

# Phenomenological aspects of the dark matter stability origin

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# DM is astonishingly stable!

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↪  $\tau_{DM} > \tau_U \sim 10^{18} \text{ sec}$

$\tau_{DM} \gtrsim 10^{26} \text{ sec}$       ↪ in most models not to produce  
 $e^+, \bar{p}, \gamma, \dots$  fluxes larger than observed

↪ in many models: stability assumed by hand



e.g. ad hoc  $Z_2$  symmetry ...

⇒ origin of the exceptional DM stability????

↪ determines the basic DM model structure with specific phenomenology

Examples of specific phenomenology

deriving from

DM stabilization mechanisms

# DM stability $\implies$ long range forces??

$\hookrightarrow$  DM stable as for the  $e^-$ : lightest charged particle under a unbroken gauged U(1)

$\hookrightarrow$  the simple adjunction of a new QED structure for a single fermion gives a viable DM candidate!!

Ackerman, Buckley, Carroll, Kamionkowski 08'; Feng, Tu, Yu 08'; Feng, Kaplinghat, Tu, Yu 09' see also Foot et al. 06'-10'

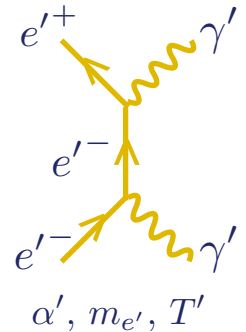
$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{QED'}$$

$$\mathcal{L}_{QED'} = \bar{\psi}_{e'}(i\not{\partial} - e\not{A}_{\gamma'} - m_{e'})\psi_{e'}$$

$\downarrow$   
 $\gamma'$

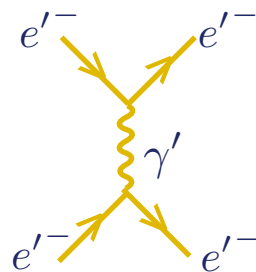
$\downarrow$   
 $e'^{\pm}$

$\implies$  stable  $\implies e'^{\pm}$  relic density:



$\implies$  long range force between charged DM  $e'^{\pm}$

- galactic halo morphology modified by DM collisions through Rutherford scatter.
- more collisions in bullet cluster through Rutherford scattering
- damping of small scale structure due to lower kinetic decoupling



+ possible communication with SM through kinetic mixing



# Lightest fermion of a secluded sector

- e.g. assume:
- a new U(1) gauge interaction
  - a charged scalar  $\phi$  breaking it
  - a charged fermion  $\psi$  (vector)
  - all SM fermions are neutral under U(1)

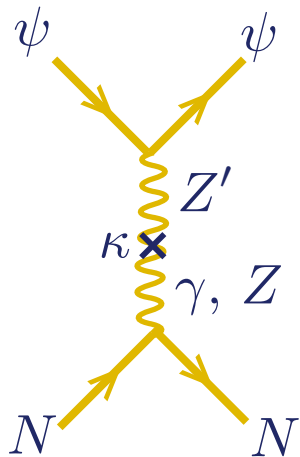
- a massive  $Z'$
- a Higgs boson
- a massive fermion  $\psi$



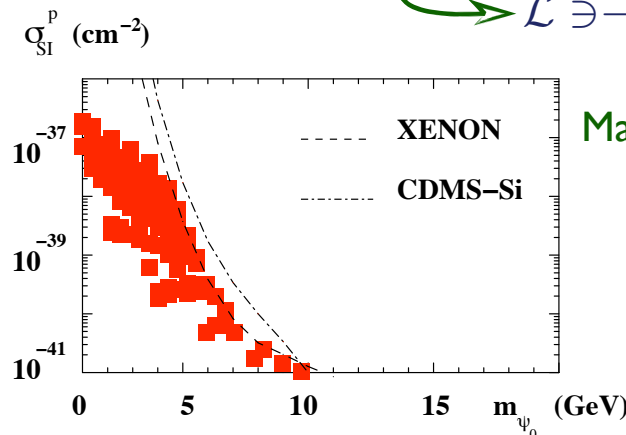
stable because lightest fermion of a secluded sector

Pospelov, Ritz, Voloshin 07'  
 Gopalakrishna, Jung, Wells 08'  
 Gopalakrishna, Lee, Wells 08'

communication with SM through Higgs portal and kinetic mixing



$\mathcal{L} \ni -\lambda \phi^\dagger \phi H^\dagger H$ 
 $\mathcal{L} \ni -\frac{1}{4} \kappa F_{\mu\nu}^Y F_{QED}^{\mu\nu}$



Mambrini 10'

$$\kappa \lesssim 10^{-(2-3)}$$

can account for DAMA-CoGeNT  
 if  $m_{DM} \sim m_{Z'}/2$  (resonance)

# Remnant global subgroup of GUT: R-symmetry

Mohapatra 86', Martin 92'

$R_m = (-1)^{3(B-L)} \Rightarrow$  R-symmetry is a  $Z_2$  subgroup of  $U(1)_{B-L}$   
 $\Downarrow$   
a subgroup of  $SO(10)$

$\Rightarrow$  if  $U(1)_{B-L}$  (or  $SO(10)$ ) is a gauge symmetry and is broken  
only by vev of fields with even B-L: R-symmetry remains as an exact symmetry

$\hookrightarrow$  10, 45, 54, 120, 126, 210, ...  $\hookrightarrow$  conserved by UV physics too

$\hookrightarrow$  high energy explanation of R-symmetry  $\leftarrow$  not experimentally testable (directly)  
but severely constrains the GUT model

Aulakh, Melfo, Rasin, Senjanovic 98'  
Aulakh, Bajc, Melfo, Rasin, Senjanovic 01'

.....

# DM stability in non-susy $SO(10)$ setups

non-susy  $U(1)_{B-L}$  (or  $SO(10)$ ) gauge theories broken by only even B-L field vev also leaves a  $Z_2$  symmetry

SM fermions are in the 16 of  $SO(10)$  which is B-L odd  
SM Higgs doublet is in the 10 of  $SO(10)$  which is B-L even

the lightest component of an extra B-L odd scalar  $SO(10)$  representation is stable

16, 144, ...

Kadastik, Kannike, Raidal 09'

can drive electroweak GUT unification

the lightest component of an extra B-L even fermion  $SO(10)$  representation is stable

10, 45, 54, 120, 126, 210, ...

