

Everything about Higgs is Puzzling

$$\mathcal{L} = yH\psi\overline{\psi} + \mu^2|H|^2 - \lambda|H|^4 - V_0 + \dots$$

- Pattern of Yukawa couplings y:
 - Flavour problem
- Magnitude of mass term μ:
 - Naturalness/hierarchy problem
- Magnitude of quartic coupling λ :
 - Stability of electroweak vacuum
- Cosmological constant term V_0 :
 - Dark energy

Higher-dimensional interactions?

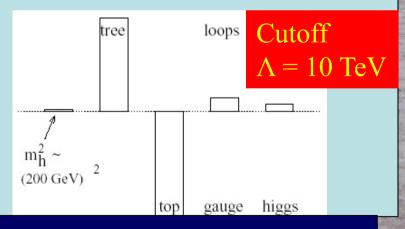
Theoretical worries about the Higgs boson

Elementary Higgs or Composite?

• Higgs field:

$$v = <0|H|0> \neq 0$$

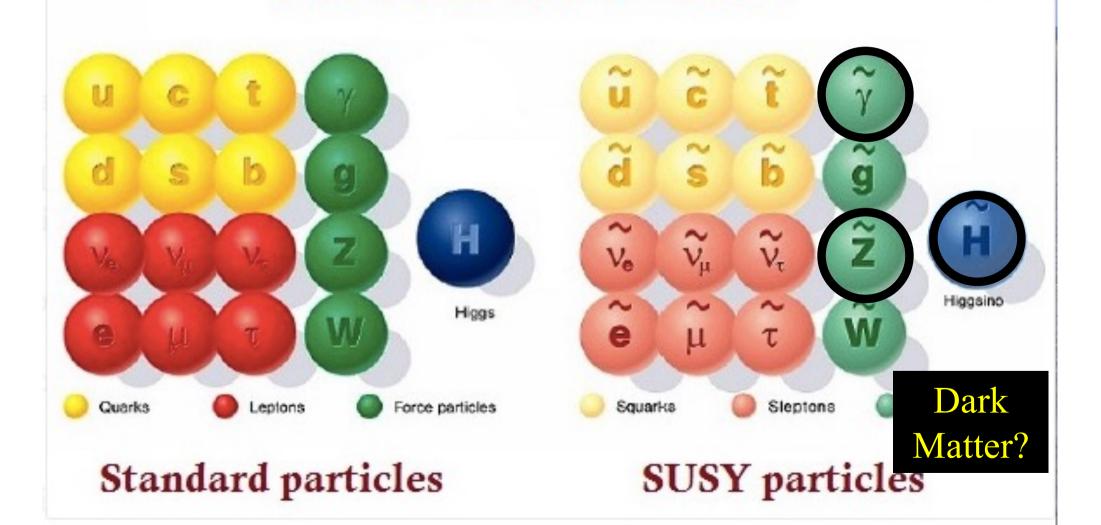
- Quantum loop problems
- M_h, v, other masses have quadratic divergences



Cut-off $\Lambda \sim 1$ TeV with Supersymmetry?

- Fermion-antifermion condensate?
- Just like π in QCD, Cooper pairs in BCS superconductivity
- Need new 'technicolour' force
- Heavy scalar resonance?
- (Problems with precision electroweak data)
- Pseudo-Nambu-Goldstone boson?

Minimal Supersymmetric Extension of the Standard Model



What lies beyond the Standard Model?

Supersymmetry

Stabilize electroweak vacuum

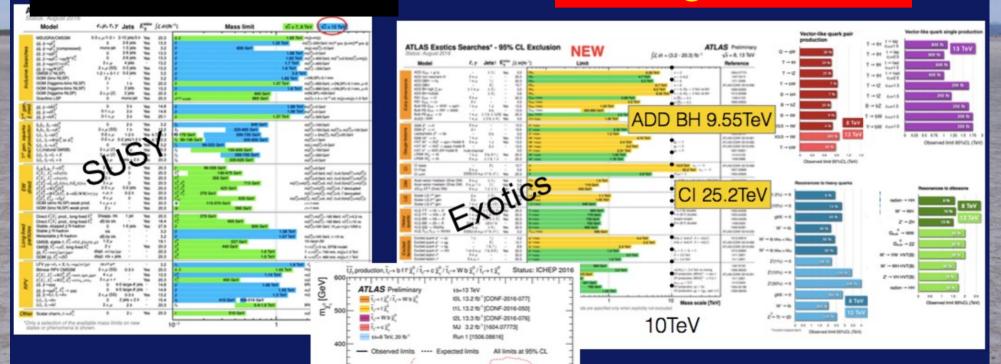
New motivations from LHC

- Successful prediction for Higgs mass
 - Should be < 130 GeV in simple models
- Successful predictions for couplings
 - Should be within few % of SM values
- Naturalness, GUTs, string, dark matter, $g_{\mu} 2$, ...

Nothing (yet) at the LHC



Nothing else, either



More of same?
Unexplored nooks?
Novel signatures?

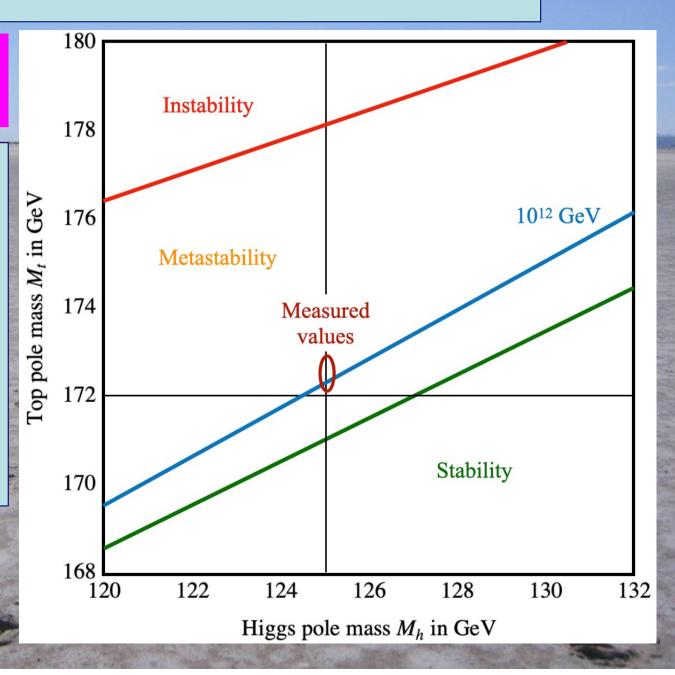
Is "Empty Space" Unstable?

Politzer & Wolfram,
Hung,
Cabibbo, Maiani, Parisi & Petronzio:

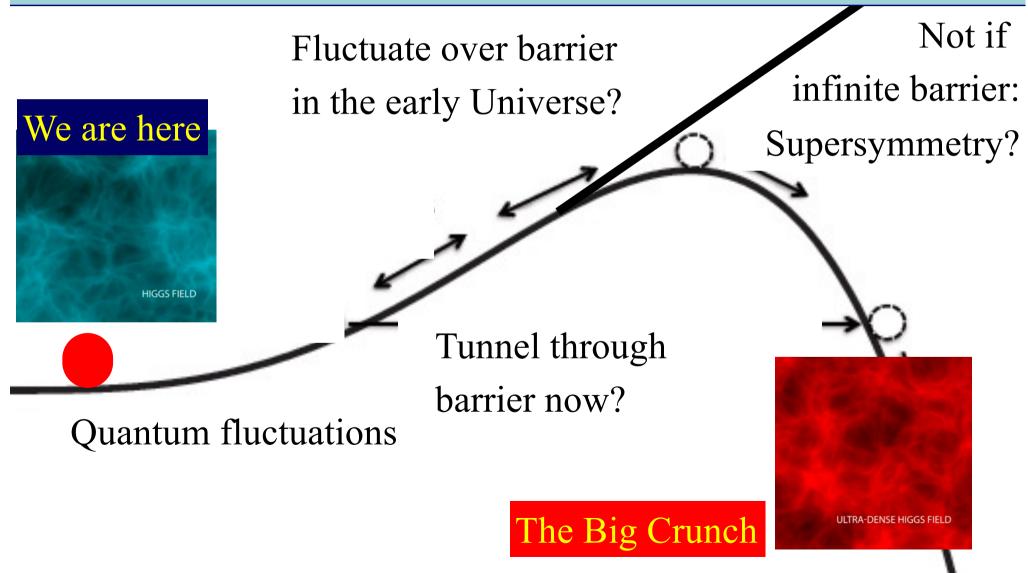
Depends on masses of Higgs boson and top quark, strong coupling

Instability scale $\sim 10^{12} \text{ GeV}$

Buttazzo et al, arXiv:1307.3536; Franceschini et al, 2203.17197



Will the Universe Collapse? Should it have Collapsed already?



Is "Empty Space" Unstable?

