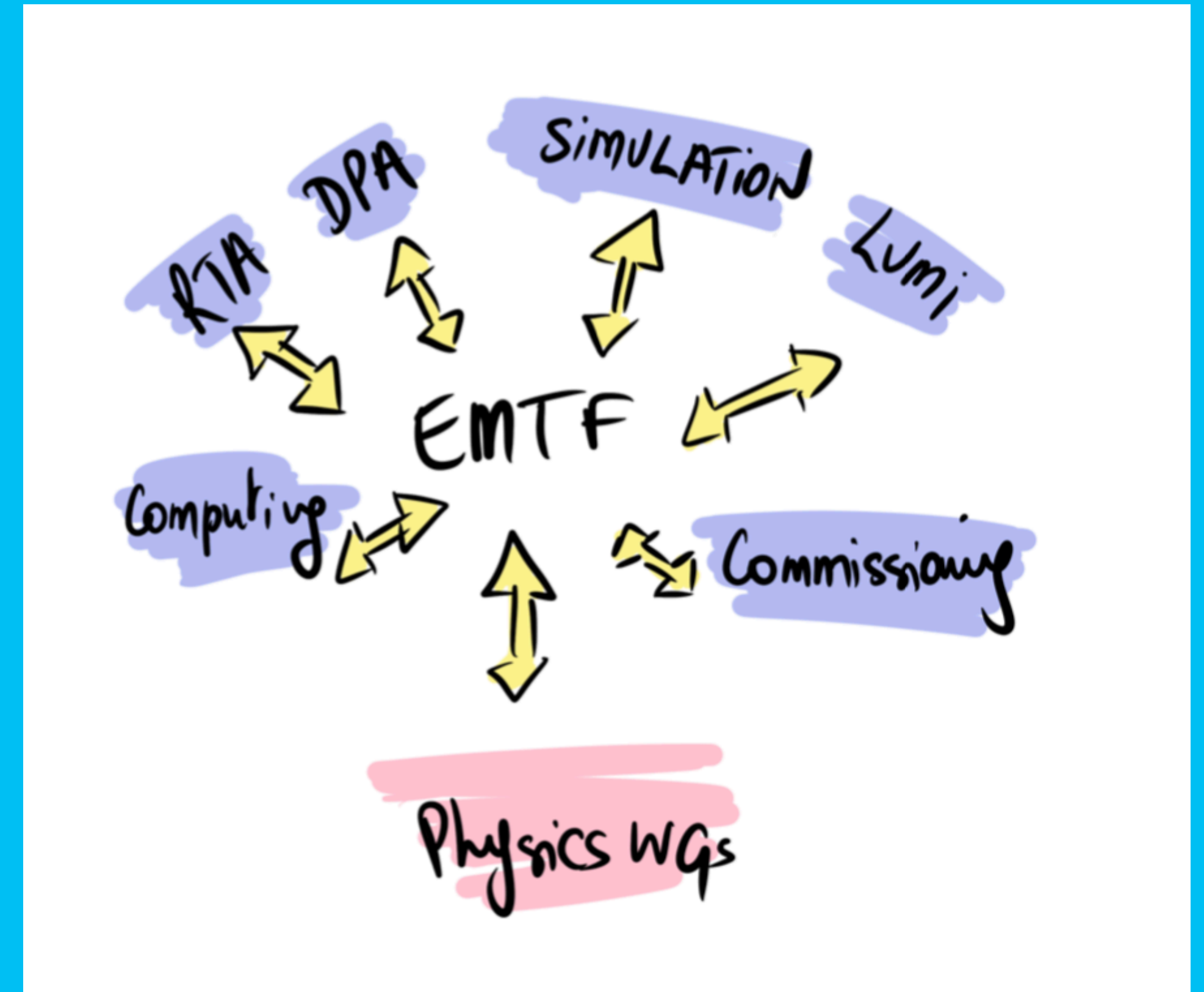


14.03.2022

EMTF 2022

Starter kit edition

Eva Gersabeck
Yasmine Amhis
Alessandro Scarabotto
Valeriia Lukashenko
Lukas Calefice