Frigerio, TH 09'

45 or 54 fermion representation: DM is the neutral component of a fermion triplet  $\Sigma^+, \Sigma^0, \Sigma^-$

relic density requires  $m_{DM} \simeq 2.7 \text{ TeV}$

Cirelli, Fornengo, Strumia 06'

stabilization mechanism "fixes" the DM mass

# DM stability from accidental symmetry: DM decay

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↪ if DM stable from accidental low energy symmetry we expect it to decay from new physics UV interactions as proton

↪ intriguing coincidence: a GUT scale induced dim 6 operator



cosmic ray fluxes from DM decay of order the observed ones



probe GUT scale physics??

Eichler; Nardi, Sannino, Strumia; Chen, Takahashi, Yanagida; Arvanitaki, Dimopoulos et al.; Bae, Kyae; Hamagushi, Shirai, Yanagida; Arina, TH, Ibarra, Weniger; ...

# A smoking gun DM decay signal: intense $\gamma$ -ray lines

↪ if DM is a massive gauge boson, abelian or non abelian

TH 08'; TH, Tytgat 09'  
Arina, TH, Ibarra, Weniger 10'



“Hidden vector”: accidental custodial symmetry



a gauged SU(2) + SSB from a scalar doublet  $\phi$

$$\mathcal{L} = -\frac{1}{4}F^{\mu\nu a}F_{\mu\nu}^a + (D^\mu\phi)^\dagger(D_\mu\phi) - \mu_\phi^2\phi^\dagger\phi - \lambda_\phi(\phi^\dagger\phi)^2$$



$\phi$  gets a vev  $v_\phi$

$$m_V = \frac{g_\phi v_\phi}{2}$$

3 massive gauge bosons  $V_i$  + a real scalar  $\eta$

$$m_\eta = \sqrt{2\lambda_\phi} v_\phi$$



stable due to residual SU(2) custodial sym.

$V_i \rightarrow \eta\eta, \dots$  forbidden:

$(V_1^\mu, V_2^\mu, V_3^\mu) = \text{triplet}$

$\eta = \text{singlet}$

+ communication with the SM through Higgs portal:  $\mathcal{L}_{Higgs\ portal} = -\lambda_m\phi^\dagger\phi H^\dagger H$



relic density, direct detection, ...

# A smoking gun DM decay signal: intense $\gamma$ -ray lines

Hidden vector: no reason that the custodial sym. not violated in the UV  $\Rightarrow$  dim 6 operators:

- (A)  $\frac{1}{\Lambda^2} \mathcal{D}_\mu \phi^\dagger \phi \mathcal{D}_\mu H^\dagger H$
- (B)  $\frac{1}{\Lambda^2} \mathcal{D}_\mu \phi^\dagger \phi H^\dagger \mathcal{D}_\mu H$
- (C)  $\frac{1}{\Lambda^2} \mathcal{D}_\mu \phi^\dagger \mathcal{D}_\nu \phi F^{\mu\nu Y}$
- (D)  $\frac{1}{\Lambda^2} \phi^\dagger F_{\mu\nu}^a \frac{\tau^a}{2} \phi F^{\mu\nu Y}$

all give 2-body radiative decays

monochromatic  $\gamma$ -rays

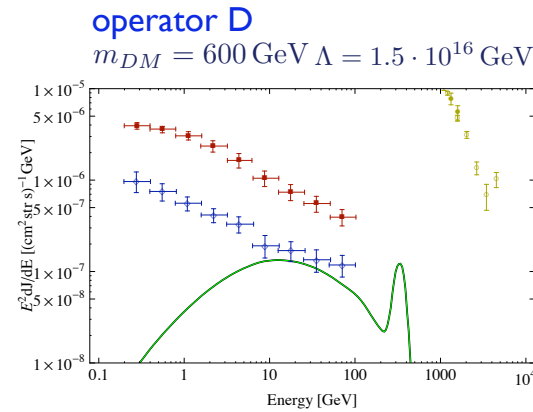
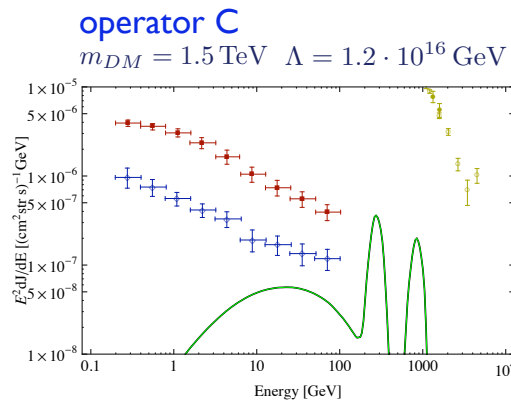
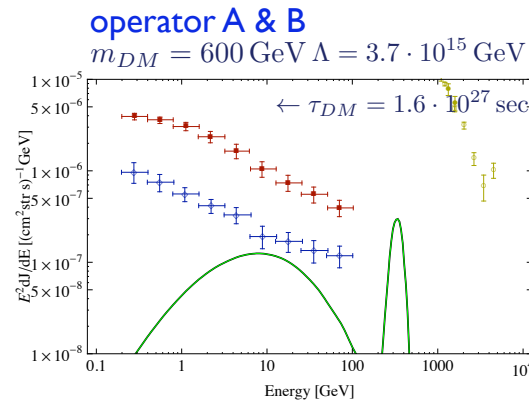
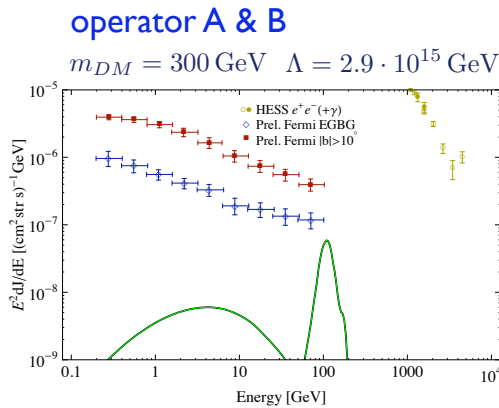
in particular off the galactic plane!

**DM smoking gun!**

(no astrophysical background!)

