The Scotos (*dark*) side of Neutrinos

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DM and ν origin story

Origin of neutrino mass and dark matter are two pressing puzzles in $PPC \rightarrow$ can be addressed by hidden sector containing new particles/symmetries

Radiative connection: long history of radiative m_{ν} generation \rightarrow are DM and m_{ν} related radiatively?

Flavor/Horizontal symmetries: Non Abelian discrete symmetries still of continuous interest to explain PMNS

Complementarity: Use DD, Colliders, Precision to constrain DM/BSM physics and take advantage of that connection between DM- ν to probe nature of m_{ν}

The Dark side of neutrinos?

Two main frameworks to connect neutrinos and *DM* in literature:

(i) DM identified with one of new particles needed to give ν mass (Mass mechanism)

	Model	Scalars	Fermions	LFV	DM	LHC
1-Loop	Zee	$(1,1,+1)_{-2}, (1,2,+1/2)_0$		 ✓ 	X	1
	Ma	$({f 1},{f 2},+1/2)_0$	$({f 1},{f 1},0)_{+1}$	1	1	1
2-Loops	Zee-Babu	$({f 1},{f 1},+1)_{-2}, ({f 1},{f 1},+2)_{-2}$		1	X	1
3-Loops	KNT	$({f 1},{f 1},+1)_{-2}$	$({f 1},{f 1},0)_{+1}$	√	√	×

Table: Phenomenological implications of radiative $SU(3)_c \times SU(2)_L \times U(1)_Y$ neutrino mass models

(*ii*) DM associated to v sector symmetries (Lepton symmetries) i.e mixing patterns and/or lepton number

The Scotogenic model

Scotogenic Model = SM + 3 singlet Fermions (N_{1,2,3}) + scalar doublet (η) + dark Z₂

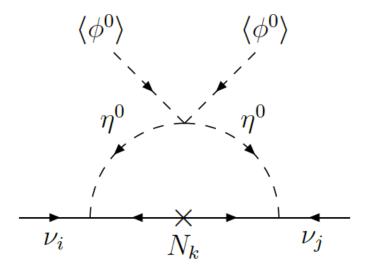
[Ma, <u>hep-ph/0601225]</u>

Dubbed the scotogenic model, from Greek word 'scotos' meaning darkness, this '06 proposal explains the smallness of m_{ν} via a loop-generated mass where DM plays a central role in completing the loop

- Particles in Loop odd under dark Z₂
- Majorana mass of N completes loop
- Mass splitting $(\lambda_5(\eta^{\dagger}\phi)^2)$ makes loop finite
- Either η_0 or *N* are *DM* candidate
- The yukawas that enters into m_{ν} generates

LFV processes however the bounds only constrain combination of yukawa coupling elements-

cancelations \rightarrow safety in Ma model



Challenges, goals and current framework

Discrete symmetry is usually put in by hand in Scotogenic theories Masses of RH neutrino has to be assumed to be at TeV scale

DM stability is arising from gauge symmetry Protect *DM* mass down to GeV-TeV scale Achieved by breaking the dark gauge symmetry around GeV-TeV scale

Anomaly-free solutions to SM with dark *U(1)* gauge symmetry Scotogenic neutrino mass: *LLHH* operator arise through loops Explore different *DM* phenomenology scenarios: (scalar *DM* (singlet-doublet mixture), Majorana/Dirac Fermion *DM*)

Anomaly free solutions with min chiral fields

- Hidden sector containing N SM-singlet chiral fermions χ(n_i) with nonzero U(1)_D charges n_i (i = 1, ..., N)
- All SM fields are neutral under $U(1)_D$
- Cancellation of gauge anomaly and mixed gauge-gravitational anomaly imposes the following constraints (Diophantine equations) on charge assignments

$$\sum_{i=1}^{N} n_i = 0$$
$$\sum_{i=1}^{N} n_i^3 = 0$$

Find **optimum** number of chiral fields (maximal charge ratio 6) + minimal scalar sector to obtain suitable Scotogenic solutions

