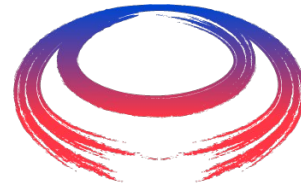


The Muon Collider Lecture 1



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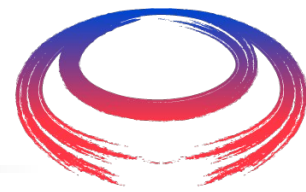
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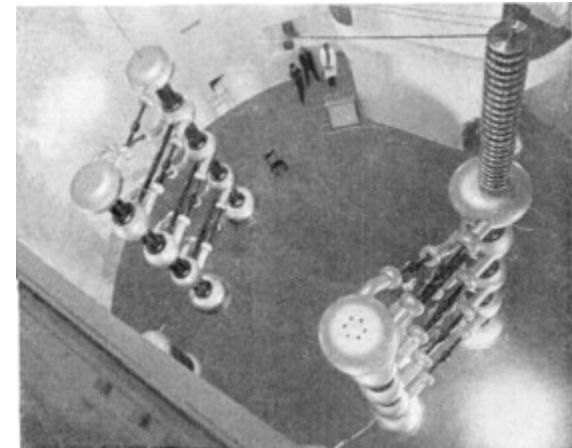
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Accelerators in Physics



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- First accelerators built in 1920s/30s
 - Accelerating protons, ions and electrons
 - Positrons in 1960s
 - Antiprotons in 1980s
- Tools for fundamental physics
- Hadron colliders
 - E.g. LHC
 - “Discovery machines”
- Electron positron colliders
 - E.g. Large Electron Positron Collider (LEP)
 - “Precision machines”
- Growing interest in building muon collider
 - Muons first accelerated in 2017 – new tech
 - Why muons? How?



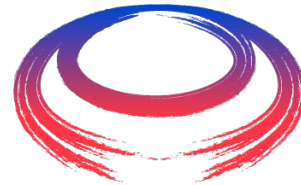
Muon Collider



- Lecture 1
 - Why are muon collisions interesting?
 - What are the ingredients required to make muons?
 - Production of low emittances
- Lecture 2
 - Rapid acceleration to fight muon lifetime
 - Extremely low β^* at the focus
 - Experimental demonstration and staging
 - how to make it happen



Why Muons?



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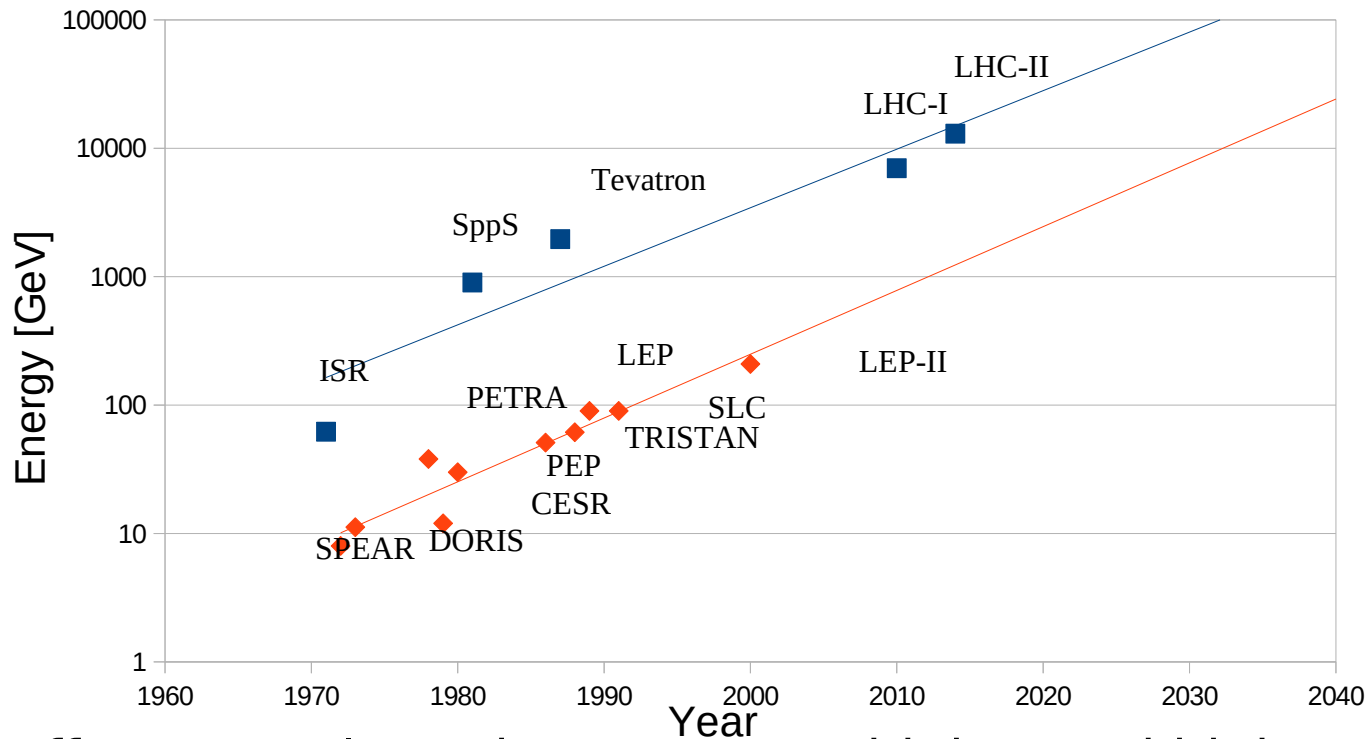
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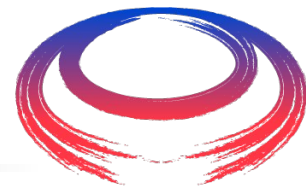
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Back to the Future...



- Effort to explore phenomena at higher and higher energies
- Corresponds to smaller scales
- Higher energy → bigger, more expensive, more power hungry

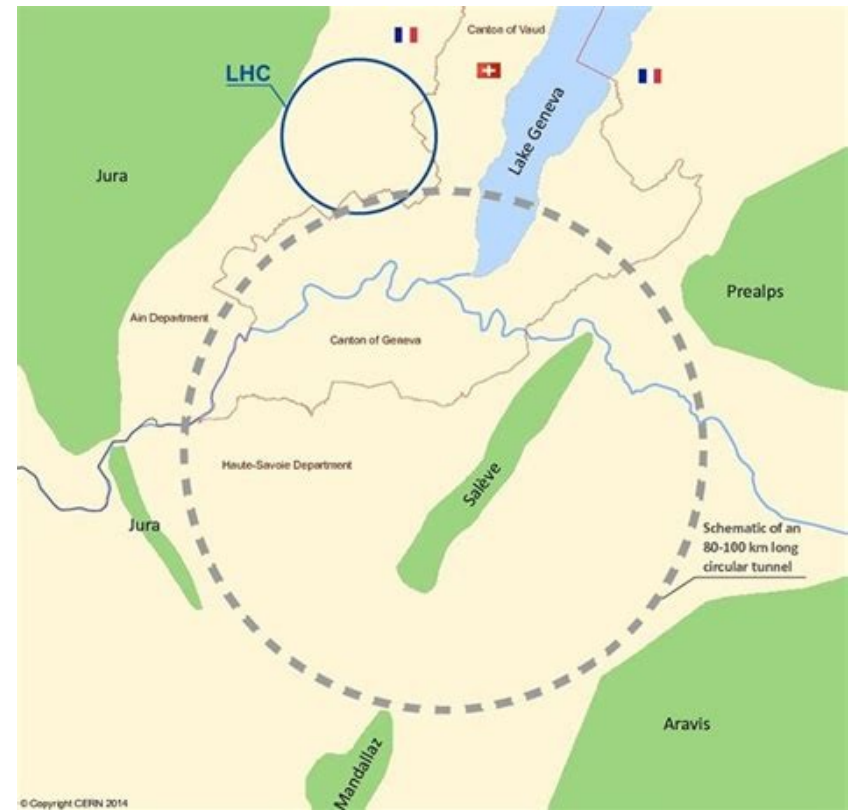
E.g. circular colliders



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- Tevatron
 - 1.96 TeV proton - antiproton
 - 6.2 km circumference
- LEP/LHC
 - 14 TeV proton proton (LHC)
 - 209 GeV e^+e^- (LEP)
 - 27 km circumference
- FCC (proposed)
 - 90 - 350 GeV e^+e^-
 - 100 TeV proton-proton
 - 90-100 km circumference

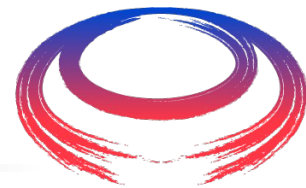
○ Tevatron



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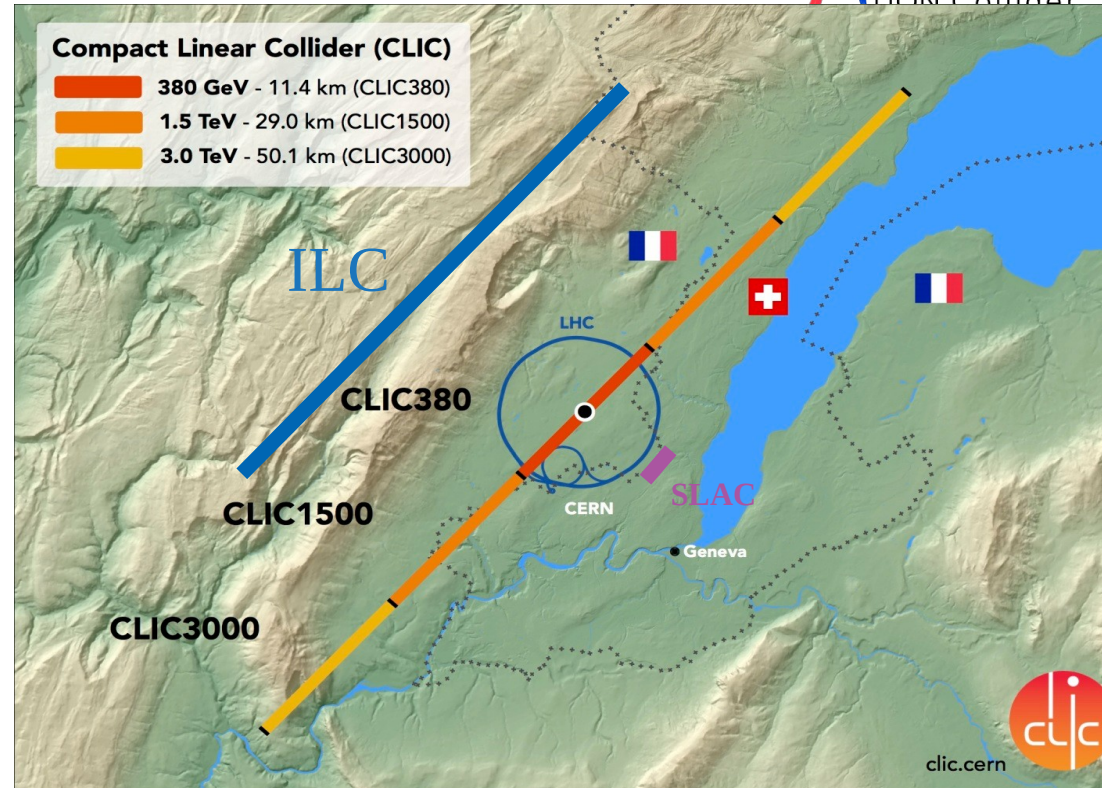
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E.g. linear colliders