Annihilation: Bergstrom, Ullio, 97' 98'; Bern, Gondolo, Perelstein 97'; Bergstrom, Bringmann, Eriksson, Gustafsson 04', 05'; Boudjema, Semenov, Temes 05'; Jackson, Servant, Shaughnessy, Tait, Taoso 09', ...  
one tree level exception: Dudas, Mambri, Pokorski, Romagnoni 09'

Decay: Buchmuller, Covi, Hamagushi, Ibarra, Tran 07'; Ibarra, Tran 07'; Ishiwata, Matsumoto, Moroi 08'; Buchmuller, Ibarra, Shindou, Takayama, Tran 09'; Choi, Lopez-Fogliani, Munoz, de Austri 09'



C.Arina, T.H., A. Ibarra, C. Weniger 10'

# Dark Matter

as a

## Pseudo Goldstone Boson

↪ stability: pseudo-Goldstone decay suppressed by large spontaneous breaking scale

↪ 2 well-known examples: - Axion  
- DM Majoron

Cadamuro, Hasenkamp talks

# The pseudo-Goldstone stabilization mechanism: Majoron case

Chikashige, Mohapatra, Peccei 81'

↪ decay width suppressed by large seesaw scale

↪  $U(1)_{B-L}$  breaking scale

- global  $U(1)_{B-L}$  spontaneous breaking scale driven by L=2 scalar field  $\phi$

$$\mathcal{L} \ni -Y_{ij} N_{Ri} L_j H - \frac{1}{2} c_{ij} \phi N_{Ri} N_{Rj}$$

$$\langle \phi \rangle \equiv f \Downarrow$$

$$-\frac{1}{2} M_{Nij} N_{Ri} N_{Rj} e^{i\theta/f} - \frac{1}{2} c_{ij} \phi' N_{Ri} N_{Rj} e^{i\theta/f}$$

Majoron



$$\phi = (\phi' + f) e^{i\theta/f}$$

$$M_{Nij} = c_{ij} f$$

Akhmedov, Berezhiani, Senjanovic, Tao 93'

- additional explicit  $U(1)_{B-L}$  breaking  $\longrightarrow$  Planck effects breaks the global sym.  $\longrightarrow$  massive Majoron

(or soft breaking terms, Gu, Ma, Sarkar 10')

$\Longrightarrow$  DM massive Majoron stable because it couples only to SM through suppressed  $\nu - N_R$  mixing

$$\Gamma(\theta \rightarrow \nu\nu) \propto (m_\nu/M_N)^4 \propto (m_\nu/f)^4$$

$$\Gamma(\theta \rightarrow e^+e^-) \propto \alpha_W (m_\nu/M_N)^4 \propto \alpha_W (m_\nu/f)^4$$

Singlet-triplet DM extension:

M. Tortola's talk

Berezinsky, Valle 93, Lattanzi, Valle 07',

Bazzocchi, Lattanzi, Riener-Sorensen, Valle 08',

Esteves, Joachim, Joshipura, Romao, Tortola, Valle 10'

$\Longrightarrow$  decay suppressed by naturally large seesaw scale



# A new pseudo-Goldstone DM framework

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Frigerio, TH, Masso, I I'

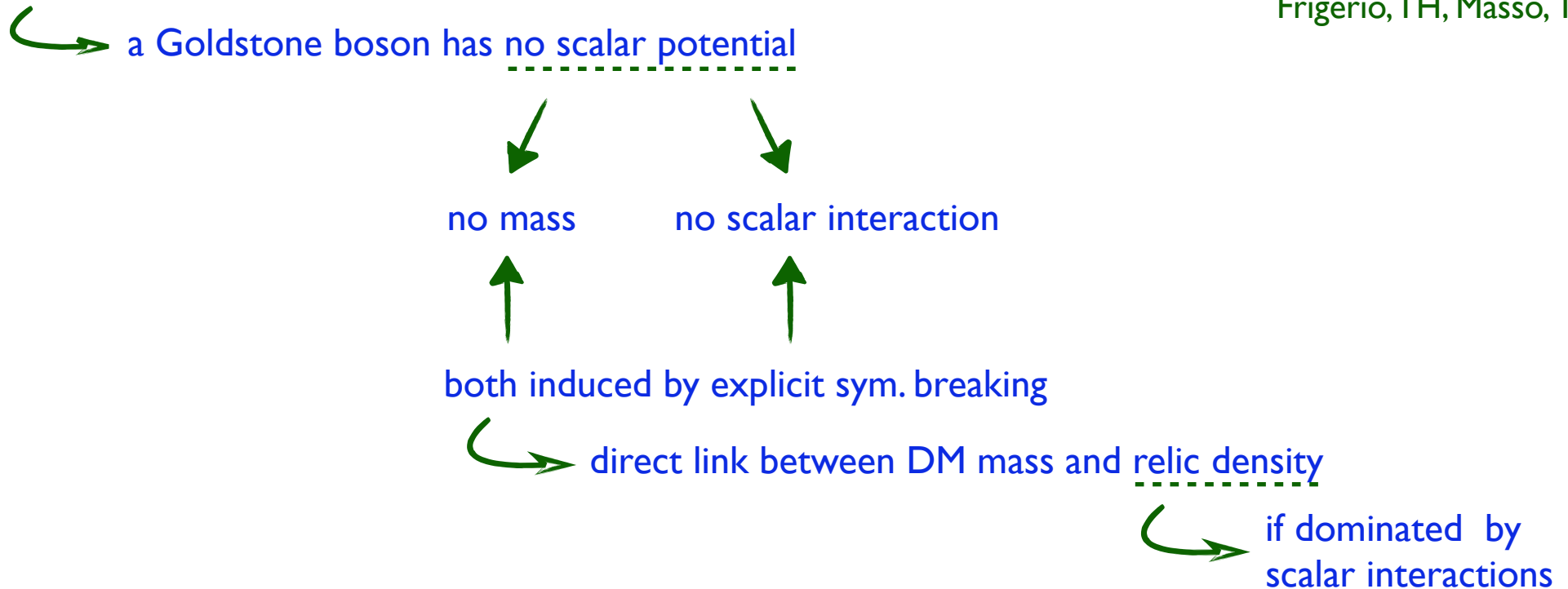
 in addition to provide the naturally large stabilizing scale the seesaw  $\mathcal{L}$  can also provide:

- the source of explicit symmetry breaking  $\Rightarrow$  of  $m_{DM}$
- the interactions producing the DM relic density
- with a justification for scalar DM at low scale with a mass radiatively “stable”

 a possibility of a direct link between DM mass and DM relic density

# A pseudo-Goldstone link between DM relic density and mass

Frigerio, TH, Masso, I I'



↪ suppose the explicit breaking term induces only a Higgs portal interaction:

$$\mathcal{L} \ni \lambda \theta \theta H^\dagger H \begin{cases} \rightarrow \text{DM mass: } m_\theta^2 = \lambda v^2 \Rightarrow \text{DM at the electroweak scale or below} \\ \rightarrow \text{DM relic density from: } \theta \theta \leftrightarrow H H^\dagger \end{cases}$$