lhcb-physics-emptf-run3@cern.ch

THE EMTF CORE TEAM



Lukas/QEE



Dylan/Charm



Federica/Charm



Fidan/B2OC



Fionn/B2OC



Holly/SL



Valeriia/B2CC



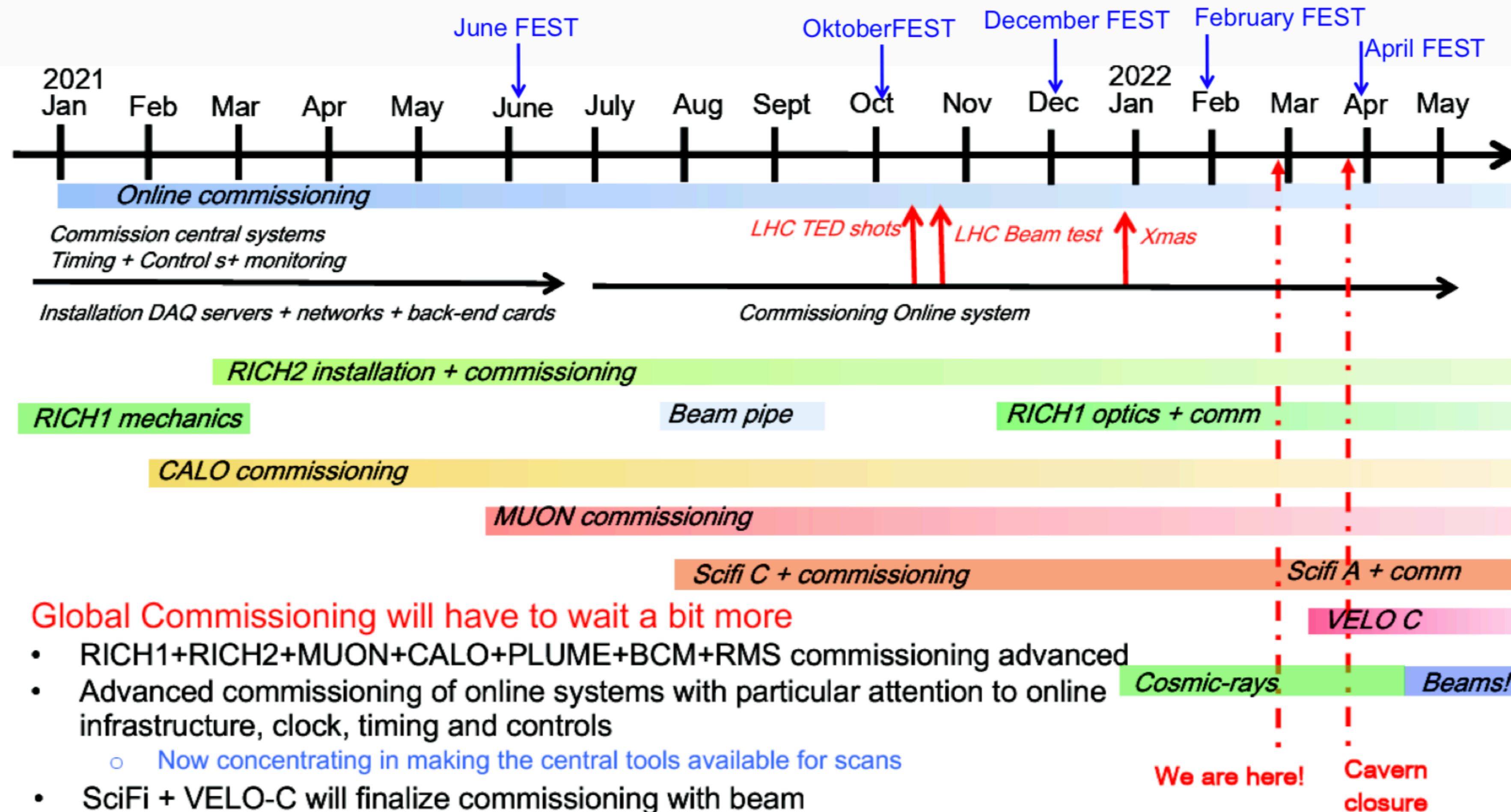
Felicia/RD



Felicia/RD

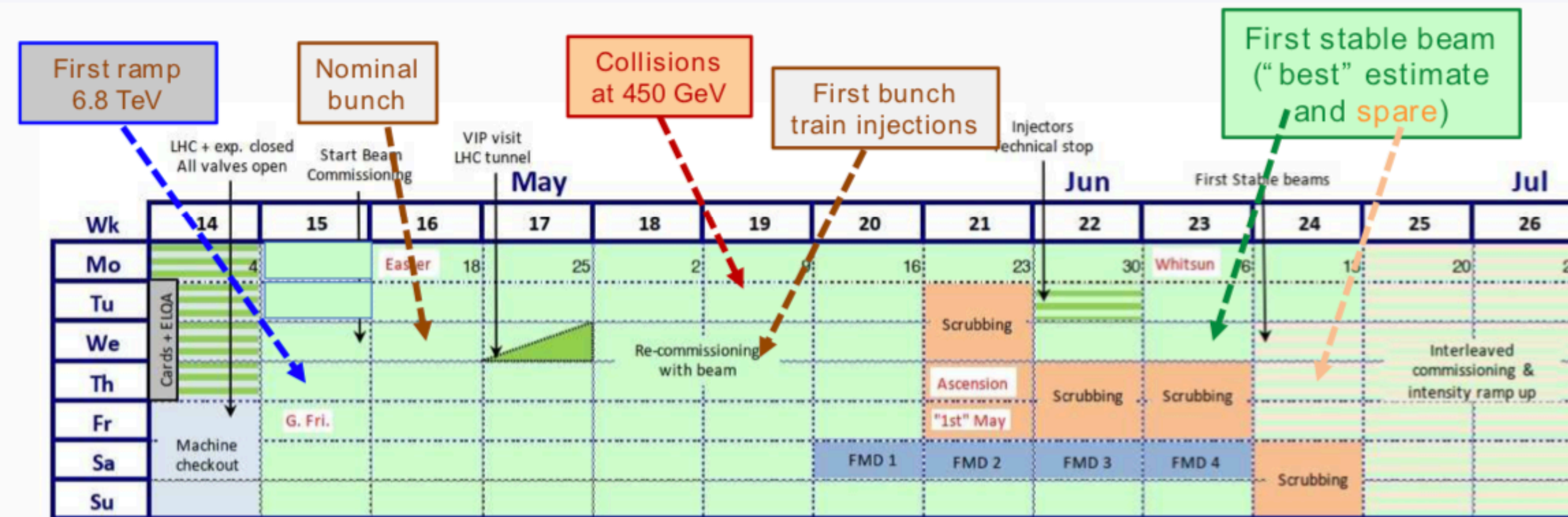


Jialu/BandQ



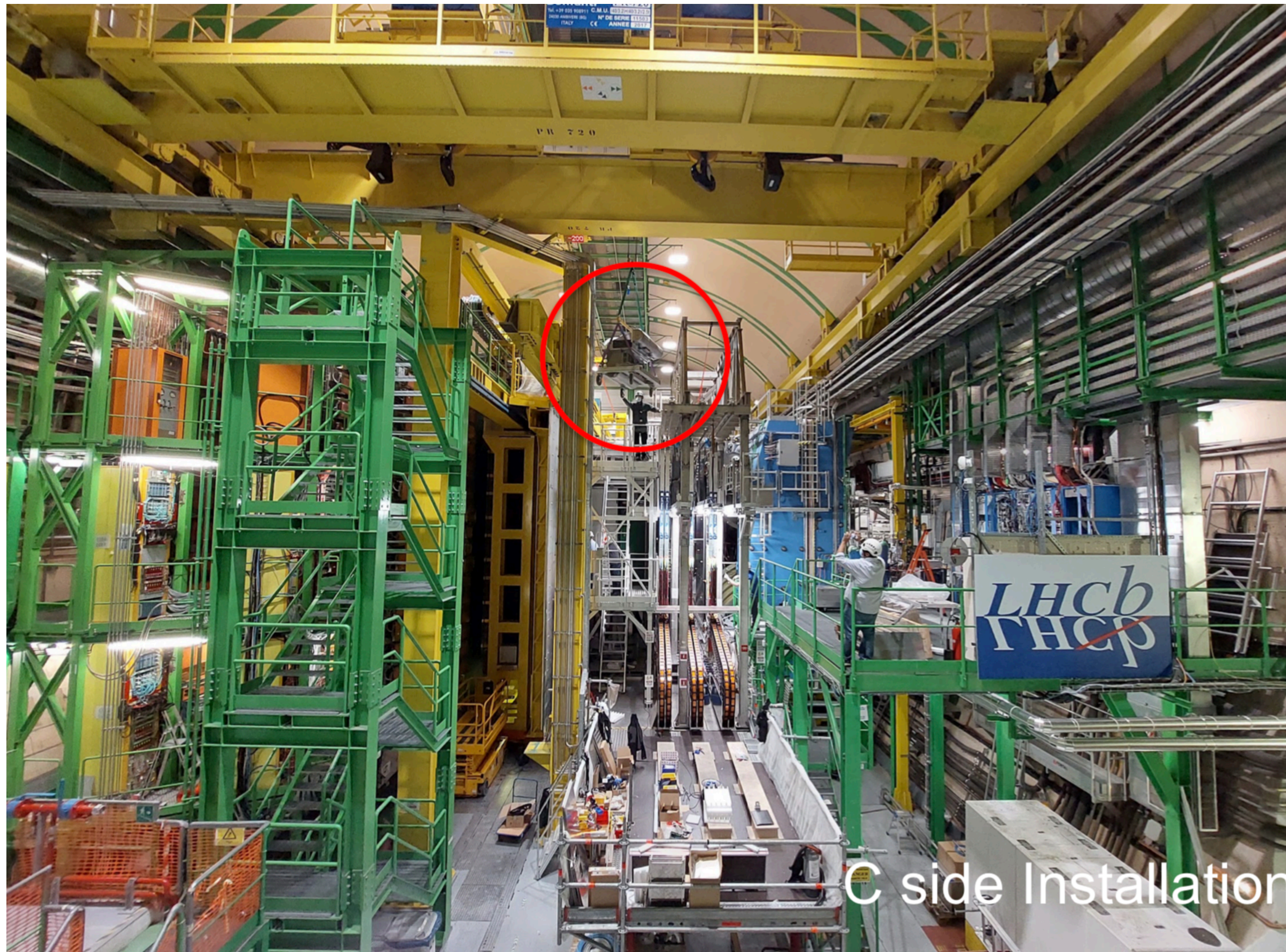
Global Commissioning will have to wait a bit more

- RICH1+RICH2+MUON+CALO+PLUME+BCM+RMS commissioning advanced
- Advanced commissioning of online systems with particular attention to online infrastructure, clock, timing and controls
 - Now concentrating in making the central tools available for scans
- SciFi + VELO-C will finalize commissioning with beam
 - Profit from occasional accesses and occasional periods with no beam
 - Establish tight coordination with LHC to perform machine-experiments interface tests
 - VELO-C work preparatory for a speedy VELO-A commissioning when installed



- commissioning new Beam Conditions Monitor (BCM) hardware, logic and connection: test to trigger dumps and calibrate threshold in May
- commissioning calibration and operational procedures with sub-detectors and collisions:
 - early 450 GeV collisions will be used to repeat time alignment with higher precision
 - test calibration and operational procedures with HV/LV ON where possible; scans, central operations, trigger validation → very final scans to be at 6.8 TeV
- commissioning SMOG2
- commissioning of LHCb luminometer (PLUME) and its luminosity communication to LHC
- Commissioning of SciFi and VELO in the “central partition”
- Need to take time to validate the “VELO closing procedure” and its Velo Safety System (VSS)

