

International
UON Collider
Collaboration



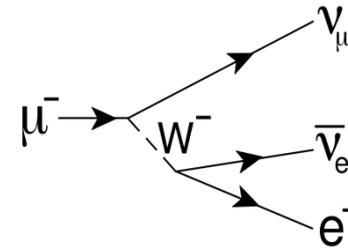
FLUKA simulations of neutrino-induced dose

G. Lerner, D. Calzolari, A. Lechner, C. Ahdida

Accelerator Design Meeting
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Context

- Neutrinos from muon $^{\pm}$ decays can lead to non-negligible dose levels



- Main features:
 - No benefit from shielding (possibly detrimental)
 - Highly collimated radiation cone (width $\sim 1/\gamma$) emerging on the earth surface
 - Neutrino cross sections dominated by Deep Inelastic Scattering and approximately proportional to energy

Strong dependence
on muon energy

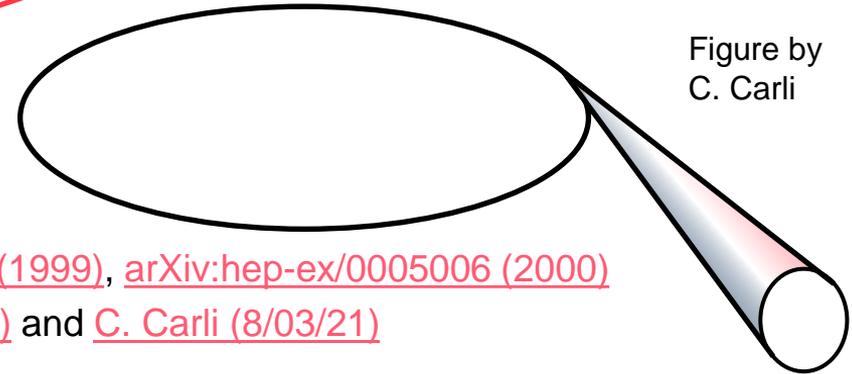


Figure by
C. Carli

- Previous work:
 - Calculations by B.J. King:, [arXiv:physics/9908017 \(1999\)](https://arxiv.org/abs/physics/9908017), [arXiv:hep-ex/0005006 \(2000\)](https://arxiv.org/abs/hep-ex/0005006)
 - More recent presentations by [D. Schulte \(18/01/21\)](#) and [C. Carli \(8/03/21\)](#)

Requirements for a full dose assessment in real scenarios

1

DOSE KERNEL: dose (or dose-equivalent) in a reference material vs longitudinal and lateral distance from fixed-point decay of monoenergetic and mono-directional muons, per unit muon decay

FOCUS OF THIS TALK
(for 5-TeV muon beam)

2

Folding with **BEAM PARAMETERS** taking into account space distribution of muon decays (e.g., along arc or straight sections), angular divergence (due to optics) and beam intensity

3

Merging the the real-world geometry to obtain a realistic **DOSE SURFACE MAP** using dedicated tools (e.g., GeoProfiler)

FLUKA simulation workflow



- A FLUKA simulation framework was set up to calculate the **dose kernel**
- Two-step approach:
 1. **MUON DECAY SIMULATION**
 - Yields as output a **list of emitted neutrinos** with their flavour, energy and angle (in the lab frame)
 - Allows to **'filter' the list of neutrinos according to the macroscopic cross section** in a reference material (e.g. soil) to obtain **distributions of interacting neutrinos**
 2. **NEUTRINO INTERACTION SIMULATION**
 - Takes as input the list of interacting neutrinos obtained in the 1st step (i.e. with cross section filter)
 - Samples uniformly the distance of the neutrino interaction with respect to the muon decay point, within a user-defined range
 - Computes the x-y-z position of the neutrino interaction from the sampled distance and the angular direction read from the input
 - Scores 3D distribution of any relevant quantity supported by FLUKA (e.g., absorbed dose, dose-equivalent, or more)

Neutrino cross sections on nucleons in FLUKA / 1



- FLUKA employs internally-computed (NUNDIS) **cross sections on nucleons**
- **4 processes** (resonant, quasi-elastic, Deep Inelastic Scattering, charm production) with total values driven by DIS at high energy
- Individual cross sections for:
 - **Neutrino flavours** (e- μ - τ) yielding negligible differences (especially at high energy)
 - **Neutrino vs antineutrino**, with antineutrino cross sections generally lower (around a factor 2)
 - **Target nucleon** (proton vs neutron)
- **Values to be compared with B. J. King's** →

B.J. King neutrino cross sections

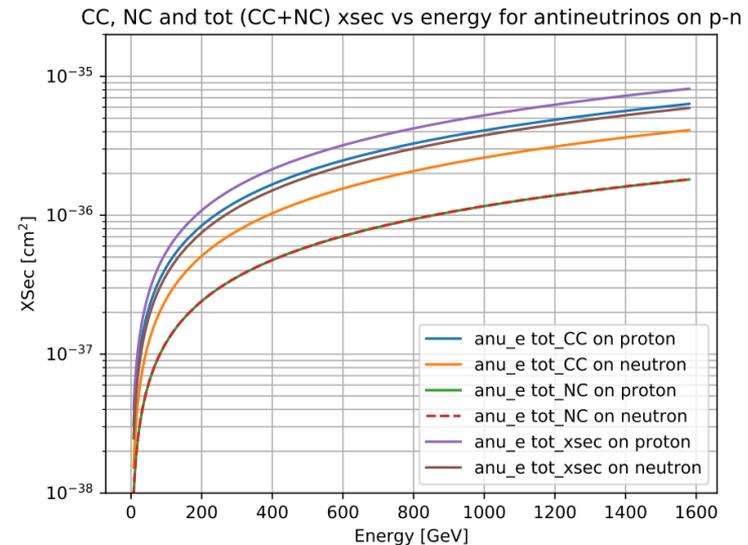
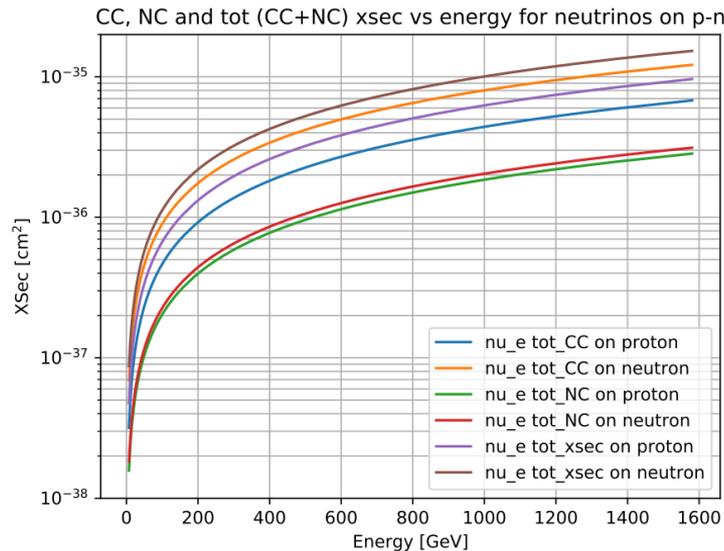
from [arXiv:physics/9908017 \(1999\)](https://arxiv.org/abs/physics/9908017)

Table 1: Contributions to the radiation dose from the different types of neutrino interactions. The reduced cross-section, σ_R^i , is specified for 100 GeV neutrinos and using a simple model for the nucleon in which the cross-section ratio for neutrinos to anti-neutrinos was assumed to be 2:1 and ignoring the small differences between the average hadronic fractions for NC and CC interactions. The reduced cross-section and product are in units of $10^{-35} \text{cm}^2/\text{TeV}$.

int., i	σ_R^i	k^i	$\langle x^2 \rangle^i$	$\sigma_R^i \cdot k^i \cdot \langle x^2 \rangle^i$
$\nu_\mu - CC$	0.722	0.458	0.533	0.176
$\nu_\mu - NC$	0.226	0.458	0.533	0.055
$\nu_e - CC$	0.722	1.000	0.400	0.289
$\nu_e - NC$	0.226	0.458	0.400	0.041
$\bar{\nu}_\mu - CC$	0.375	0.292	0.533	0.058
$\bar{\nu}_\mu - NC$	0.131	0.292	0.533	0.020
$\bar{\nu}_e - CC$	0.375	1.000	0.400	0.150
$\bar{\nu}_e - NC$	0.131	0.292	0.400	0.015
SUM				0.804

Neutrino cross sections on nucleons in FLUKA /2

- Cross sections vs energy for electron neutrinos and antineutrinos
- Comparison with B. J. King's table (using numbers at 1TeV): generally similar, but individual values can deviate, and differences such as neutron vs proton targets are not considered by B. J. King