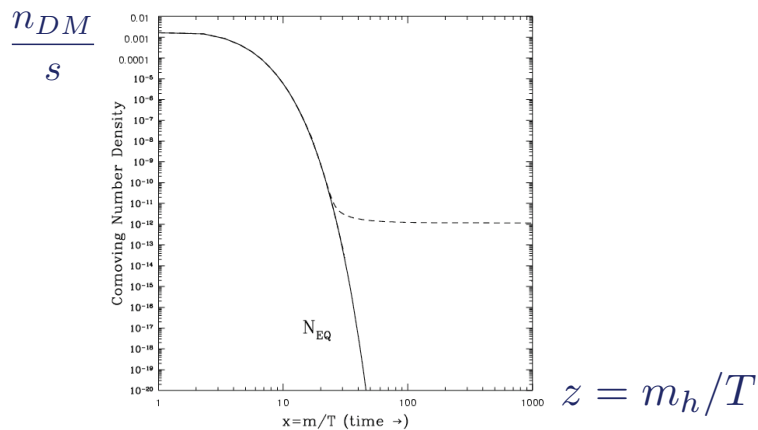
# DM relic density from the Higgs portal: $\mathcal{L} \ni \lambda \theta \theta H^\dagger H$

$$m_\theta^2 = \lambda v^2$$

2 regimes

large  $\lambda$  coupling: freezeout  
of DM annihilation

$$\theta\theta \rightarrow hh, WW, ZZ, f\bar{f}$$

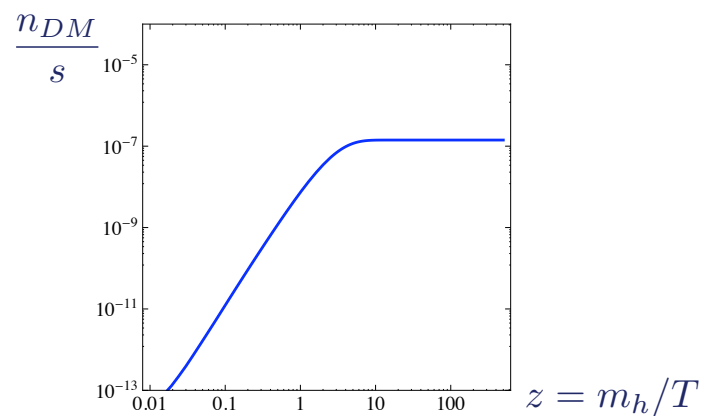


$$\Rightarrow m_{DM} = 55 \text{ GeV} \quad (m_h = 120 \text{ GeV})$$

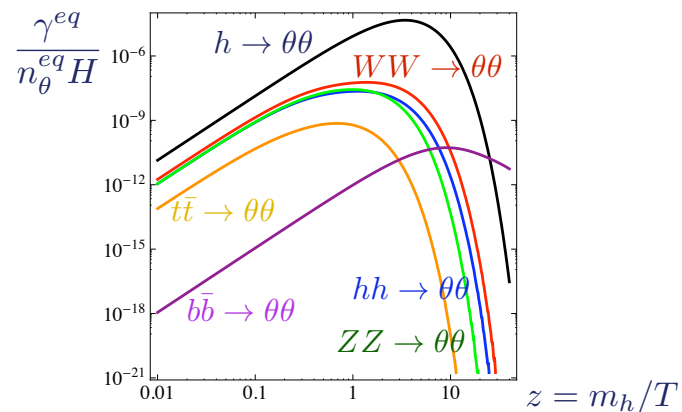
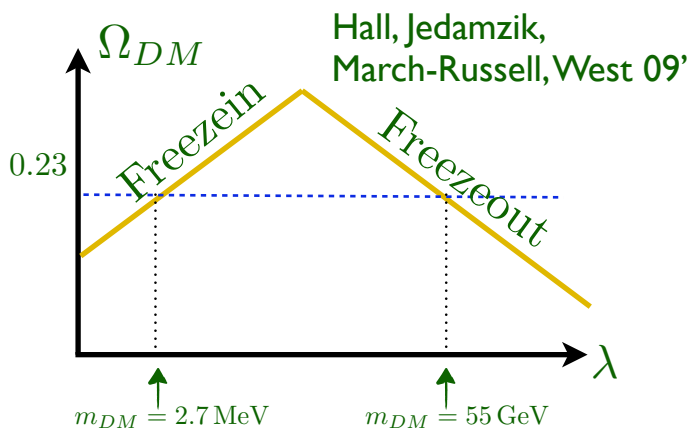
Farina, Pappadopulo, Strumia 10'

small  $\lambda$  coupling: freezein  
of DM pair creation

$$hh, WW, ZZ, f\bar{f} \rightarrow \theta\theta \text{ and } h \rightarrow \theta\theta$$



$$\Rightarrow m_{DM} = 2.7 \text{ MeV} \quad \text{Frigerio, TH, Masso, 11'}$$



# Collective breaking of a U(1) flavour symmetry

collective breaking of global symmetry à la Hill and Ross '88

consider  $U(1)_X$  flavour symmetry (instead of  $U(1)_{B-L}$ )

induces a  $\theta$  mass only if several  $M_{Nij} \neq 0$

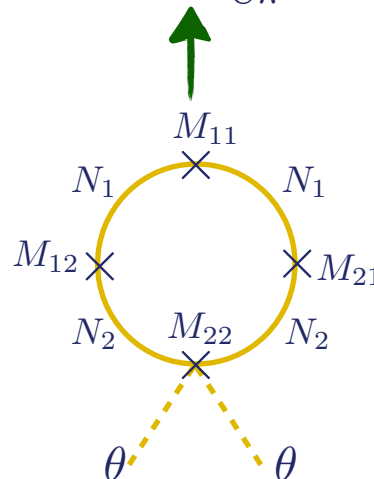
example:  $N_{1,2} : Q_{N_1}^X = 0$   
 $Q_{N_2}^X = 1$   
 $Q_\phi^X = 2$

$$M_N = \begin{pmatrix} M_{11} & M_{12} \\ M_{12} & M_{22}e^{i\theta/f} \end{pmatrix}$$

allowed by  $U(1)_X$  (pointing to  $M_{11}$ )  
 explicit  $U(1)_X$  breaking (pointing to  $M_{12}$ )  
 from  $U(1)_X$  SSB (pointing to  $M_{22}e^{i\theta/f}$ )

$$m_\theta^2 \sim \frac{1}{8\pi^2} \frac{M_{11}M_{12}M_{22}M_{21}}{f^2} \ln \frac{\Lambda^2}{\mu^2}$$

if any of  $M_{ij} = 0$  we recover an exact  $U(1)$



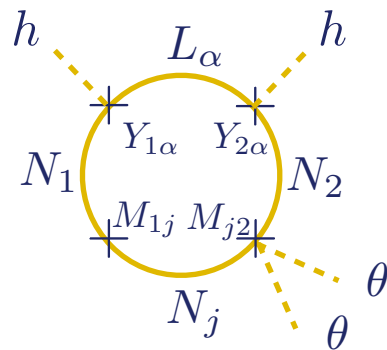
logarithmic only

mass not quadratically sensitive to the cutoff

# Inducing the pseudo-Goldstone-Higgs portal

Frigerio, TH, Masso, II

other possible realization of collective breaking involving both Majorana and Dirac terms:



$\Rightarrow$  Higgs portal:  $\lambda = \frac{1}{4\pi^2} \frac{M_{12}(M_{11} + M_{22})}{f^2} Y_{1\alpha} Y_{2\alpha} \log \frac{\Lambda^2}{\mu^2}$

$m_\theta^2 = \lambda v^2 = \frac{1}{4\pi^2} \frac{M_{12}(M_{11} + M_{22}) m_{1\alpha}^D m_{2\alpha}^D}{f^2} \log \frac{\Lambda^2}{\mu^2}$

$\mathcal{L} \ni \lambda \theta \theta H^\dagger H$

example:  $Q_{N_1}^X = -1, Q_{N_2}^X = 1, Q_\phi^X = 2, Q_\nu^X = 1$

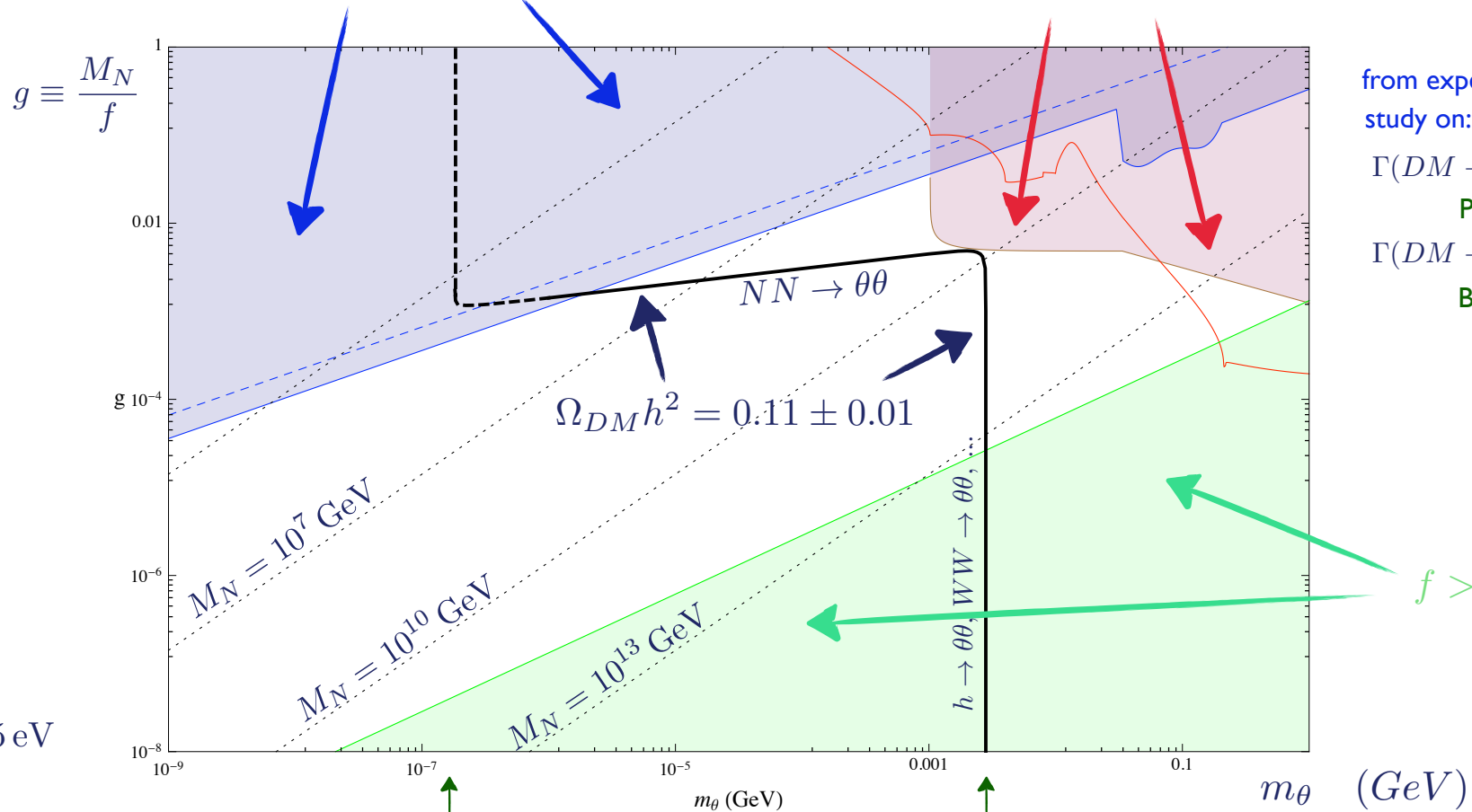
- NB:** - to avoid generation of a DM tadpole, the easiest way is to assume CP symmetry
- if reheating temperature very high and  $\phi NN$  couplings large enough the relic density can be produced also from  $NN \rightarrow \theta\theta$  pair production

# Compilation of constraints

Frigerio, TH, Masso, II

$\Gamma(\theta \rightarrow \nu\nu) \propto m_\theta m_\nu^4 / f^4$  too large

$\Gamma(\theta \rightarrow e^+e^-) \propto \alpha_W m_\theta m_e^2 m_\nu^2 M^2 / f^2$  too large



from experimental constraint study on:

$\Gamma(DM \rightarrow \nu\nu)$  :

Palomares-Ruiz 08'

$\Gamma(DM \rightarrow e^+e^-)$  :

Bell, Galen, Petraki 10'

$f > M_{Planck}$

$m_\nu = 0.05 \text{ eV}$

$m_{DM} = 0.15 \text{ KeV}$

$m_{DM} = 2.7 \text{ MeV}$

unique value if  $T_{reheating} < M_N$  or  $g \equiv M_N/f$  small

# Summary

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- The origin of particle DM stability is a fundamental question!  
Each UV or low energy scenario requires a very specific pattern in term of type of particle needed, energy scale, ..., which leads to a specific phenomenology
  - long range dark force
  - intense flux of cosmic rays,  $\gamma$ -lines, ...
  - GUT specific realizations
  - specific DM mass
  - relic density/DM mass link
  - ....
- Pseudo-Goldstone setup: seesaw interactions assuming a U(1) flavour symmetry:
  - can provide the explicit breaking source  $\implies m_{DM}$
  - generate the DM relic density  $\implies m_{DM} \propto v_{EW}$
  - with a one-to-one link between  $m_{DM}$  and  $\Omega_{DM} \implies m_{DM} = 2.7 \text{ MeV}$
  - that can be probed by  $DM \rightarrow \nu\nu$  and  $DM \rightarrow e^+e^-$  searches

