

Cosmic Strips Telescope Simulation with Allpix²

Report on a work in progress

Maximilian Felix Caspar
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Cosmic Strips Telescope Simulation with Allpix²

Agenda

01 The ATLAS ITk Strips System Test

- ATLAS ITk
- System Test

02 ITk Strips Endcap Modules

- Petals

03 Cosmic Muon Flux

- Distribution and East-West Effect

04 The DepositionCosmics module

- The CRY framework
- DepositionCosmics
- Example Usage

05 Current Results

- Tracking Resolution of a Strips Telescope
- Flux Distribution for CRY Particles
- Conclusion & Outlook

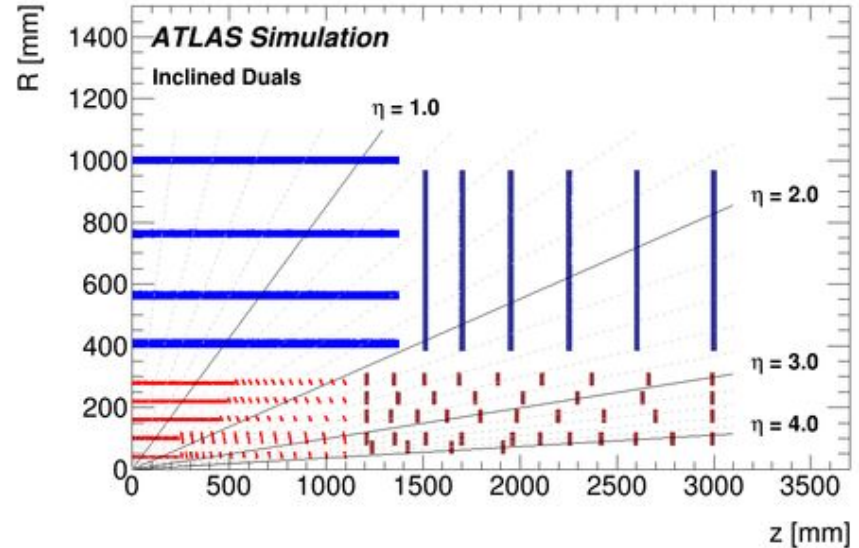
06 Backup Slides

The ATLAS ITk Strips System Test

The ATLAS ITk Strips System Test

ATLAS ITk

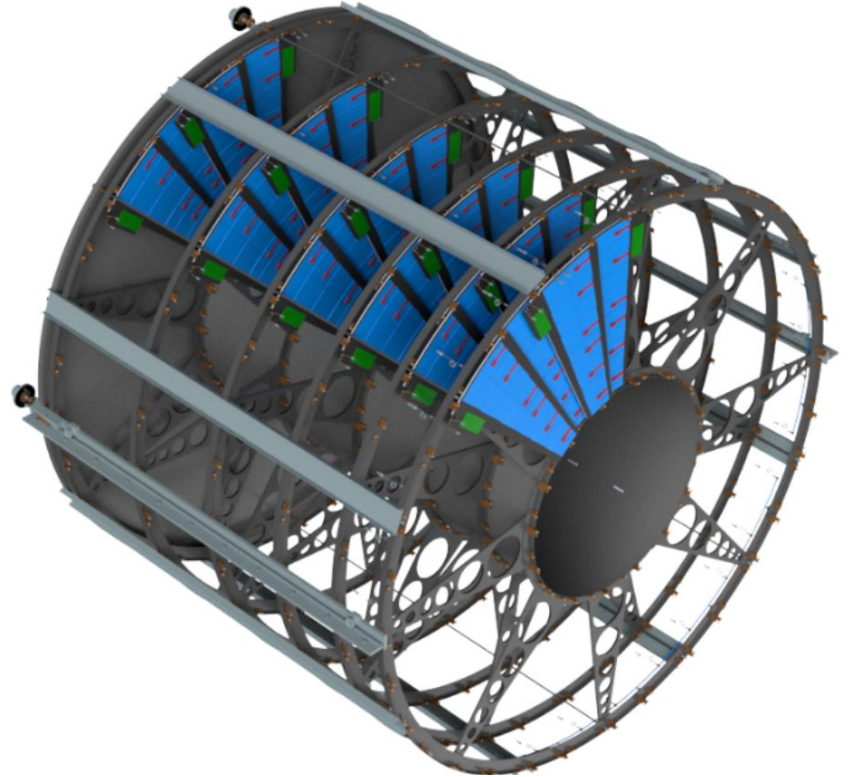
- All silicon tracking detector closest to the interaction point
- A pixel detector measures the tracks close to the initial vertex
- A strip detector is added for more tracking info
- The strip detector consists of a barrel section and two disc-shaped endcap sections



