}$$

$$m_t = 163.643 + 7.557 + 1.617 + 0.501 + 0.197 + \dots \text{ GeV}$$

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- Calculation in the large  $b_0$  limit:

$$\begin{split} m_p &= \overline{m}(\overline{m}) \sum_{i=0} c_i \alpha^i \text{ till } \mathcal{O}(\alpha_{\rm S}^4) \quad \text{[Marquard, Smirnov}^2, \text{ Steinhauser '15]} \\ m_c &= 1.270 + 0.212 + 0.205 + 0.289 + 0.529 + \dots \text{ GeV} \\ m_b &= 4.180 + 0.398 + 0.198 + 0.144 + 0.135 + \dots \text{ GeV} \\ m_t &= 163.643 + 7.557 + 1.617 + 0.501 + 0.197 + \dots \text{ GeV} \\ \text{[Ball, Beneke, Braun '95]} \end{split}$$

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 $\Delta m_t = 7.557 + 2.345 + 0.584 + 0.241 + 0.127 + 0.085 + 0.067 + 0.063 + 0.067 + \dots \text{ GeV}$ 

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Asymptotic formula known

[Beneke, Braun '94]

$$c_{n+1} \to N \,\overline{m} \,(2b_0)^n \, \frac{\Gamma \,(1+n+b)}{\Gamma(1+b)} \left(1 + \sum_{k=1}^\infty \frac{s_k}{n}\right) \quad \text{with } b = \frac{b_1}{2b_0^2}, s_i = s_i(b_0, b_1, \ldots)$$

• Fitting N from the exact relation

[Beneke, Marquard, Nason, Steinhauser, '16]

 $\Delta m_t = \underbrace{7.577 + 1.617 + 0.501 + 0.197}_{\text{exact}} + 0.112 + 0.079 + 0.066 + \underbrace{0.064}_{\text{exact}} + 0.071 + \dots \text{ GeV}$ 



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### Single-top production and decay: reconstructed-top mass

Modulo finite top width effects, the physical renormalon which affects the observable definition, largely cancels when using the pole mass

		$\left  \left\langle M \right\rangle = \sum_{i=0}^{\infty} d^{i}$	$c_i lpha_{ m s}^i;$ [MeV]
i	$m_{ m p} - \overline{m}(\mu)$	pole, $R=1.5$	$\overline{\mathrm{MS}},\ R = 1.5$
3	+430	+ 14(1)	+438(1)
4	+171	-6(1)	+163(1)
5	+89	-10(1)	+79(1)
6	+60	-11(1)	+49(1)
7	+47	-11(1)	+35(1)
8	+44	-12(1)	+31(1)
9	+46	-15(1)	+31(1)
10	+55	-19(1)	+36(1)

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## Single-top production and decay: leptonic observables

A	finite $\Gamma_t$	chang	jes significantly
T	$\sigma'(\lambda)$ for	$\langle {f E}_{f W}  angle$	

$\Gamma_t$	slope ( <mark>pole</mark> )	slope ( $\overline{\mathrm{MS}}$ )
NWA	0.53(2)	0.46(2)
10 GeV	0.058(8)	0.004(8)
20 GeV	0.061(2)	0.001(2)

 $E_W$  (in the lab frame) in the  $\overline{\rm MS}$  scheme has a linear renormalon only in NWA. (Top frame:  $\lambda^2$  because of OPE)

$c_i lpha_{ m S}^i$ [MeV]	pole	$\overline{\mathrm{MS}}$
i = 4	-94(6)	-78(6)
i = 5	-44(5)	-35(5)
i = 6	-22(4)	-17(4)
i=7	-13(4)	-8(4)
i = 8	-9(4)	-4(4)
i = 9	-7(4)	-2(4)
i = 10	-6(5)	-1(5)
i = 11	-7(6)	0(6)
i = 12	-9(9)	1(9)

 $\Gamma_t = 1.33 \text{ GeV}$  To see the linear renormalon screening provided by  $\Gamma_t$  in the  $\overline{\text{MS}}$  scheme, you need to be sensitive to  $\Gamma_t = m_t e^{1-i}$ , *i.e.*  $i \approx 6$ .

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cic WWWAR	NING!		
Leptonic correcti scheme!	<b>observa</b> l on even w	<b>bles can l</b> rith a well-	have a linear power -defined renormalisation ing 1e
$\frac{i}{i} = 8$ $i = 9$	-9(4) -7(4)	-4(4) -2(4)	MS scheme, you need to be sensitive
i = 10 i = 11 i = 12	$ \begin{array}{r} -6(5) \\ -7(6) \\ -9(9) \end{array} $	-1(5) 0(6) 1(9)	to $\Gamma_t = m_t e^{1-i}$ , i.e. $i \approx 6$ .