• Dependence of instability scale on masses of Higgs boson and top quark, and strong coupling:

$$\log_{10} \frac{\Lambda}{\text{GeV}} = 10.5 - 1.3 \left(\frac{m_t}{\text{GeV}} - 172.6 \right) + 1.1 \left(\frac{m_H}{\text{GeV}} - 125.1 \right) + 0.6 \left(\frac{\alpha_s(m_Z) - 0.1179}{0.0009} \right)$$

• New CMS value of m_t :

$$m_t = 171.77 \pm 0.38 \text{GeV}$$

Buttazzo et al, arXiv:1307.3536; Franceschini et al. 2203.17197

CMS Collaboration, April 2022

• Particle Data Group values:

$$m_H = 125.25 \pm 0.17$$
 GeV, $\alpha_s(m_Z) = 0.1179 \pm 0.0009$

• Instability scale:

$$Log_{10} \frac{\Lambda}{GeV} = 11.7 \pm 0.8$$

• Dominant uncertainties those in $lpha_S$ and m_t

Comments on Dark Energy

• Many orders of magnitude smaller than expected contributions from 'known' physics:

today: 10^{-48} GeV^4

QCD: $\Lambda_{\rm OCD}^4 \sim 10^{-4} \, {\rm GeV}^4$

Higgs: $m_W^4 \sim 10^8 \text{ GeV}^4$

Broken susy: $m_{susy}^{4} \sim 10^{12} \text{ GeV}^{4}$

GUT: $m_{GUT}^{4} \sim 10^{64} \text{ GeV}^{4}$

Quantum Gravity: $m_P^4 \sim 10^{76} \text{ GeV}^4$

- Need new physics!
- A great challenge for string theory

Weinberg: Anthropic Estimate of the Cosmological Constant

The cosmological constant problem

"... the laws of nature should allow the existence of intelligent beings that can ask about the laws of nature ..."

Steven Weinberg

Theory Group, Department of Physics, University of Texas, Austin, Texas 78712

Astronomical observations indicate that the cosmological constant is many orders of magnitude smaller than estimated in modern theories of elementary particles. After a brief review of the history of this problem, five different approaches to its solution are described.

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As I was going up the stair, I met a man who wasn't there. He wasn't there again today, I wish, I wish he'd stay away.

II. EARLY HISTORY

After completing his formulation of general relativity in 1915-1916, Einstein (1917) attempted to apply his new theory to the whole universe. His guiding principle was that the universe is static: "The most important fact that we draw from experience is that the relative velocities of the stars are very small as compared with the velocity of light." No such static solution of his original equations could be found (any more than for Newtonian gravitation), so he modified them by adding a new term involving a free parameter λ , the cosmological constant:²

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R - \lambda g_{\mu\nu} = -8\pi G T_{\mu\nu} . \qquad (2.1)$$

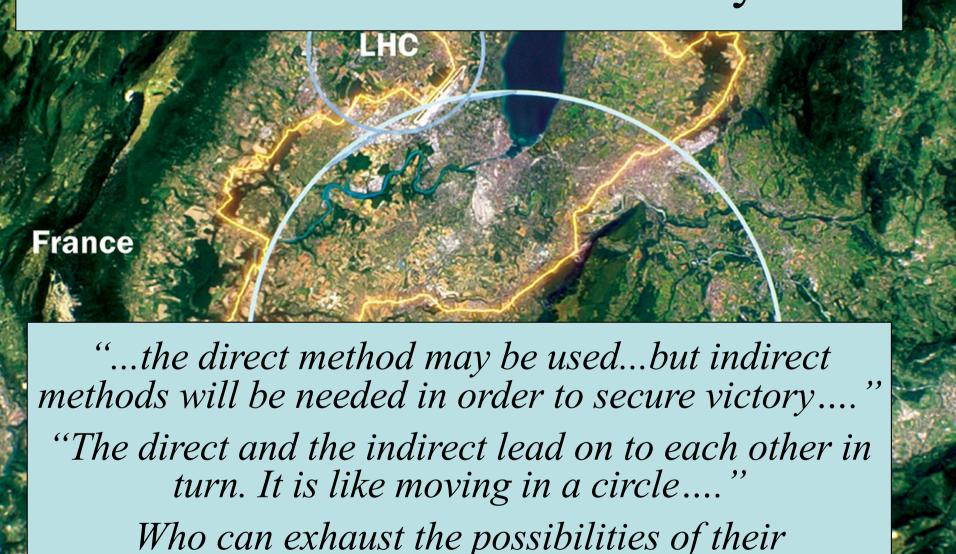
Now, for $\lambda > 0$, there was a static solution for a universe filled with dust of zero pressure and mass density

$$\rho = \frac{\lambda}{8\pi G} \ . \tag{2.2}$$

Its geometry was that of a sphere S_3 , with proper circumference $2\pi r$, where

Hughes Mearns





combination?"

Sun Tzu

Effective Field Theories (EFTs)

a long and glorious History

- 1930's: "Standard Model" of QED had
- Fermi's four-fermion theory of the weak force
- Dimension-6 operators: form = S, P, V, A, T?
 - Due to exchanges of massive particles?
- V-A → massive vector bosons → gauge



- Due to exchanges of mesons? → pions
- Chiral dynamics of pions: $(\partial \pi \partial \pi)\pi\pi$ clue \rightarrow QCD

Standard Model Effective Field Theory: a powerful way to analyze the data

- Assume the Standard Model Lagrangian is correct (quantum numbers of particles) but incomplete
- Look for additional interactions between SM particles due to exchanges of heavier particles
- Analyze Higgs data together with electroweak precision data and top data
- Most efficient way to extract largest amount of information from LHC and other experiments
- Model-independent way to look for physics beyond the Standard Model (BSM)

Summarize Analysis Framework

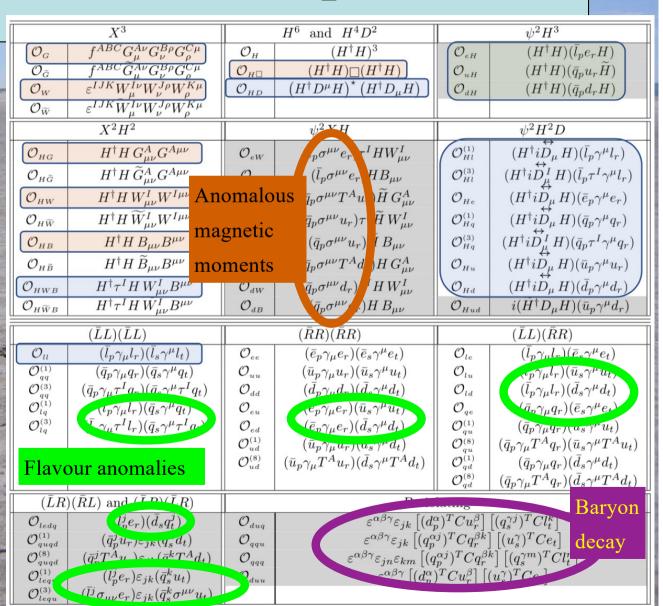
• Include all leading dimension-6 operators?

$$\mathcal{L}_{\mathrm{SMEFT}} = \mathcal{L}_{\mathrm{SM}} + \sum_{i=1}^{2499} \frac{C_i}{\Lambda^2} \mathcal{O}_i$$

- Simplify by assuming flavour SU(3)⁵ or SU(2)² x SU(3)³ symmetry for fermions
- Work to linear order in operator coefficients, i.e. $O(1/\Lambda^2)$
- Use G_F, M_z, α as input parameters

Dimension-6 SMEFT Operators

- Including bosonic,2- and 4-fermion operators
- Different colours for different data sectors
- Grey cells violate SU(3)⁵ symmetry
- Important when including top observables



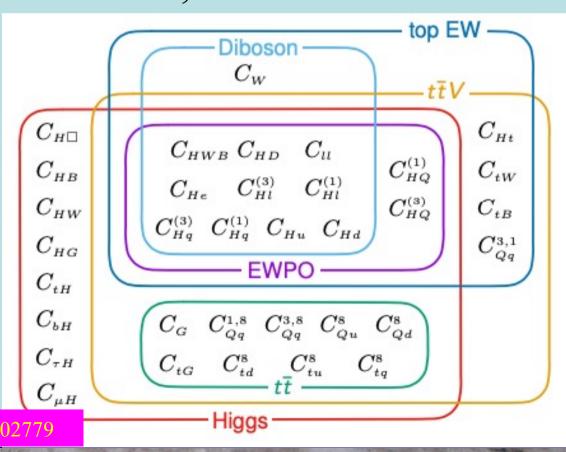
Global SMEFT Fit

to Top, Higgs, Diboson, Electroweak Data

• Global fit to dimension-6 operators using precision electroweak data, W+W- at LEP, top, Higgs and diboson data from LHC Runs 1, 2