Solutions with $N \leq 8$ fields

Π	Model	Chiral Fermions: Multiplicity*(Charge)	Minimal Higgs Sector: (Charge)
1	1	$3^{*}(1)+2^{*}(-4)+1^{*}(5)$	
	2	$1^{(1)+1^{(-2)+1^{(-3)+2^{(-5)+1^{(-6)}}}}$	
	3	$1^{(-1)+2^{(-2)+3^{(-2)}+3^{(-4)}}$	
	4	$2^{*}(1)+1^{*}(2)+1^{*}(3)+2^{*}(-4)+1^{*}(-5)+1^{*}(6)$	(2)
Í	5	$1^{(-1)+3^{(-2)}+1^{(3)}+1^{(5)}+1^{(6)}+1^{(-7)}}$	(4)
	6	$2^{(-1)+1^{(-2)+1^{(-3)}+1^{(-4)+1^{(-4)}+1^{(-3)}+1^{(-8)}}$	(5)
	7	$1^{(-1)+2^{(-2)+1^{(-4)+2^{(5)+1^{(7)+1^{(-8)}}}}$	(3)
	8	$1^{(1)+1^{(2)+2^{(3)+1^{(-4)+1^{(-6)+1^{(-8)+1^{(9)}}}}}$	(5)
	9	$1^{(2)+1^{(-3)+1^{(-4)}+1^{(-5)+1^{(-6)}+2^{(7)}+1^{(-8)}}}$	(1)
	10	$1^{(1)+1^{(2)}+1^{(3)}+1^{(5)}+2^{(-6)}+1^{(-9)}+1^{(10)}$	(4)
	11	$1^{(-1)+1^{(-2)+1^{(-4)+1^{(-5)+2^{(7)+1^{(8)+1^{(-10)}}}}}$	(3)
	12	$1^{(-1)+1^{(-2)+2^{(-4)+1^{(5)+1^{(7)+1^{(10)+1^{(-11)}}}}}$	(6)
	13	$2^{(1)+1^{(3)}+1^{(-4)}+1^{(6)}+1^{(-8)}+1^{(-10)}+1^{(11)}$	(7)
	14	$1^{(-1)+1^{(3)}+1^{(5)+1^{(6)}+1^{(-7)}+2^{(-8)}+1^{(10)}}$	(2)
	15	$1^{(-2)+2^{(-3)}+1^{(-4)}+1^{(5)}+1^{(9)}+1^{(10)}+1^{(-12)}}$	(7)
	16	$1^{(3)+1(-4)+1(-5)+1(6)+1(-8)+2(9)+1(-10)}$	(1)
	17	$1^{(-1)+1^{(-2)+1^{(-4)}+1^{(-6)}+2^{(7)+1^{(11)}+1^{(-12)}}}$	(5)
	18	$2^{(-2)+1^{(-3)+1^{(-5)+1^{(-6)+1^{(10)+1^{(-13)}+1^{($	(8)
	19	$1^{(1)+1^{(2)+2^{(5)+1^{(-6)+1^{(-8)+1^{(-12)+1^{(13)}}}}}$	(7)
	20	$2^{(-1)+1^{(-4)+1^{(-5)+1^{(-8)+1^{(-10)+1^{(-13)+1^{(-14)}}}}$	(9)
	21	$1^{(1)+1^{(2)}+1^{(5)}+1^{(7)}+2^{(-8)}+1^{(-13)}+1^{(14)}$	(6)
	22	$1^{(-2)+1^{(-3)}+1^{(-4)}+1^{(-7)}+2^{(9)}+1^{(-12)}+1^{(-14)}}$	(5)
	23	$1^{(1)+1^{(3)}+1^{(6)}+1^{(7)}+2^{(-10)}+1^{(-11)}+1^{(14)}$	(4)
	24	$1^{(-1)+1^{(-2)+2^{(-6)+1^{(7)}+1^{(9)+1^{(14)+1^{(-15)}}}}$	(8)
	25	$1^{(4)+1^{(-5)}+1^{(-6)}+1^{(-1)}+1^{(-10)}+2^{(11)}+1^{(-12)}}$	(1)
	26	$1^{(1)+1^{(-4)+1^{(-7)+1^{(-8)+1^{(-10)+2^{(11)+1^{(-14)}}}}}$	(3)
	27	$1^{(-2)+1^{(-4)+2^{(-5)+1^{(7)+1^{(11)+1^{(14)+1^{(-16)}}}}}$	(9)
	28	$1^{(-1)+1^{(-2)}+1^{(-6)}+1^{(-8)}+2^{(9)}+1^{(-15)}+1^{(-16)}}$	(7)
	29	$1^{(-3)+1^{(5)}+1^{(7)}+1^{(-9)}+1^{(10)}+2^{(-12)}+1^{(14)}$	(2)
	30	$1^{(-3)+2^{(-4)}+1^{(-6)}+1^{(7)}+1^{(13)}+1^{(14)}+1^{(-17)}}$	(10)

