Novel Manifestations of Primordial Black Holes



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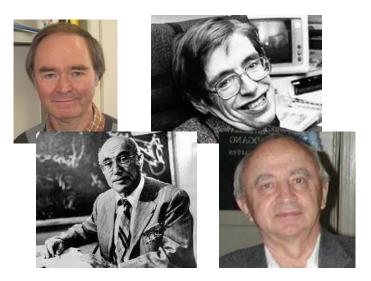
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Black Holes

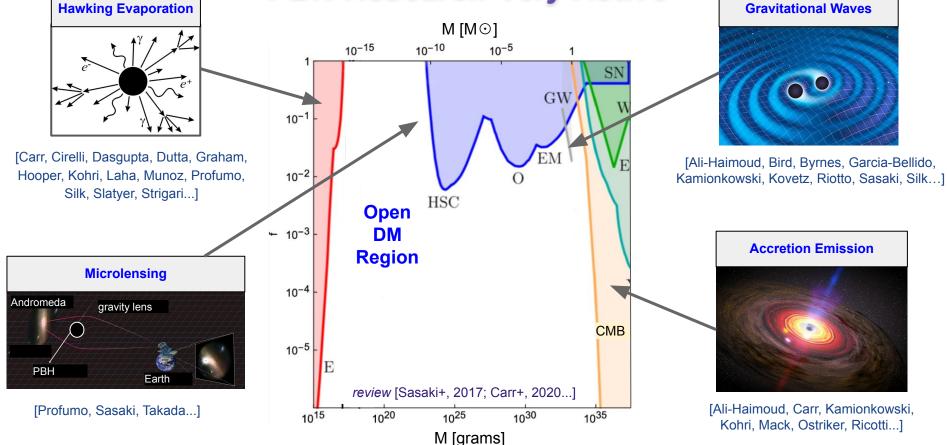
 ${
m EXIST} \ M_{
m BH} \sim 10^6 M_{\odot}$



...might appear from early Universe (primordial), and contribute to dark matter



PBH Research Very Active

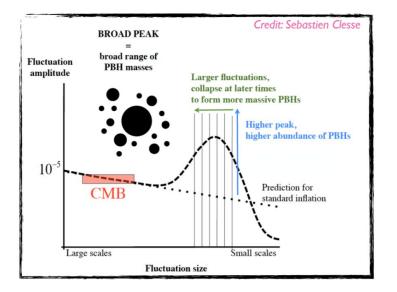


"Standard" PBH Formation

● Big perturbations enter horizon → collapse [Carr, Kamionkowski, Kawasaki, Sasaki, Silk, Yanagida...]

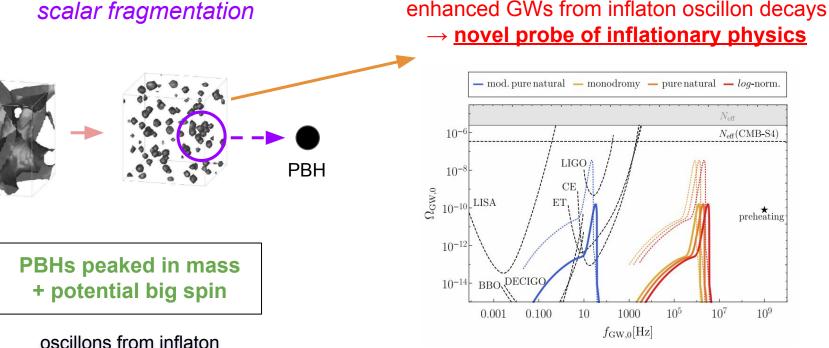
$$M_H \approx \frac{c^3 t}{G} = 10^{15} \,\mathrm{g}\left(\frac{t}{10^{-23} \,\mathrm{s}}\right)$$

- Need to fine tune inflaton potential
 → sensitive to restrictions on field behavior
 - Example: "string swampland conjectures" [Kawasaki, VT, PRD, 1810.02547]



→ many formation mechanisms: e.g. domain walls & string loops [Deng, Ferrer, Garriga, Khlopov, Hawking, Vilenkin...]