P.Collins



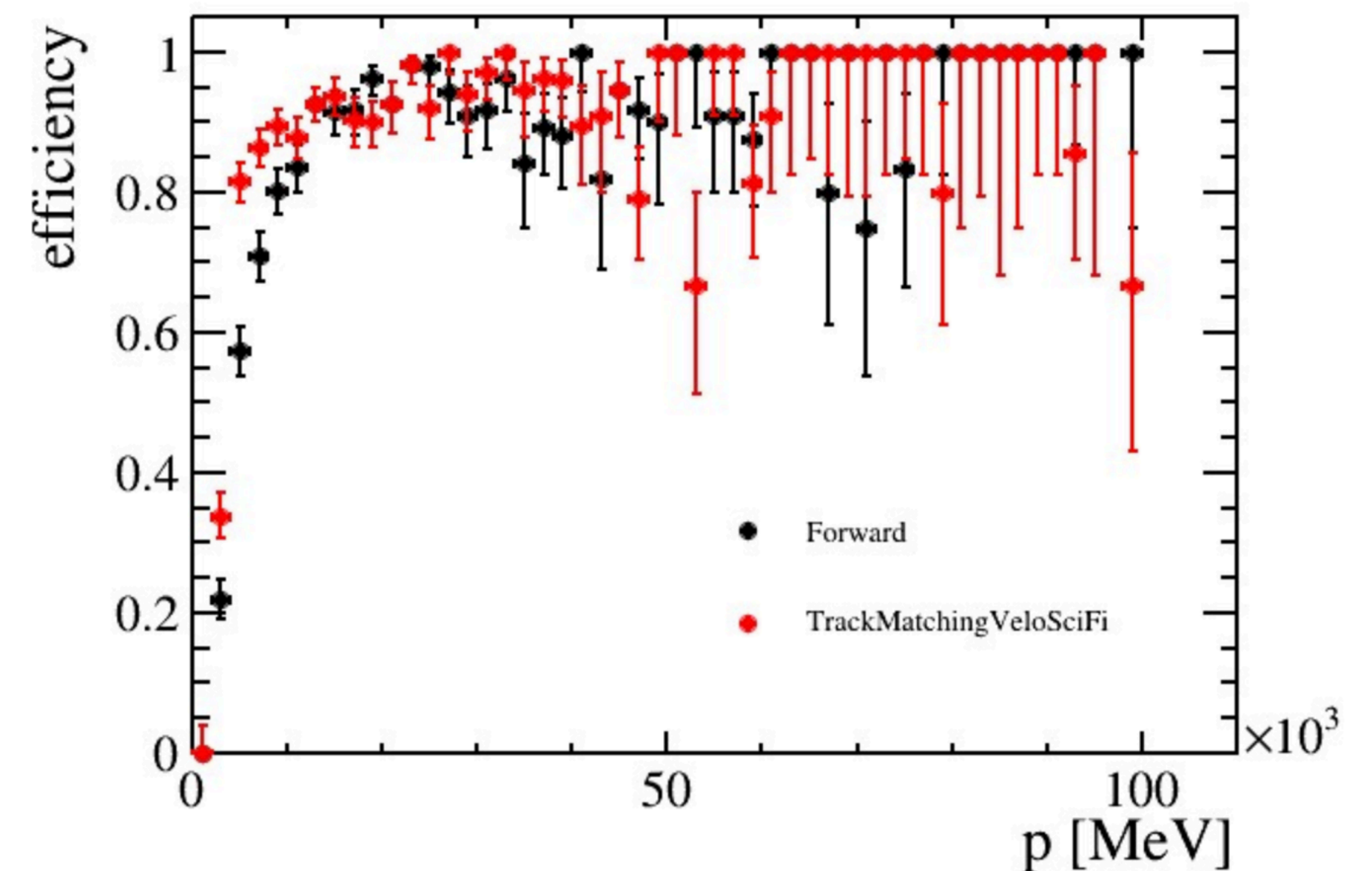
MC REQUESTS

- Encourage people to make “small” requests for their analyses with: **VeloClosed+noUT** configuration.

This will help to exercise the machinery and hopefully be the closest to what we will have for 2022.

MC REQUEST: STEP BY STEP

0. Check if your decay was already produced: [list](#)
1. Check if your EventType exists in the DecFiles project: [link1](#) or [link2](#)
 - If not, follow the [guide](#) to submit your decay.dec
2. Define your request (events number mainly)
3. Send an email to vladimir.gligorov@cern.ch and alessandro.scarabotto@cern.ch specifying the decay (with EventTypes) and how many events per polarity
4. Follow the status of your request on Dirac (you will receive also automatic emails)