 $\textbf{Table 1: } U(1)_D \subset SU(3) \times SU(2)$

Representative model

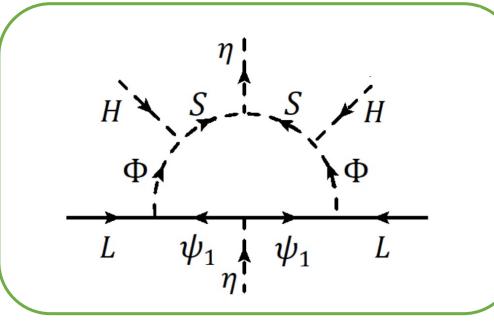
Anomaly-free charge assignments (multiplicity*(charge))

Model: $2 \times \{1\} + 2 \times \{-4\} + 1 \times \{2\} + 1 \times \{3\} + 1 \times \{-5\} + 1 \times \{6\}$

Two Weyl fermions with $U(1)_D$ charges 1 and -4, one state each with charge 2, 3, -5 and 6 Minimally extended Higgs sector:

scalar singlet field η (1, 0, 2), singlet S (1, 0, 1), doublets H (2, $\frac{1}{2}$, 0) and $\Phi(2, \frac{1}{2}, 1)$

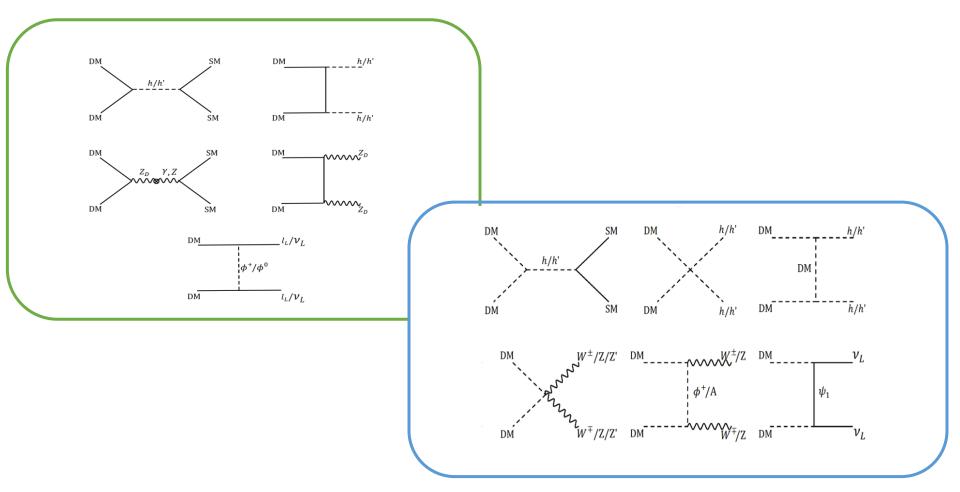
Topology for 1 loop generation of m_{ν}



Lightest field inside the loop could be DM

Dark Matter

 $U(1)_D$ gauge symmetry broken down to discrete Z_2 ensuring the stability of *DM* I) Fermionic *DM* (Dirac and Majorana) II) Scalar *DM* of singlet-doublet mixture



Ref: Relevant Feynman diagrams that contribute to the annihilation of the Fermion DM (L) and Scalar DM (R)

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Fermionic DM phenomenology(Majorana)

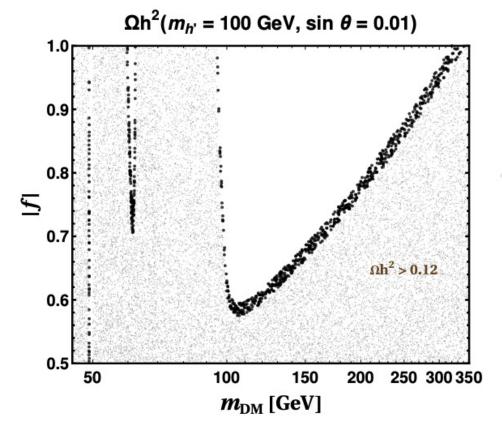


Fig: Parameter space in Yukawa coupling (f) vs. m_{DM} plane consistent with Planck constraint for Majorana *DM*