Distinct PBH Features Possible

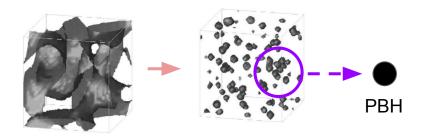


[Lozanov, VT, 2204.07152]

Cotner, Kusenko, VT, PRD, 1801.03321; Cotner, Kusenko, Sasaki, VT, JCAP, 1907.10613]

Distinct PBH Features Possible

scalar fragmentation



PBHs peaked in mass + potential big spin

oscillons from inflaton [Cotner, Kusenko, VT, PRD, 1801.03321; Cotner, Kusenko, Sasaki, VT, JCAP, 1907.10613]

vacuum bubble "multiverse"

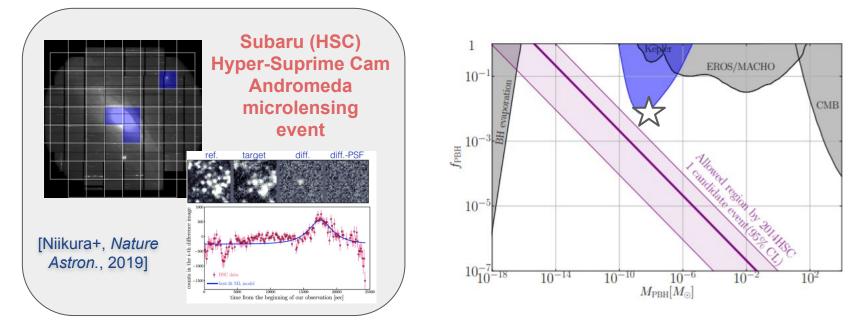


PBHs broadly distributed in mass

[Deng, Sasaki, Vilenkin...; [Kusenko, Sasaki, Sugiyama, Takada, **VT**, Vitagliano, *PRL*, 2001.09160]

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PBH DM from Bubble Multiverse: Detected by HSC ?!



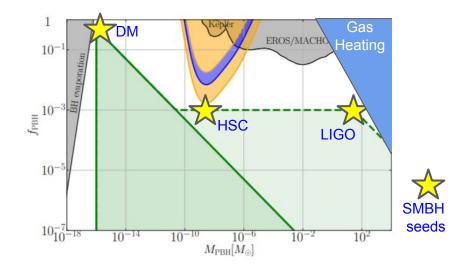
- PBH DM from bubble multiverse consistent with detected HSC event !
 - \rightarrow tail of broad PBH distribution allows for indirect test of open DM window

[Kusenko, Sasaki, Sugiyama, Takada, VT, Vitagliano, PRL, 2001.09160]

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PBH DM from Bubble Multiverse: Detected by HSC ?!

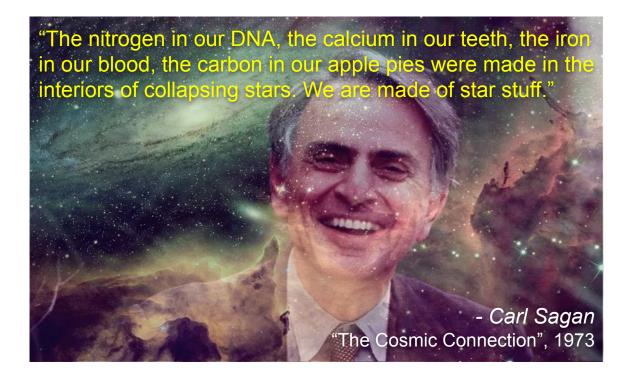
Generalized model explains many observables simultaneously (DM, LIGO, SMBH seeds...)



*** indirectly test NANOGrav GWs with HSC via PBHs [Sugiyama, VT+, PLB, 2010.02189]

Will be <u>definitively</u> tested with new HSC data !

[Kusenko, Sasaki, Sugiyama, Takada, VT, Vitagliano, PRL, 2001.09160]

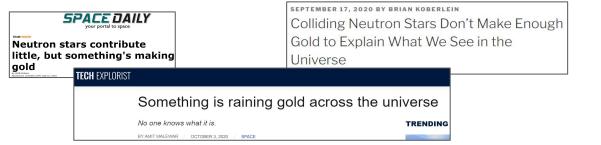


Where do heavy elements (gold) come from? \rightarrow major problem

Making Gold with Tiny PBHs

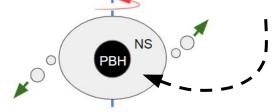
• Origin of heavy elements (gold) long-standing problem

→ neutron star mergers great, but might not be enough e.g. [Kobayashi+, 2020]