The ATLAS ITk Strips System Test

ITk Strips Endcap

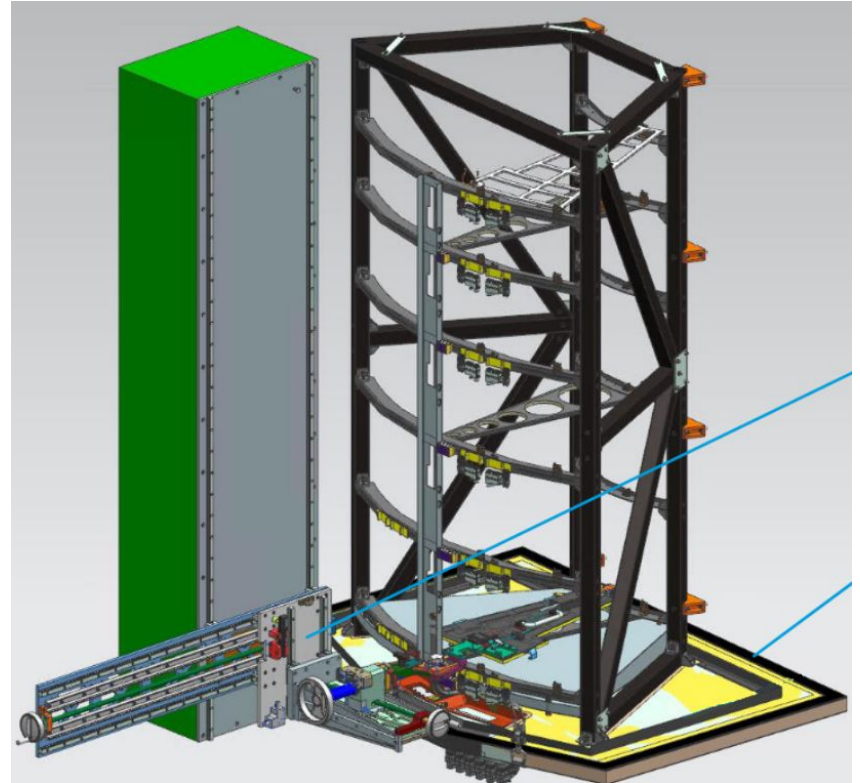
- The endcap consists of six layers of petals
 - Petals are double sided trapezoidal structures with strip modules
- The EC has a carbon fiber mechanical structure, CO₂ cooling and optical fiber readout



The ATLAS ITk Strips System Test

System Test

- The entire setup needs to be tested:
 - Do readout, tracking, cooling, ... work?
 - Can we see cosmic muons when taking data?
- We're building a $\frac{1}{8}$ endcap section to do these tests (called the System Test)
- This is currently being built at the DESY DAF (Detector Assembly Facility)