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Linear Collider

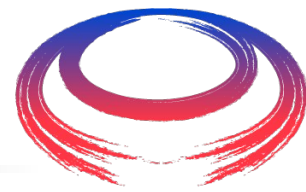
- SLAC (California)
 - 3 km length
 - 90 GeV e^+e^-
- ILC (proposed)
 - 31 km
 - 500 GeV e^+e^-
- CLIC (proposed)
 - 380 GeV e^+e^-
 - 11 km



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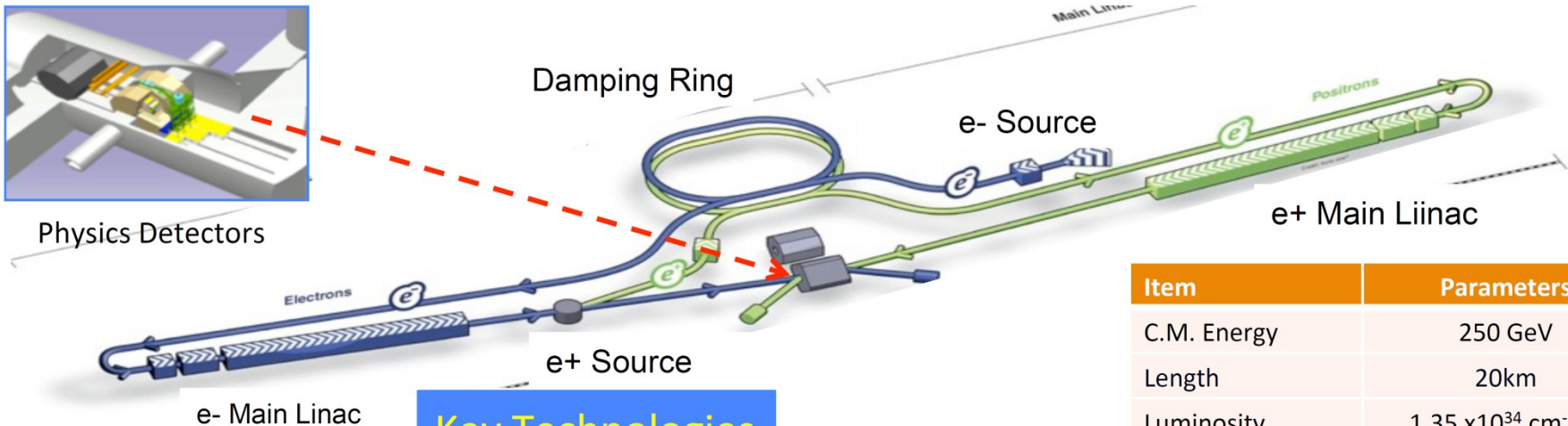
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Electron-positron colliders

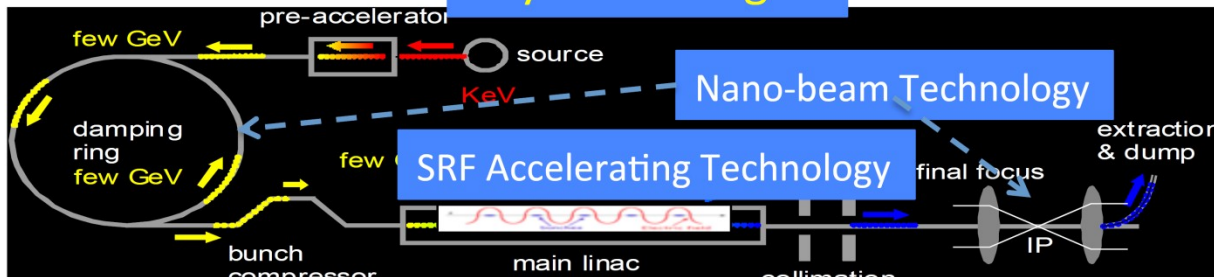


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- Circular machines limited by synchrotron radiation
 - Power emitted $\sim E^4/m^4$
 - Practically limits centre-of-mass energy to \sim low 100s GeV
- Linear machines limited by available RF acceleration
 - Practically limits centre-of-mass energy to \sim 100s GeV (TeV)

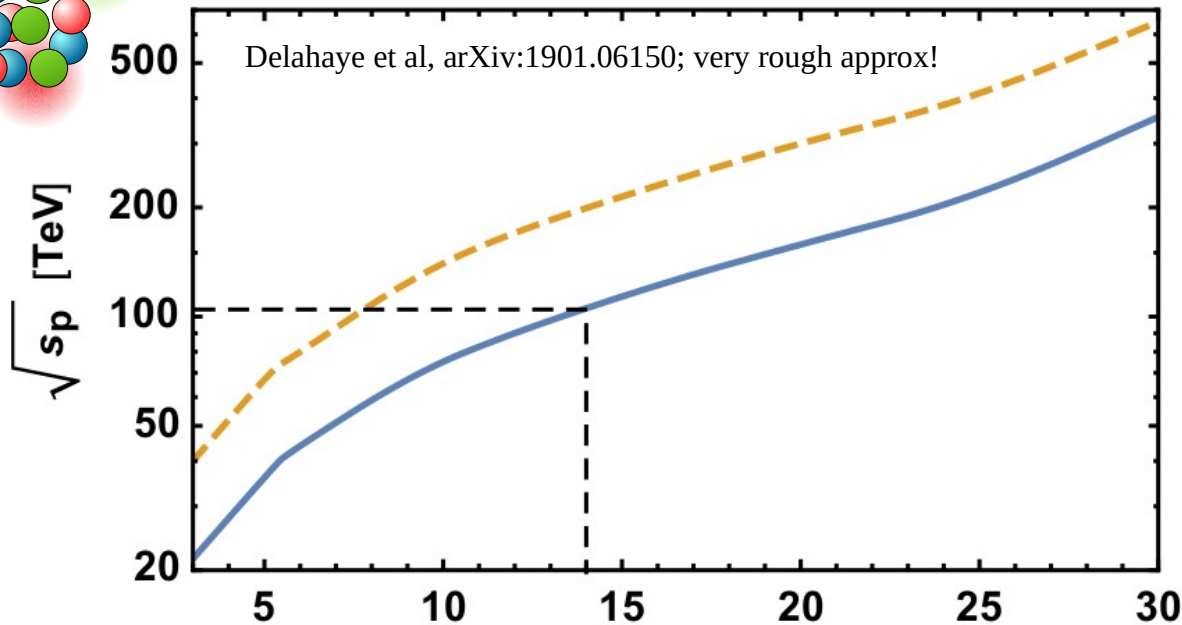
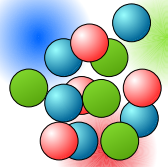


Key Technologies



Item	Parameters
C.M. Energy	250 GeV
Length	20km
Luminosity	$1.35 \times 10^{34} \text{ cm}^{-2}\text{s}^{-1}$
Repetition	5 Hz
Beam Pulse	0.73 ms
Beam Current	5.8 mA (in pulse)
Beam size (y) at FF	7.7 nm@250GeV
SRF Cavity G.	31.5 MV/m

Muons Physics Reach



Energy at which
cross-section is equal

- Assuming equal Feynman amplitude (EW)
- Assuming factor 10 enhancement in pp (EW+QCD)

$\sqrt{s_\mu}$ [TeV] ●

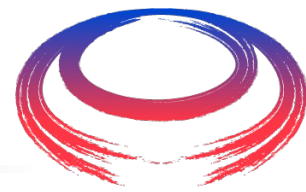
- Seek a particle which
 - Is not so low mass as an electron
 - Is a fundamental particle
- **Muons!**

Muons



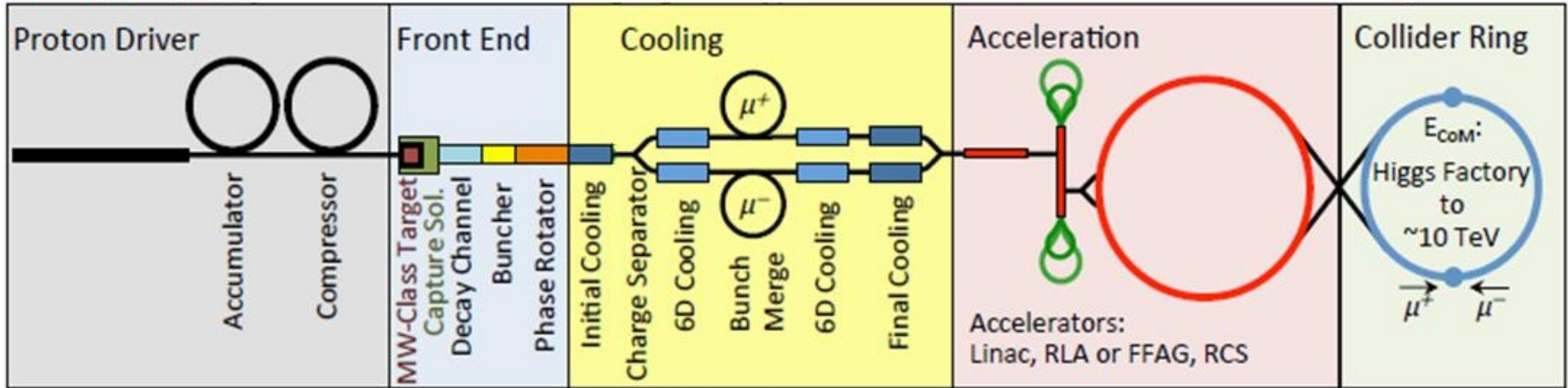
- Muon
 - Half-life 2.2 μs
 - Mass 105.658 MeV/c
 - 207 times electron mass
- What would a muon collider look like?