Macroscopic neutrino cross sections

- The key normalization parameter of the simulations is the **macroscopic cross section Σ** , i.e., the probability of a neutrino interaction per unit distance travelled
- Our approach: compute it for the material (or compound) of interest as:

$$\Sigma(E) = n_p \cdot \sigma_p(E) + n_n \cdot \sigma_n(E)$$

Number density of protons and neutrons in the material (compound)

Neutrino cross section on protons and neutrons

- This approach considers the exact density of protons/neutrons (including breakdown of isotopes)
- Neglected: Fermi motion (changing effective E), Pauli exclusion principle

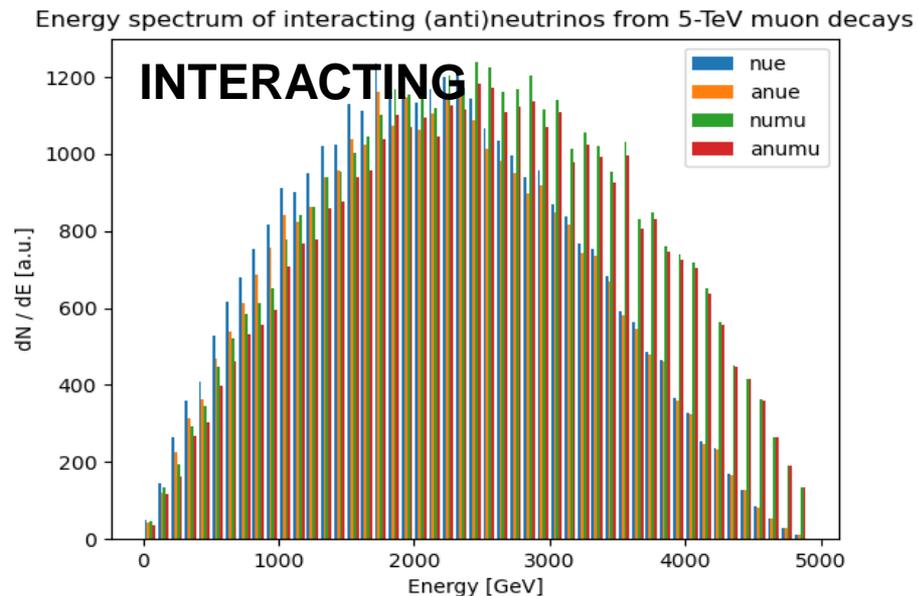
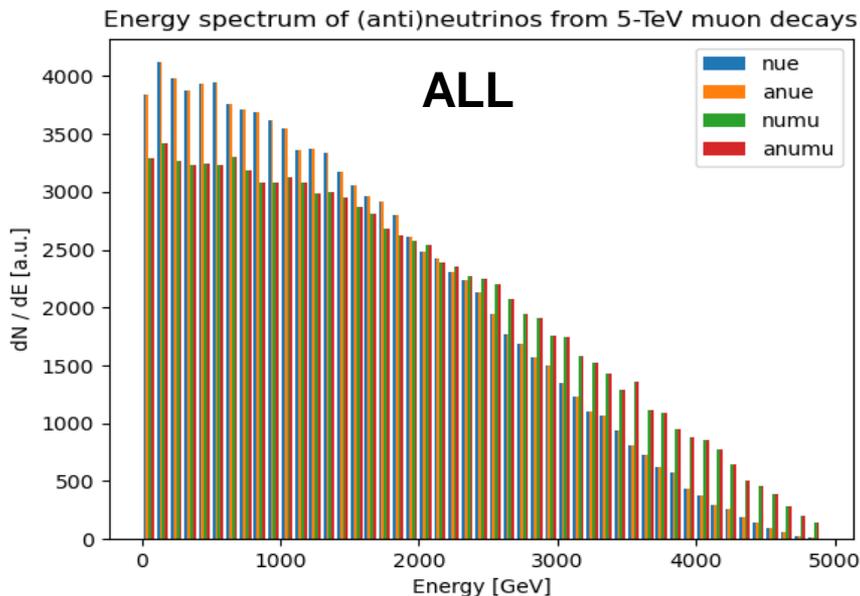
FLUKA simulation: 1st step (muon decay)



- Muons with fixed energy (5 TeV for the examples presented in this talk) moving along z are forced to decay at $(x,y,z)=(0,0,0)$
- The neutrino info (id, energy, angle cosines) is written in dump files with probability proportional to the energy-dependent macroscopic cross section in a user-selected material relative to 'max-E' neutrinos (i.e. neutrinos with the energy of the decaying muons)
→ *example: for 5-TeV muons, a neutrino with energy $X < 5\text{TeV}$ is retained with probability $\Sigma_X / \Sigma_{5\text{TeV}}$*
- The above cross sections are calculated automatically in the routine for any material or compound, using the formula in slide 7

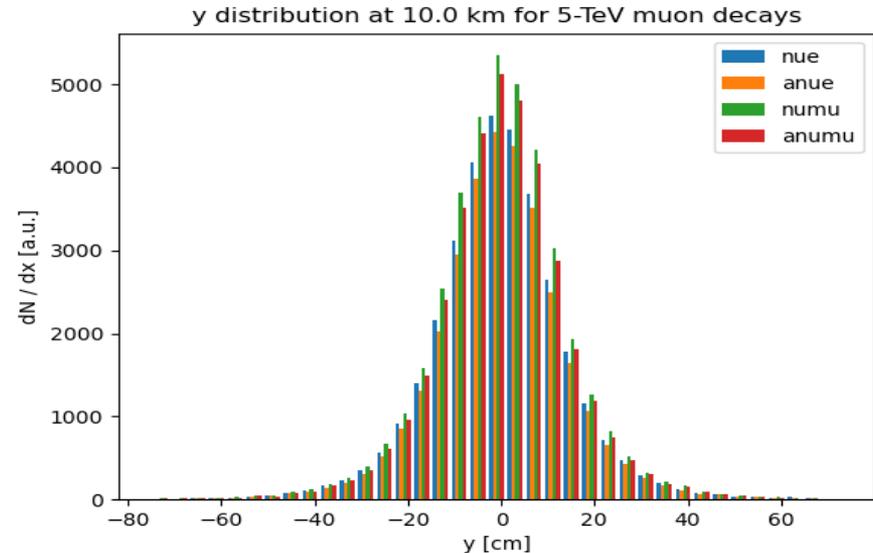
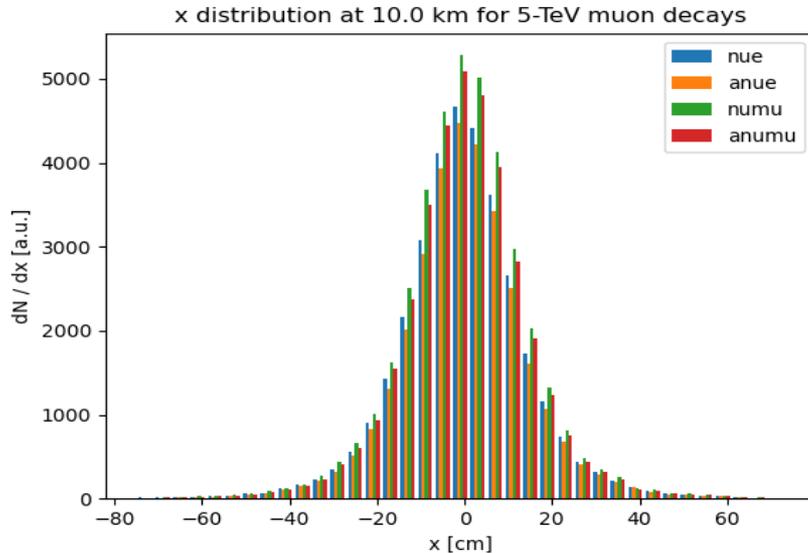
Energy distribution of total and interacting neutrinos

- (anti)neutrino energy spectra before (all) and after (interacting) xsec filtering
- The spectrum of interacting neutrinos peaks at around half the muon energy
- Muon (anti)neutrinos are relatively more energetic than the electron ones



Width of neutrino cone at 10km

- At 10km from the muon decay point, the neutrino beam width is of the order of few tens of cm (i.e., the angular aperture is of the order of the 10^{-5} rad)
→ in practice, a **narrow cone** shall be expected at any distance

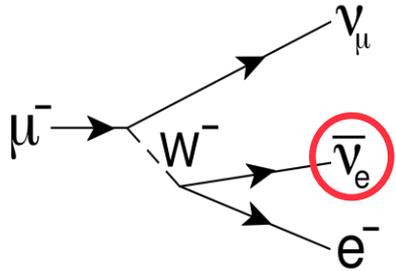


FLUKA simulation: 2nd step (neutrino interactions)