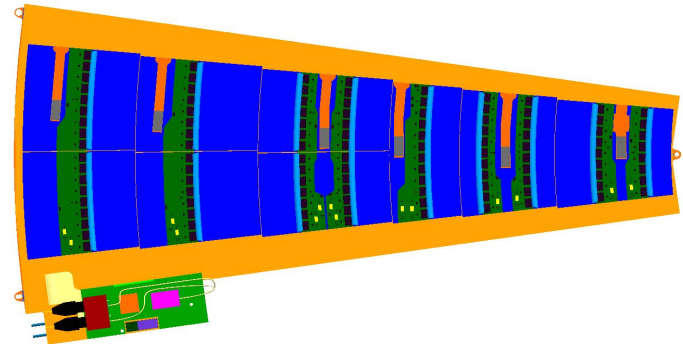
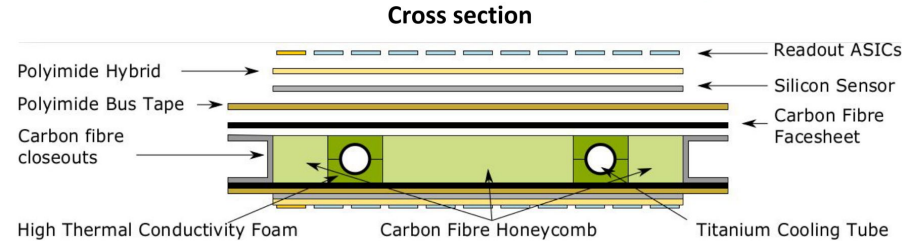


ITk Strips Endcap Modules

ITk Strips Endcap Modules

Petals

- Modules are glued to a carbon fiber structure called “Petal”
- Petals are populated on both sides
 - Due to the stereo angle, this allows for tracking
- The structure also combines electrical links and CO₂ cooling loops

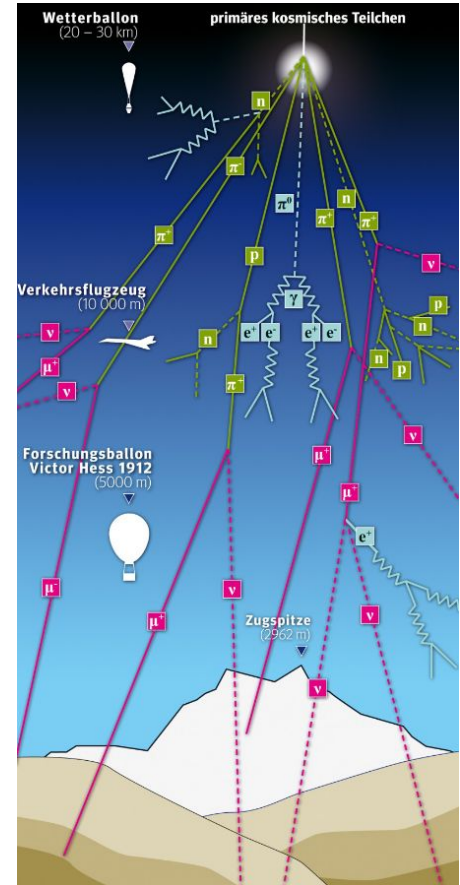
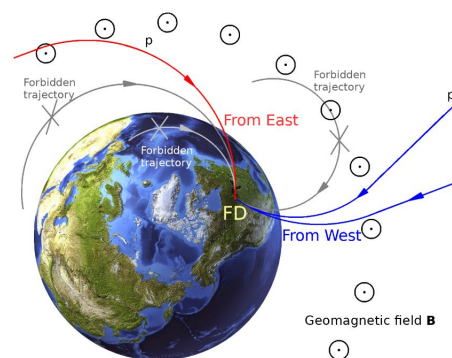