MC REQUESTS

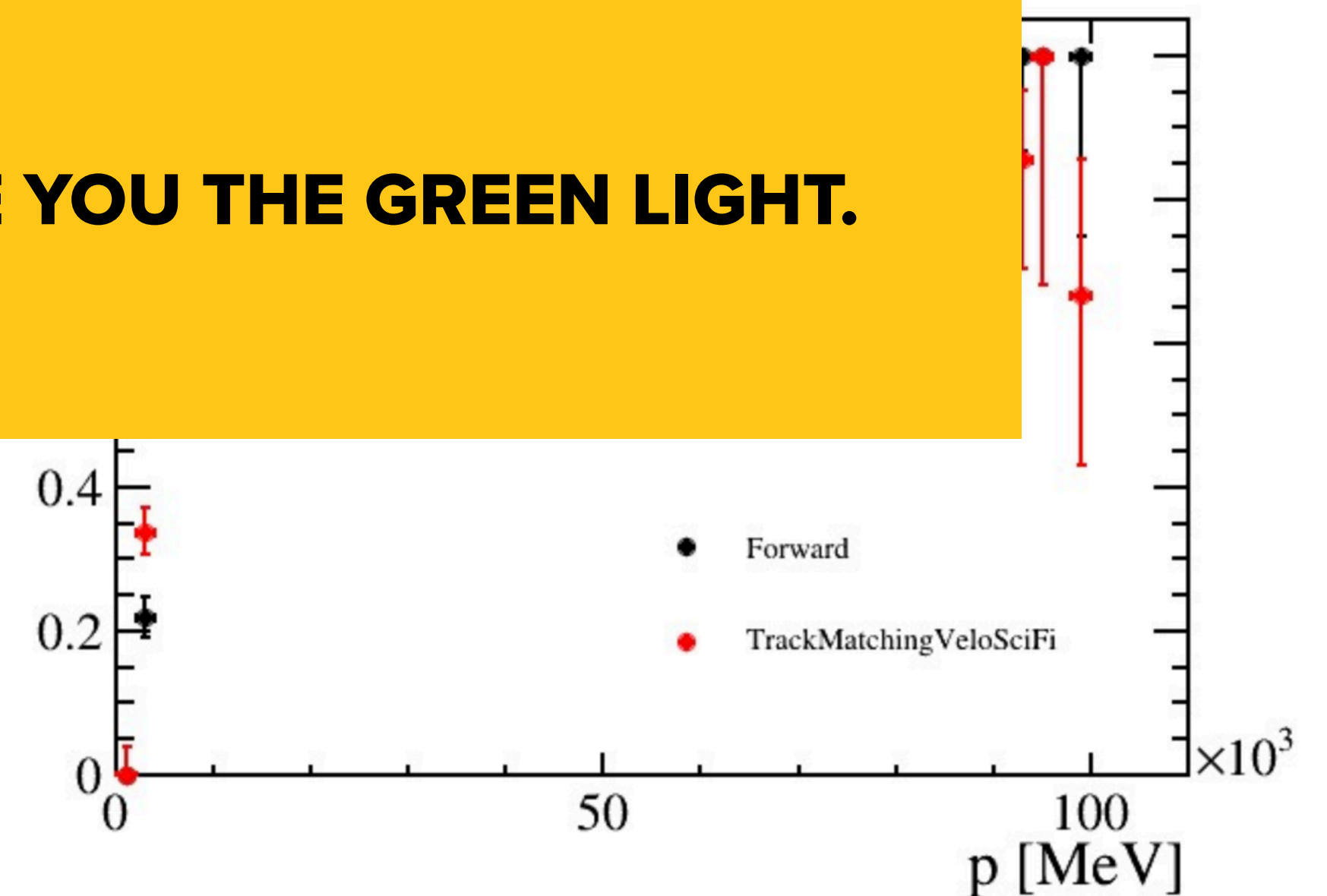
- Encourage people to make “small” requests for their analyses with: **VeloClosed+noUT** configuration.

This will help to exercise the machinery and hopefully be the closest to what we will have for 2022.

WE WILL DISCUSS WITH SIMULATION AND GIVE YOU THE GREEN LIGHT.

0. Check if you have the configuration
1. Check if you have the configuration
 - If not, follow the instructions
2. Define your request (events number mainly)
3. Send an email to vladimir.gligorov@cern.ch and alessandro.scarabotto@cern.ch specifying the decay (with EventTypes) and how many events per polarity
4. Follow the status of your request on Dirac (you will receive also automatic emails)