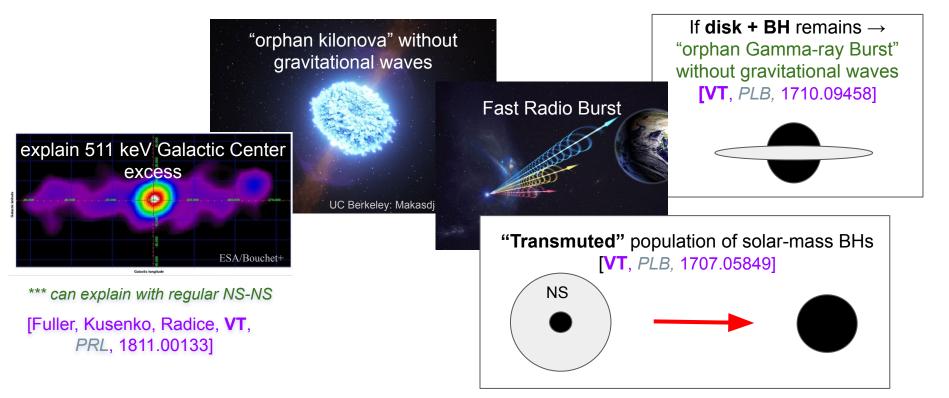
Elegant solution: asteroid-mass PBHs making DM captured by neutron stars, small PBHs eat & explode them → "r-process nucleosynthesis" factories





[Fuller, Kusenko, VT, PRL, 1704.01129] + Viewpoint Highlight by H.-T. Janka

Neutron Stars (+ White Dwarfs) as PBH Laboratories

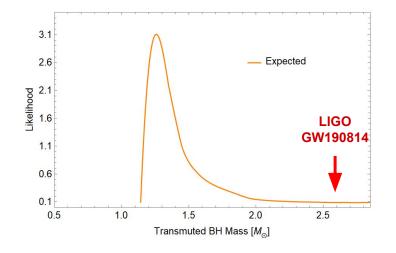


[Fuller, Kusenko, VT, PRL, 1704.01129; VT, PLB, 1707.05849; VT, PLB, 1710.09458]

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Origin of Solar-mass Black Holes

- Solar-mass (~1-2.5 M☉) BHs unexpected in astrophysics → PBHs ? Transmuted?
- LIGO detected candidate event [Abbott+, ApJL, 2020...] ...how to tell BH origin ?
- **Solution:** *transmuted* BHs from PBHs (or particle) DM eating NSs follow NS mass distribution

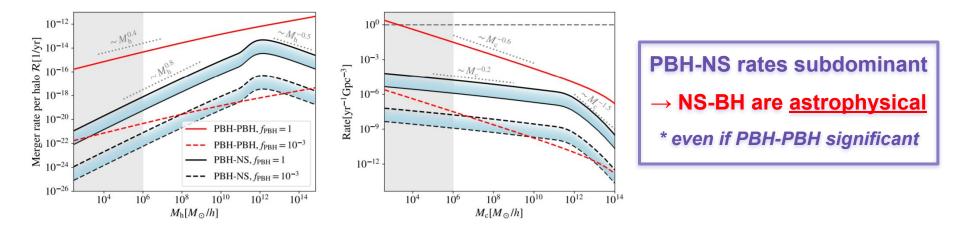


Large (> 1.5 M^o) candidates unlikely to be transmuted BHs!

[**VT**+, *PRL*, 2008.12780]

Identifying Black Hole - Neutron Star (BH-NS) Mergers

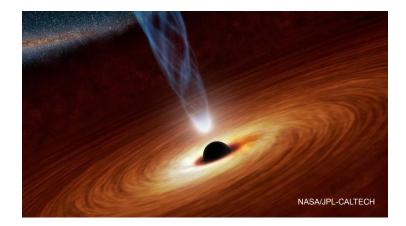
- PBH-PBH been linked with LIGO BH-BH GW events [Byrnes, Kamionkowski, Riotto, Sasaki ...]
- First reported BH-NS candidates by LIGO [Abbott+, ApJL, 2021...]from PBHs?
- Unlike PBH-PBH, PBH-NS can only form after star formation



[Sasaki, VT, Vardanyan, Zhang, ApJ, 2110.09509]

Are Intermediate-mass BHs Primordial ?