Yukawas entering LFV process are distinct from yukawas needed here to satisfy DM bounds

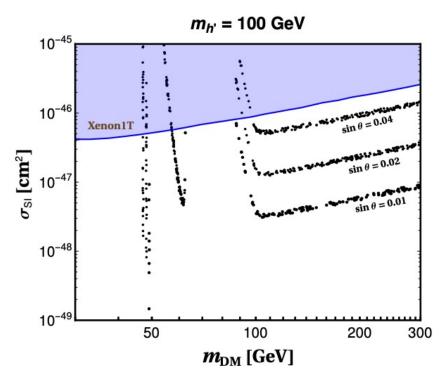


Fig: Spin independent cross section vs. m_{DM} consistent with DD constraints for Majorana DM

Fermion DM phenomenology (Dirac)

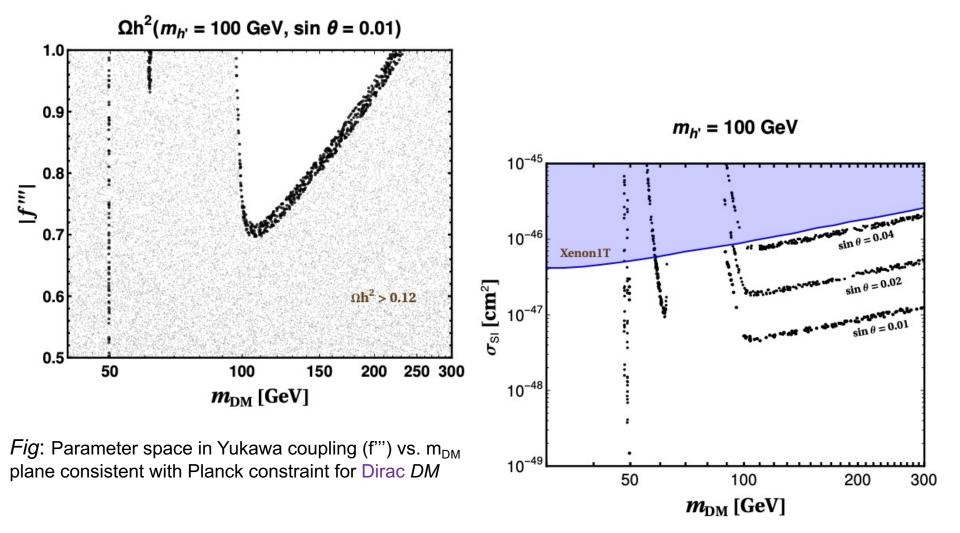


Fig: Spin independent cross section vs. m_{DM} consistent with DD constraints for Dirac DM

Fermion DM phenomenology (gauge portal)

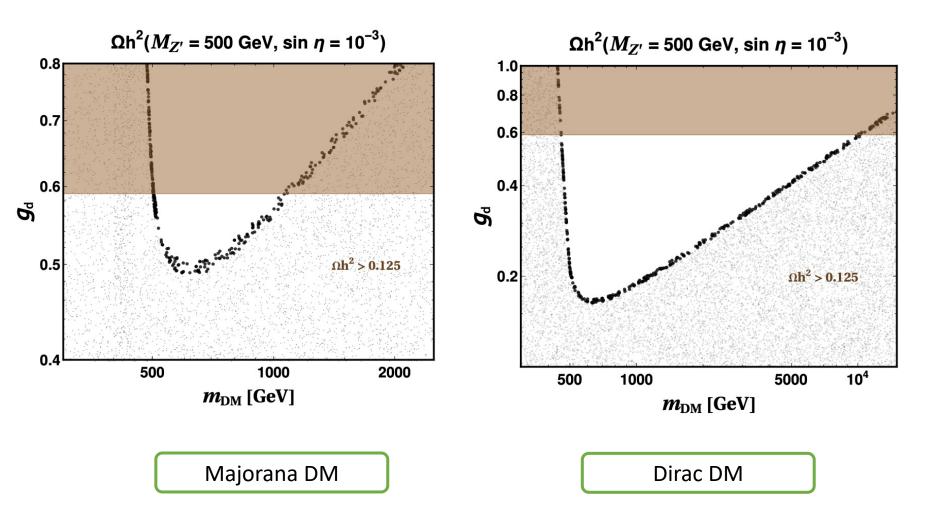


Fig: Parameter space in gauge coupling (g_d) vs. m_{DM} plane consistent with DM relic density constraint. $m_{H'}$ is 2.1× m_{DM} and $sin\eta$ represents kinetic mixing