Muon Collider



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MUON Collider

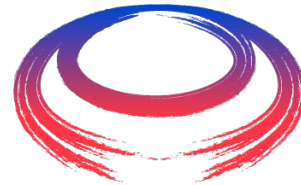
Muon Collider



- MW-class proton driver \rightarrow target
- Pions produced; decay to muons
- Muon capture and cooling
- Acceleration to TeV & Collisions
- Designed for high energy while **maximising luminosity**
 - Luminosity is key



It's All About Luminosity



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Collaboration

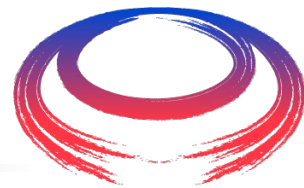
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Luminosity



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- **Luminosity** is key challenge
 - Number of events per cross-section per time
 - Diffuse beam → low chance of particles colliding → low luminosity
- Change in integrated luminosity per beam j^{th} crossing

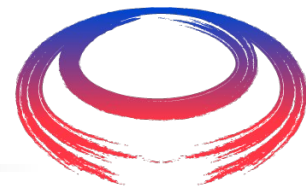
$$\Delta \mathcal{L} = \frac{N_{+,j} N_{-,j}}{4\pi \sigma_{\perp}^2}$$

$N_{+/-}$ = Number of μ^+ or μ^- on j^{th} crossing

σ_{\perp} = size of the beam in x/y
... assume cylindrical symmetry

- What drives luminosity? Can we relate luminosity to
 - Repetition rate of accelerator
 - Efficiency of muon creation
 - Proton beam parameters
 - Etc

Luminosity - stored muons (1)



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$$\Delta \mathcal{L} = \frac{N_{+,j} N_{-,j}}{4\pi\sigma_{\perp}^2}$$

- The number of particles is always falling due to muon decay

Number of particles entering collider

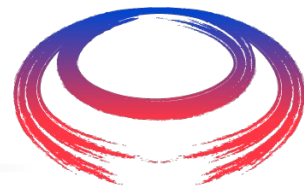
Distance between collisions i.e. collider circumference

$$N_{\pm,j} = N_{\pm} \exp\left(\frac{-2\pi Rj}{c\gamma T_{\mu}}\right)$$

Muon speed

Muon lifetime in the lab

Luminosity – stored muons (2)



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- Luminosity (change in integrated luminosity per time)

$$\mathcal{L} = f_r n_b \sum_{j=0}^{\infty} \Delta \mathcal{L}$$

Number of bunches per acceleration cycle

Repetition rate (number of acceleration cycles per second)

$$\Delta \mathcal{L} = \frac{N_{+,j} N_{-,j}}{4\pi\sigma_{\perp}^2}$$

$$N_{\pm,j} = N_{\pm} \exp\left(\frac{-2\pi R j}{c\gamma\tau_{\mu}}\right)$$

$$\mathcal{L} = f_r n_b \frac{N_{+} N_{-}}{4\pi\sigma_{\perp}^2} \sum_{j=0}^{\infty} \exp\left(-\frac{4\pi R}{\gamma c\tau_{\mu}} j\right)$$

Luminosity – stored muons (3)

$$\mathcal{L} = f_r n_b \frac{N_+ N_-}{4\pi\sigma_{\perp}^2} \sum_{j=0}^{\infty} \exp\left(-\frac{4\pi R}{\gamma c \tau_{\mu}} j\right)$$

- Assuming muon lifetime is long compared to ring time-of-flight

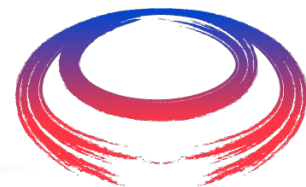
$$\mathcal{L} \approx f_r n_b \frac{N_+ N_-}{(4\pi)^2 \sigma_{\perp}^2} \frac{\gamma c \tau_{\mu}}{R}$$

$R = p/(e\bar{B}) \approx \gamma m_{\mu} c / (e\bar{B})$

- So

$$\mathcal{L} \approx f_r n_b \frac{N_+ N_-}{(4\pi)^2 \sigma_{\perp}^2} \frac{\tau_{\mu} e \bar{B}}{m_{\mu}}$$

Luminosity – Facility efficiency



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$$\mathcal{L} \approx f_r n_b \frac{N_+ N_-}{(4\pi)^2 \sigma_{\perp}^2} \frac{\tau_{\mu} e \bar{B}}{m_{\mu}}$$

$$N_{\pm} = \frac{\eta_{\tau} \eta_{\pm} P_p}{n_b f_r}$$

Efficiency of
muon acceleration

Number of muons per
proton beam power



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Luminosity - σ_{\perp} (1)

$$\mathcal{L} \approx f_{\tau} n_b \frac{N_+ N_-}{(4\pi)^2 \sigma_{\perp}^2} \frac{\tau_{\mu} e \bar{B}}{m_{\mu}}$$

Transverse beam size

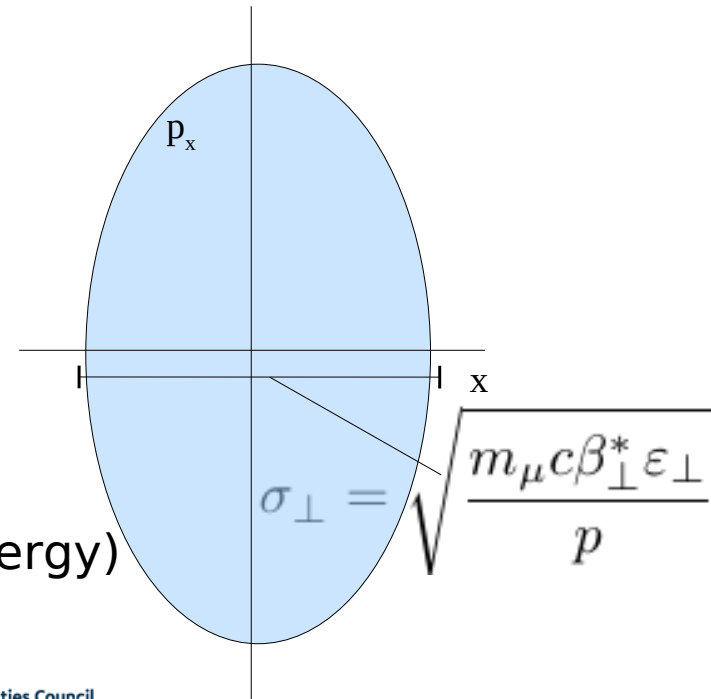
- Definition of emittance

$$\sigma_{\perp} = \sqrt{\frac{m_{\mu} c \beta_{\perp}^* \epsilon_{\perp}}{p}}$$

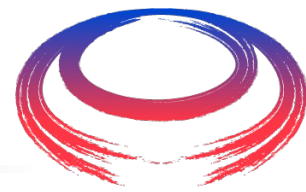
Twiss function

Emittance

- Need very tight focusing!
- Limits:
 - Focusing strength of magnets
 - Chromaticity (focusing depends on energy)
 - Hour glass effect (next slide)



Luminosity - σ_{\perp} (2)



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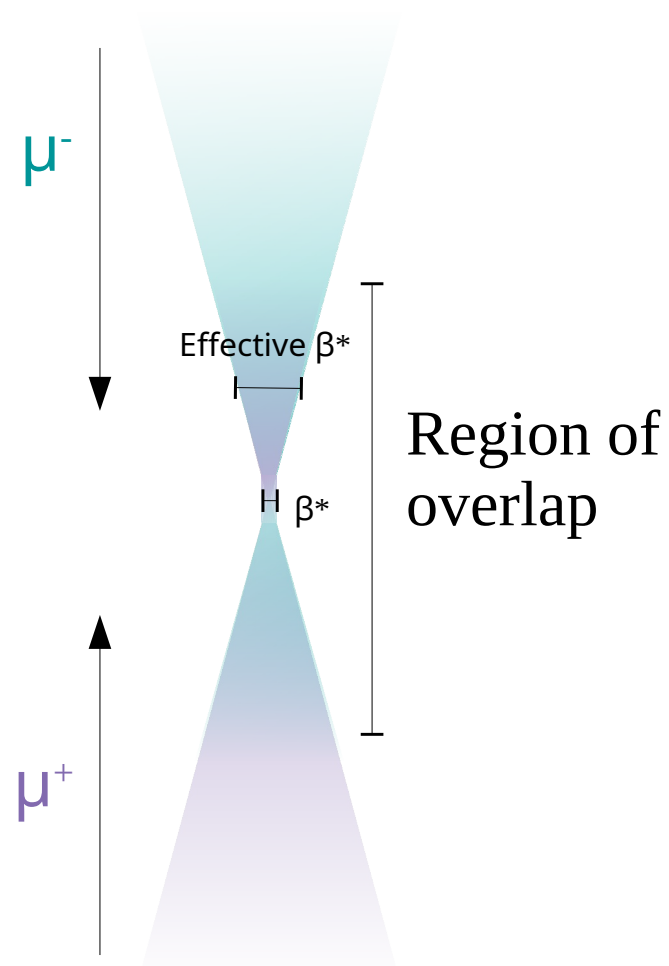
- Hour glass effect
 - Even for collider ring with super small β^*
 - Small β^* means short focal length
 - Region of overlap is very short
 - Bunch needs to be short as well!