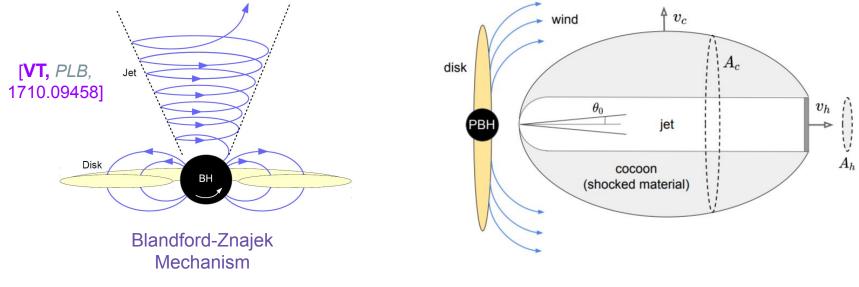
- GW190521 event ~ 150 M^o merger mass [Abbott+, PRL, 2020], first definitive IMBH detection
- New general cosmology-independent observable: interactions and heating of gas
- Gas heating mechanisms:
 - gravitational drag (dynamical friction)
 - accretion disk photons
 - accretion outflows / winds
- Great testing site: dwarf galaxies (Leo T)



[Lu, **VT**+, *ApJL*, 2007.02213; **VT**, Lu+, *JCAP*, 2105.06099]

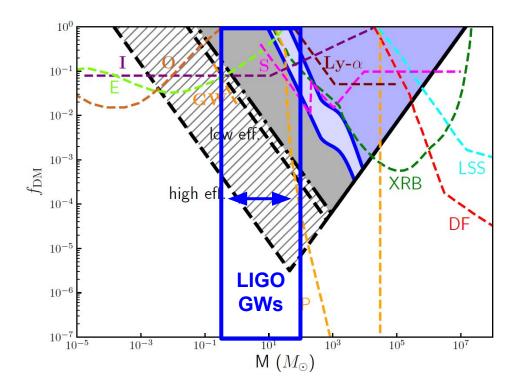
PBH Outflow Winds and Jets

• Outflow winds and powerful jets (especially for spinning PBH) expected to deposit efficiently significant energy via shock heating $L\sim\epsilon\dot{M}$



[VT, Lu, Murase, Inoue, Gelmini, 2111.08699]

PBH Outflow Winds and Jets



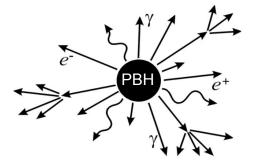
* gas heating from evaporating PBHs

[Laha, Lu, **VT**, *PLB*, 2009.11837] (also [Kim, 2020])

[**VT**, Lu, Murase, Inoue, Gelmini, 2111.08699]

Exploring Evaporating PBHs with GWs

 Evaporating PBHs with mass ≤ 10⁹ g unconstrained, <u>how to explore scenarios ?</u>



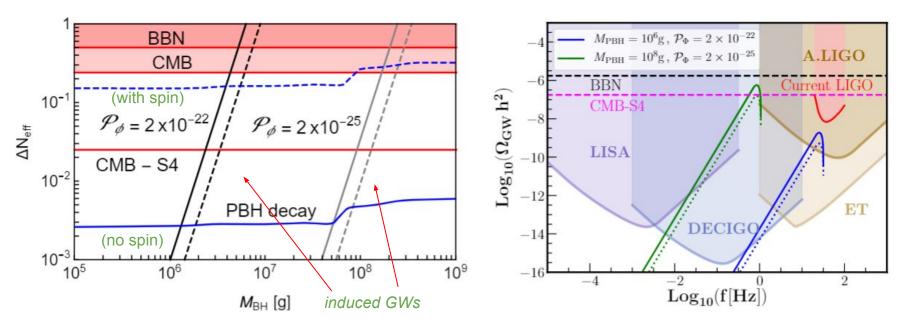
- Evaporating PBH emission products \rightarrow "dark radiation" \rightarrow change ΔN_{eff}
 - PBH spin distribution can significantly modify [Hooper, H

[Hooper, Krnjaic, March-Russell, McDermott, 2020; Arbey, Auffinger+, 2021; Masina, 2021]

- Rapid evaporation of PBHs dominating Universe \rightarrow induced GWs \rightarrow change ΔN_{eff}
 - PBH *mass* distribution can significantly modify

[Inomata, Kohri+, 2019; Papanikolaou, Vennin+, 2020; Domenech, Lin, Sasaki, 2020...]

Exploring Evaporating PBHs with GWs



 \rightarrow Coincidence signals allow probing many scenarios over broad mass-range !

[Domenech, VT, Sasaki, PLB, 2105.06816]

Summary

- Renaissance era in PBH research
- Strong synergy with observational (especially multi-messenger) astrophysics
- Many new ideas emerging for PBH production, signals, solutions to old puzzles



