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Advancing Globular Cluster Constraints on the Axion-Photon Coupling

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For decades stellar evolution has been a rich source of constraints on physics beyond the Standard Model. In this talk I will discuss our recent improvement of the upper bound on the axion-photon coupling from stellar evolution, which has been derived using the R_2 parameter, the ratio of stellar populations on the Asymptotic Giant Branch to Horizontal Branch in Globular Clusters. I will compare observed limits on R_2 with data from simulations using the stellar evolution code MESA which include the effects of axion production. The benefits of considering R_2 over the traditionally employed R -parameter will be discussed, with particular attention given to quantifying in detail the effects of uncertainties on these parameters due to the modelling of convective core boundaries. Using a semiconvective mixing scheme we constrain the axion-photon coupling to be $g_{a\gamma\gamma} < 0.47 \times 10^{-10} \text{ GeV}^{-1}$. This rules out new regions of QCD axion and axion-like particle parameter space. Complementary evidence from asteroseismology suggests that this could improve to as much as $g_{a\gamma\gamma} < 0.34 \times 10^{-10} \text{ GeV}^{-1}$ as the uncertainties surrounding mixing across convective boundaries are better understood.

Author: HISKENS, Frederick (The University of Melbourne)

Co-authors: Prof. DOLAN, Matthew (University of Melbourne); Prof. VOLKAS, Raymond (The University of Melbourne)

Presenter: HISKENS, Frederick (The University of Melbourne)

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