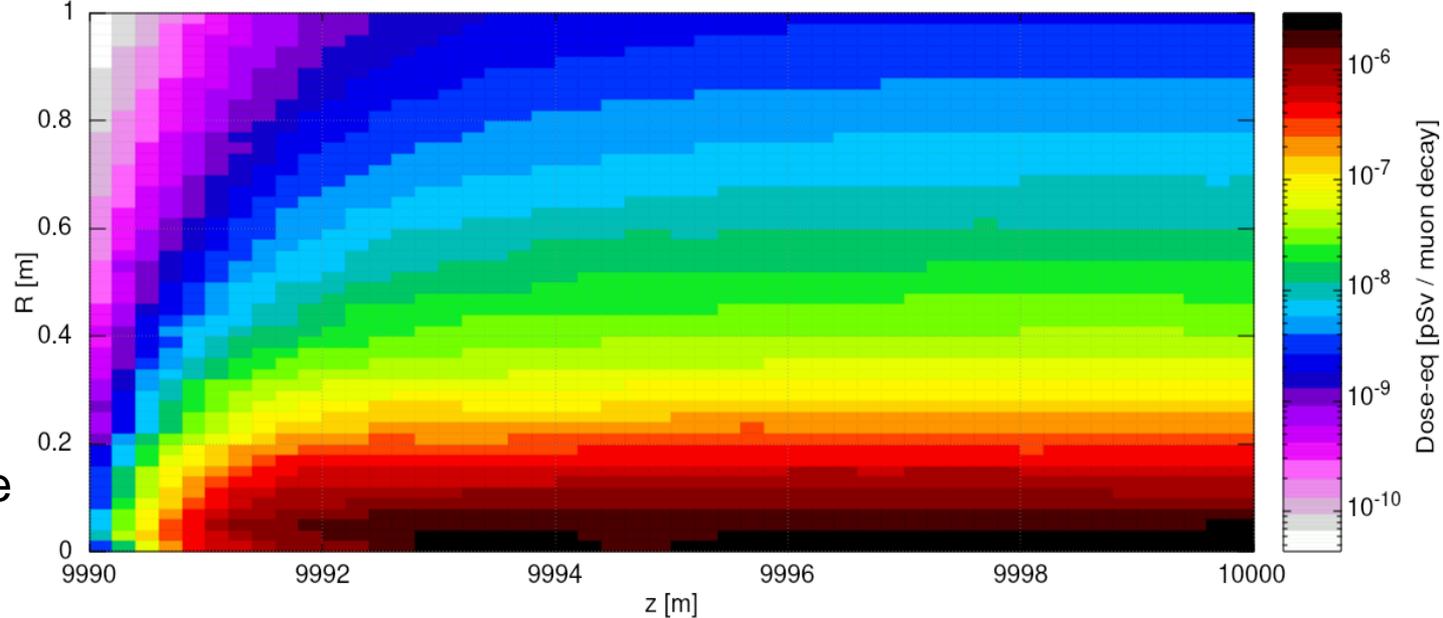
- The 2nd step simulation uses as source the interaction of the neutrinos written in the output file of the 1st step, **sampling the longitudinal position L of the interactions in a user-selected $[L_{\min}, L_{\max}]$ interval**
- With this method the neutrino interactions are correctly sampled proportionally to the energy-dependent cross sections, and the correlation between energy and angles is kept
- The exact position of each interaction positions is a function of the sampled L and of the angle cosines written in the output file: $(x,y,z) = R^* (tx,ty,tz)$
- *Our example:*
 - *neutrino interactions forced in soil (standard FLUKA 'earth' compound used for many other applications)*
 - *Scoring of **dose-equivalent (in pSv) with EDWORST coefficients** (as advised by C. Ahdida)*

Dose-eq build-up over 10m in soil: antineutrino-e interactions

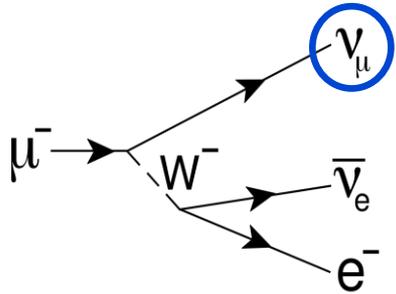


- Forcing neutrino interactions in [9990,10000]m range
- The majority of the dose-eq build-up occurs over few metres

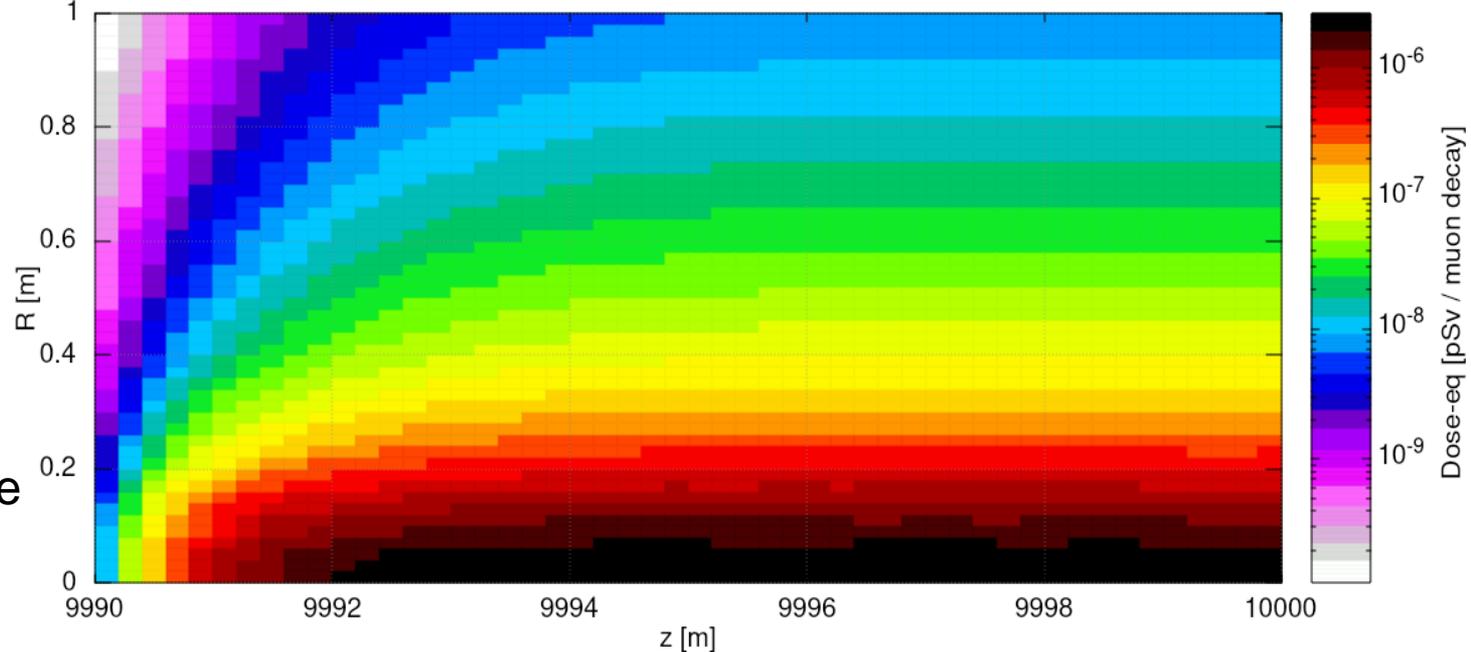
2D Dose-eq map from electron antineutrino interactions in a 10m-thick soil layer at 10km from 5-TeV muon- decay point



Dose-eq build-up over 10m in soil: neutrino-mu interactions

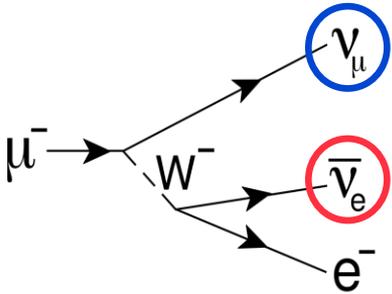


2D Dose-eq map from muon neutrino interactions in a 10m-thick soil layer at 10km from 5-TeV muon- decay point

