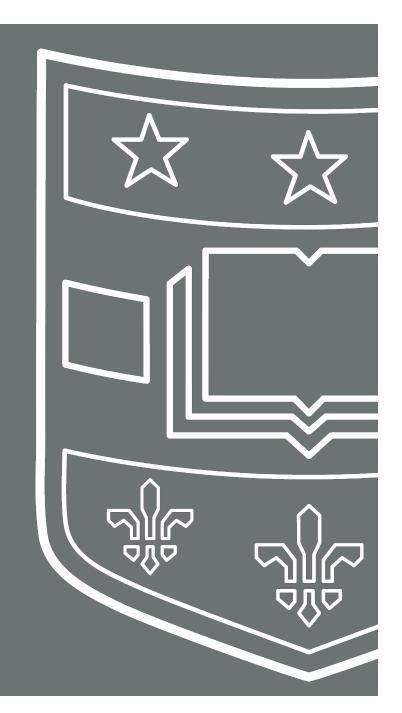
# Generative Adversarial Network for Approximating the Chameleon Scalar Field

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## Chameleon Gravity



- Chameleon scalar particle: proposed by Khoury and Weltman (2004) as a possible dark energy candidate
  - 5<sup>th</sup> force screened by mass, so not seen in locally dense regions
- Chameleon equation of motion in the Einstein frame has form:

 $\nabla^2\phi=V_{eff}'(\phi)$ 

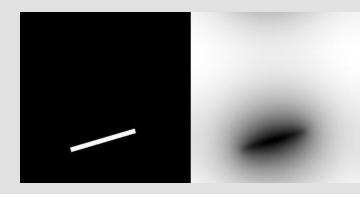
 $V_{eff}(\phi) = V(\phi)A(\phi)\rho_m = \lambda^4 + \lambda^5/\phi + \rho_m/M$ 



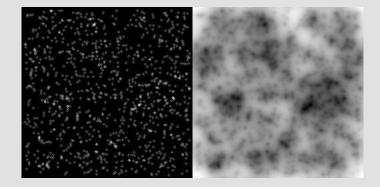
# Chameleon Gravity



- Good representative model for family of other types of modified gravity (symmetron gravity)
  - Relatively easy to work with, proof of concept
- The chameleon is a scalar-tensor modification of gravity that mediates a fifth force
- Numerics:
  - Solving for a fifth force potential like the chameleon scalar field is expensive
  - Want to integrate into n-body simulations



$$abla^2 \phi = -\lambda^5/\phi^2 + 
ho_m/M$$



## Chameleon Gravity

• One ok numerical method – matrix inversion

$$\begin{split} \phi &= \phi_0 + \delta \phi \\ \nabla^2(\phi_0 + \delta \phi) &= -\lambda^5(\phi_0 + \delta \phi)^{-2} + \rho_m / M \\ \nabla^2(\phi_0 + \delta \phi) &= -\lambda^5(1/\phi_0^2 - 2\delta \phi / \phi_0^3) + \rho_m / M \\ [\nabla^2 - 2\lambda^5(1/\phi_0^3)] \delta \phi &= -\nabla^2 \phi_0 - \lambda^5 / \phi_0^2 + \rho_m / M \\ \delta \phi &= -[\nabla^2 - 2\lambda^5(1/\phi_0^3)]^{-1} [\nabla^2 \phi_0 - \lambda^5 / \phi_0^2 + \rho_m / M ] \end{split}$$

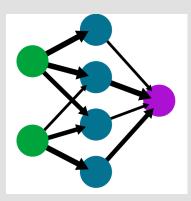
- $\exists$  alternatives (like typical relaxation method)
  - Also expensive
- Computational expense not unique to chameleon

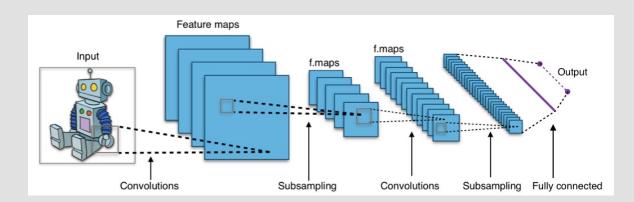


# Machine Learning

- Physicists are lazy
- Neural nets: cheap, clever, black box
- Dense neural network:

• Convolutional neural network:

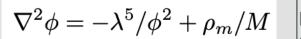




 Backpropagation: process of calculating each weight's contribution to the total loss, and updating it

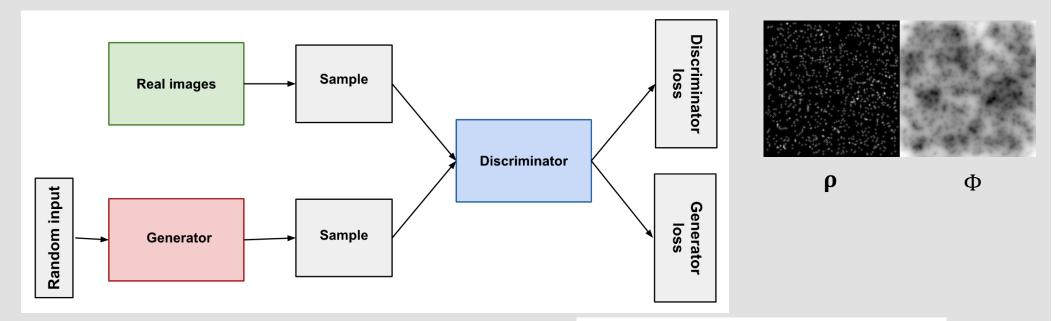


## Machine Learning





Generative adversarial network:



- Generator loss: L1 plus KL divergence  $D_{\text{KL}}(P \parallel Q) = \int_{-\infty}^{\infty} p(x) \log\left(\frac{p(x)}{q(x)}\right) dx$
- Discriminator loss: classification error

# GAN Training Results

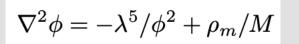
 Generator and discriminator
 losses over a few rounds of training on different data sets



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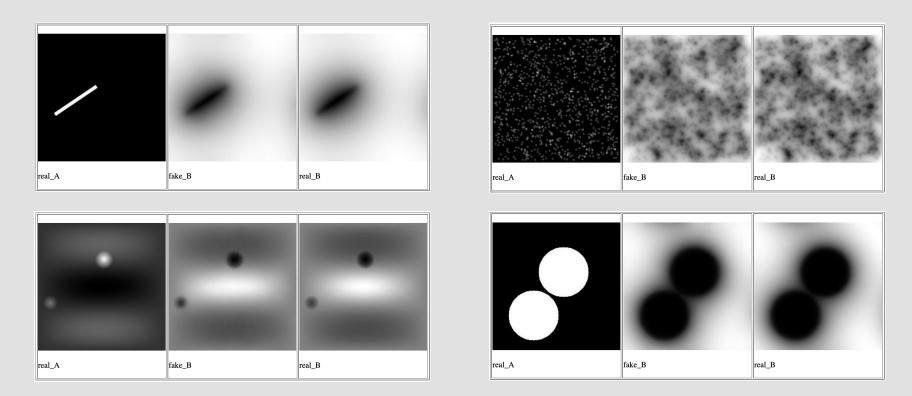
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## Seeing is Believing





• Testing on unseen data - no overfitting



• Pixel-to-pixel error between 0.1% and 5%

## Running Time

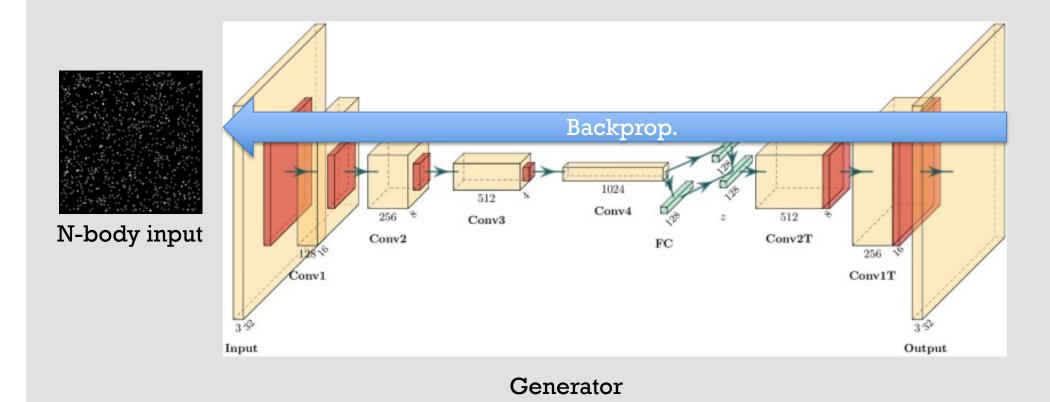
- 2.3 GHz 8-Core Intel Core i9:
  - GAN:
    - 0.19 ± 0.03 seconds per evaluation
  - Matrix Inversion:
    - 30.42 ± 5.80 seconds per evaluation
- ~150x speed up for 5% error



#### **Experimental Motivation**



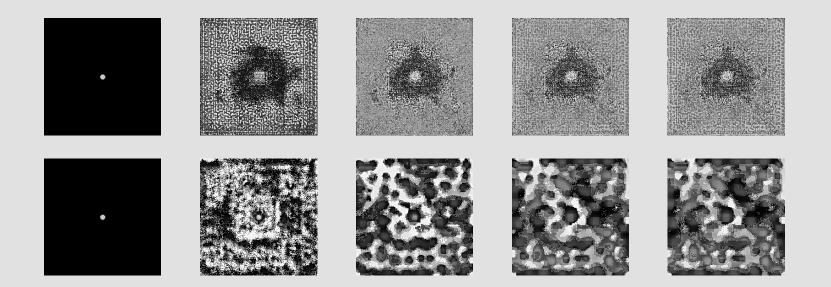
• Best part: we can backpropagate all the way through the input



#### Experimental Motivation & Results



- We choose a new loss function we would like to find a continuous, smooth mass distribution which maximizes observable 5th forces then update the input distribution  $\rho$  to optimize output
  - Help to direct experimental searches for chameleon / other scalar fields



Objective function: maximize mean of scalar field  $\Phi$ 

Objective function: minimize gradients in  $\rho$ , and maximize [gradients in  $\Phi$  + mean of  $\Phi$ ]

#### References



- Khoury, Justin; Weltman, Amanda (2004). "Chameleon cosmology". Physical Review D. 69 (4): 044026. arXiv:astro-ph/0309411
- T. P. Waterhouse: "An Introduction to Chameleon Gravity", 2006; arXiv:astroph/0611816
- Isola, P., Zhu, J.-Y., Zhou, T., & Efros, A. A. (2016). Image-to-image translation with conditional adversarial networks.

#### Thank you! I will take any questions at this time.