**Backup**



# DM stability in non-susy SO(10) setups: scalar case

Kadastik, Kannike, Raidal 09'

↪ add a 16 scalar representation:

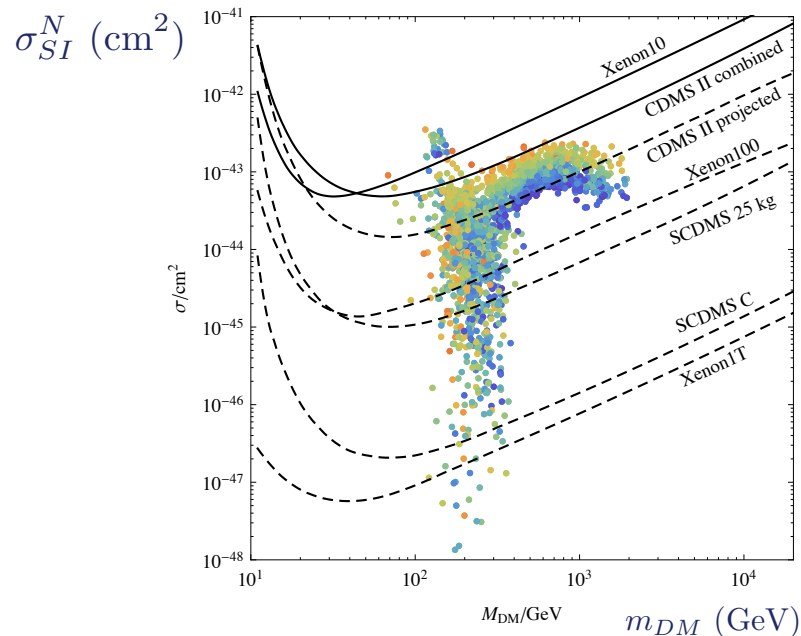
DM is a combination of a scalar doublet and a scalar singlet

⇓  
inert doublet

↘ similar phenomenology  
with additional constraint:  
 $m_{DM} \gtrsim 100 \text{ GeV}$

↪ not for DAMA, CoGeNT

direct detection:



Huitu, Kannike, Raccioppi, Raidal 10'

+ long lived DM partners at LHC

# DM stability in non-susy SO(10) setups: fermion case

Frigerio, TH 09'

↪ add a 45 or 54 fermion representation:

DM is the neutral component of a fermion triplet  $\Sigma^+, \Sigma^0, \Sigma^-$



advantage that the DM triplet can drive gauge coupling unification



as in split susy but without susy

↪ low energy pheno is as for a generic fermion triplet:

- relic density requires  $m_{DM} \simeq 2.7 \text{ TeV}$     Cirelli, Fornengo, Strumia 06'
- $\sigma_{SI}^N \sim 10^{-45} \text{ cm}^2$
- indirect detection:
  - too many antiprotons for explaining  $e^+$  excess of Pamela
  - $DM DM \rightarrow \gamma\gamma$  expected to give  $\gamma$ -lines with a rate than can be probed at atmospheric Cerenkov telescopes

# DM stability from unbroken U(1) gauge group

→ as for the  $e^-$ : stable because lightest charged particle under a U(1)

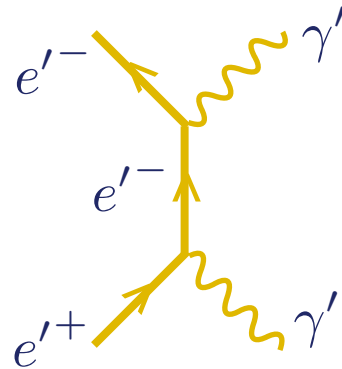
→ the simple adjunction of a new QED structure for a single fermion gives a viable DM candidate!!

Ackerman, Buckley, Carroll, Kamionkowski 08'  
 Feng, Tu, Yu 08'; Feng, Kaplinghat, Tu, Yu 09'  
 Foot et al. 06'-10'

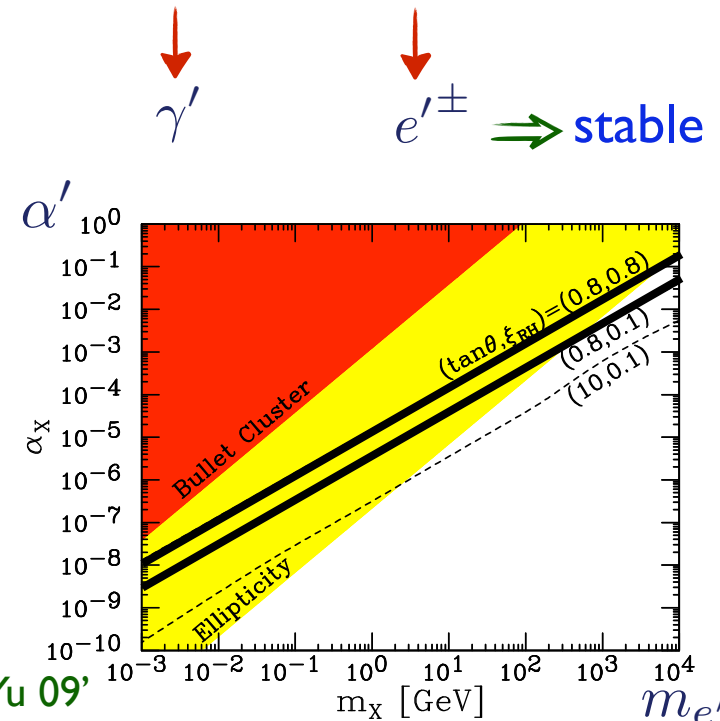
$$\mathcal{L} = \mathcal{L}_{SM} + \mathcal{L}_{QED'}$$

$$\mathcal{L}_{QED'} = \bar{\psi}_{e'}(i\not{\partial} - e A_{\gamma'} - m_{e'})\psi_{e'}$$

⇒  $e'^{\pm}$  relic density



→ depends on  $m_{e'}$ ,  $\alpha'$ ,  $\xi \equiv T_{\gamma'}/T_{\gamma}$



Feng, Kaplinghat, Tu, Yu 09'

# DM stability from accidental symmetry: Minimal Dark Matter

as the  $p^+$  in the SM

Cirelli, Fornengo, Strumia 06'  
Cirelli, Strumia, Tamburini 07'

without adding any new gauge group, large  $SU(2)_L$  multiplet cannot have any renormalizable interactions with SM fields due to  $SU(2)_L$  gauge invariance (or dimension-5)

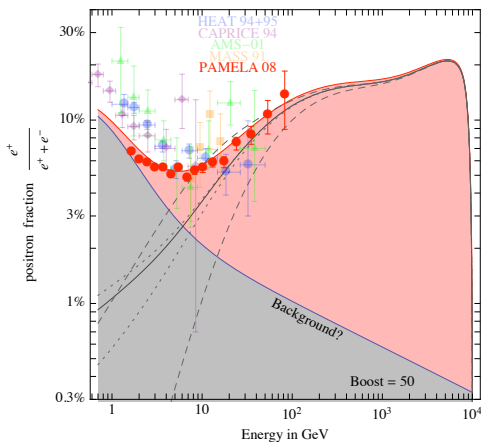
a fermion quintuplet, septuplet, ...

a scalar septuplet, nonuplet, ...