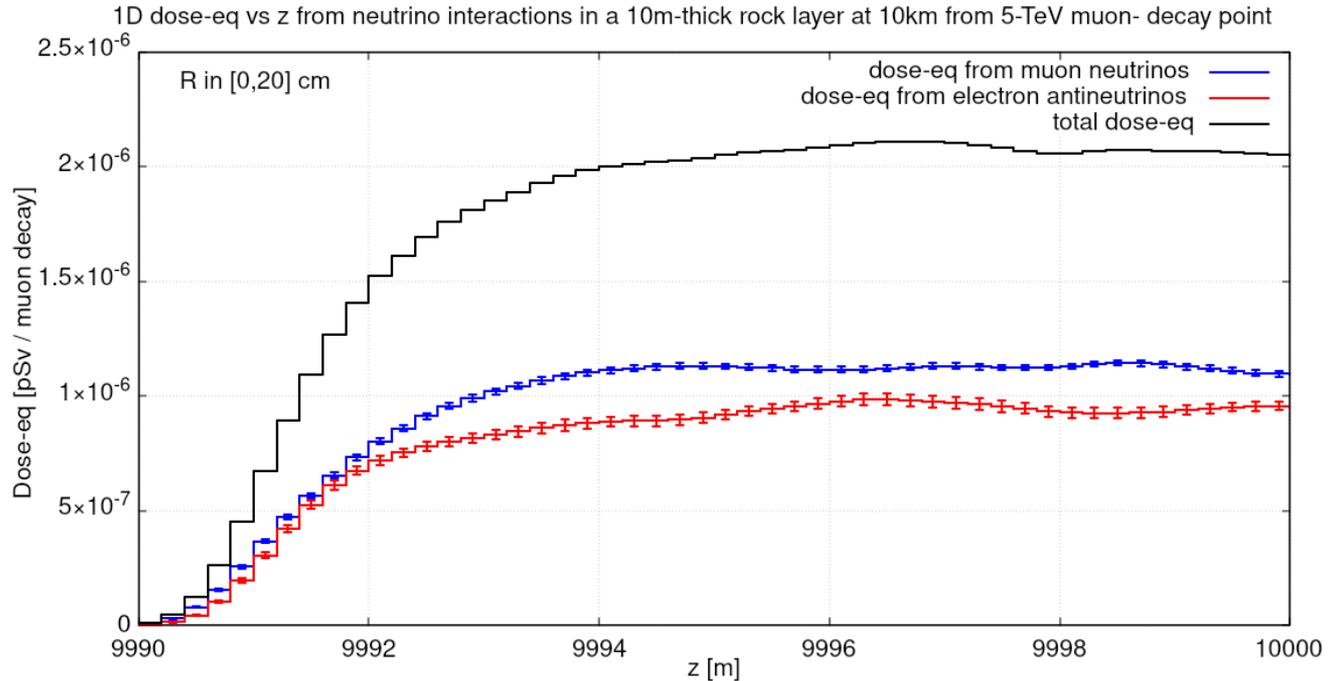


- Forcing neutrino interactions in [9990,10000]m range
- Again, most of the dose-eq build-up occurs over few metres

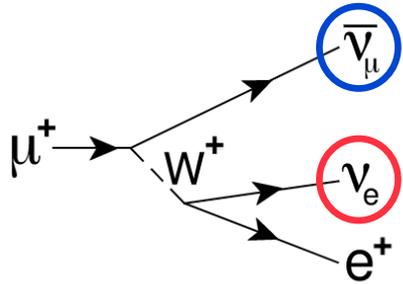
Dose-eq build-up over 10m in soil: 1D projections vs z



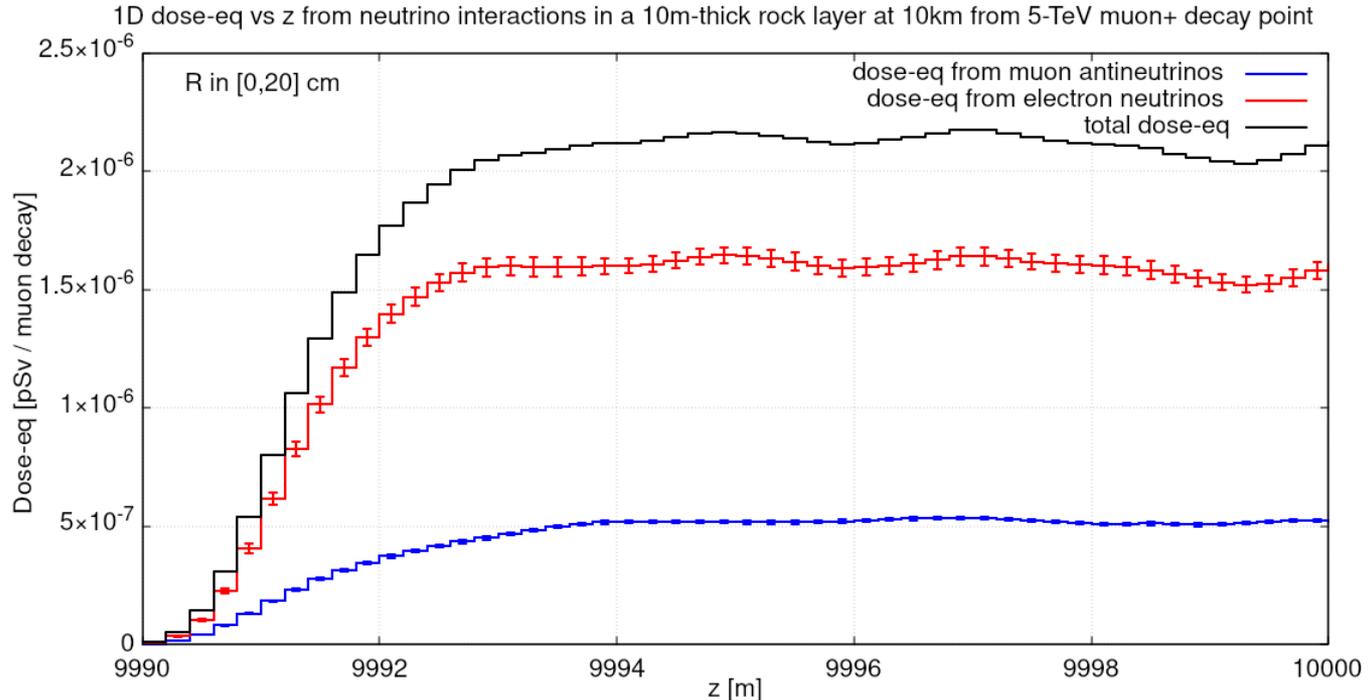
- Ignoring statistical fluctuations and minor effects, after around $\sim 5\text{m}$ in soil the dose-eq reaches plateau conditions



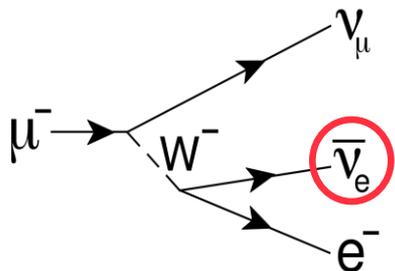
Dose-eq build-up over 10m in soil: 1D projections vs z



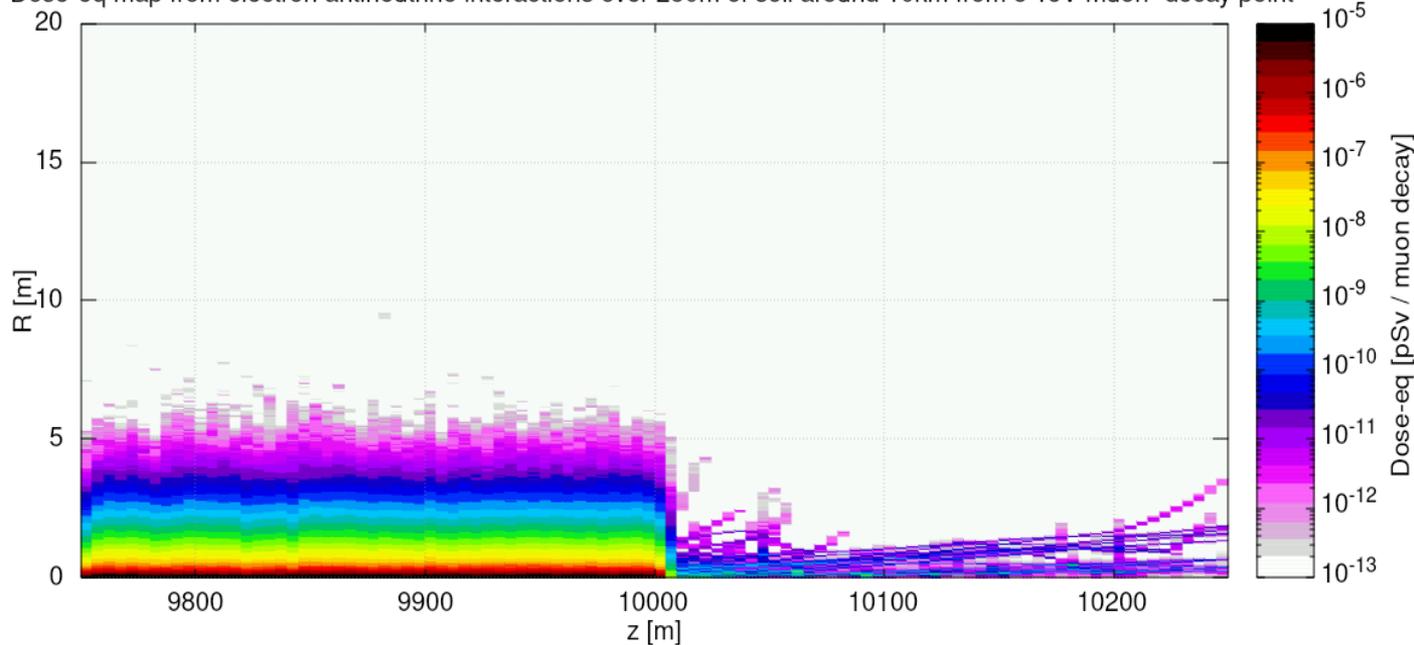
- In the case of the muon+ decay, the dose-eq at the plateau is dominated by the electron neutrino component, but the total is similar