Cosmic Muon Flux

Cosmic Muon Flux

Distribution and East-West Effect

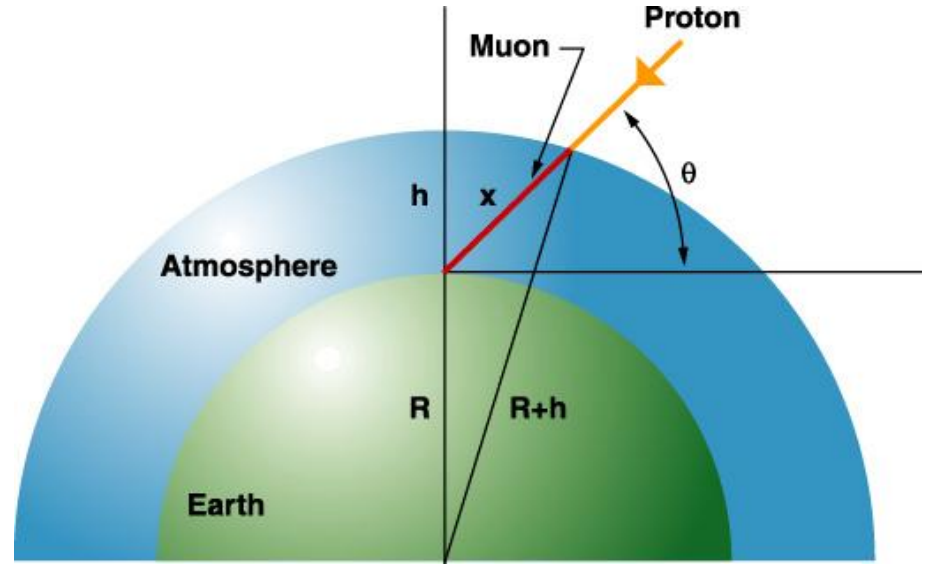
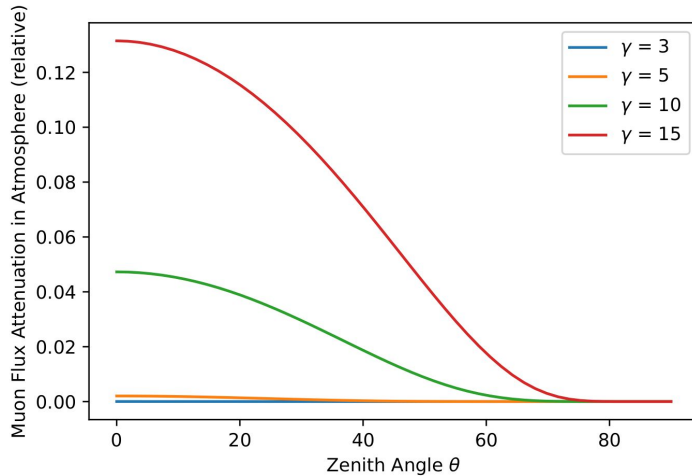
- Cosmic Rays colliding with the atmosphere produce pions, which decay into muons
- Zenith angle dependent flux tends to be $\propto \cos^2\theta$
- Because of earth's magnetic field, there tend to be more muons coming from the west (since cosmic ray particles tend to be positively charged)



Cosmic Muon Flux

Distribution and East-West Effect

- Travel distance after creation depends on θ
- Some muons won't make it due to their short lifetime (even with time dilation!)



- This introduces a dependence on the muon energy spectrum at creation
- Usually, a $\cos^2\theta$ approximation works fine

Cosmic Muon Flux

Flux Calculation from Track Angles

- Azimuthal and zenith angles are calculated from reconstructed tracks in your detector
- For each bin, the flux is calculated as

$$\Phi = \frac{\partial N}{\partial t \partial A \partial \Omega}$$

where $\Omega = \int_{\varphi_{\text{low}}}^{\varphi_{\text{high}}} \int_{\theta_{\text{low}}}^{\theta_{\text{high}}} \sin(\theta') d\theta' d\varphi'$