Fermionic DM phenomenology (low mass)

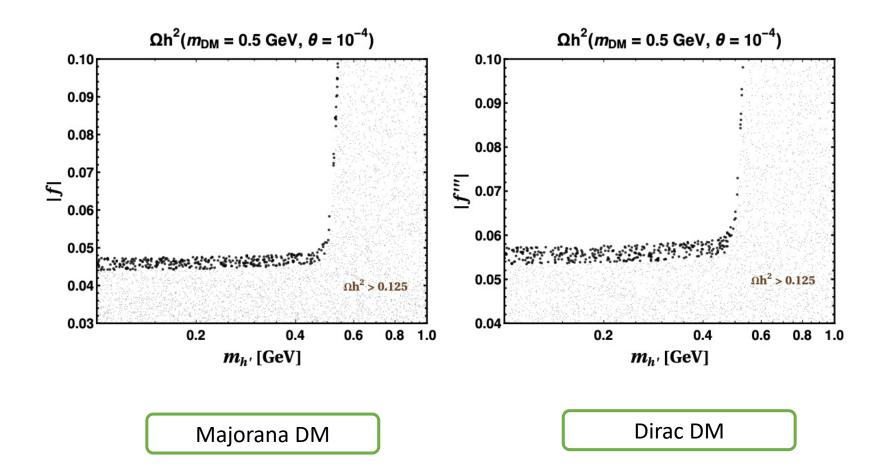


Fig: Parameter space in Yukawa coupling (f, f ") vs. scalar mass (m_h) plane consistent with the Planck constraint for the Majorana (left) and Dirac *DM* for low mass region

Scalar DM phenomenology

Scalar DM consists of singlet-doublet mixture

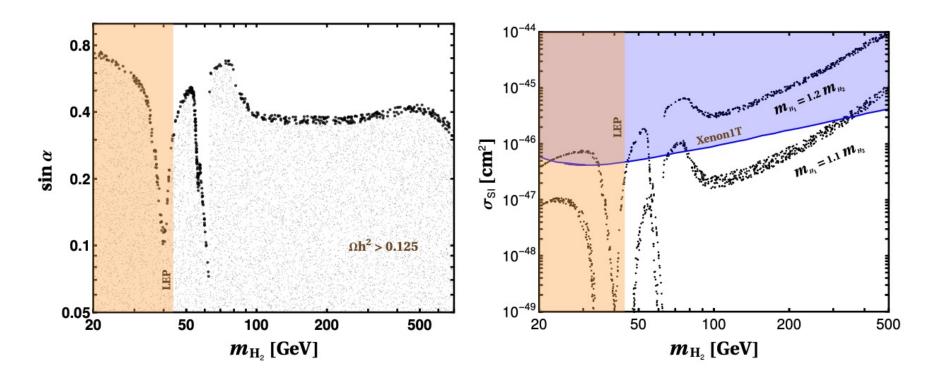


Fig (L): Parameter space in mixing angle sin α (mixing for CP even scalar fields H_1 and H_2) vs. m_{DM} plane for m_{H1}=1.1 (1.2)×M_{DM} consistent with Planck constraint for scalar *DM*. Fig (R): *Spin independent* cross section vs. m_{DM} consistent with LEP bound and Xenon1T bound

Summary

Present complete set of relevant anomaly-free solutions with add^n of $U(1)_D$ chiral fermions with minimalized scalar sector to address *DM* and m_v

Model of the dark sector includes m_{ν} and *DM* originating from Scotogenic mechanism

DM stability is protected by the residual of $U(1)_D$ symmetry with the scalar singlet η generating masses to the all dark fermions

DM phenomenology explored for both fermionic type (Majorana and Dirac) and scalar type (singlet-doublet) with DM masses ranging from 100 MeV to TeV scale

Thank you! Questions?