Alessandro's slides



Christina's Slides

Analysis preservation

<https://gitlab.cern.ch/lhcb/100pbchallenge>

We will provide templated and examples:

<https://gitlab.cern.ch/lhcb-dpa/wp6-analysis-preservation-and-open-data/>

People are vividly(*) encouraged to:

- ☑ use **snakemake**
- ☑ use **analysis production** tools (dedicated session this week).
- ☑ Follow up REANA developments.



(*) Apply British understatement here.

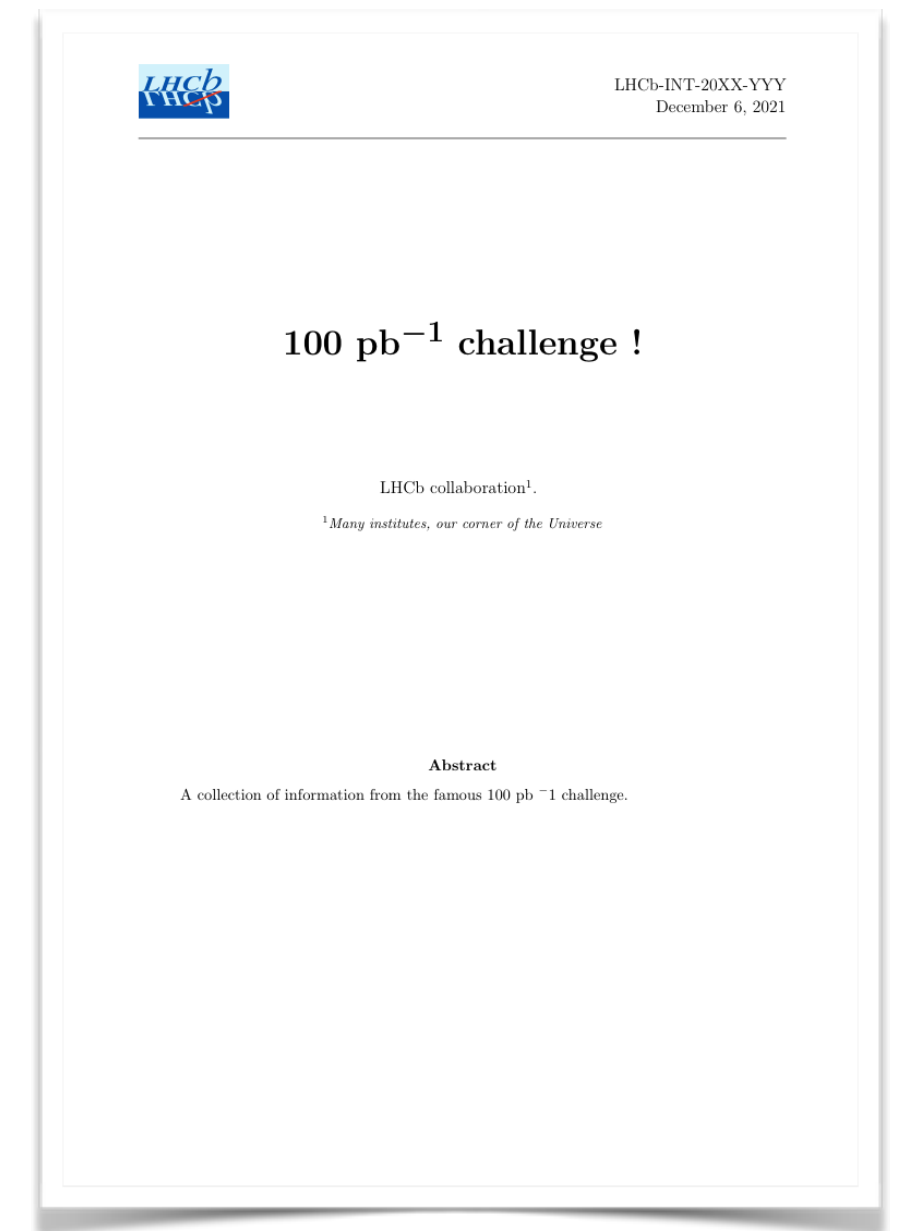
USEFUL LINKS

```
# 100 pb-1 challenge

## Name of your team (copy this template below)

* Analysis goal.
* Decays involved
* List of MC needed/used.
* List of HLT1/2, sprucing lines.
* Selection/fit
* Promised plots !

sign up below it's free !
```