- Introduce hour-glass factor f_{hg}
 - Relates the effective to lattice β^*

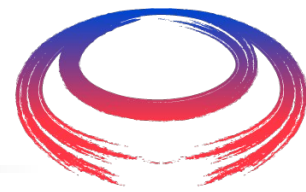
$$\beta_{eff}^* = \beta^* f_{hg}$$

- If $\sigma_z = \beta^*$
 - Hour-glass factor is 0.76

$$\sigma_{\perp} = \sqrt{\frac{m_{\mu} c \sigma_z \epsilon_{\perp}}{p f_{hg}}}$$



Luminosity - σ_{\perp} (3)



- Definition of longitudinal emittance

$$\epsilon_l = \gamma m_{\mu} c^2 \sigma_{\delta} \sigma_z$$

- So

$$\sigma_{\perp} = \sqrt{\frac{m_{\mu} c \sigma_z \epsilon_{\perp}}{p f_{hg}}}$$

$$\sigma_{\perp}^2 = \frac{\epsilon_l \epsilon_{\perp}}{p f_{hg} \gamma c \sigma_{\delta}}$$

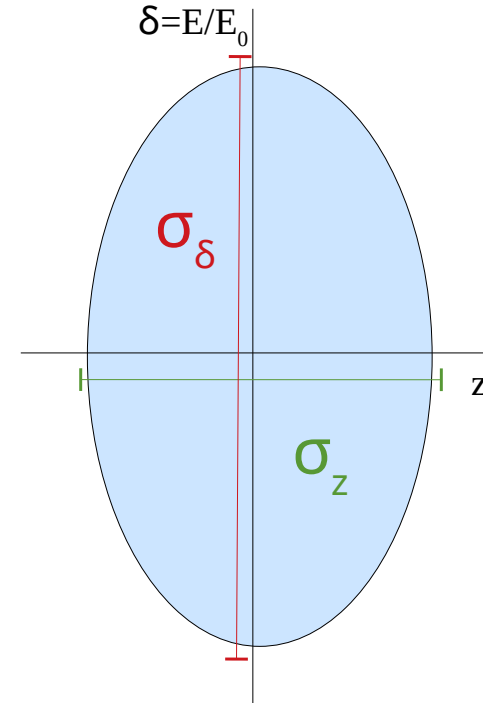
- Recalling the expression for luminosity and N

$$\mathcal{L} \approx f_r n_b \frac{N_+ N_-}{(4\pi)^2 \sigma_{\perp}^2} \frac{\tau_{\mu} e \bar{B}}{m_{\mu}}$$

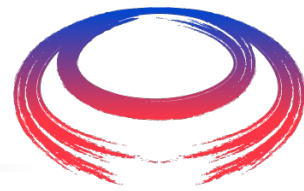
$$N_{\pm} = \frac{\eta_{\tau} \eta_{\pm} P_p}{n_b f_r}$$

- Bringing everything together

$$\mathcal{L} \approx \underbrace{\frac{e \tau_{\mu}}{(4\pi m_{\mu} c)^2}}_{K_L} \frac{f_{hg} \sigma_{\delta} \bar{B}}{\epsilon_{\perp} \epsilon_L n_b f_r} \underbrace{\eta_+ \eta_- (\eta_{\tau} P_p \gamma m_{\mu} c^2)^2}_{P_+ P_-}$$



Recap



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$$\mathcal{L} \approx \underbrace{\frac{e\tau_\mu}{(4\pi m_\mu c)^2}}_{K_L} \underbrace{\frac{f_{hg}\sigma_\delta \bar{B}}{\varepsilon_\perp \varepsilon_L n_b f_r}}_{\substack{4 \\ 3}} \underbrace{\eta_+ \eta_- (\eta_\tau P_p \gamma m_\mu c^2)^2}_{P_+ P_-} \quad \substack{2 \\ 4 \\ 5 \\ 1}$$

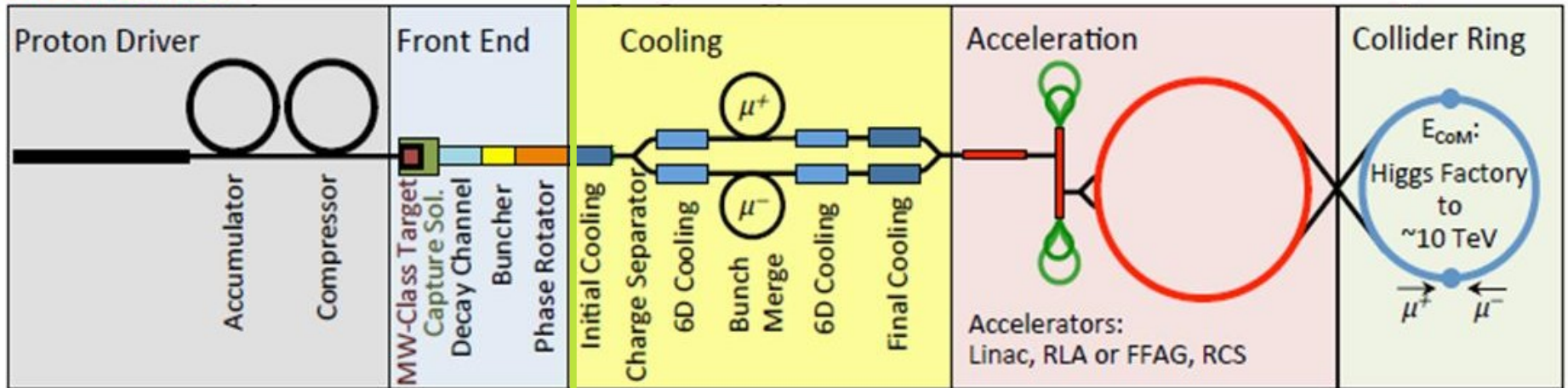
- 1) Luminosity increases with the square of muon energy/power
 - Number of collisions per bunch increases as muon lifetime increases
 - Beam size decreases as energy increases (geometric emittance)
 - 2) High field, low circumference collider ring → more luminosity
 - Shorter path length, more collisions before muon decay
 - 3) Low repetition rate, few bunches is best
 - Assume that the bottleneck is in the number of protons
 - Fewer collisions, but each collision is more intense
 - 4) High quality muon source is essential
 - Low emittance, good capture efficiency
 - 5) Good efficiency acceleration is essential
 - High voltage systems
- The whole muon collider is designed to maximise luminosity!

The Facility - From protons to muons

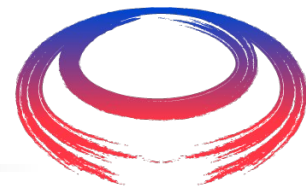


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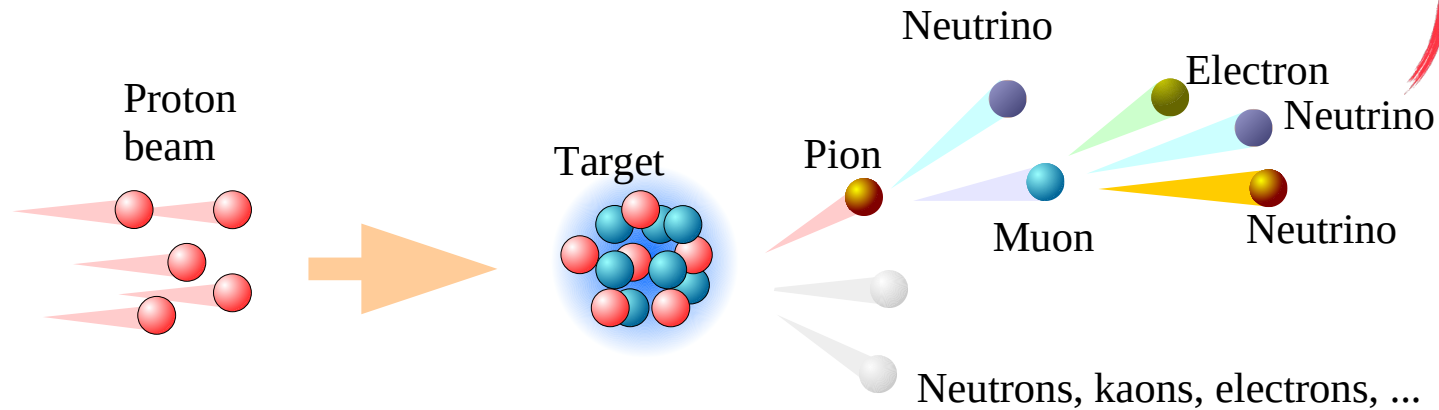
Muon Collider



Artificial Muons



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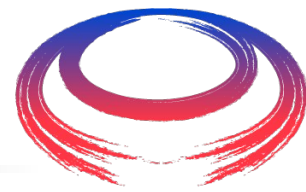
- Muons produced by putting protons onto target
- Pions come out
- Pions decay radioactively to muons
- Enables an intense muon source



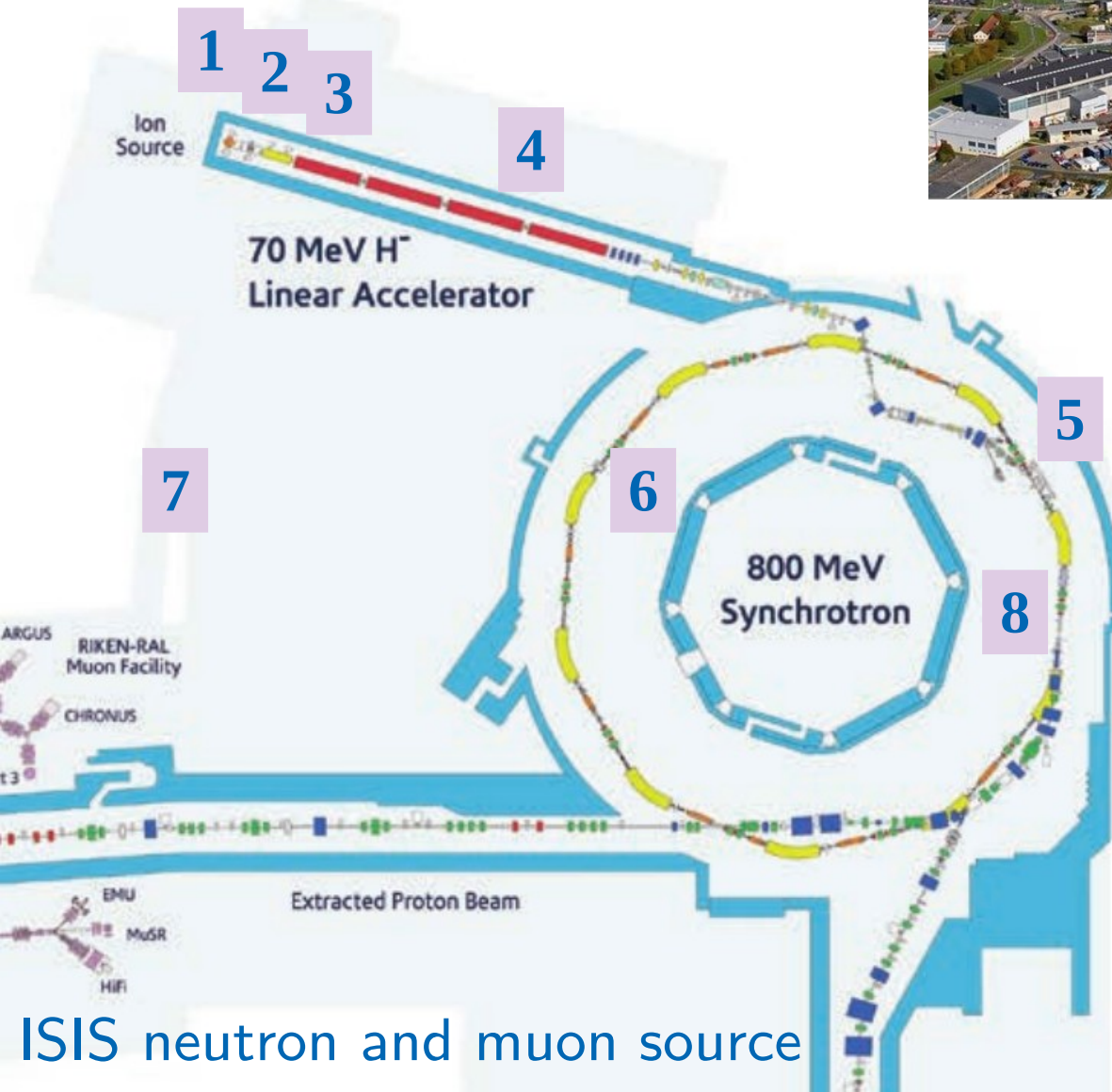
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Proton Source