Search for BSM

- Constraints on BSM
 - At tree level
 - At loop level



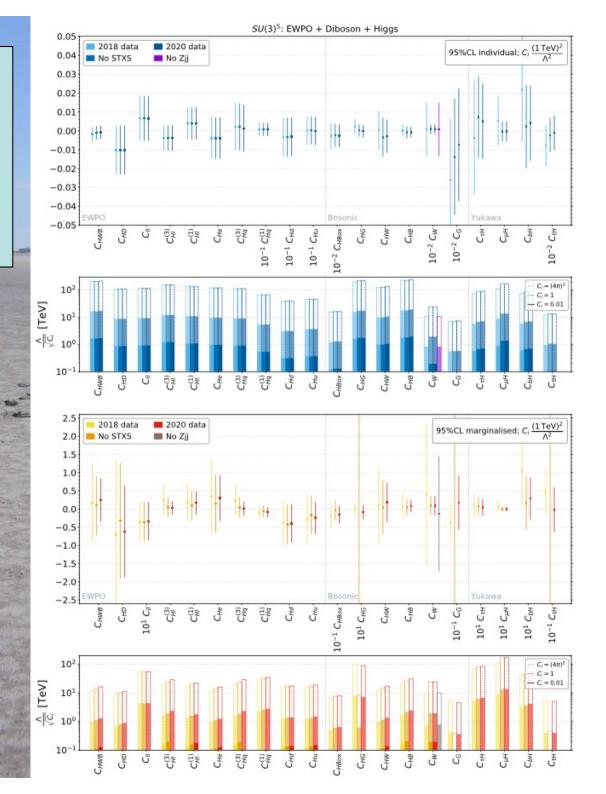
JE, Madigan, Mimasu, Sanz & You, arXiv:2012.02779

Data included in Global Fit

	EW precision observables			D-£			
	Precision electroweak measurem	LHC Run 2 Higgs	Tevat	ron & Run 1 top	$n_{ m obs}$	Ref.	
	Γ_Z , $\sigma_{\rm had.}^0$, R_ℓ^0 , A_{FB}^ℓ , $A_\ell({\rm SLD})$, A	ATLAS combination of		on combination of differential tt forward-backward asymmetry,	4	[7]	
ł	Combination of CDF and D0 W	including ratios of bra					-
	LHC run 1 W boson mass measi	Signal strengths coars	$ATLA$ $d\sigma$			$n_{\mathbf{obs}}$	Ref.
	Erro run i vi boson mass measi	CMS LHC combination	$\overline{dm_{t\bar{t}}}$ ATLA	CMS $t\bar{t}$ differential distributions in the dilepton channel.		6	[36,
	Diboson LEP & LHC	Production: ggF , VB		$\overline{dm_{t\bar{t}}}$			231]
	W^+W^- angular distribution me	Decay: $\gamma \gamma$, ZZ , $W^+ W$	CMS t	CMS $t\bar{t}$ differential distributions in the ℓ +jets channel.		10	[37]
	W^+W^- total cross section meas	CMS stage 1.0 STXS	$\frac{d\sigma}{dm_{t\bar{t}}}$	$\overline{dm_{t\bar{t}}}$ ATLAS measurement of differential $t\bar{t}$ charge asymmetry, $A_C(t)$)	5	loci
	final states for 8 energies	13 parameter fit 7 pa	CMS 1	ATLAS measurement of differential tt charge asymmetry, $A_C(t)$ ATLAS $t\bar{t}W$ & $t\bar{t}Z$ cross section measurements. $\sigma_{t\bar{t}W} \sigma_{t\bar{t}Z}$	$n_{t\bar{t}}$).	2	[38]
	W ⁺ W ⁻ total cross section meas	CMS stage 1.0 STXS	ATLA	CMS $t\bar{t}W$ & $t\bar{t}Z$ cross section measurements. $\sigma_{t\bar{t}W} \sigma_{t\bar{t}Z}$		1 1	[40]
	qqqq final states for 7 energies	CMS stage 1.1 STXS	dilepte	CMS $t\bar{t}V$ differential distributions.		4 4	[41]
	W^+W^- total cross section mean	CMS differential cross	ATLA	$d\sigma = \int d\sigma$		-114	[**]
	& qqqq final states for 8 energies	tion in the $WW^* \to \ell$	$A_C(m)$	$dp_Z^T = d\cos\theta^*$ CMS measurement of differential cross sections and charge rat	ios for t-	5 5	[42]
	ATLAS W+W- differential cre	$\frac{d\sigma}{dn_{jet}}$ $\frac{d\sigma}{dp_H^T}$	$\frac{d\sigma}{dm_{i\bar{i}}dy}$	channel single-top quark production.		-10	[]
	$p_T > 120 \text{ GeV}$ overflow bin	ATLAS $H \to Z\gamma$ sign	ATLA	$\frac{d\sigma}{dp_{t+\bar{t}}^T} \mid R_t \left(p_{t+\bar{t}}^T \right)$			
200	ATLAS W^+W^- fiducial differen	ATLAS $H \rightarrow \mu^{+}\mu^{-}$ si	decay.	CMS measurement of t -channel single-top and anti-top cross se	ections.	4	[43]
	$\frac{d\sigma}{dp_{I}^{T}}$		ATLA	$\sigma_t, \sigma_{\bar{t}}, \sigma_{t+\bar{t}} \& R_t.$. ,
		1	f_0, f_L CMS	CMS measurement of the t -channel single-top and anti-top cross	sections.	1 1 1 1	[44]
	ATLAS $W^{\pm}Z$ fiducial differentia	I cross section in the ℓ^+	f_0, f_L	$\sigma_t \mid \sigma_{\bar{t}} \mid \sigma_{t+\bar{t}} \mid R_t.$			
100	$\frac{dp_Z^T}{dp_Z^T}$		ATLA	CMS t-channel single-top differential distributions.		4 4	[45]
	CMS $W^{\pm}Z$ normalised fiducial d	lifferential cross section	CMS ($\frac{d\sigma}{dp_{t+\bar{t}}^T} \mid \frac{d\sigma}{d y_{t+\bar{t}} }$, [
	channel, $\frac{1}{\sigma} \frac{d\sigma}{dp_Z^T}$		$\frac{d\sigma}{dp_t^T}$	ATLAS tW cross section measurement. 341 m	neasu	remer	its _
	ATLAS Zjj fiducial differential c	ross section in the $\ell^+\ell^-$	$\frac{dp_t^I}{\text{CMS}}$	CMS tZ cross section measurement.			
Œ,			CMS t	CMS tW cross section measurement.	clude	din	
	LHC Run 1 Higgs		$\frac{d\sigma}{dp_{t+\bar{t}}^T}$		Clude	tu III	
	ATLAS and CMS LHC Run 1 co	mbination of Higgs sign	CMS 1	CMS $tZ(Z \to \ell^+\ell^-)$ cross section measurement			_
	Production: ggF, VBF, ZH, WH & ttH			$\sigma_t \mid \sigma_{\bar{t}} \mid \sigma_{t+\bar{t}} \mid R_t$. ATLAS s-channel single-top cross section measurement.			
7	December 77 W+W+ V- bb			S s-channel single-top cross section measurement. W cross section measurement.		1001	
	ATLAS inclusive $Z\gamma$ signal stren	gth measurement		S tW cross section measurement in the single lepton channel	1	[34]	Sec.
-	Marie Committee of the			S tW cross section measuremen JE, Madigan, Mimasu, Sanz	& You, a	rXiv:2012.02	2779
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Dimension-6 Constraints with Flavour-Universal SU(3)⁵ Symmetry

- Individual operator coefficients
- Marginalised over all other operator coefficients



JE, Madigan, Mimasu, Sanz & You, arXiv:2012.02779

Dimension-6 Constraints with Top-Specific $SU(2)^2 \times SU(3)^3$

0.02

-0.01 -0.02

-0.03 -0.04

^ [TeV]

- Individual operator coefficients
- Marginalised over all other operator coefficients

JE, Madigan, Mimasu, Sanz & You

95%CL individual; C: (1TeV)2

PARTICLE PHYSICS

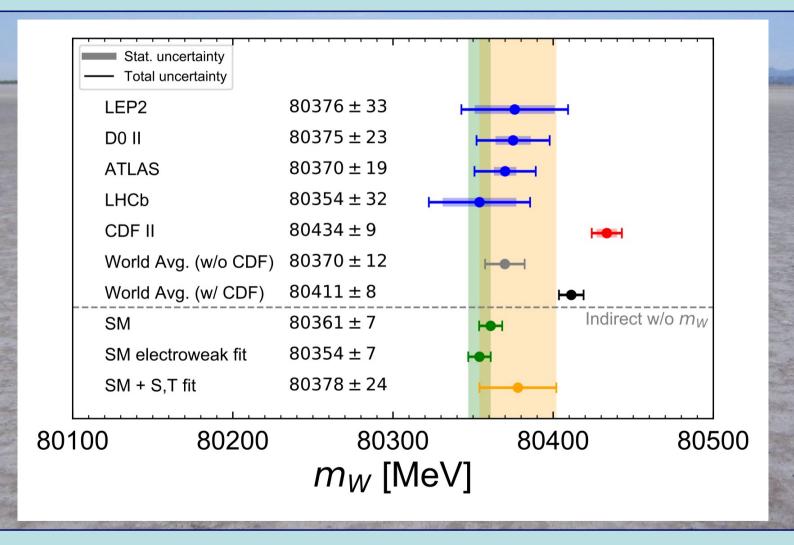
High-precision measurement of the W boson mass with the CDF II detector