[egroup-lhcb-physics-emtf-100pbchallenge](#)

[HTTPS://CODIMD.WEB.CERN.CH/100PBCHALLENGE](https://codimd.web.cern.ch/100PBCHALLENGE)



A very good opportunity to understand the performances of the detector
across the LHCb physics phase space.

WELCOME TO THE LHCB STARTERKIT!

The organisers of the Run 3 edition:



Alessandro Scarabotto
(LPNHE)



Valeriia Lukashenko
(NIKHEF/KINR)



Lukas Calefice
(Dortmund/LPNHE)



With a lot of help from Yasmine and Eva & all teachers and helpers of the sessions!

THE SCHEDULE

Time (CET)	Monday	Tuesday	Wednesday	Thursday	Friday
11:00-12:00				HLT Efficiency Checker	Allen & HLT1
12:00-12:30		HLT2 line development	DaVinci & FunTuple	Ross	Roel
12:30-13:00		Marian	Abhijit & Davide		
13-14	Early Measurements EMTF				
14:00-14:30					
14:30-15:00		Run 3 data flow @ the Tuesday meeting zoom here			
15-16		Sprucing			Dylan
16-17		Nicole			Analysis Productions

- Instructions on monitoring of HLT2 lines are postponed as the configuration is still in development

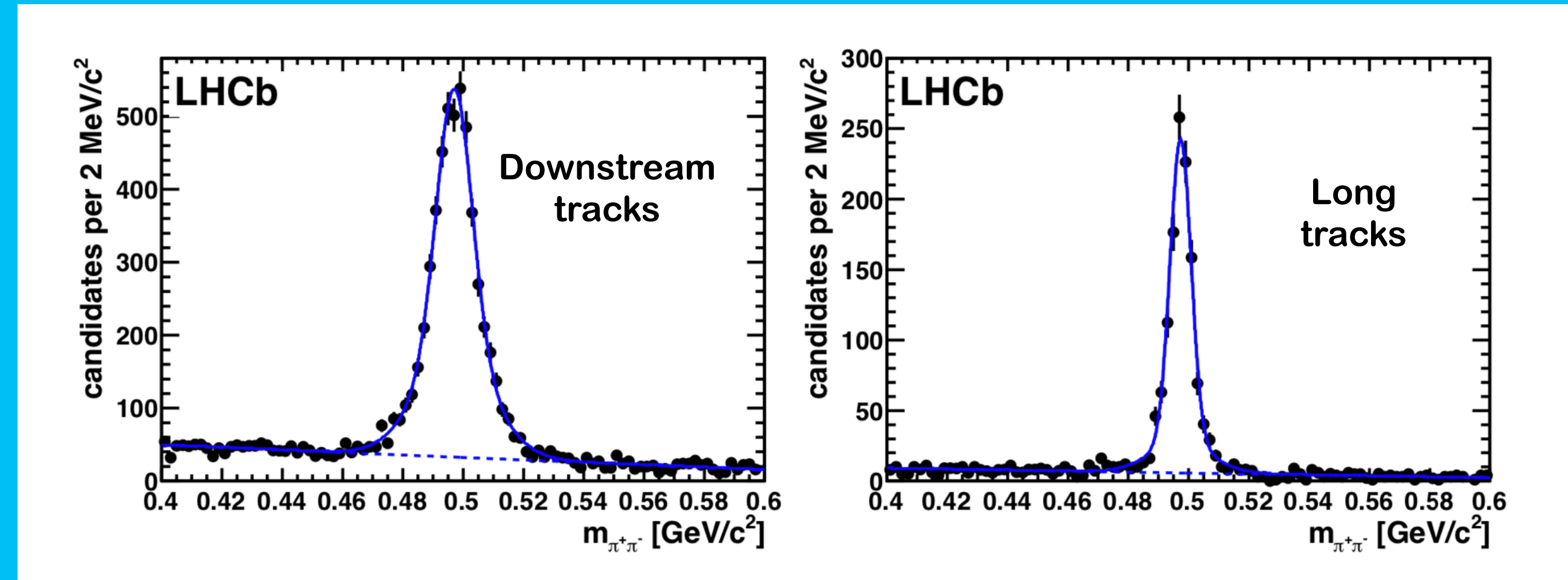
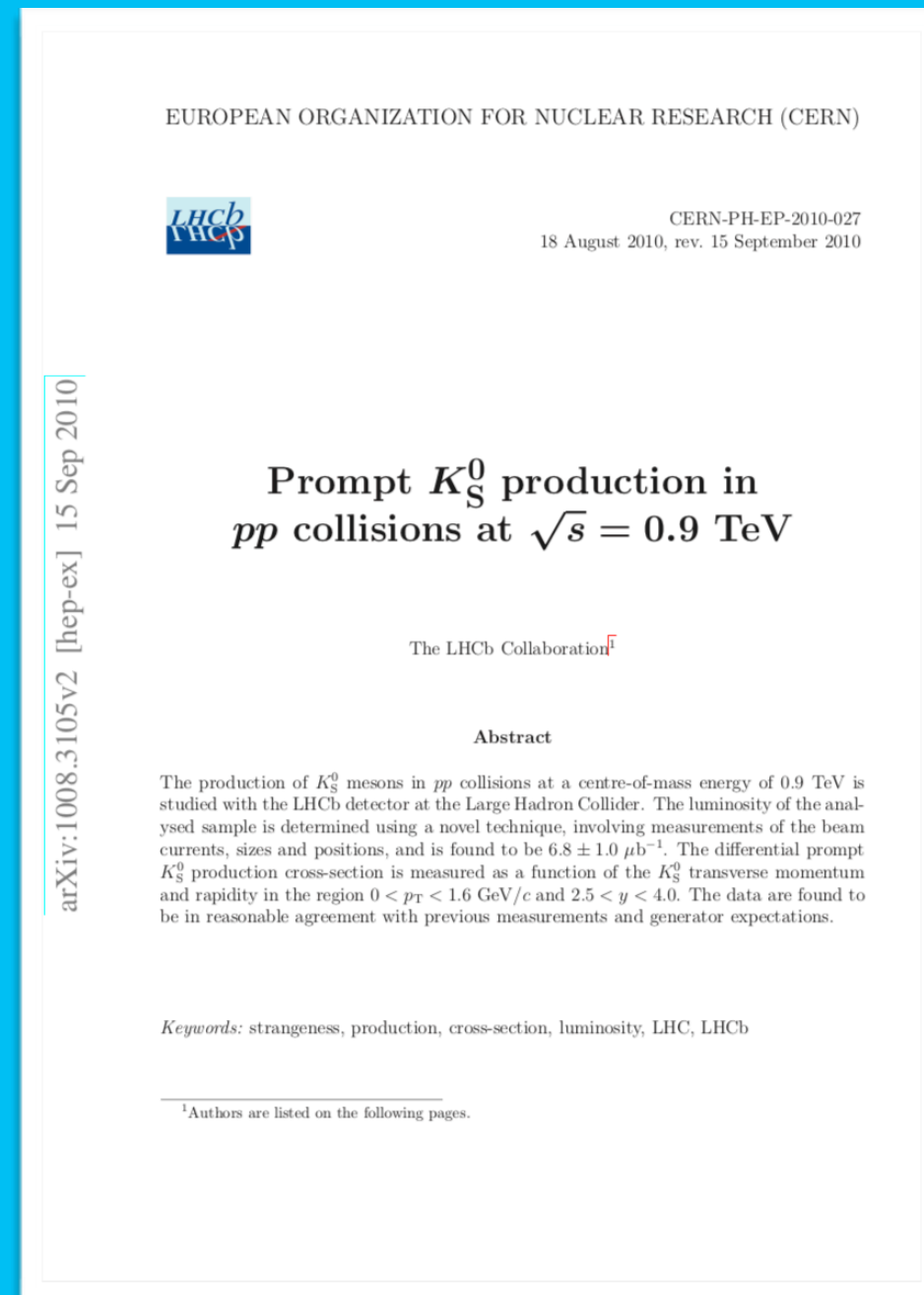
PRACTICAL INFORMATION

- Questions that are relevant for everyone should be ask in the zoom session.
- Join our Helpdesk mattermost channel:
<https://mattermost.web.cern.ch/starterkit/channels/run-3-starterkit-2022-helpdesk>
→ debugging questions or technical issues can be posted there
- All sessions will be recorded and be made available on CDS Videos
- the link to the recordings will be shared via indico mail
- Download stack build tarball from:
[/eos/lhcb/user/there-will-be-a-link/stack.tar.gz](#)
→ need to run `tar -zxvf stack.tar.gz` & make Moore afterwards
- See this collection of useful documentation references:
<https://codimd.web.cern.ch/QCc26-TmQgCaP6AXDgVYmA>