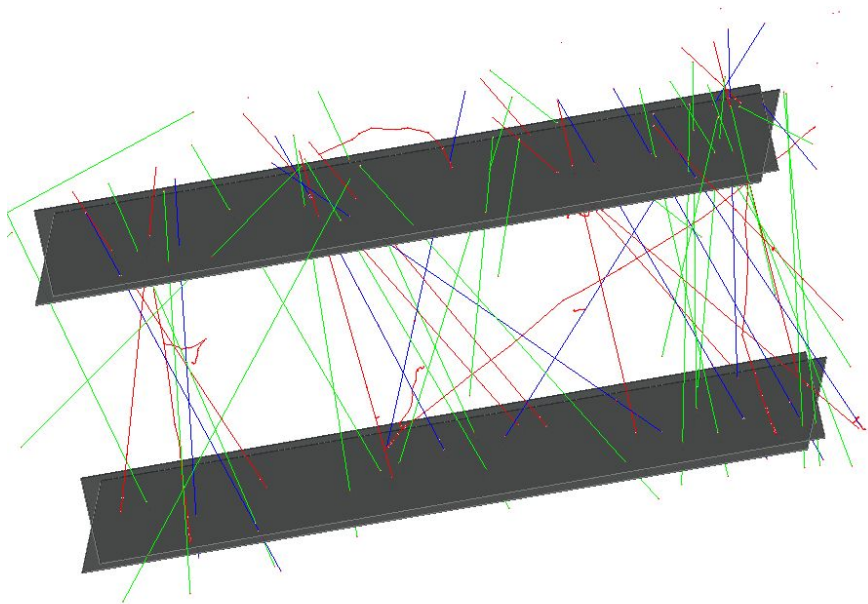
- The effective surface area of the detector has to be found using MC simulations

The DepositionCosmics module

The DepositionCosmics module

The CRY framework

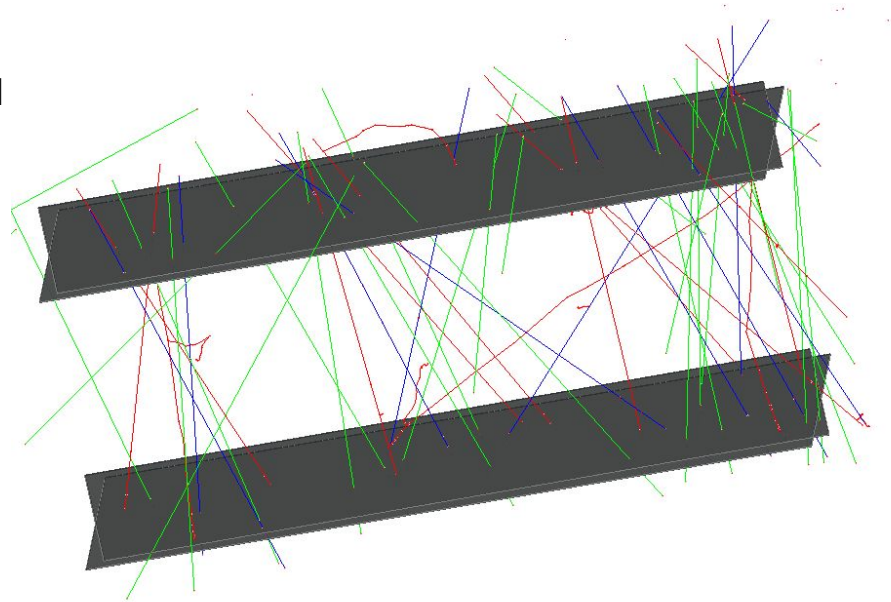
- Generates correlated cosmic-ray particle showers as an input to Geant4
- Uses tables of pre-computed full simulations of CRs that have been checked against data
- Yields a time duration for every event generated



The DepositionCosmics module

DepositionCosmics

- Simon came up with an interface to Allpix Squared
- CRY can now be invoked similar to the DepositionGeant4 module
- Size of the output area is chosen based on the detector size
 - All particles are generated on a square area heading in negative z direction
 - Side lengths 1, 3, 10, 30, 100, and 300 m are available



The DepositionCosmics module

Example Usage

- Usage is similar to DepositionGeant4
- Particles are generated in the xy plane and move in negative z direction
- The total simulation time is printed at the end of the run