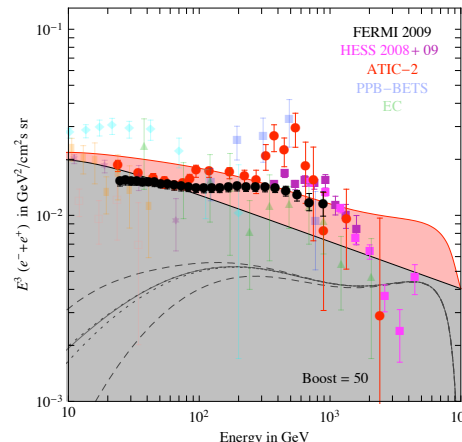
$$m_{DM} = (9.6 \pm 0.2) \text{ TeV}$$

$$\sigma_N^{SI} = 1.2 \cdot 10^{-44} \text{ cm}^2$$

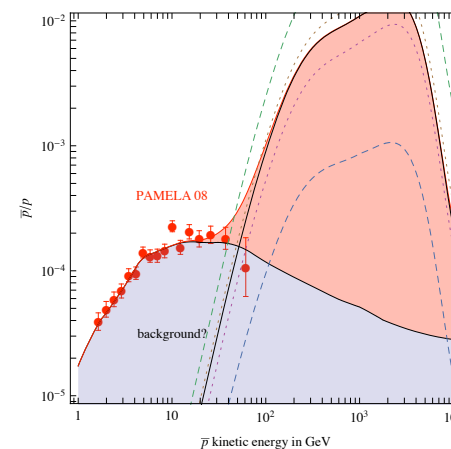
nice fit of Pamela  $e^+$  excess



$e^+ + e^-$  total flux



$\bar{p}$  flux



(large Sommerfeld resonance boost + need of  $\sim 50$  astro boost)

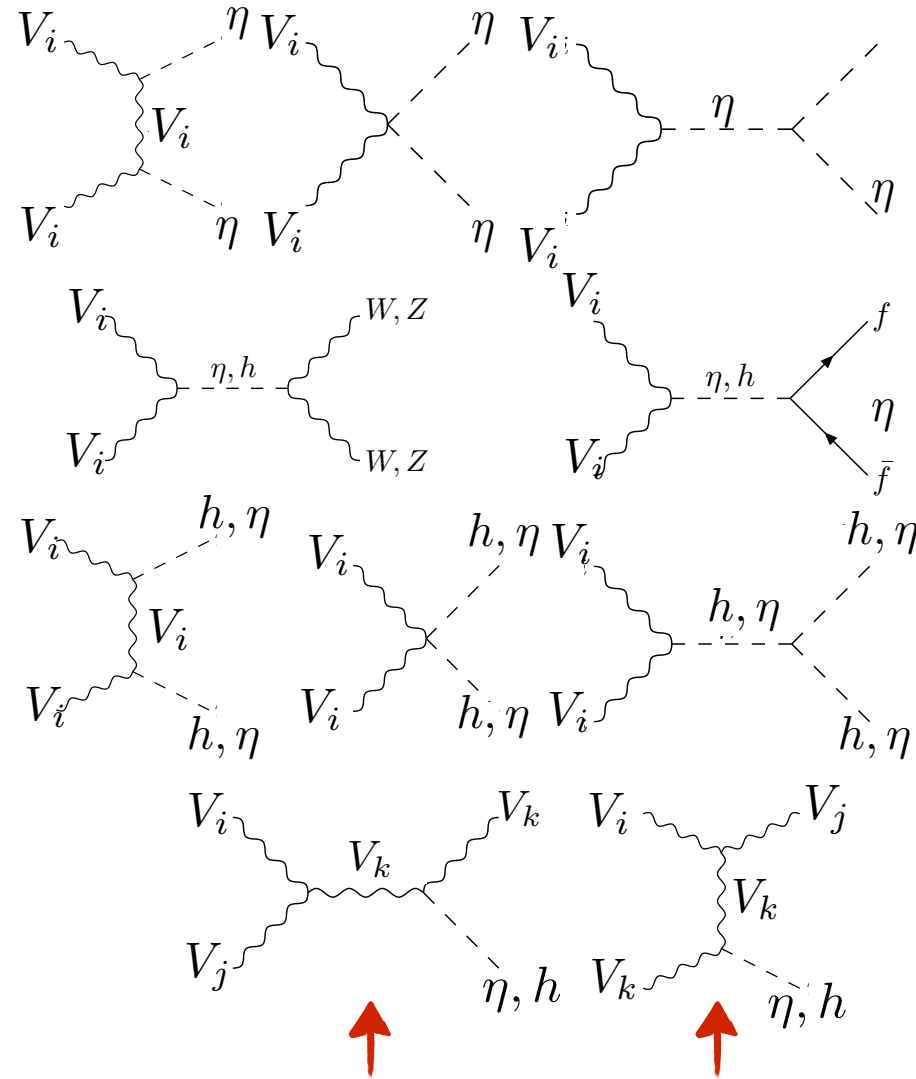
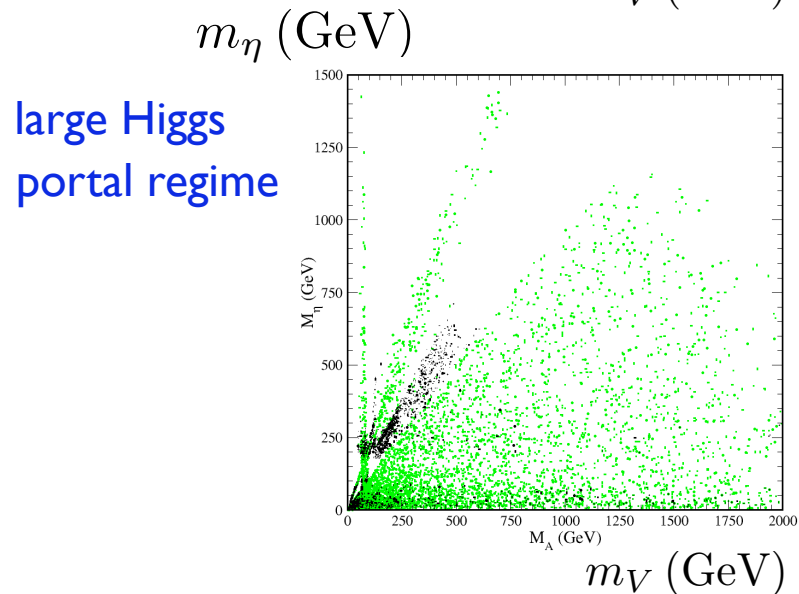
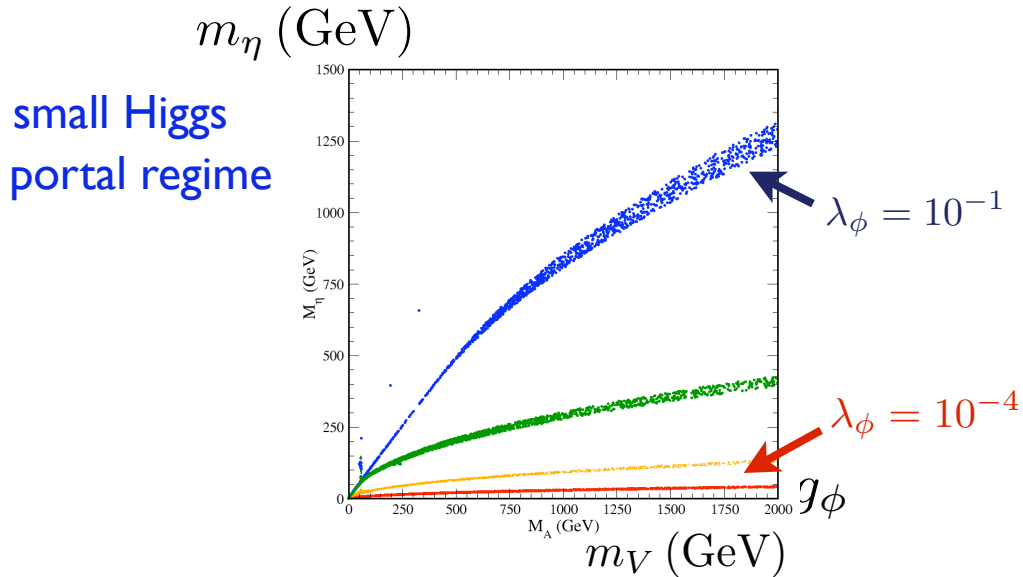
( $m_{DM}$  too high for HESS cutoff)