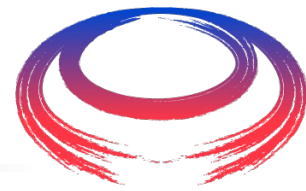
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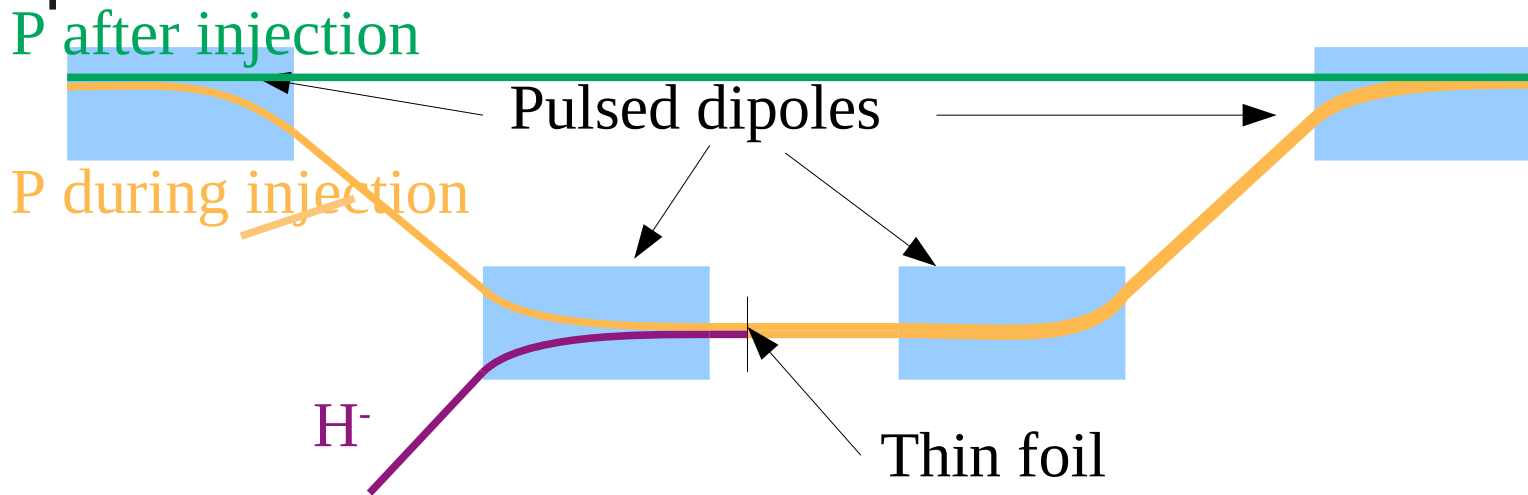
- 1) Ion source: spark across H gas to make H⁻ ions
- 2) Accelerate and focus in Radiofrequency Quadrupole
- 3) Chop into pulsed beam using fast/slow kicker
- 4) Accelerate in linac
- 5) Inject into a ring through a foil
- 6) Accelerate some more (maybe)
- 7) Compress the proton bunch to very short length
- 8) Extract and bring onto a target

ISIS neutron and muon source

Charge Exchange Injection

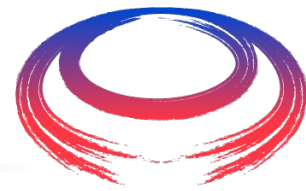


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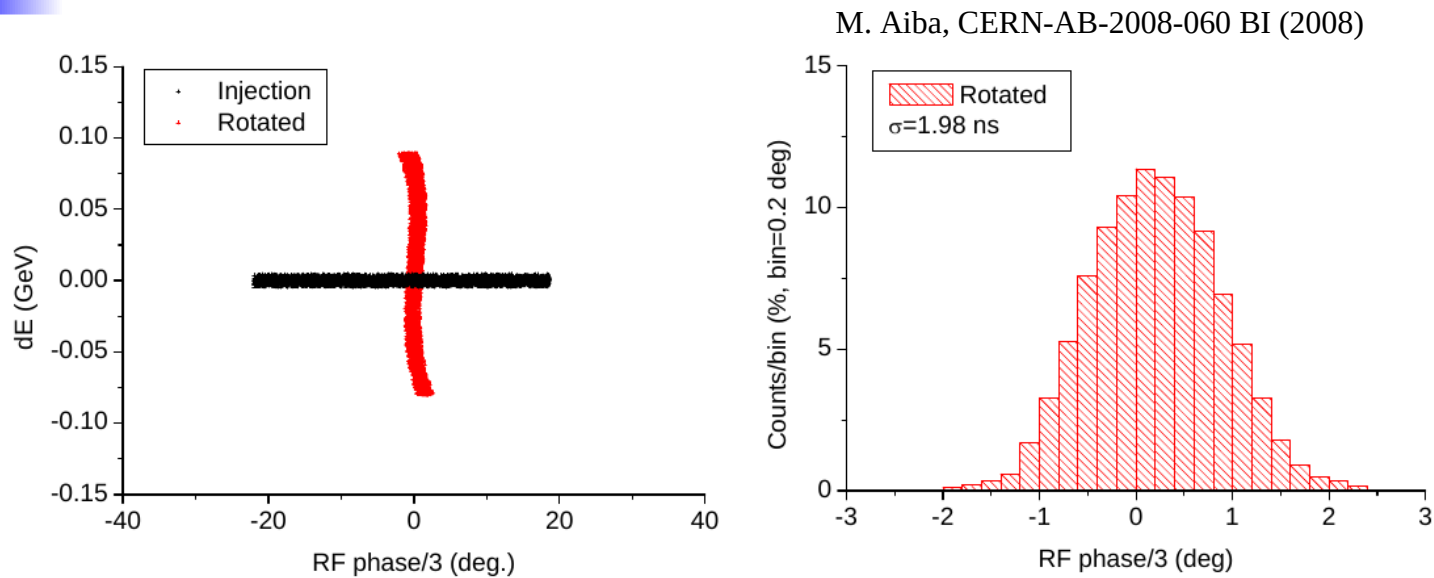


- High current → accumulate beam over many turns
 - Charge exchange injection of H^- ions through a thin foil
 - Foil removes electrons
 - Issues: Scattering and energy loss of protons in foil
- Painting of beam into synchrotron acceptance using fast “bumper” magnets
 - Move recirculating/injected beam phase space
- Foil lifetime is critical limit
- Space charge at injection is critical limit

Bunch Compression

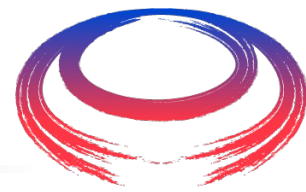


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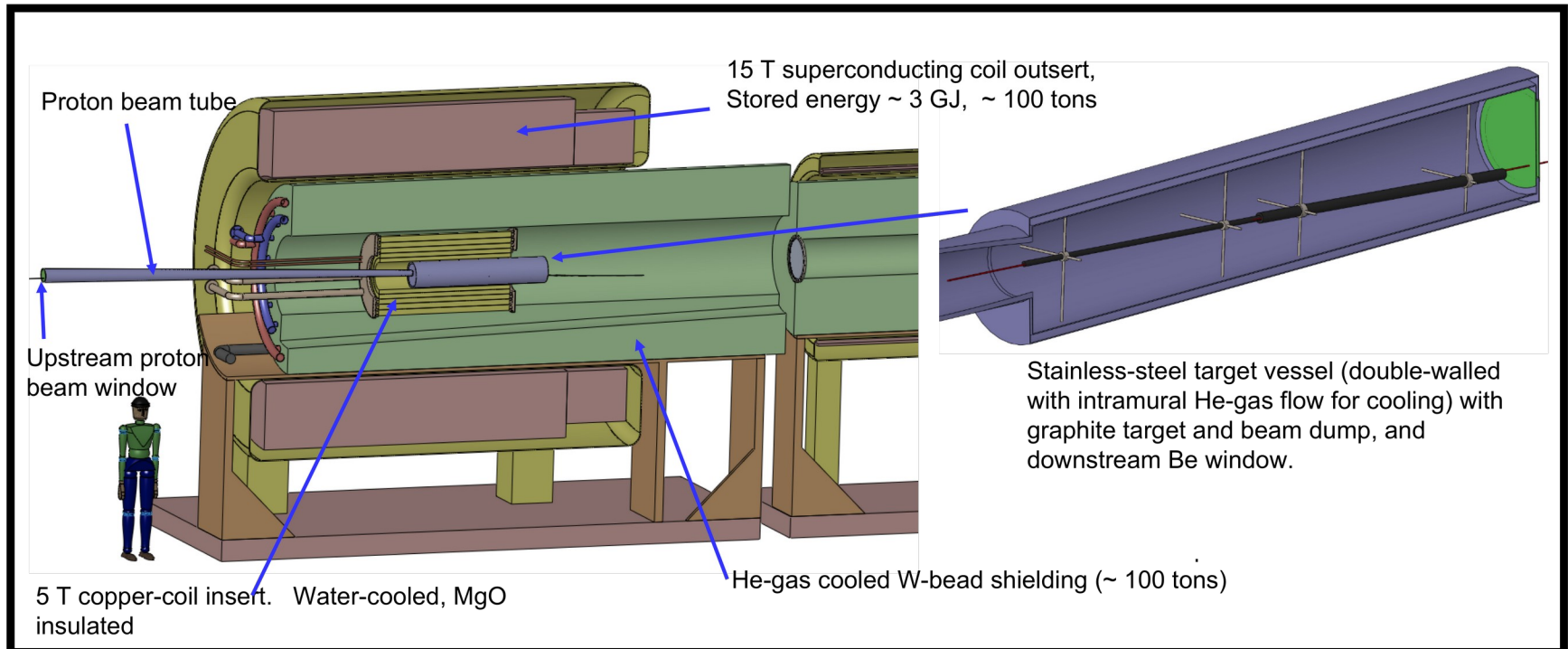
- Aim is to rotate the beam in longitudinal phase space
 - Short proton bunch \rightarrow short muon bunch
 - Reduce longitudinal emittance of the muons
- Achieve bunch compression by rotation in the RF bucket
- Limitations:
 - Microwave instability \rightarrow higher energy
 - Space charge \rightarrow higher energy