Instantiate Module

Types of particles

Data path
(usually not necessary)

```
# Define the cosmic ray source:
[DepositionCosmics]
physics_list = FTFP_BERT_LIV
number_of_particles = 9
max_step_length = 10.0um

return_neutrons = false
return_protons = false
return_gammas = false
return_pions = false
return_kaons = false
return_electrons = false
return_muons = true

data_path = "/afs/desy.de/user/c/caspar/dep-cosmics"
```

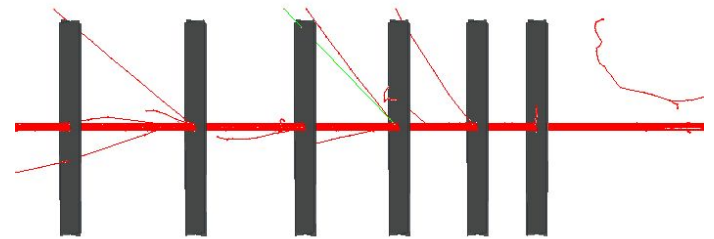
```
15:30:17.745| (STATUS) Finished run of 3830 events
15:30:17.745| (STATUS) [F:DepositionCosmics] Total simulated time in CRY: 99.3428s
```


Current Results

Current Results

Tracking Resolution of a Strips Telescope

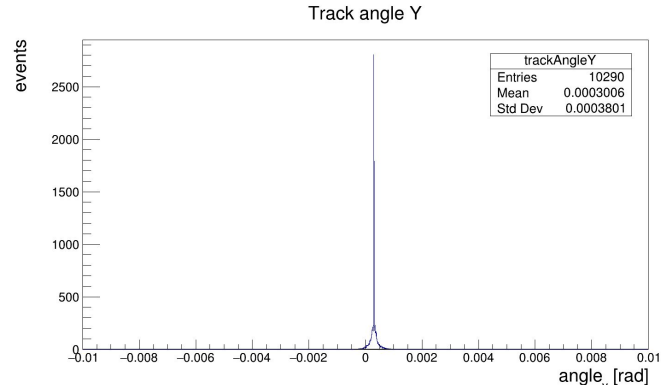
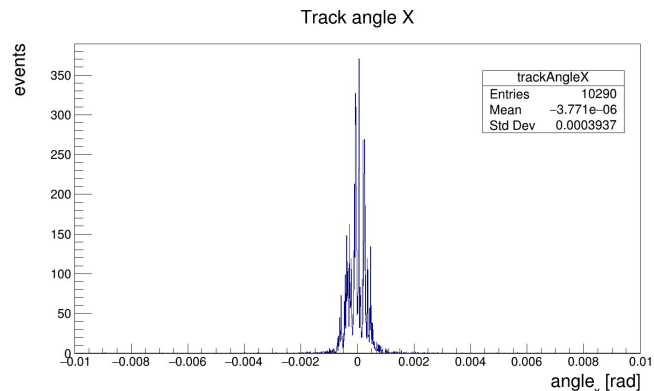
- Use six layers of double strips sensors
 - Geometry inspired by the ITk barrel section
 - 1280 x 28 (400 μm x 24.10 mm) Pixels, 40 mrad stereo angle
- Beam of 10 GeV muons to test the tracking
 - 1.013 tracks per event



Current Results

Tracking Resolution of a Strips Telescope

- Use six layers of double strips sensors
 - Geometry inspired by the ITk barrel section
 - 1280 x 28 (400 μm x 24.10 mm) Pixels, 40 mrad stereo angle
- Beam of 10 GeV muons to test the tracking
 - 1.013 tracks per event
 - Angular resolution in X = 0.4 mrad, in Y = 0.4 mrad (for perpendicular tracks)