CDF Collaboration, Science 376 (2022) p170

CDF Collaboration + ±, T. Aaltonen 1.2, S. Amerio 3.4, D. Amidei 5, A. Anastassov 6, A. Annovi 7, J. Antos 8.9, G. Apollinari 6, J. A. Appel 6, T. Arisawa 10, A. Artikov 11, J. Asaadi¹², W. Ashmanskas⁶, B. Auerbach¹³, A. Aurisano¹², F. Azfar¹⁴, W. Badgett⁶, T. Bae^{15,16,17,18,19,20,21}, A. Barbaro-Galtieri²², V. E. Barnes²³, B. A. Barnett²⁴, P. Barria^{25,26}, P. Bartos^{8,9}, M. Bauce^{3,4}, F. Bedeschi²⁵, S. Behari⁶, G. Bellettini^{25,27}, J. Bellinger²⁸, D. Benjamin²⁹, A. Beretvas⁶, A. Bhatti³⁰, K. R. Bland³¹, B. Blumenfeld²⁴, A. Bocci²⁹, A. Bodek³², D. Bortoletto²³, J. Boudreau³³, A. Boveia³⁴, L. Brigliadori^{35,36}, C. Bromberg³⁷, E. Brucken¹², J. Budagov¹¹8, H. S. Budd³², K. Burkett⁶, G. Busetto^{3,4}, P. Bussey³⁸, P. Butti^{25,27}, A. Buzatu³⁸, A. Calamba³⁹, S. Camarda⁴⁰, M. Campanelli⁴¹, B. Carls⁴², D. Carlsmith²⁸, R. Carosi²⁵, S. Carrillo⁴³§, B. Casal⁴⁴, M. Casarsa⁴⁵, A. Castro^{35,36}, P. Catastini⁴⁶, D. Cauz^{45,47,48}, V. Cavaliere⁴², A. Cerri²², L. Cerrito⁴¹, Y. C. Chen⁴⁹, M. Chertok⁵⁰, G. Chiarelli²⁵, G. Chlachidze⁶, K. Cho^{15,16,17,18,19,20,21}, D. Chokheli¹¹, A. Clark⁵¹, C. Clarke⁵², M. E. Convery⁶, J. Conway⁵⁰, M. Corbo⁶, M. Cordelli⁷, C. A. Cox⁵⁰, D. J. Cox⁵⁰, M. Cremonesi²⁵, D. Cruz¹², J. Cuevas⁴⁴, R. Culbertson⁶, N. d'Ascenzo⁶, M. Datta⁶, P. de Barbaro³², L. Demortier³⁰, M. Deninno³⁵§, M. D'Errico^{3,4}, F. Devoto^{1,2}, A. Di Canto^{25,27}, B. Di Ruzza⁶, J. R. Dittmann³¹, S. Donati^{25,27}, M. D'Onofrio⁵³, M. Dorigo^{45,54}, A. Driutti^{45,47,48}, K. Ebina¹⁰, R. Edgar⁵, A. Elagin³⁴, R. Erbacher⁵⁰, S. Errede⁴², B. Esham⁴², S. Farrington¹⁴, J. P. Fernández Ramos⁵⁵, R. Field⁴³, G. Flanagan⁶, R. Forrest⁵⁰, M. Franklin⁴⁶, J. C. Freeman⁶, H. Frisch³⁴, Y. Funakoshi¹⁰, C. Galloni^{25,27}, A. F. Garfinkel²³, P. Garosi^{25,26}, H. Gerberich⁴², E. Gerchtein⁶, S. Giagu⁵⁶, V. Giakoumopoulou⁵⁷, K. Gibson³³, C. M. Ginsburg⁶, N. Giokaris⁵⁷§, P. Giromini⁷, V. Glagoley¹¹, D. Glenzinski⁶, M. Gold⁵⁸, D. Goldin¹², A. Golossanov⁶, G. Gomez⁴⁴, G. Gomez-Ceballos⁵⁹, M. Goncharov⁵⁹, O. González López⁵⁵, I. Gorelov⁵⁸, A. T. Goshaw²⁹, K. Goulianos³⁰, E. Gramellini³⁵, C. Grosso-Pilcher³⁴, J. Guimaraes da Costa⁴⁶, S. R. Hahn⁶, J. Y. Han³², F. Happacher⁷, K. Hara⁶⁰, M. Hare⁶¹, R. F. Harr⁵², T. Harrington-Taber⁶, K. Hatakevama³¹, C. Havs¹⁴, J. Heinrich⁶², M. Herndon²⁸, A. Hocker⁶, Z. Hong¹², W. Hopkins⁶, S. Hou⁴⁹, R. E. Hughes⁶³, U. Husemann⁶⁴, M. Hussein³⁷, J. Huston³⁷, G. Introzzi^{25,65,66}, M. Iori^{56,67}, A. Ivanov⁵⁰, E. James⁶, D. Jang³⁹, B. Javatilaka⁶, E. J. Jeon^{15,16,17,18,19,20,21}, S. Jindariani⁶, M. Jones²³, K. K. Joo^{15,16,17,18,19,20,21}, S. Y. Jun³⁹, T. R. Junk⁶, M. Kambeitz⁶⁸, T. Kamon^{15,16,17,18,19,20,21,12}, P. E. Karchin⁵², A. Kasmi³¹, Y. Kato⁶⁹, W. Ketchum³⁴, J. Keung⁶², B. Kilminster⁶, D. H. Kim^{15,16,17,18,19,20,21}, H. S. Kim⁶, J. E. Kim^{15,16,17,18,19,20,21}, M. J. Kim⁷, S. H. Kim⁶⁰, S. B. Kim^{15,16,17,18,19,20,21}, Y. J. Kim^{15,16,17,18,19,20,21}, Y. K. Kim³⁴, N. Kimura¹⁰, M. Kirby⁶, K. Kondo¹⁰§, D. J. Kong^{15,16,17,18,19,20,21}, J. Konigsberg⁴³, A. V. Kotwal²⁹*, M. Kreps⁶⁸, J. Kroll⁶², M. Kruse²⁹, T. Kuhr⁶⁸, M. Kurata⁶⁰, A. T. Laasanen²³, S. Lammel⁶, M. Lancaster⁴¹, K. Lannon⁶³, G. Latino^{25,26}, H. S. Lee^{15,16,17,18,19,20,21}. J. S. Lee 15,16,17,18,19,20,21, S. Leo 42, S. Leone 25, J. D. Lewis 6, A. Limosani 29, E. Lipeles 62, A. Lister 51, O. Liu 23, T. Liu 6, S. Lockwitz 64, A. Loginov 64 8. D. Lucchesi^{3,4}, A. Luca^{7,6}, J. Lueck⁶⁸, P. Lujan²², P. Lukens⁶, G. Lungu³⁰, J. Lys²²§, R. Lysak^{8,9}, R. Madrak⁶, P. Maestro^{25,26}, S. Malik³⁰, G. Manca⁵³, A. Manousakis-Katsikakis⁵⁷, L. Marchese³⁵, F. Margaroli⁵⁶, P. Marino^{25,70}, K. Matera⁴², M. E. Mattson⁵², A. Mazzacane⁶, P. Mazzanti³⁵, R. McNulty⁵³, A. Mehta⁵³, P. Mehtala^{1,2}, A. Menzione²⁵§, C. Mesropian³⁰, T. Miao⁶, E. Michielin^{3,4}, D. Mietlicki⁵, A. Mitra⁴⁹, H. Miyake⁶⁰, S. Moed⁶, N. Moggi³⁵, C. S. Moon^{15,16,17,18,19,20,21}, R. Moore⁶, M. J. Morello^{25,70}, A. Mukherjee⁶, Th. Muller⁶⁸, P. Murat⁶, M. Mussini^{35,36}, J. Nachtman⁶, Y. Nagai⁶⁰, J. Naganoma¹⁰, I. Nakano⁷¹, A. Napier⁶¹, J. Nett¹², T. Nigmanov³³, L. Nodulman¹³, S. Y. Noh^{15,16,17,18,19,20,21}, O. Norniella⁴², L. Oakes¹⁴, S. H. Oh²⁹, Y. D. Oh^{15,16,17,18,19,20,21}, T. Okusawa⁶⁹, R. Orava^{1,2}, L. Ortolan⁴⁰, C. Pagliarone⁴⁵, E. Palencia⁴⁴, P. Palni⁵⁸, V. Papadimitriou⁶, W. Parker²⁸, G. Pauletta^{45,47,48}, M. Paulini³⁹, C. Paus⁵⁹, T. J. Phillips²⁹, G. Piacentino⁶, E. Pianori⁶², J. Pilot⁵⁰, K. Pitts⁴², C. Plager⁷², L. Pondrom²⁸, S. Poprocki⁶, K. Potamianos²², A. Pranko²², F. Prokoshin¹¹, F. Ptohos⁷, G. Punzi^{25,27}, I. Redondo Fernández⁵⁵, P. Renton¹⁴, M. Rescigno⁵⁶, F. Rimondi³⁵§, L. Ristori^{25,6}, A. Robson³⁸, T. Rodriguez⁶², S. Rolli⁶¹, M. Ronzani^{25,27}, R. Roser⁶, J. L. Rosner³⁴, F. Ruffini^{25,26}, A. Ruiz⁴⁴, J. Russ³⁹, V. Rusu⁶, W. K. Sakumoto³², Y. Sakurai¹⁰, L. Santi^{45,47,48}, K. Sato⁶⁰, V. Saveliev⁶, A. Savoy-Navarro⁶, P. Schlabach⁶, E. E. Schmidt⁶, T. Schwarz⁵, L. Scodellaro⁴⁴, F. Scuri²⁵, S. Seidel⁵⁸, Y. Seiya⁶⁹, A. Semenov¹¹, F. Sforza^{25,27}, S. Z. Shalhout⁵⁰, T. Shears⁵³, P. F. Shepard³³, M. Shimojima⁶⁰, M. Shochet³⁴, I. Shreyber-Tecker⁷³, A. Simonenko¹¹, K. Sliwa⁶¹, J. R. Smith⁵⁰, F. D. Snider⁶, H. Song³³, V. Sorin⁴⁰, R. St. Denis³⁸S, M. Stancari⁶, D. Stentz⁶, J. Strologas⁵⁸, Y. Sudo⁶⁰, A. Sukhanov⁶, I. Suslov¹¹, K. Takemasa⁶⁰, Y. Takeuchi⁶⁰, J. Tang³⁴, M. Tecchio⁵, P. K. Teng⁴⁹, J. Thom⁶, E. Thomson⁶², V. Thukral¹², D. Toback¹², S. Tokar^{8,9}, K. Tollefson³⁷, T. Tomura⁶⁰, S. Torre⁷, D. Torretta⁶, P. Totaro³, M. Trovato^{25,70}, F. Ukegawa⁶⁰, S. Uozumi^{15,16,17,18,19,20,21}, F. Vázquez⁴³, G. Velev⁶, K. Vellidis⁵⁷, C. Vernieri^{25,70}, M. Vidal²³, R. Vilar⁴⁴, J. Vizán⁴⁴, M. Vogel⁵⁸, G. Volpi⁷, P. Wagner⁶², R. Wallny⁶, S. M. Wang⁴⁹, D. Waters⁴¹, W. C. Wester III⁶, D. Whiteson⁶², A. B. Wicklund¹³, S. Wilbur⁵⁰, H. H. Williams⁶², J. S. Wilson⁵, P. Wilson⁶, B. L. Winer⁶³, P. Wittich⁶, S. Wolbers⁶, H. Wolfmeister⁶³, T. Wright⁵, X. Wu⁵¹, Z. Wu³¹, K. Yamamoto⁶⁹, D. Yamato⁶⁹, T. Yang⁶, U. K. Yang^{15,16,17,18,19,20,21}, Y. C. Yang^{15,16,17,18,19,20,21}, W.-M. Yao²², G. P. Yeh⁶, K. Yi⁶, J. Yoh⁶, K. Yorita¹⁰, T. Yoshida⁶⁹, G. B. Yu^{15,16,17,18,19,20,21}, I. Yu^{15,16,17,18,19,20,21}, A. M. Zanetti⁴⁵, Y. Zeng²⁹, C. Zhou²⁹, S. Zucchelli^{35,36}