**You can contact us directly on
mattermost or via email if you have
organisational questions or
comments!**

**Enjoy Run 3
Starterkit!**

A BIT OF ARCHEOLOGY



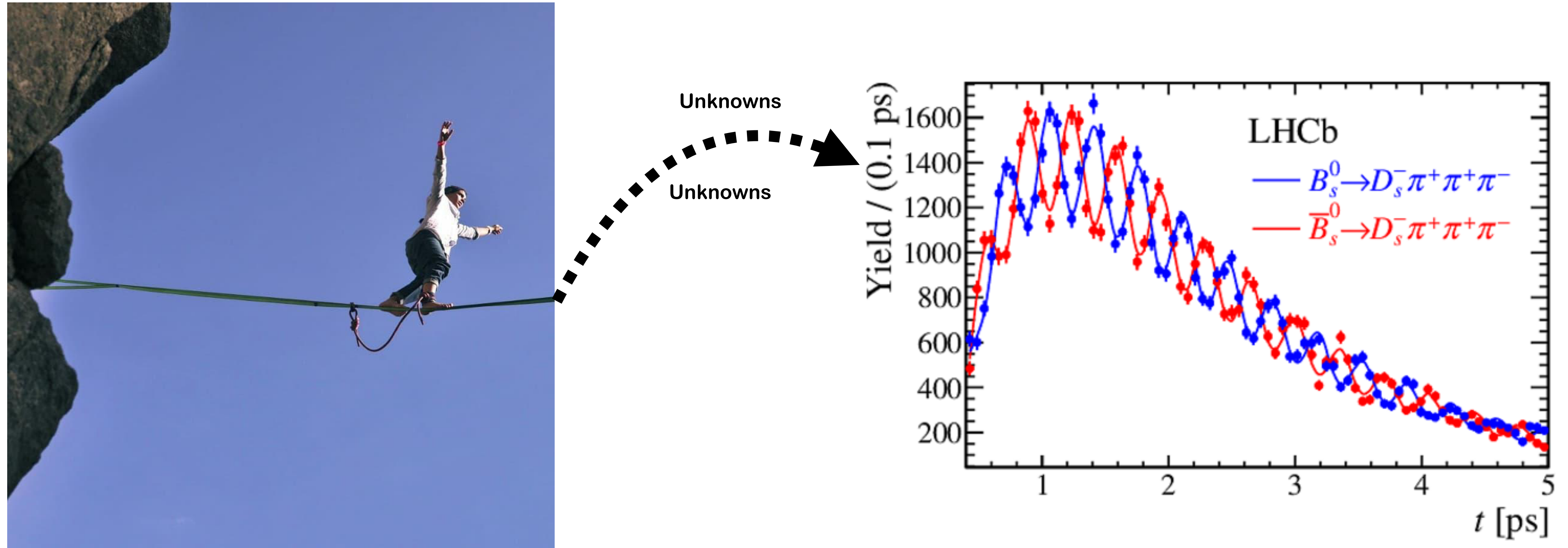
Integrated luminosity 6.9 micro barns.
Center of mass energy 0.9 TeV

During the 2009 run, low intensity beams collided in LHCb at the LHC injection energy, corresponding to a total energy of 0.9 TeV. Due to the dipole magnetic field the beams have a crossing angle that results in the pp centre-of-mass frame moving with velocity $0.0021c$ in the $-x$ direction. Both the beam sizes and crossing angle were larger than those designed for high-energy collisions. In order not to risk the safety of the VELO, the 2009 data were recorded with the two VELO halves positioned 15 mm away from their nominal data-taking position (VELO partially open), resulting in a reduced azimuthal coverage. For this run, the magnetic dipole field was pointing downwards.

“Bah, 2022 will be a **commissioning** year”

“Well, yes but no, not just.”

“Goal is to reach steady **running** by Autumn ”



We might need to iterate some steps.
But progress is not linear.

WHERE TO FIND US

- Our Twiki Page.
- Our Mattermost

OUR ROLE

- Follow closely the first data taking (very soon) and ensure that “everything” will be in place for early measurements and further key channels.



THAT'S ALL FOR TODAY!



Revised 2023 offline resources

- Strategy for 2022+2023
 - We have been **pledged nearly all** the 2022 resources we requested for, **assuming a full data taking year**
 - In 2022 we will take data for **less time** and **lower luminosity**
 - Storage requirement **decrease**
 - CPU requirements **decrease** (less data → less events to simulate)
 - Nevertheless, data taking in **2022 is going to be “bumpy”**
 - We have **enough contingency** for commissioning
 - A good chunk of data will be **“transient”**
- **(re-)use** part of 2022 storage in 2023
- Procurement can be delayed, if helpful for funding agencies
- Effectively delaying the “big” step to 2024

SPECIAL RUNNING CONDITIONS

- We contacted the PWG to know if for their program there is a need of a special running condition such as a value of μ . If so please let us know.

DIFFERENT OUTPUT FILES

- Definition of the different output file types:
 - SIM: only the Gauss simulation part
 - DIGI: samples with Boole digitization
 - XDIGI: extended info from Gauss saved in a DIGI file (you can re-run Boole for example)
 - (μ)DST: sample with reconstruction information (Moore or Brunel)
 - XDST: extended info saved from Gauss and Boole (possible to re-run Boole or Moore/Brunel)
 - LDST: sample with reconstruction and additional linker tables stored (useful for tracking studies for example pattern recognition, clusters info, ...)
 - MDF: samples emulating the real data taking samples (used for software development studies for example in Allen)

The track matching algorithm

- Code adapted from PrMatchNN in HLT2 (thank you Sevda and Michel!)
- Two main inputs: **SciFi** and **VELO** seeds
- Algorithm approach:
 - “Kink” approximation: Velo/SciFi seeds extrapolated to matching position as straight lines
 - Magnetic field and bending in y parametrised with truth simulation to calculate $z_{\text{match}}(x,y)$