Broader shower in soil: **antineutrino-e** int. (dose-eq)

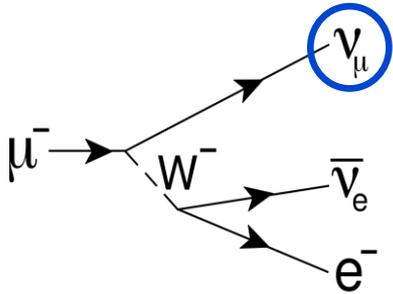


2D Dose-eq map from electron antineutrino interactions over 250m of soil around 10km from 5-TeV muon- decay point

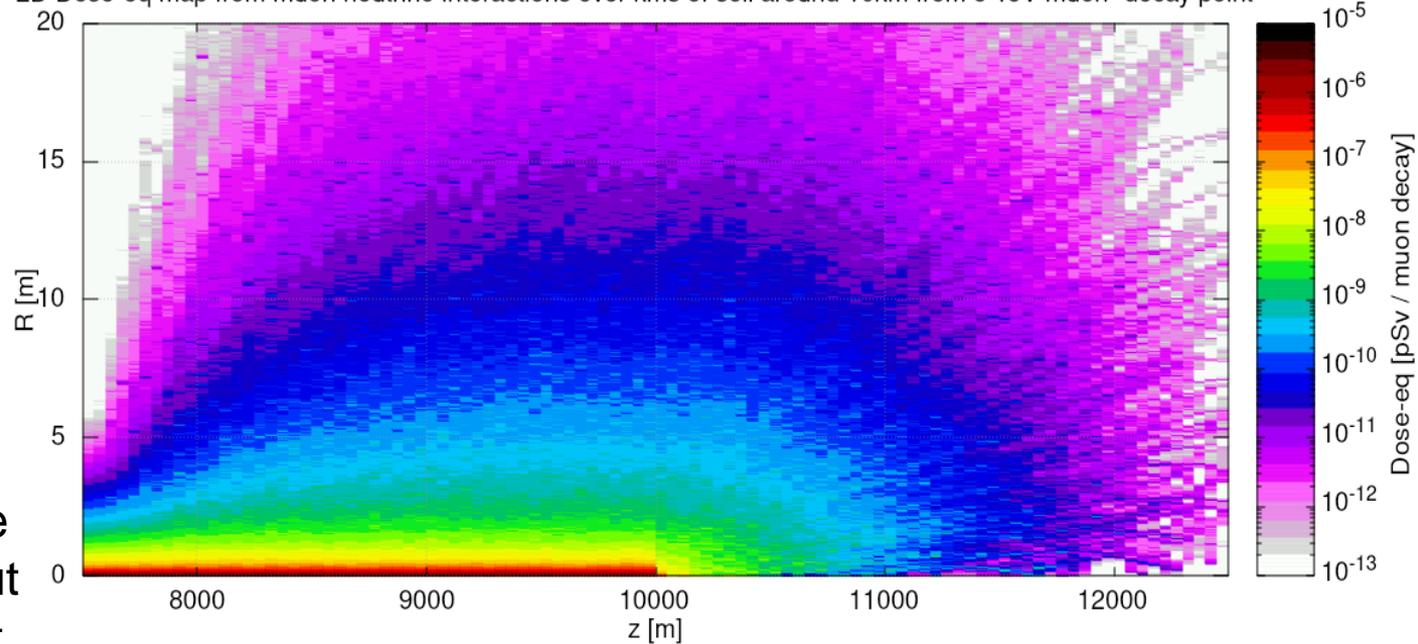


- Forcing neutrino interactions in [9750, 10000]m range
- Relatively narrow shower, only few low-E muons leaking

Broader shower in soil: neutrino-mu int. (dose-eq)

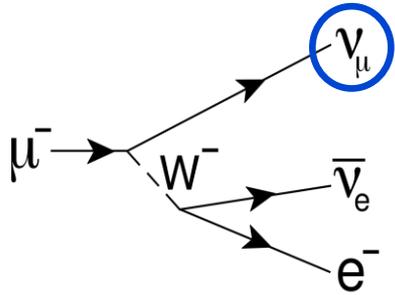


2D Dose-eq map from muon neutrino interactions over kms of soil around 10km from 5-TeV muon- decay point



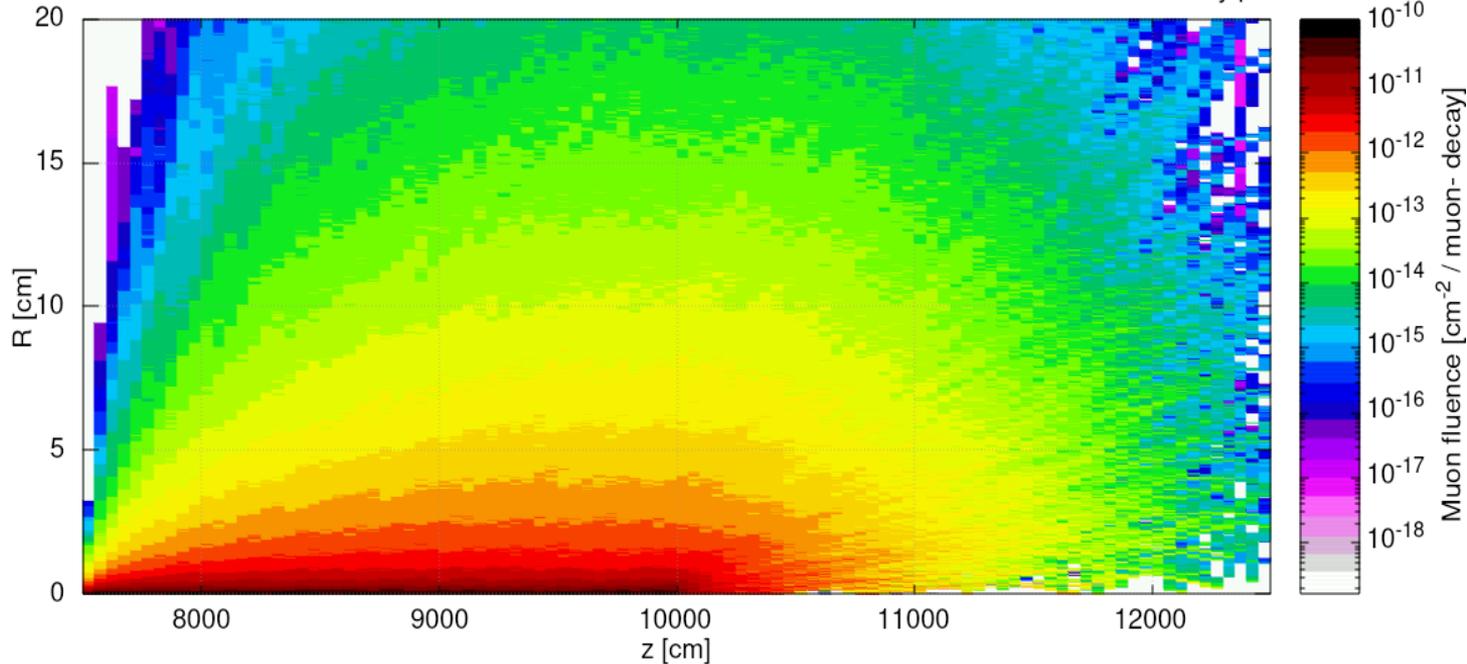
- Forcing neutrino interactions in [9750,10000]m range
- Broader shower due to high-E muons, but dose-eq peak still at low R

Broader shower in soil: neutrino-mu int. (muon fluence)



- 2D map of **muon fluence**, confirming that the dose-eq at high distance from neutrino interactions is due to long-range muons

2D muon fluence from muon neutrino interactions over kms of soil around 10km from 5-TeV muon- decay point