CDF Measurement of mw

compared with previous measurements



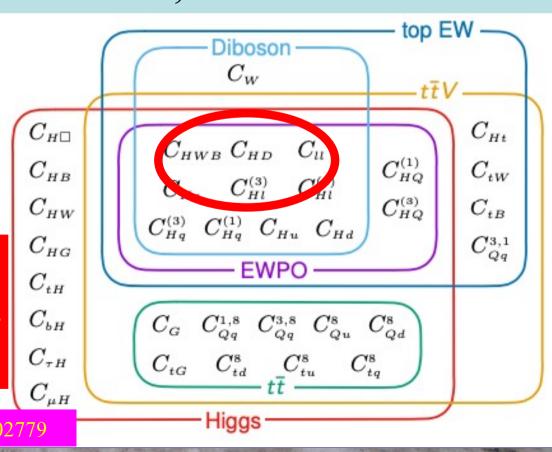
• Tension: $7-\sigma$ discrepancy with Standard Model?

Global SMEFT Fit

to m_W, Top, Higgs, Diboson, Electroweak Data

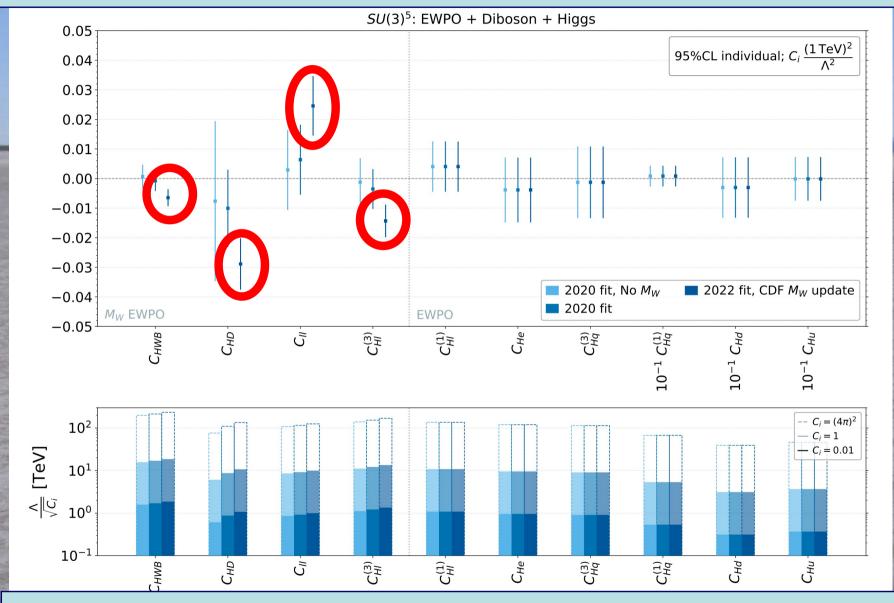
- Global fit to dimension-6 operators using precision electroweak data, W+W- at LEP, top, Higgs and diboson data from LHC Runs 1, 2
- Search for BSM
- Constraints on BSM
 - At tree level
 - At loop level

Positive contributions to mw



JE, Madigan, Mimasu, Sanz & You, arXiv:2012.02779

SMEFT Fit with the Mass of the W Boson



• Non-zero coefficients for any of four operators can fit W mass

Single-Field Extensions of the Standard Model

Namo	Spin	SU(3)	SU(2)	U(1)	Name	Spin	SU(3)	SU(2)	U(1)
S	0	1	1	0	Δ_1	$\frac{1}{2}$	1	2	$-\frac{1}{2}$
S_1	0	1	1	1	Δ_3	$\frac{1}{2}$	1	2	$-\frac{1}{2}$
φ	0	Spin ze	ero 2	$\frac{1}{2}$	Σ	$\frac{1}{2}$	1	3	0
Ξ	0	1	3	0	Σ_1	$\frac{1}{2}$	1	3	-1
-1	0	1	3	1	U	$\frac{1}{2}$	3	1	$\frac{2}{3}$
P	1	1	1	0	D	$\frac{1}{2}$	3	1	$-\frac{1}{3}$
B_1	1	Voctor	1	1	Q_1	$\frac{1}{2}$	3	2	$\frac{1}{6}$
W	1	Vector	3	0	Q_5	$\frac{1}{2}$	3	2	$-\frac{5}{6}$
W_1	1	1	3	1	Q_7	$\frac{1}{2}$	3	2	$\frac{7}{6}$
M	$\frac{1}{2}$	1	1	0	T_1	$\frac{1}{2}$	3	3	$-\frac{1}{3}$
E	$\frac{1}{2}$	1	1	-1	T_2	$\frac{1}{2}$	3	3	$\begin{array}{c c} \frac{2}{3} \\ \hline \frac{1}{6} \end{array}$
T	$\frac{1}{2}$	3	1	$\frac{2}{3}$	TB	$\frac{1}{2}$	3	2	$\frac{1}{6}$