MC Target



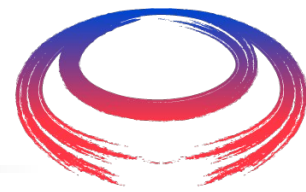
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X. Ding et al, Carbon and Mercury target system for muon colliders and neutrino factories, IPAC16



- Protons on target → pions → muons
 - Heavily shielded, very high field solenoid captures π^+ and π^-
- Challenge: Energy deposition on solenoid
- Challenge: Solid target lifetime

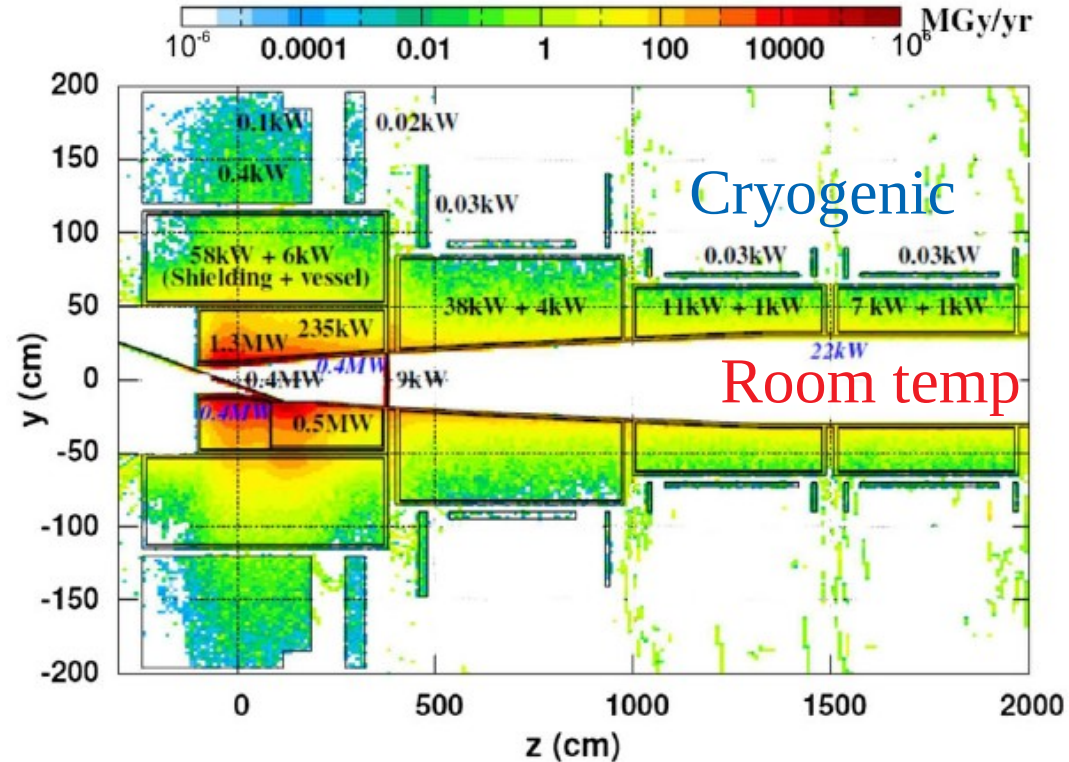
Radiation issues (magnet)



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- Radiation load significant issue
 - Degrades insulation/glue
 - Requires more cooling
 - 1 kW heat → O(200) kW electricity
- Shield at room temperature
- Magnet at superconducting temperature
 - HTS → warmer, more efficient

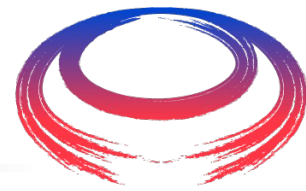
Neutrino factory, Bogomilov et al, PRSTAB 17 (2014)



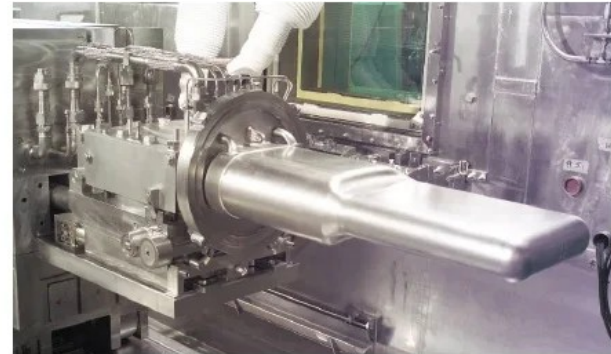
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Radiation issues (target)

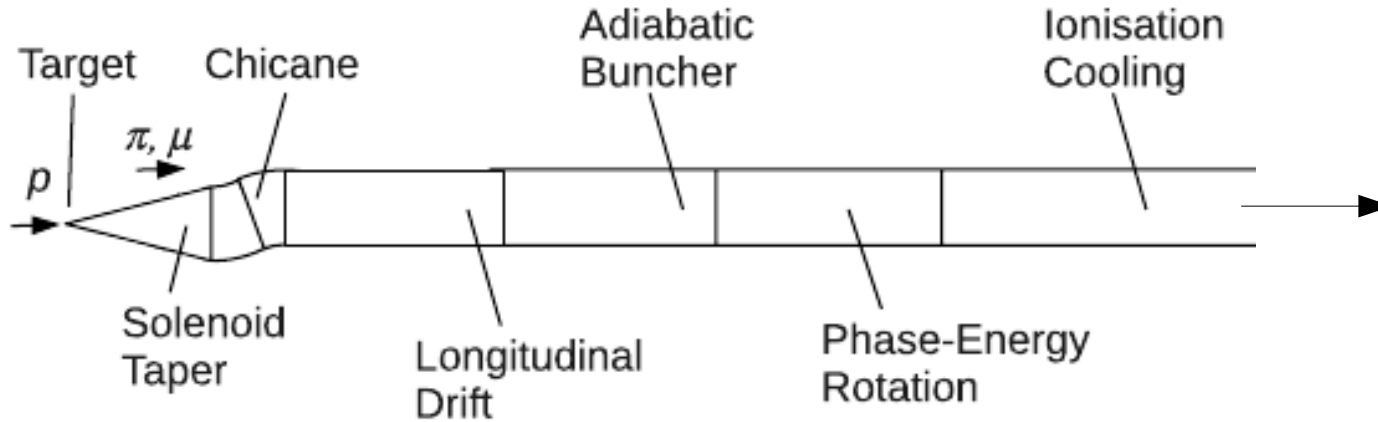


- Radiation on target can make an issue
 - Instantaneous shock
 - Long term radiation damage
- Liquid metal targets (Pb)
 - Cavitation issues
 - Specific issues around Hg
- Flowing/moving solid targets
 - Geometry issues
 - Target wheels - e.g. PSI
 - Fluidised powder



Parameter	CNGS	Muon Collider 1.5MW
Proton fluence [p+/cm ²]	5.77E+22	1.70E+21
PoT	1.27E+20	1.32E+21
Beam size [mm]	0.53	5
Extractions	5.29E+06	5.51E+07
Integrated Op time [days]	183	128
DPA	1.5	

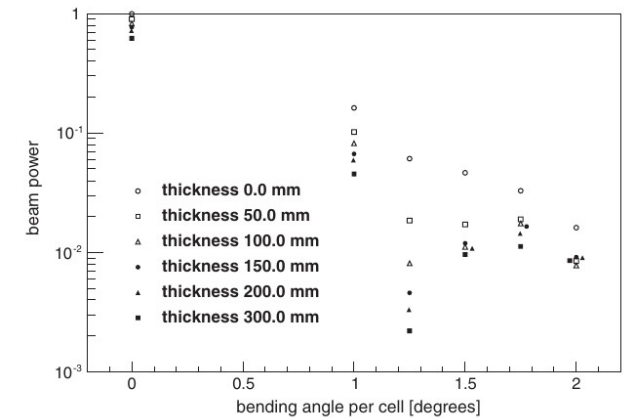
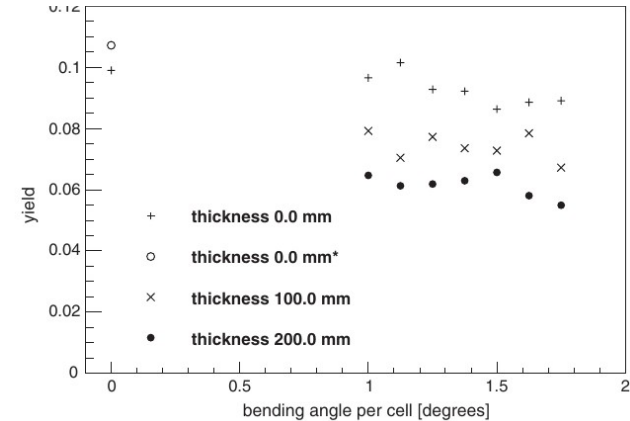
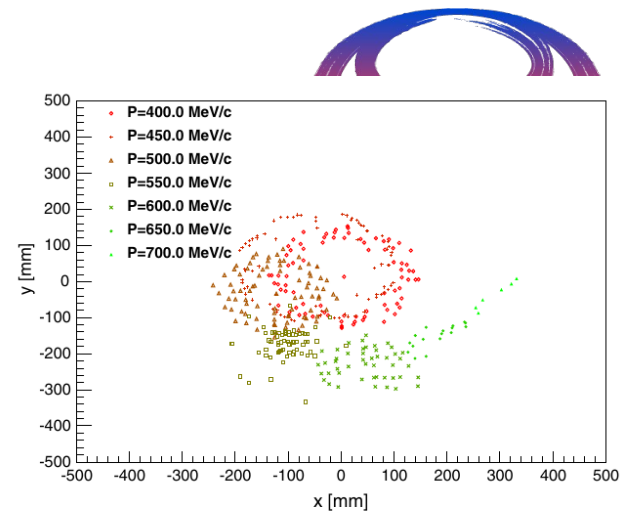
Muon front end