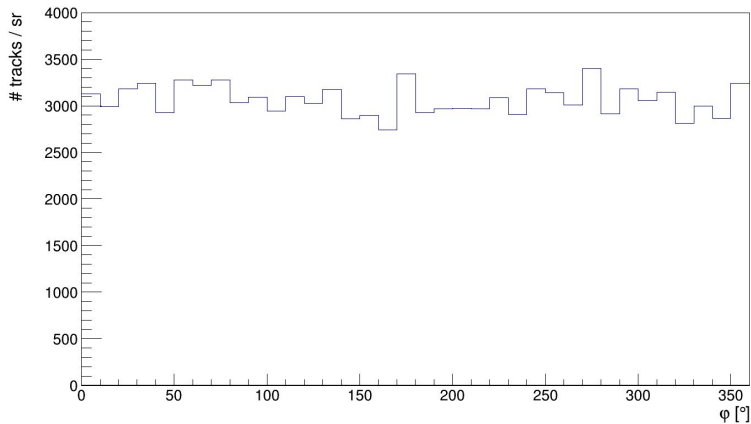
Current Results

Flux Distribution for CRY Particles

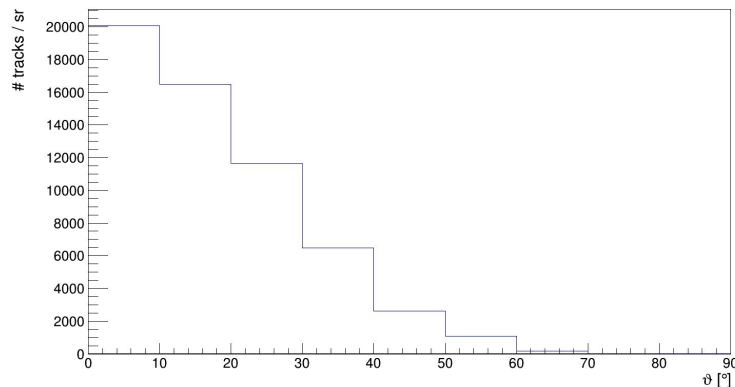
- Same geometry as before
- Only generate Muons using CRY
 - Have a 3 m x 3 m x 10 cm concrete wall before the telescope to simulate being in a building
- Zenith and azimuthal fluxes are calculated

$$\Phi = \frac{\partial N}{\partial \Omega}$$

Azimuthal flux distribution



Zenith angle flux distribution



Current Results

Conclusion & Outlook

- We can generate cosmic ray particles in Allpix² and get reasonable results
- Simulating regular strip detectors also works
- Next steps:
 - Combine this with Radek's results to get a full system test simulation
 - Normalize the flux to the effective detector area to allow for comparison with other experiments

Thank you

Backup Slides

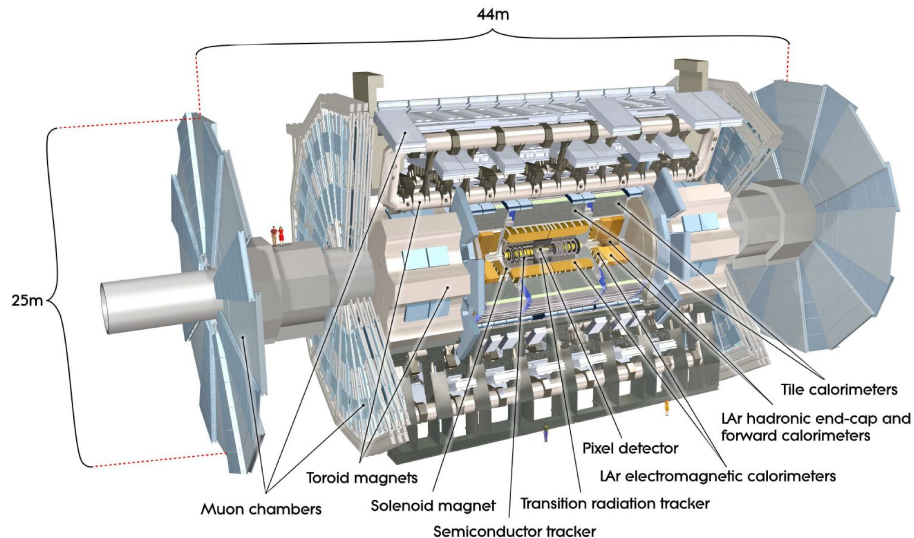
Corryvreckan module: AnalysisParticleFlux