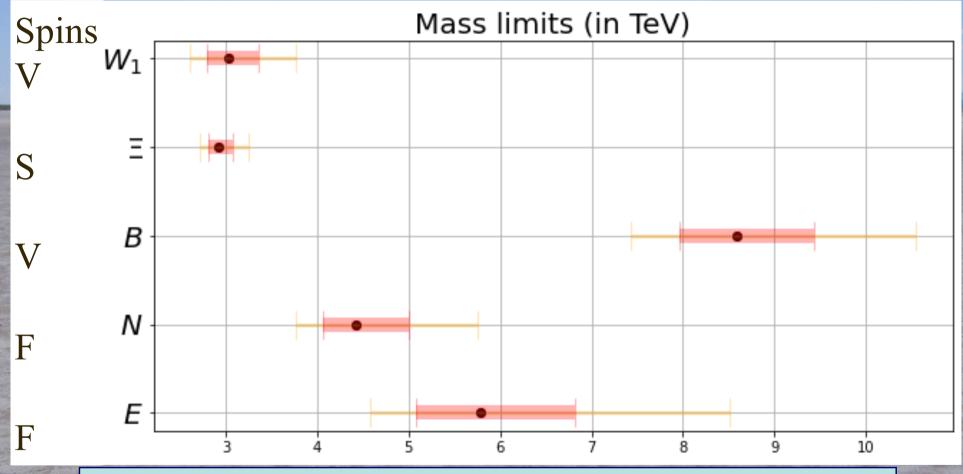
Single-Field Models that can Contribute to W Mass

Model	C_{HD}	C_{ll}	$C_{\mathrm{H}^{\prime}}^{(3)}$	$C_{Hl}^{(1)}$	C_{He}	$C_{H\square}$	$C_{ au H}$	C_{tH}	C_{bH}
S_1		X							
Σ	Wrong	sion	*	$\frac{3}{16}$			$\frac{y_{ au}}{4}$		^
Σ_1	Wrong		$\frac{1}{16}$	$-\frac{3}{16}$			$\frac{y_{ au}}{8}$		
N			$-\frac{1}{4}$	$\frac{1}{4}$					
E			$-\frac{1}{4}$	$-\frac{1}{4}$			$\frac{y_{ au}}{2}$,
B_1	X	D . 1				$-\frac{1}{2}$	$-\frac{y_{ au}}{2}$	$-\frac{y_t}{2}$	$-\frac{y_b}{2}$
B	-2	Kıgl	nt sign				$-y_{ au}$	$-y_t$	$-y_b$
Ξ	-2					$\frac{1}{2}$	$y_{ au}$	y_t	y_b
W_1	$-\frac{1}{4}$					$-\frac{1}{8}$	$-\frac{y_{\tau}}{8}$	$-\frac{y_t}{8}$	$-\frac{y_b}{8}$
W	X					$-\frac{1}{2}$	$-y_{ au}$	$-y_t$	$-y_b$

Bagnaschi, JE, Madigan, Mimasu, Sanz & You, arXiv:220

Operators contributing to mw

Models Fitting the Mass of the W Boson



- 68 and 95% CL ranges of masses assuming unit coupling
- Masses proportional to couplings
- Large masses consistent with SMEFT approximation

Models Fitting W Mass

	Model	Pull	Best-fit mass	1 - σ mass	2 - σ mass	$1-\sigma \text{ coupling}^2$
Spins			(TeV)	range (TeV)	range (TeV)	range
V	W_1	6.4	3.0	[2.8, 3.6]	[2.6, 3.8]	[0.09, 0.13]
S	B	6.4	8.6	[8.0, 9.4]	[7.4, 10.6]	[0.011, 0.016]
V	Ξ	6.4	2.9	[2.8, 3.1]	[2.7, 3.2]	[0.011, 0.016]
F	N	5.1	4.4	[4.1, 5.0]	[3.8, 5.8]	[0.040, 0.060]
F	E	3.5	5.8	[5.1, 6.8]	[4.6, 8.5]	[0.022, 0.039]

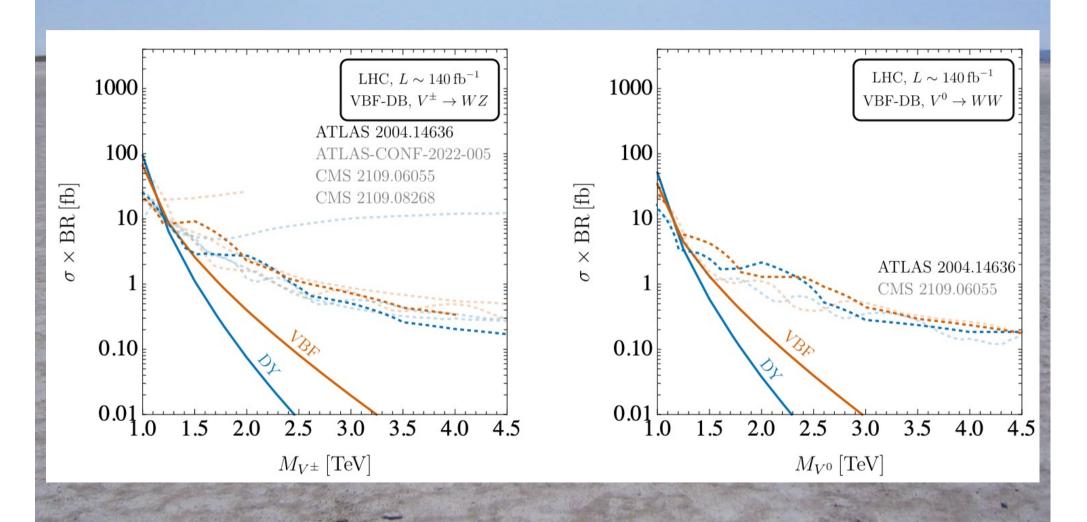
Best-fit, 68 and 95% CL ranges of masses assuming unit couplings

68% CL ranges of couplings for 1 TeV

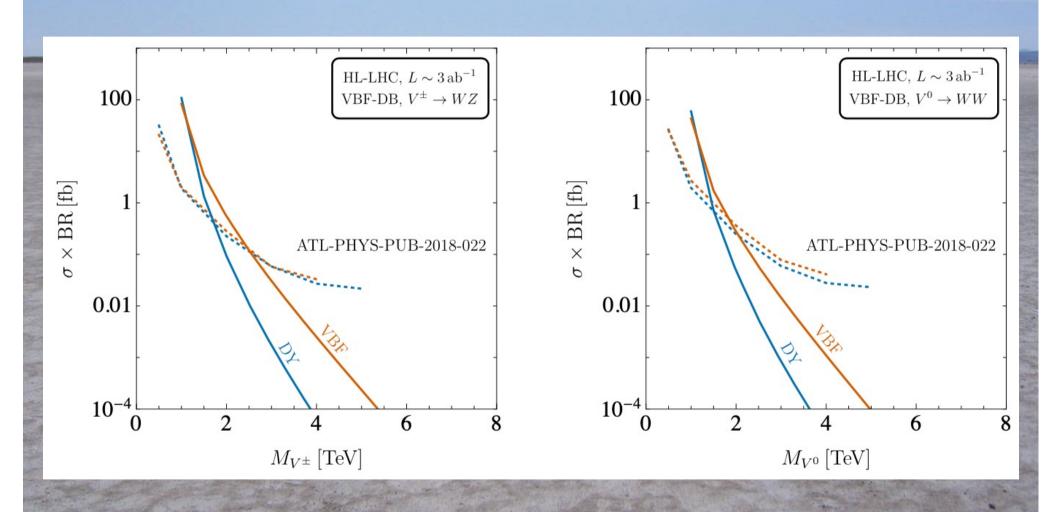
Searching for Models Fitting the Mass of the W Boson

- W: Isotriplet vector boson, mass ~ 3 TeV x coupling, electroweak production, accessible at LHC?
- B: Singlet vector boson, mass ~ 8 TeV x coupling, phenomenology depends on fermion couplings, too heavy for LHC?
- E: Isotriplet scalar boson, mass ~ 3 TeV x coupling, detectable in LHC searches for heavy Higgs bosons?
- N: Isosinglet neutral fermion, mass \sim 4 TeV x coupling, similar to (right-handed) singlet neutrino
- E: Isosinglet charged fermion, mass ~ 6 TeV x coupling, similar to (right-handed) singlet electron

