

Dark Matter Effective Field Theory

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Outline

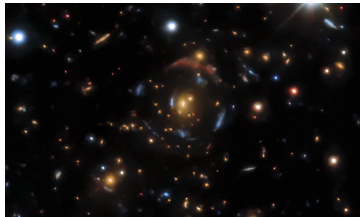
- 1 Motivation
- 2 SM effective field theories
- 3 DM effective field theories
- 4 Phenomenological example
- 5 Summary

based on: [2202.06968](#), in collaboration with Wolfgang Altmannshofer, Elizabeth Jenkins and Aneesh Manohar

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Dark matter



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SM Effective Theory (SMEFT)

Symmetry

$$SU(3)_C \times SU(2)_L \times U(1)_Y$$

Fields

$$u, d, c, s, b, t, \ell, \nu_\ell, g, W, Z, H$$

Poincaré invariance

Dim 6 operators

$$\mathcal{L}_{\text{SMEFT}} = \mathcal{L}^{(4)} + \frac{1}{\Lambda} \mathcal{C}^{(5)} \mathcal{O}^{(5)} + \frac{1}{\Lambda^2} \sum_i \mathcal{C}_i^{(6)} \mathcal{O}_i^{(6)} + \mathcal{O}\left(\frac{1}{\Lambda^3}\right)$$

Weak effective theory (WET)

Symmetry

$$SU(3)_C \times U(1)_{em}$$

Fields

$$u, d, c, s, b, \ell, \nu_\ell, g, \gamma$$

Poincaré invariance

Dim 6 operators

Assumptions

SMEFT/WET

No new light degrees of freedom

Fields

SM content

New Physics

Only at higher scales

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New particles

DM fields	ϕ	χ	X_μ
Spin	0	1/2	1
Number	n_ϕ	n_χ	n_X

Table: DM fields: ϕ is a real scalar field, χ is a *right-handed* fermion, and X_μ is a vector field. The DM fields are assumed to be singlets under the SM gauge group, so all DM fields are electrically neutral.

Dark EFT

Dark SMEFT (DSMEFT) & Dark WET (DWET)

SMEFT/WET + singlet scalar ϕ , right-handed fermion χ , vector X

Operators

DM: 26

DSMEFT: 77

DWET: 116

Matching

DSMEFT \rightarrow DWET

Example: Operators

DSMEFT: dimension 6

$\Delta B = \Delta L = 0$

($d_{\text{SM}}, d_{\text{DM}}$)	Name	Operator	Number
(2,4)	\mathcal{Q}_{BX^2}	$B_\mu^\nu X_{a\nu}^\alpha X_{b\alpha}^\mu$	$\binom{n_X}{2}$
	$\mathcal{Q}_{\tilde{B}X^2}$	$\tilde{B}_\mu^\nu X_{a\nu}^\alpha X_{b\alpha}^\mu$	$\binom{n_X}{2}$
	$\mathcal{Q}_{BX\phi^2}$	$B_{\mu\nu} X_a^{\mu\nu} \phi_b \phi_c$	$\binom{n_\phi+1}{2} n_X$
	$\mathcal{Q}_{\tilde{B}X\phi^2}$	$\tilde{B}_{\mu\nu} X_a^{\mu\nu} \phi_b \phi_c$	$\binom{n_\phi+1}{2} n_X$
	\mathcal{Q}_{HX}	$(H^\dagger H) X_{a\mu\nu} X_b^{\mu\nu}$	$\binom{n_X+1}{2}$
	$\mathcal{Q}_{H\tilde{X}}$	$(H^\dagger H) \tilde{X}_{a\mu\nu} X_b^{\mu\nu}$	$\binom{n_X+1}{2}$
	$\mathcal{Q}_{H^2\phi^4}$	$(H^\dagger H) \phi_a \phi_b \phi_c \phi_d$	$\binom{n_\phi+3}{4}$
	$\mathcal{Q}_{H\chi\phi}$	$(H^\dagger H) (\chi_a^T C \chi_b) \phi_c + \text{h.c.}$	$n_\phi \binom{n_\chi+1}{2} + \text{h.c.}$
	$\mathcal{Q}_{B\chi\phi}$	$B_{\mu\nu} (\chi_a^T C \sigma^{\mu\nu} \chi_b) \phi_c + \text{h.c.}$	$n_\phi \binom{n_\chi}{2} + \text{h.c.}$

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Example: Two dark vectors $X_{1,2}$

Assumptions

SM & two vectors $X_{1,2}$, dark parity

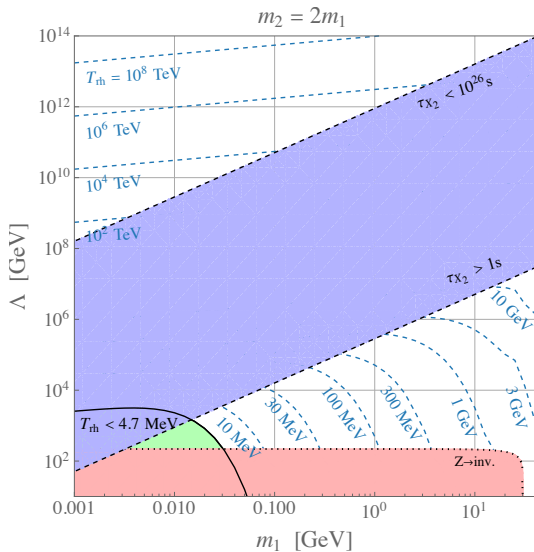
Lagrangian

$$\mathcal{L} = \mathcal{L}_{\text{SM}} - \frac{1}{4} X_1^{\mu\nu} X_{1\mu\nu} + \frac{m_1^2}{2} X_1^\mu X_{1\mu} - \frac{1}{4} X_2^{\mu\nu} X_{2\mu\nu} + \frac{m_2^2}{2} X_2^\mu X_{2\mu} \\ + C_{BX^2} B_\mu^\nu X_{1\nu}^\alpha X_{2\alpha}^\mu + C_{\tilde{B}X^2} \tilde{B}_\mu^\nu X_{1\nu}^\alpha X_{2\alpha}^\mu$$

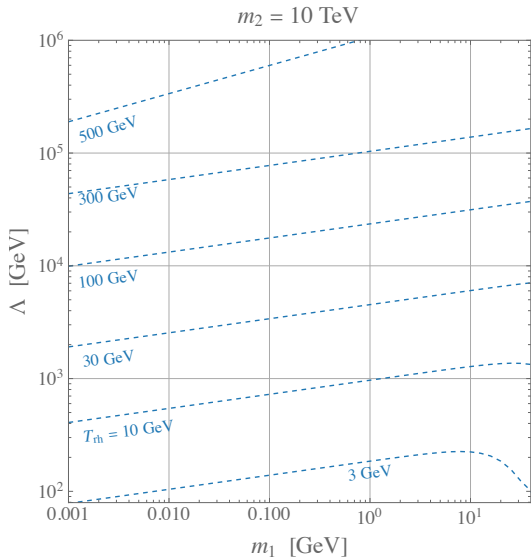
Freeze-in scenario

CMB and BBN, direct searches

Dark vector masses



Dark vector masses



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Summary

DM

Astronomical observations

DSMEFT/DWET

SMEFT/WET + $\phi + \chi + X_\mu$

Example

Two dark vectors $X_{1,2}$