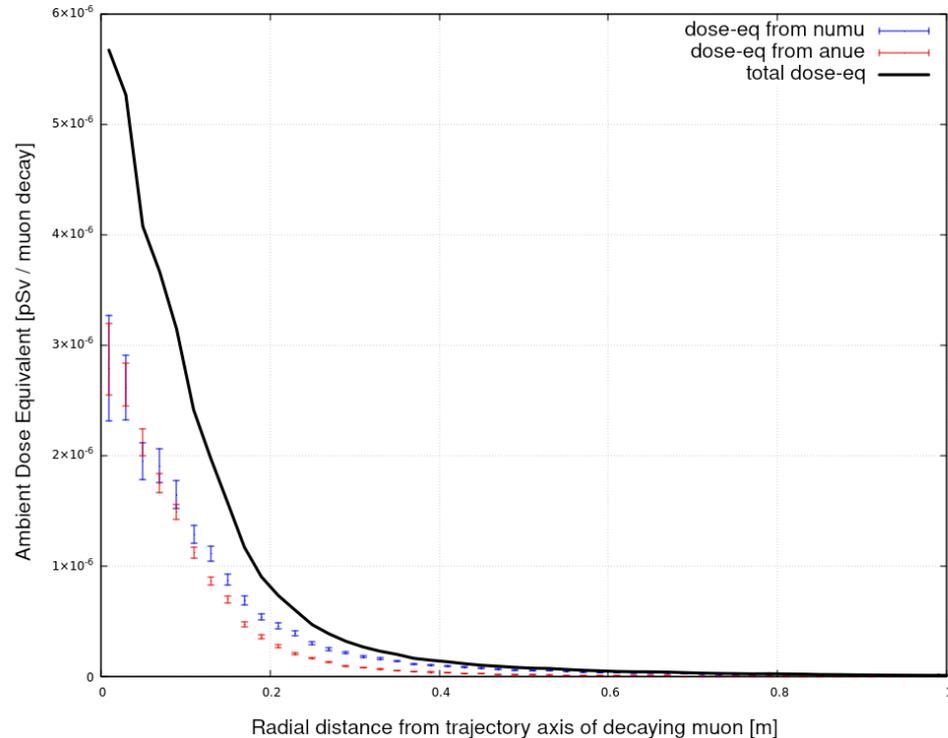
Key **dose kernel** quantity: dose-eq vs R at plateau

- We have seen that the dose-equivalent begins to ‘plateau’ already after few metres of soil
- Other quantities, such as the muon fluence, take longer distances to plateau (due to the longer muon range) but they yield a relatively small contribution to the dose-eq
- **To obtain the dose kernel we compute the profile vs R of the dose-equivalent at different distances from the muon decay point, taking the ‘plateau’ values as reference**

Example/1: dose-eq vs R at plateau from muon- at 10km

- Very narrow dose-eq peak, as expected from the highly collimated neutrino beam
- Wider dose-eq shape for muon neutrinos compared from electron antineutrinos, with similar contribution at the peak
- Note: the 1D projections vs z in slides 14-15 were showing average dose-eq between R=0 and R=20cm

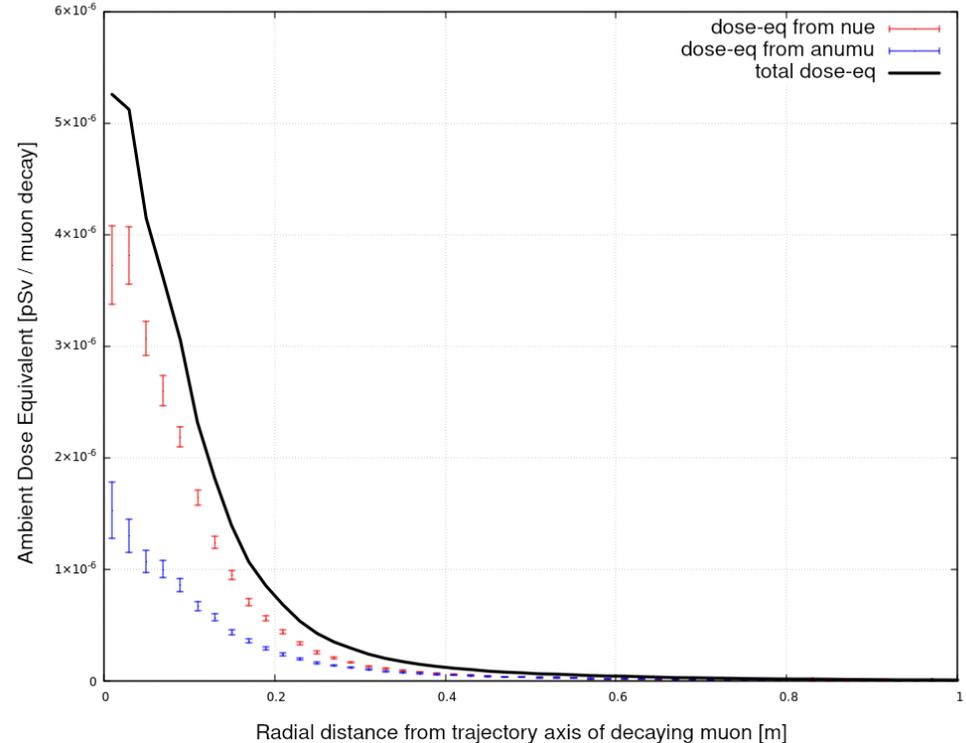
Ambient dose equivalent lateral profile at 10 km from mu- decay



Example/1: dose-eq vs R at plateau from muon+ at 10km

- Similar lateral width also for the muon+ decays
- In this case the dose peak is dominated by the electron neutrinos, as already seen in slide 15.

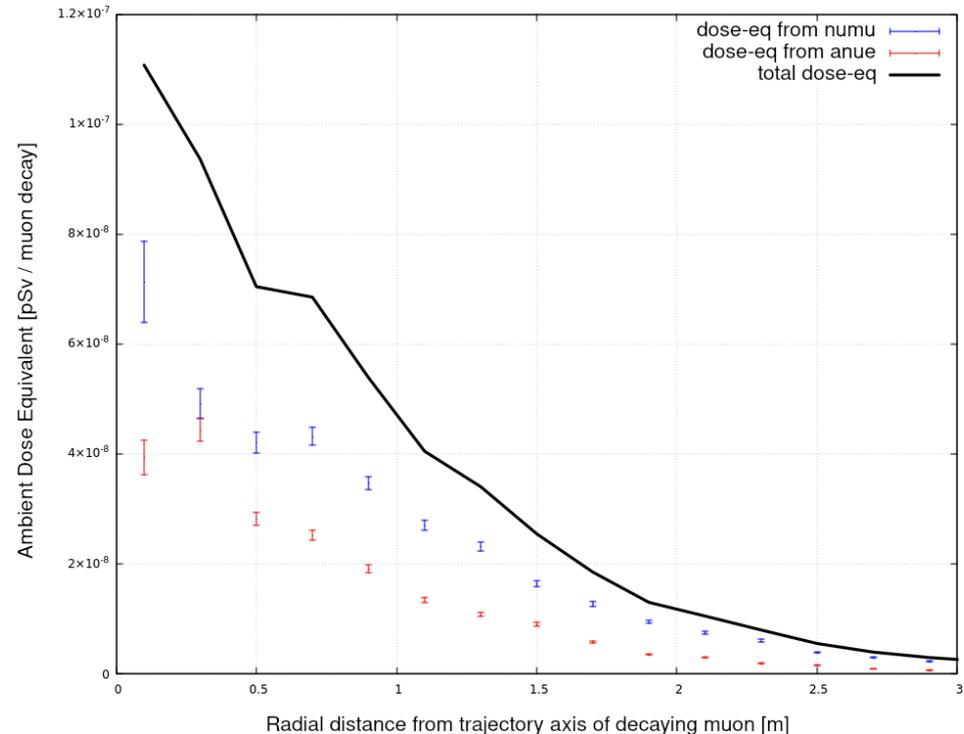
Ambient dose equivalent lateral profile at 10 km from mu+ decay



Example/2: dose-eq vs R at plateau from muon- at 100km

- More statistical fluctuations at 100km
- The profile is wider, as expected (notice the different R range in the figure)
- The contribution to the dose-eq from muon neutrinos is slightly larger, but electron antineutrinos aren't negligible, especially at $R=0$