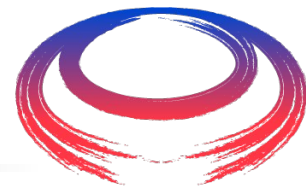
- Muon front-end to capture muon beam
- Solenoid taper
- Solenoid chicane removes high momentum particles
- Beryllium plug removes low momentum impurities
- Longitudinal capture system
 - Adiabatically bunch beam
 - Phase rotate

Chicane/proton absorber

- Solenoid chicane
 - No dipoles!
 - Vertical dispersion → low pass filter
 - Excellent transport properties within acceptance
- Beryllium plug
 - Protons stop more quickly than muons/pions
 - Removes low momentum protons

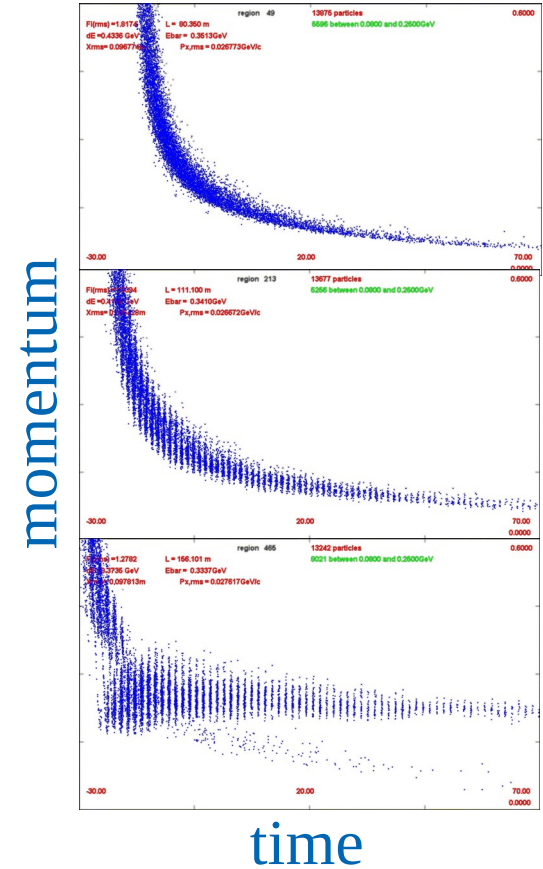


Buncher/Phase Rotator

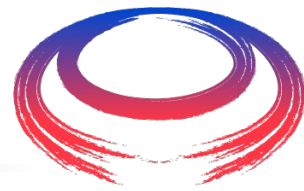


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- Drift to develop energy-time relation
- Buncher adiabatically ramp RF voltages
- Phase rotator → misphase RF
 - High energy bunches decelerated
 - Low energy bunches accelerated



Luminosity consideration



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$$\mathcal{L} \approx \underbrace{\frac{e\tau_\mu}{(4\pi m_\mu c)^2}}_{K_L = 4.4e36 \text{ MeV MW}^{-2} \text{ T}^{-1} \text{ s}^{-2}} \frac{f_{hg}\sigma_\delta \bar{B}}{\varepsilon_\perp \varepsilon_L n_b f_r} \underbrace{\eta_+ \eta_- (\eta_\tau P_p \gamma m_\mu c^2)^2}_{P_+ P_-}$$

$$N_\pm = \frac{\eta_\tau \eta_\pm P_p}{n_b f_r}$$

Number of muons per proton beam power

Proton beam power

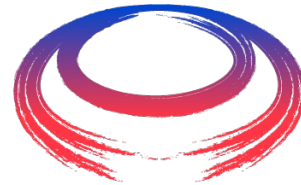
Efficiency of muon acceleration

Number of bunches

Rep rate

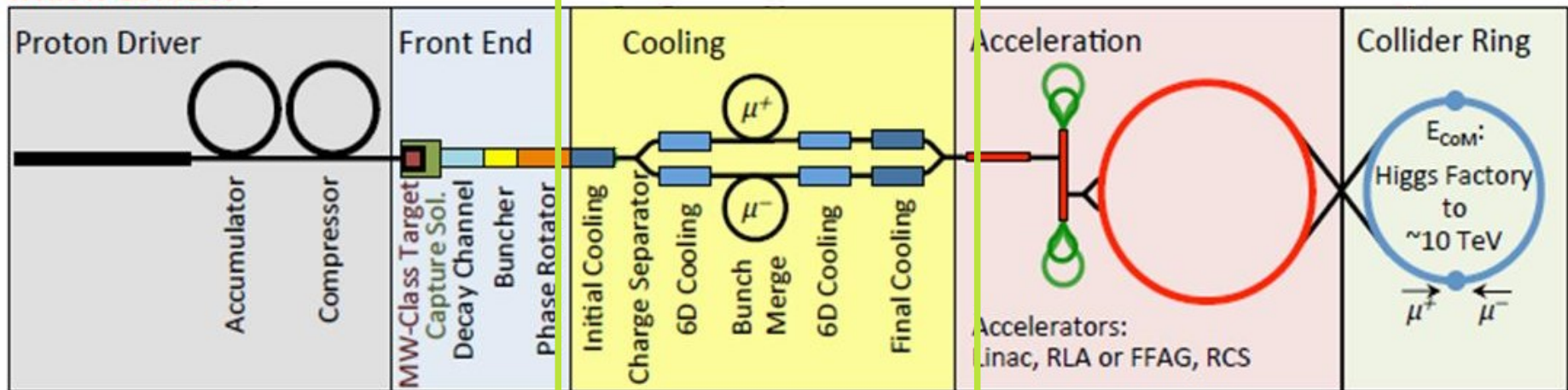
- Proton beam power ~ 1-2 MW → (FNAL, JPARC, SNS)
- Approx 0.1 $\mu^{+/-}$ per 8 GeV proton → O(1e14) muons per MW
- BUT: muon front end produces multiple bunches (about 20)
- Rep rate is between 60 Hz (SNS) and 0.1 Hz (JPARC)
- Emittance is huge

The Facility - Ionisation Cooling



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Muon Collider





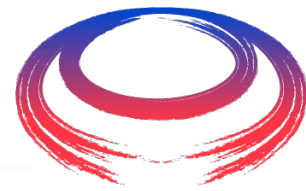
Ionisation Cooling - intro



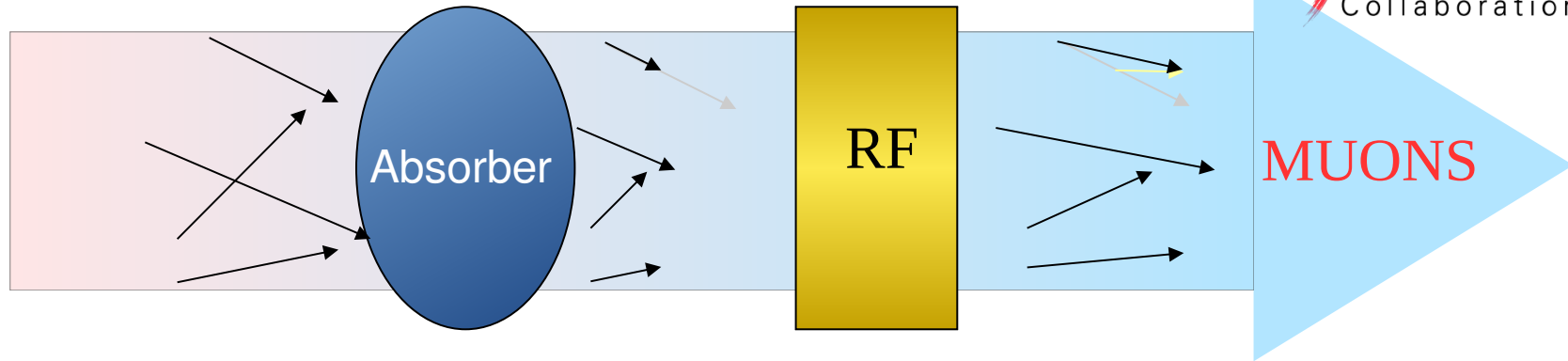
- Muon front end produces huge flux of muons
- Muons have too large emittance at the source
- How can we reduce beam emittance? COOLING!
 - Laser cooling
 - Stochastic cooling
 - Electron cooling
 - **Too slow**
- Ionisation cooling (and Frictional cooling)



Ionisation Cooling



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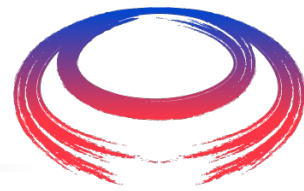
- Beam loses energy in absorbing material
 - Absorber removes momentum in all directions
 - RF cavity replaces momentum only in longitudinal direction
 - End up with beam that is more straight
- Multiple Coulomb scattering from nucleus ruins the effect
 - Mitigate with tight focussing
 - Mitigate with low-Z materials
 - Equilibrium emittance where MCS completely cancels the cooling



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Beam emittance in 4D



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- Normalised RMS beam emittance in 2D
 - area of ellipse aligned with beam

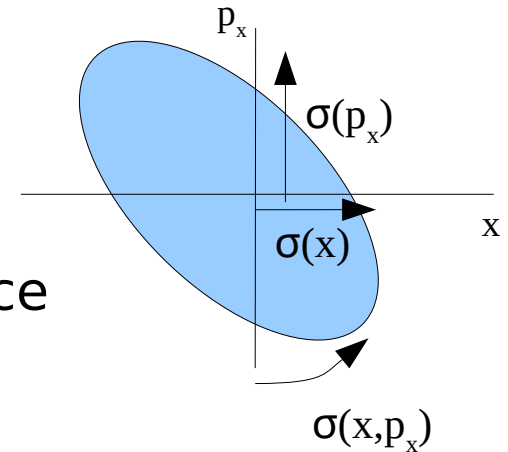
$$\varepsilon_{2d} = \frac{1}{m} \sqrt{\sigma^2(x)\sigma^2(p_x) - \sigma^2(x, p_x)}$$