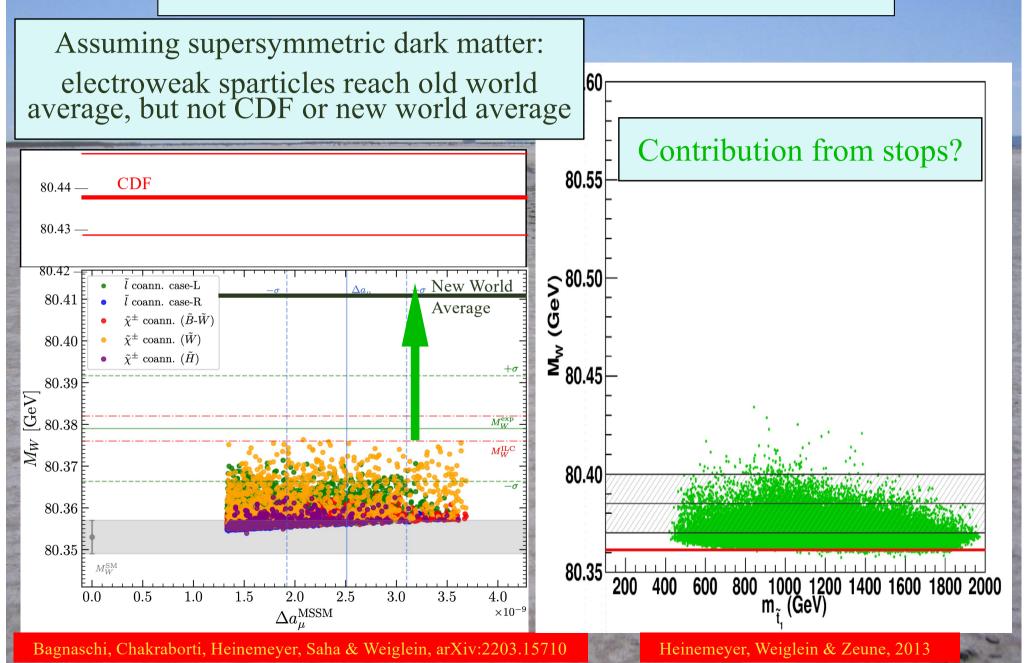
LHC Searches for Triplet Vector Boson



LHC Searches for Triplet Vector Boson



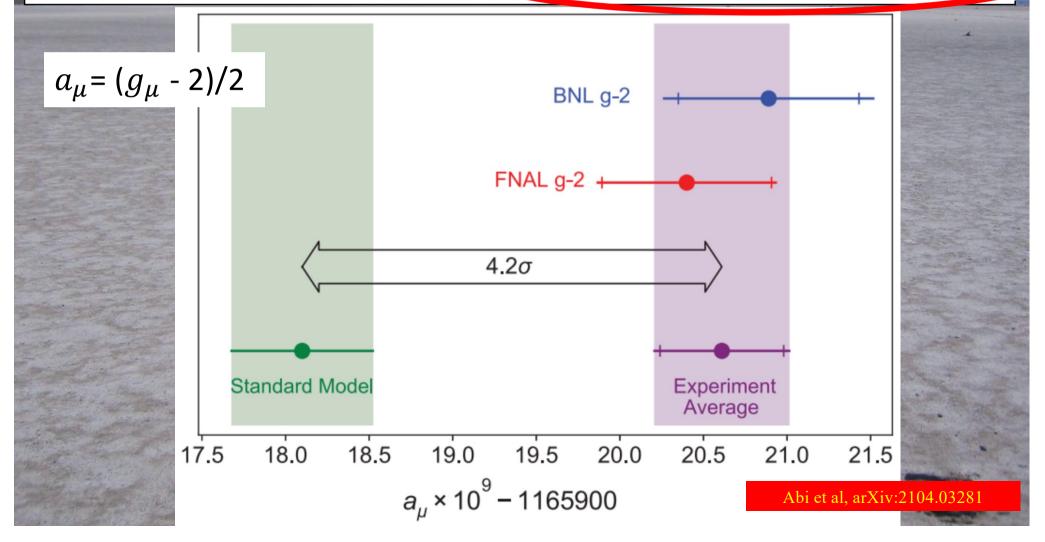
W Mass in Supersymmetry?





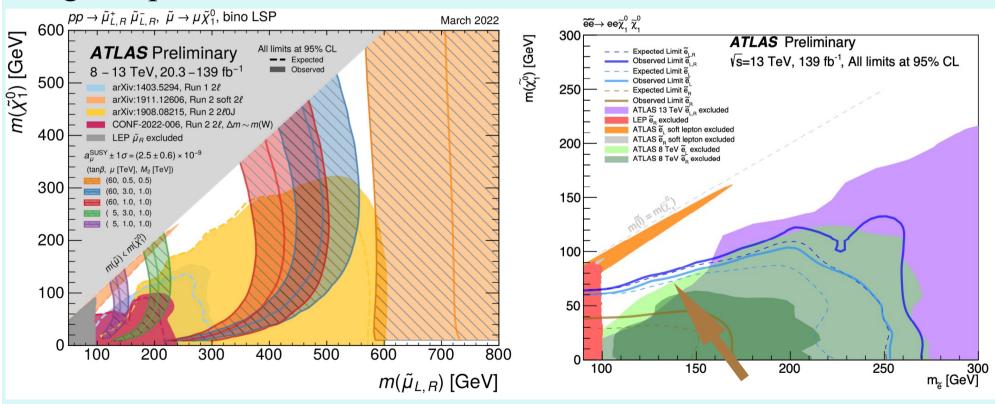
Fermilab Measurement of $g_{\mu} - 2$

```
FNAL result a_{\mu}(\text{FNAL}) = 116\,592\,040(54) \times 10^{-11} (0.46 ppm)
Combined result: a_{\mu}(\text{Exp}) = 116\,592\,061(41) \times 10^{-11} (0.35 ppm)
Difference from Standard Model a_{\mu}(\text{Exp}) - a_{\mu}(\text{SM}) = (251 \pm 59) \times 10^{-11}
```



LHC vs Supersymmetry

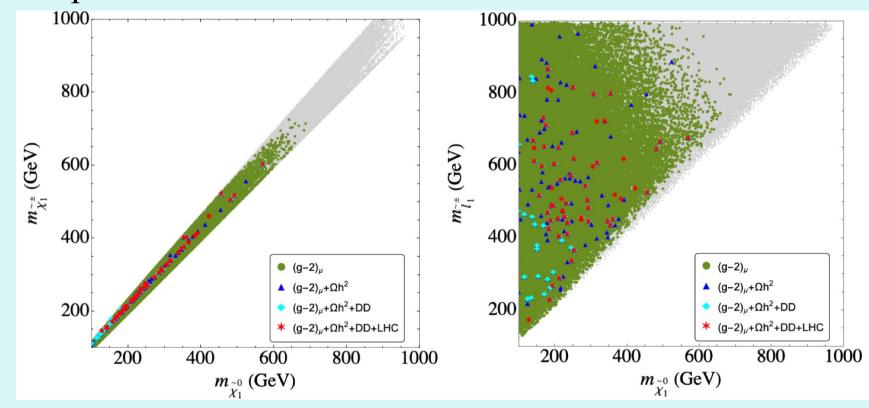
- LHC favours squarks & gluinos > 2 TeV (but loopholes)
- Does not exclude lighter electroweakly-interacting particles, e.g., sleptons



• Most models have $m_{\tilde{\mu}_L} > m_{\tilde{\mu}_R}$ but $m_{\tilde{\mu}_R} \simeq m_{\tilde{e}_R}$: relevant constraint

Supersymmetry

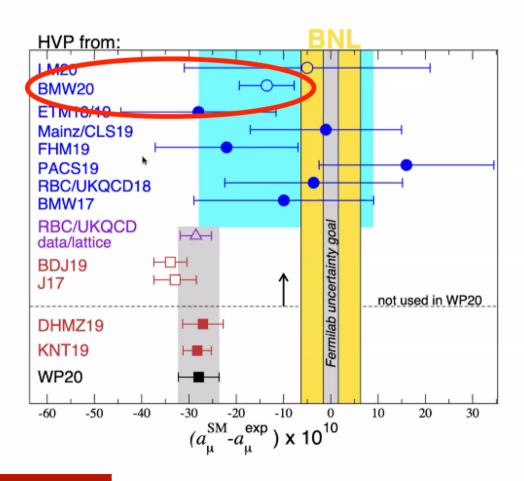
• g_{μ} – 2-friendly scenario with light neutralino, chargino & slepton



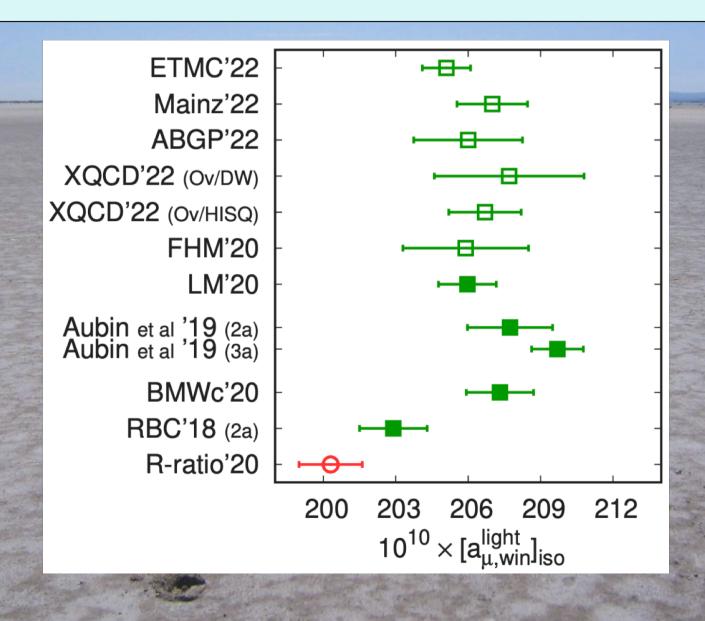
• Red star points include all relevant LHC, dark matter density and direct scattering constraints

Comparison with Lattice Calculations of Hadronic Vacuum Polarization

$$\left[a_{\mu}^{ ext{HVP}} + \left[a_{\mu}^{ ext{QED}} + a_{\mu}^{ ext{Weak}} + a_{\mu}^{ ext{HLbL}}
ight] > \ a_{\mu}^{ ext{SM}}$$



Update on Lattice Calculations



General Interest in Antimatter Physics

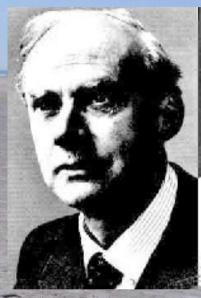


Physicists cannot make enough for Star Trek or Dan Brown!