- Histograms:
 - Zenith angle vs. azimuthal angle of tracks and associated 1D histograms
 - Same histograms divided by bin solid angle
 - Zenith and azimuth dependent efficiencies for each detector
 - 2D efficiency map for each detector
- TODO:
 - Implement the calculation of the effective area (need the track angles from MC for this)

Backup Slides

The ATLAS Detector

- Largest experiment at LHC
- General purpose experiment to test (beyond) Standard Model physics
- Due to luminosity upgrades, we're rebuilding most of the detector
- The Inner Detector (ID) is going to be completely replaced

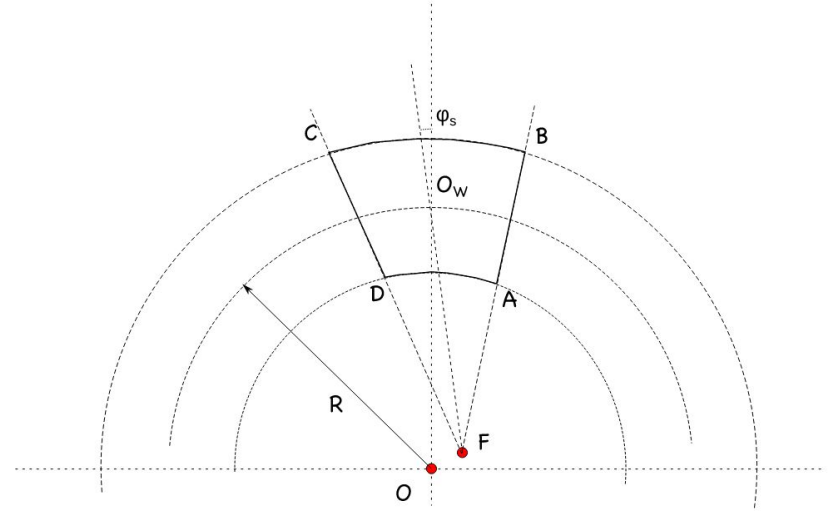
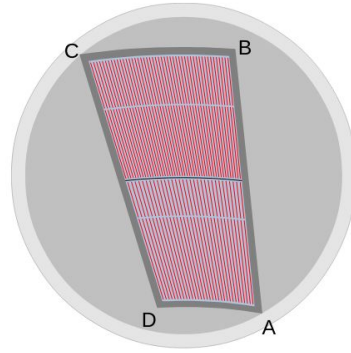


Backup Slides

Geometry and Readout of Endcap Modules

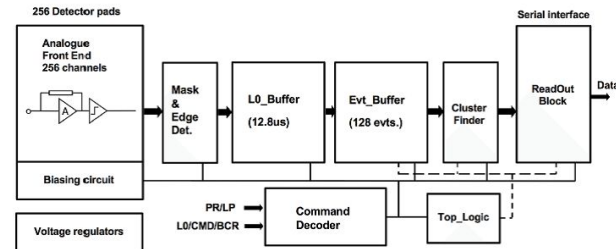
Geometry

- Rather complex geometry:
 - The strips are oblique with respect to the radii
- The geometry is currently implemented in Allpix by Radek Privara



Readout

- Binary readout via the ABC* chip (glued on top of the sensor)
- Data is collected from all FE chips via the HCC* chip



Contact

Deutsches Elektronen-
Synchrotron DESY

www.desy.de

Maximilian Felix Caspar

ATLAS

maximilian.caspar@desy.de

+49 40 8998 2734