- $\sigma^2(u_i)$ and $\sigma^2(u_i, u_j)$ are variance and covariance
 - Also written as $\langle u_i u_j \rangle$
- Can be written as

$$\varepsilon_{2d} = \frac{1}{m} \sqrt{|\mathbf{V}_{2d}|}$$

- In higher dimensions the definition generalises

$$\varepsilon_{2nd} = \frac{1}{m_\mu} \sqrt[n]{|\mathbf{V}|}$$



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Transverse cooling (1)



- Say we pass through some material at a focus
 - P decreases due to ionisation
 - Multiple Coulomb Scattering increases angular spread
- For a cylindrically symmetric beam with angular divergence Θ_x

$$\begin{aligned}\sqrt{|\mathbf{V}_\perp|} &= (\langle x^2 \rangle \langle p_x^2 \rangle - \langle xp_x \rangle^2 - \langle xp_y \rangle^2) \\ &\approx p_z^2 (\langle x^2 \rangle \langle \Theta_x^2 \rangle - \langle x\Theta_x \rangle^2 - \langle x\Theta_y \rangle^2)\end{aligned}$$

The change in emittance is given by

$$\frac{d\epsilon_n}{dz} = \frac{1}{2m^2\epsilon_n} \frac{d\sqrt{|\mathbf{V}_\perp|}}{dz}$$

$$\epsilon_{2nd} = \frac{1}{m_\mu} \sqrt[n]{|\mathbf{V}|}$$

Transverse cooling (2)

$$\frac{d\epsilon_n}{dz} = \frac{1}{2m^2\epsilon_n} \frac{d\sqrt{|\mathbf{V}_\perp|}}{dz}$$

- Only p_z and $\langle \Theta_i^2 \rangle$ change; applying product rule

$$\frac{d\epsilon_n}{dz} \approx \frac{1}{2m^2\epsilon_n} \left(2 \frac{dp_z}{dz} \frac{\sqrt{|\mathbf{V}_\perp|}}{p_z} + \langle x^2 \rangle p_z^2 \frac{d\langle \Theta_x^2 \rangle}{dz} \right)$$

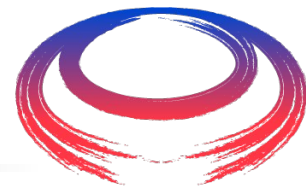
- Use (from $E^2 + p^2 = m^2$) $E dE/dz \approx p_z dp_z/dz$

- Use standard formula $\beta_\perp = \frac{\langle x^2 \rangle p}{m\epsilon_n}$

- Use scattering (from atomic physics) $\frac{d\langle \Theta_x^2 \rangle}{dz} \approx \frac{13.6^2}{(p\beta_{rel})^2 L_R}$.

- Gives $\frac{d\epsilon_n}{dz} \approx \frac{1}{\beta_{rel}^2 E} \left\langle \frac{dE}{dz} \right\rangle \epsilon_n + \frac{1}{2m^2\epsilon_n} \langle x^2 \rangle \frac{13.6^2}{(\beta_{rel})^2 L_R}$.

Transverse cooling (3)



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- Rearranging
$$\frac{d\epsilon_n}{dz} \approx \frac{1}{E} \left\langle \frac{dE}{dz} \right\rangle \epsilon_n + \frac{1}{2m} \frac{13.6^2}{L_R} \frac{\beta_{\perp}}{\beta_{rel}^3 E}$$

dE/dz is negative!
Cooling

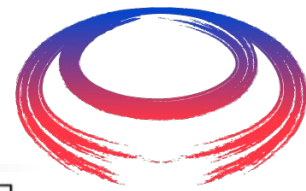
Heating

- There exists an equilibrium emittance where the two terms balance (no emittance change)

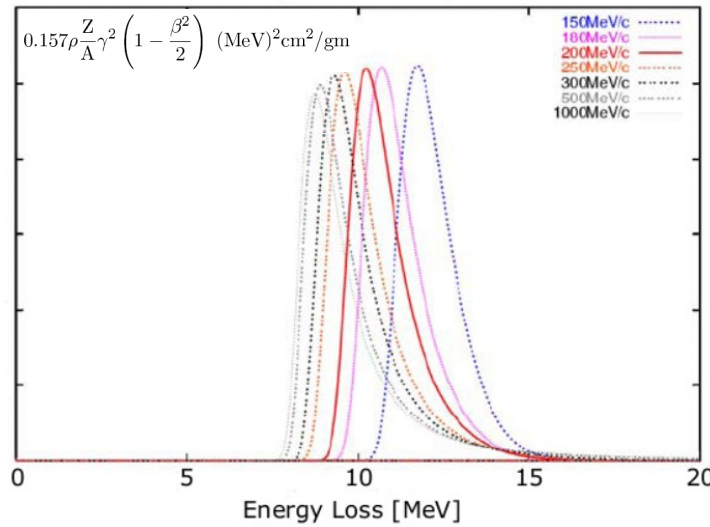
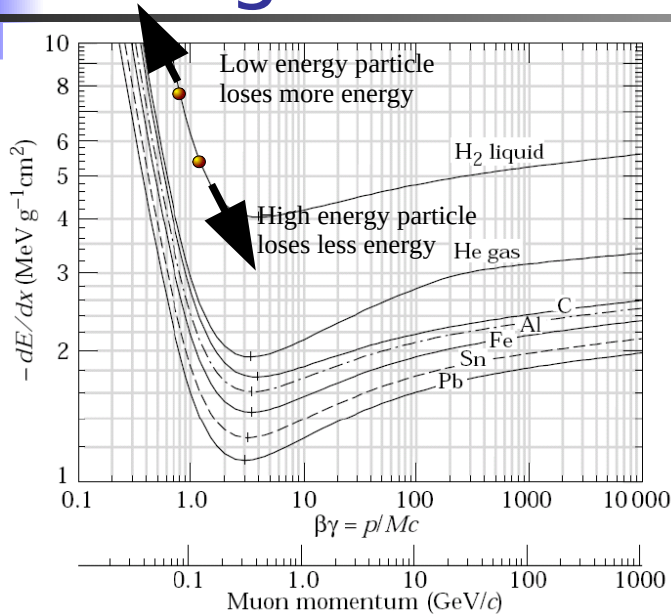
$$\epsilon_n(\text{equilibrium}) = \frac{1}{2m} \frac{13.6^2}{L_R} \frac{\beta_{\perp}}{\beta_{rel} \left\langle \frac{dE}{dz} \right\rangle}$$



Longitudinal Heating



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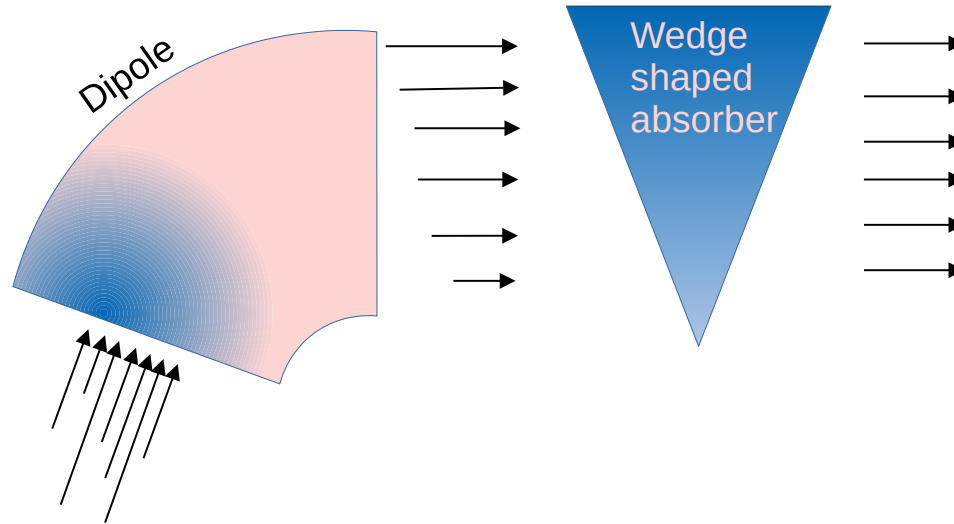


- In longitudinal phase space, the beam is usually heated
 - Heating due to random noise in the energy loss I.e. “straggling”
 - Heating due to curvature in energy loss (heating or weak cooling)

$$\frac{d \langle E^2 \rangle}{dz} = \left(2 \frac{d}{dE} \frac{dE}{dz} \right) \langle E^2 \rangle + \left(\frac{d \langle E^2 \rangle}{dz} \right)_{Vlasov}$$

- Mitigate using emittance exchange
 - Move emittance from longitudinal to transverse phase space

Emittance exchange



- Initial beam is narrow with some momentum spread
 - Low transverse emittance and high longitudinal emittance
- Beam follows curved trajectory in dipole
 - Higher momentum particles have higher radius trajectory
 - Beam leaves dipole wider with energy-position correlation
- Beam goes through wedge shaped absorber
 - Beam leaves wider without energy-position correlation
 - High transverse emittance and low longitudinal emittance

Emittance exchange

- Longitudinal emittance change becomes

$$\frac{d \langle E^2 \rangle}{dz} = \left(2 \frac{d}{dE} \frac{dE}{dz} \right) \langle E^2 \rangle + \left(\frac{d \langle E^2 \rangle}{dz} \right)_{Vlasov}$$

$$\frac{\partial \frac{dE}{ds}}{\partial E} \Rightarrow \frac{\partial \frac{dE}{ds}}{\partial E} \Big|_0 + \frac{dE}{ds} \frac{\eta \rho'}{\beta c p \rho_0}$$

dispersion

Effective density
Variation with position

- Transverse emittance change becomes

$$\frac{d\epsilon_n}{dz} \approx \frac{1}{\beta_{rel}^2 E} \left\langle \frac{dE}{dz} \right\rangle \epsilon_n + \frac{1}{2m^2 \epsilon_n} \langle x^2 \rangle \frac{13.6^2}{(\beta_{rel})^2 L_R}$$

$$\frac{1}{\beta^2 E} \frac{dE}{ds} \left(1 - \frac{\eta \rho'}{\rho_0} \right) \epsilon_N$$



Summary - Lecture 1



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Final Word – Part 1



- The muon collider is an exciting prospect
- An opportunity to build a sustainable path in HEP
 - Wall plug power
 - Energy reach
- Many challenges remain on the road
 - Luminosity is the key
 - Capture and cooling of the muon beam
 - Ionisation cooling
- Next lecture, explore
 - How ionisation cooling may be realised in practice
 - Acceleration
 - Collision
 - The path to a muon collider