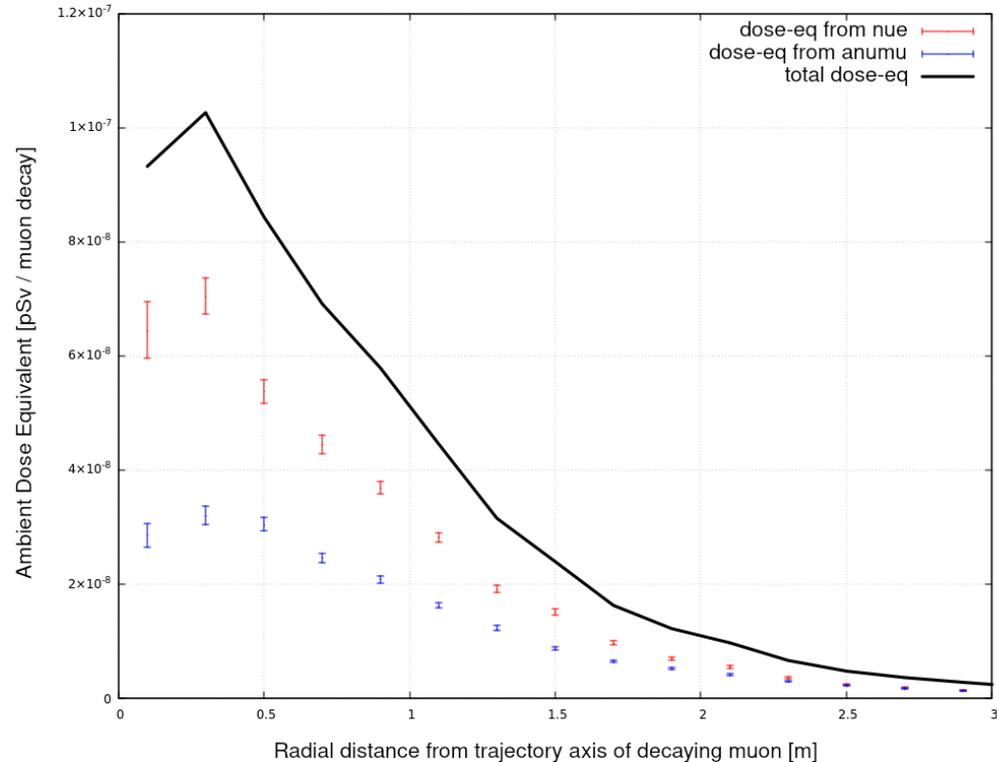
Ambient dose equivalent lateral profile at 100 km from mu- decay



Example/2: dose-eq vs R at plateau from muon+ at 100km

- Similar dose-eq profile at 100km from muon+ decays too
- In this case, most of the dose-eq near R=0 comes from electron neutrino interactions

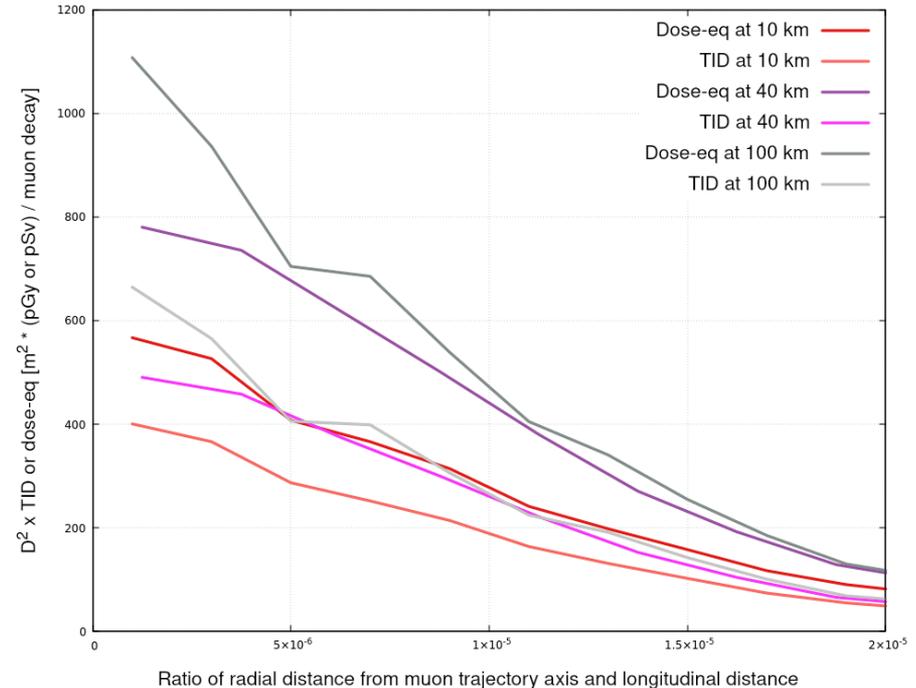
Ambient dose equivalent lateral profile at 100 km from mu+ decay



Dose vs dose-eq

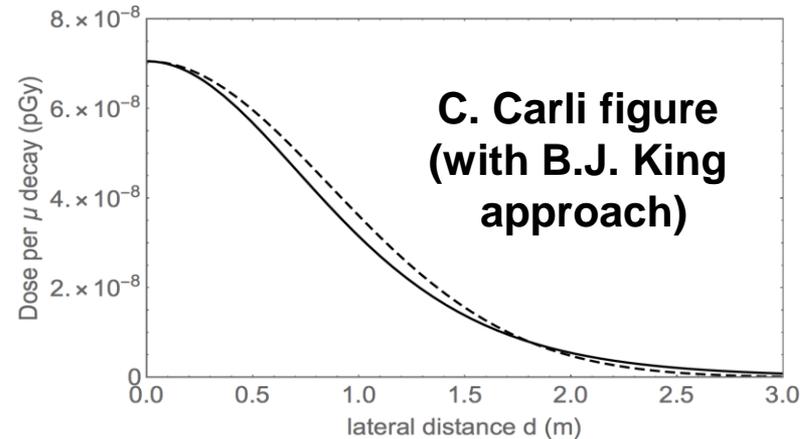
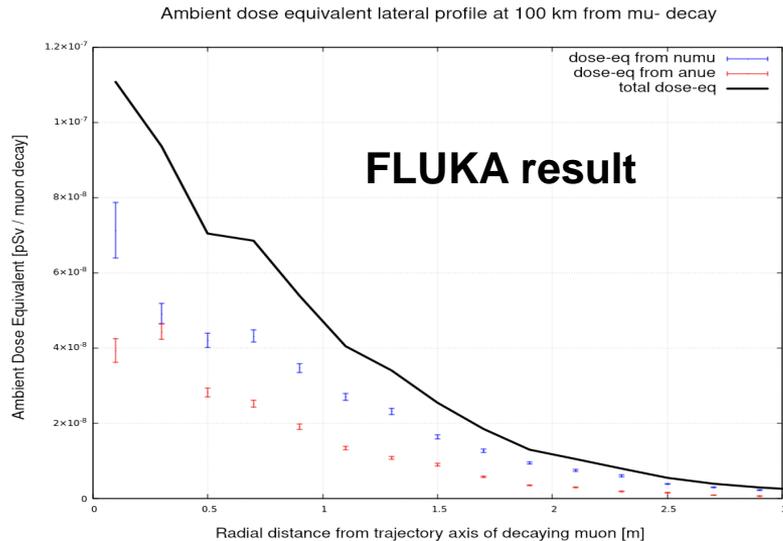
- To benchmark the FLUKA results with previous calculations, we need to compare the scoring of **dose-equivalent** (shown so far) with the **absorbed dose**, simply defined as energy per unit mass (in units of Gy = J/kg)
- Note that the plot is showing $L^2 \cdot D$ vs R/L to better compare different distances
- **The dose-eq (expressed in Sv) is about 50% higher**

Lateral profile of $D^2 \times$ (TID or dose-equivalent) from 5 TeV mu- decay



Comparing FLUKA and B. J. King / C. Carli calculations

- The FLUKA dose-eq peak is $\sim 1.1 \cdot 10^{-7}$ pSv/muon
- C. Carli's calculation yields a peak of $\sim 7 \cdot 10^{-8}$ pGy/muon, i.e., $\sim 50\%$ less, consistently with the expected difference due to TID vs dose-eq
→ **excellent agreement!**