How do Matter and Antimatter Differ?

Dirac predicted the existence of antimatter:
same mass
opposite internal properties:
electric charge, ...
Discovered in cosmic rays
Studied using accelerators

Used in PET scanners





Matter and antimatter not quite equal and opposite: WHY?

Why does the Universe mainly contain matter, not antimatter?

Experiments at LHC and elsewhere looking for answers

How to Create the Matter in the Universe?

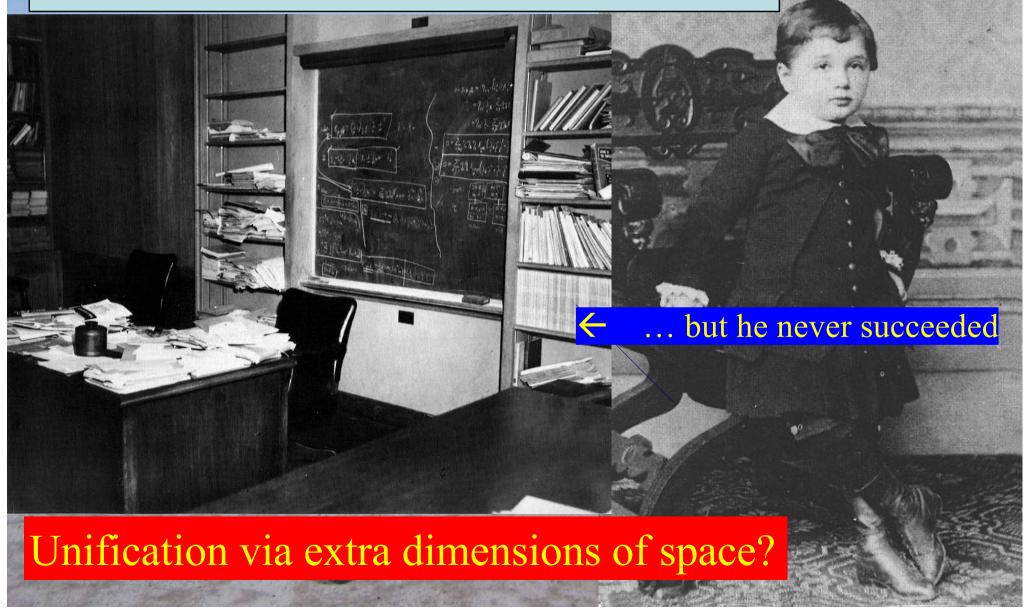
Sakharov

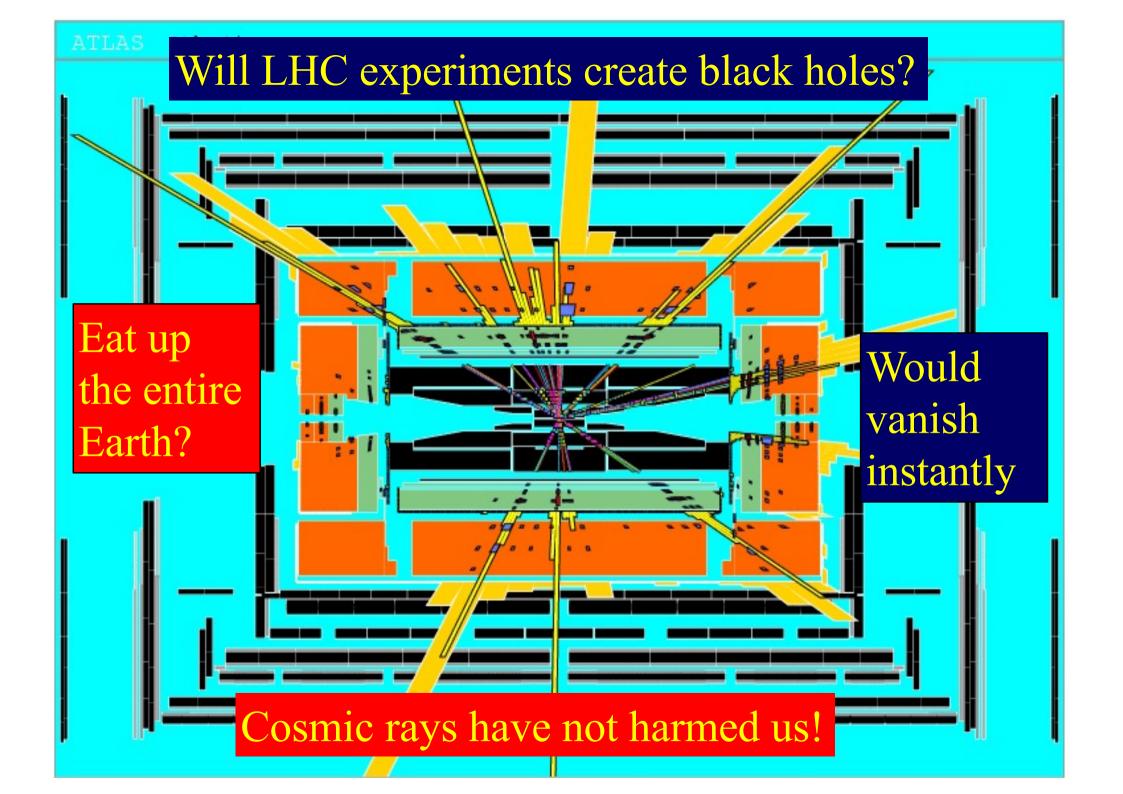
- Need a difference between matter and antimatter observed in the laboratory
- Need interactions able to create matter predicted by theories not yet seen by experiment
- Need the expansion of the Universe a role for the Higgs boson?



Will we be able to calculate using laboratory data?







Standard Model Particles:

Years from Proposal to Discovery

Electron

Photon

Muon

Electron neutrino

Muon neutrino

Down

Strange

Up

Charm

Tau

Bottom

Gluon

W boson

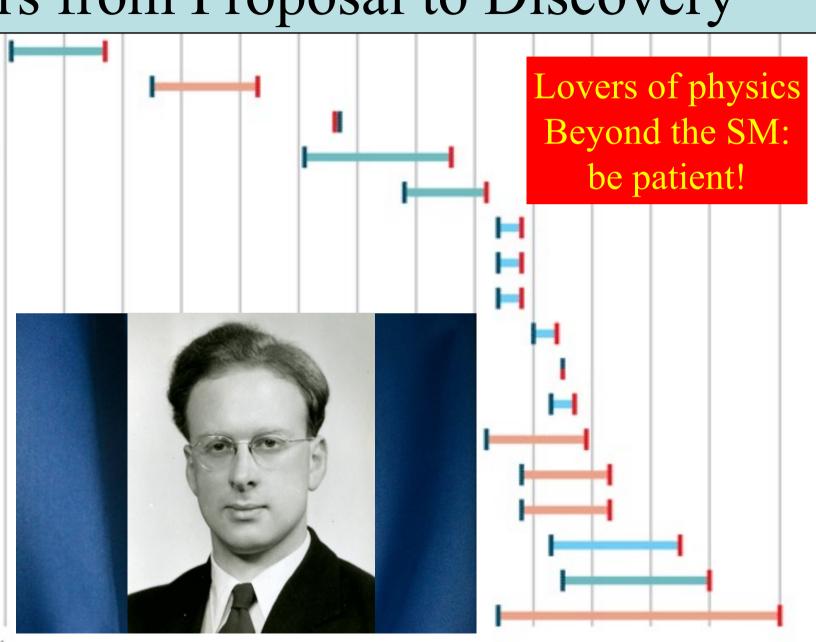
Z boson

Top

Tau neutrino

HIGGS BOSON

Source: The Economist



Higgstorical Summary

- Speculation
- Hypothesis
- Theory
- Search
- Discovery
- Building-block