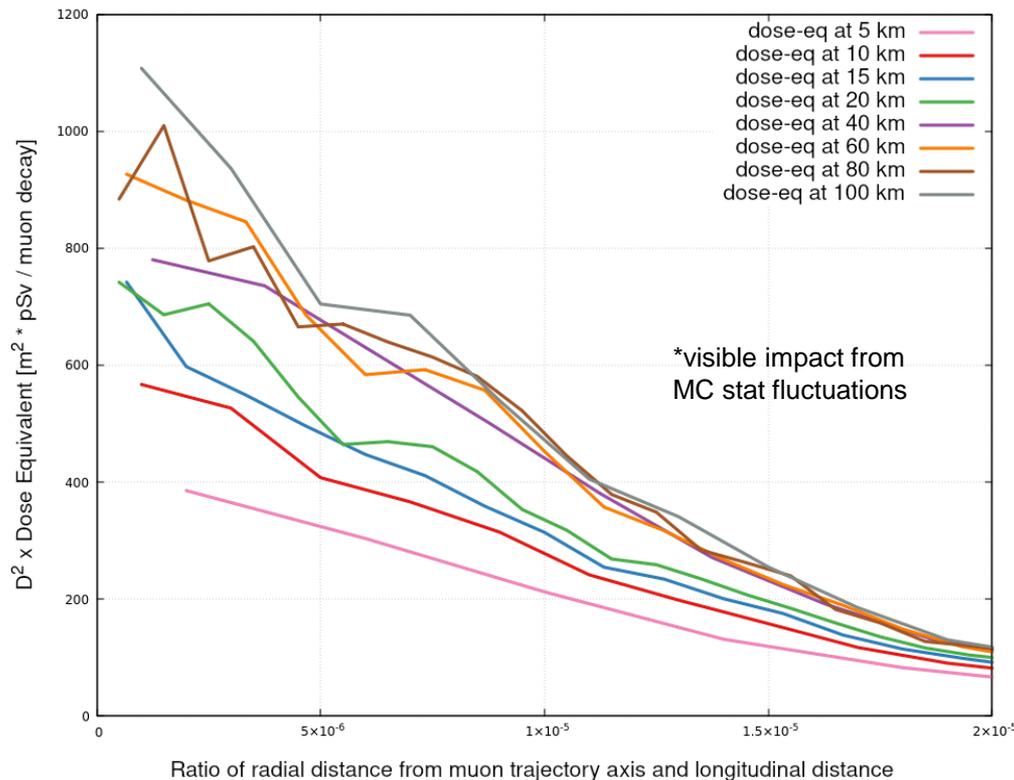


10 TeV c.o.m. energy, $L_s = 100$ km

Summary: dose-eq vs R at plateau from muon-

- Summary figure of dose-equivalent vs R at different distances from the **muon-** decay point
- Showing again $L^2 \cdot D$ vs R/L to better compare different distances

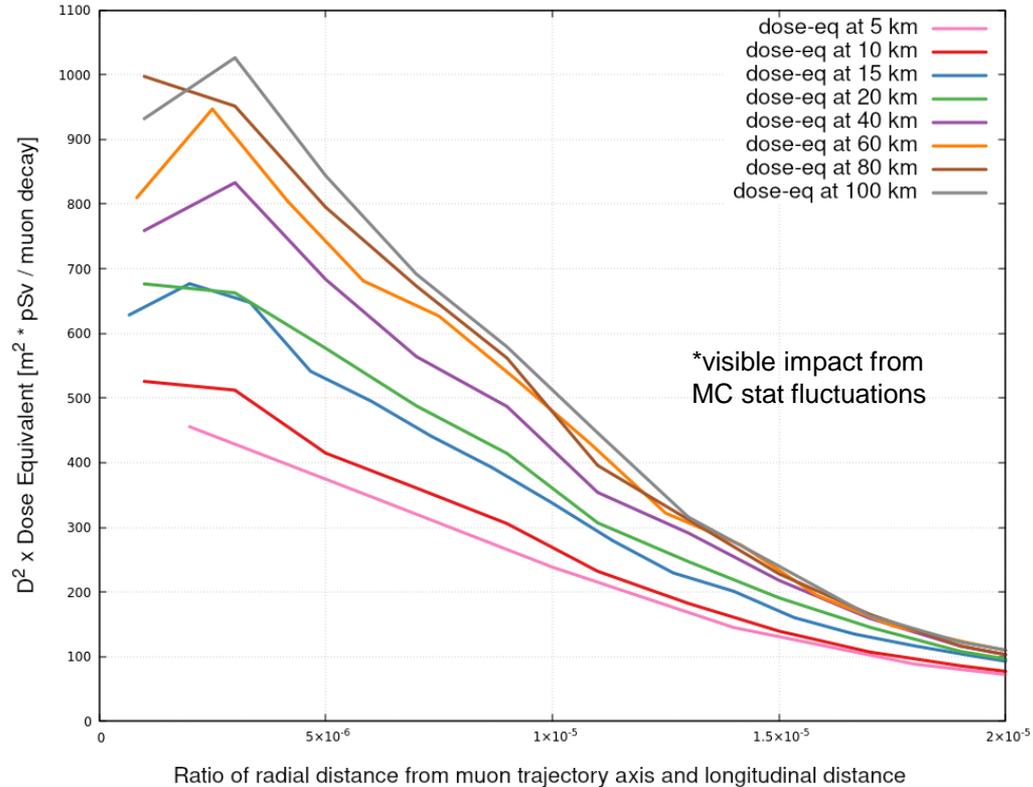
Lateral profile of $D^2 \times$ dose-equivalent from 5 TeV mu- decay



Summary: dose-eq vs R at plateau from muon+

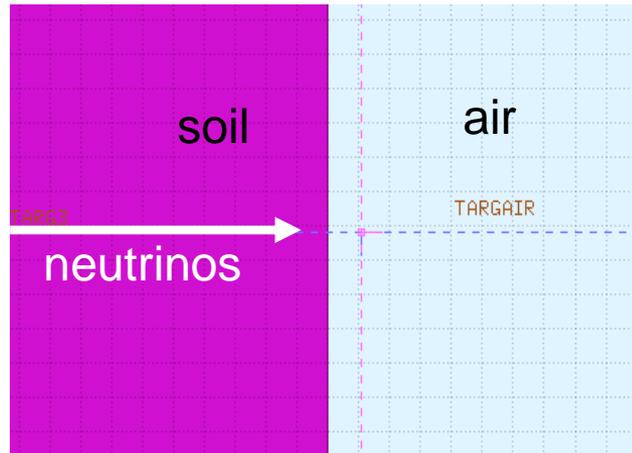
Lateral profile of D^2 x dose-equivalent from 5 TeV mu+ decay

- Summary figure of dose-equivalent vs R at different distances from the **muon+** decay point
- Showing again $L^2 \cdot D$ vs R/L to better compare different distances



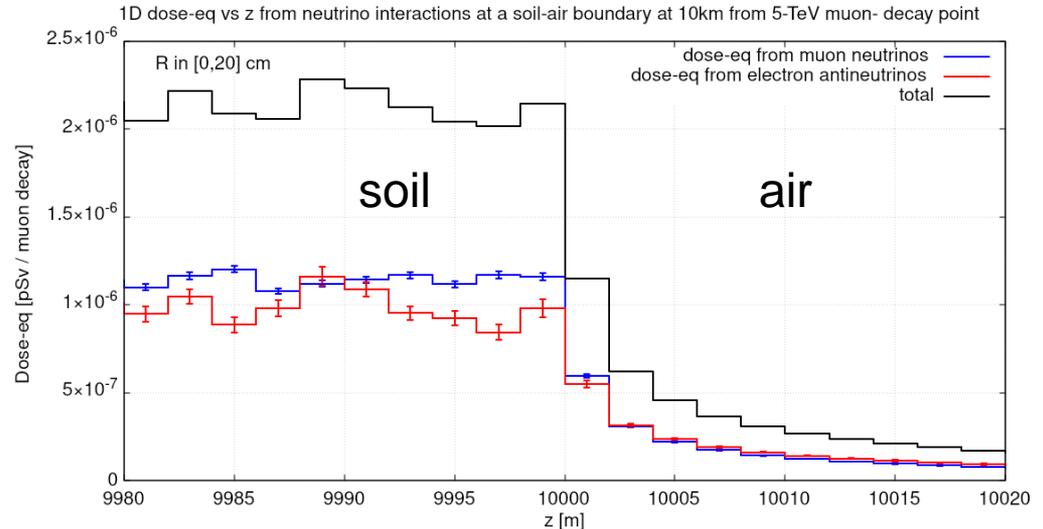
Extra: neutrino-induced dose-eq at soil-air boundary

- Simplified geometry: neutrino beam exiting orthogonally* a soil-air surface



*in realistic scenarios small angles should be considered, but this first calculation can already yield interesting insights

- Study the dose-eq in air from the neutrino interactions in soil only (neglecting the few neutrino interactions in air)
- In the first few metres of air the dose-eq is still relatively high, as expected



Summary and Outlook



- I presented the results of FLUKA simulations of neutrino-induced radiation from point-like decays of 5-TeV muons, without angular divergence
- The **dose-eq vs R** maps shall serve as **dose kernels** for full dose predictions, requiring a convolution with machine parameters (arc/LSS lengths, intensity) and considering the geography of the region
- Good agreement is found with previous calculations by B. J. King and C. Carli, after taking into account a ~50% difference between dose-eq and TID
- The build-up of the dose-eq in soil and the dose-eq decay in air occur within ~few metres (neglecting neutrino interactions upstream and in air respectively)
- Next steps:
 - simulate more cases (e.g., more muon energies) and quantities
 - study the dose-eq breakdown in individual particle types (e.g., electrons, hadrons..)
 - boost the Monte Carlo statistics as much as possible