( $m_{DM}$  high enough to avoid low energy excess)

prefers an isothermal profile for compatibility with galactic center and dwarf galaxy  $\gamma$  flux

# Hidden vector: relic density

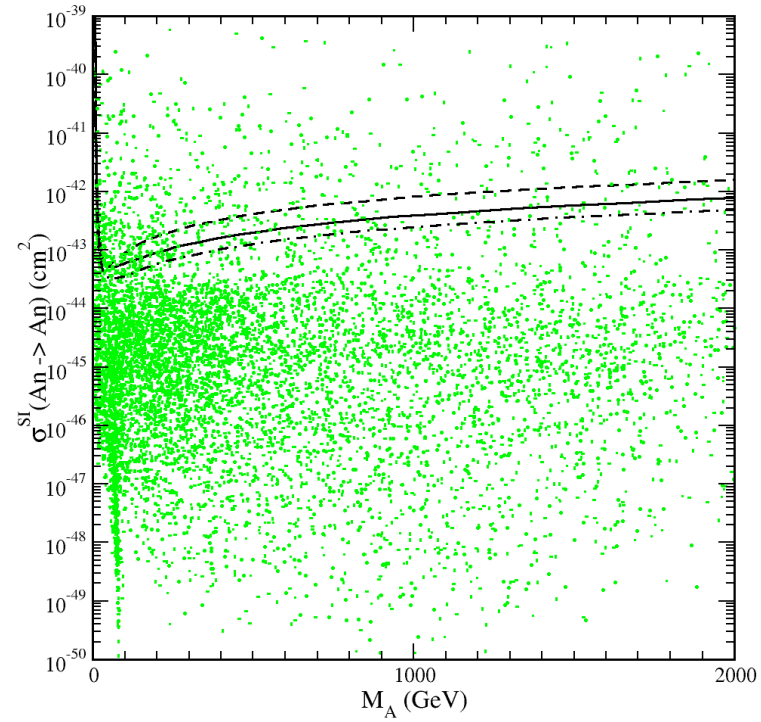
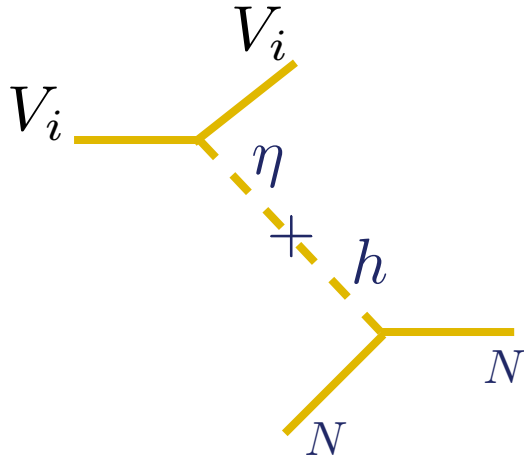
relic density from thermal freezeout




two DM to one DM particle annihilation

$$1 \text{ MeV} < M_{DM} < 10 \text{ TeV}$$

# Hidden vector: direct detection




# Monochromatic $\gamma$ -ray lines: a smoking gun for DM

  $DM DM \rightarrow \gamma\gamma, \gamma Z$  annihilation leads to a monochromatic  $\gamma$ -ray line  
(not expected in astrophysics background)

 e.g. obtained at one loop level  $\Rightarrow$  rather suppressed

Bergstrom, Ullio, 97' 98'; Bern, Gondolo, Perelstein 97';  
Bergstrom, Bringmann, Eriksson, Gustafsson 04', 05';  
Boudjema, Semenov, Temes 05';  
Jackson, Servant, Shaughnessy, Tait, Taoso 09', ...  
one tree level exception: Dudas, Mambrini, Pokorski,  
Romagnoni 09'

  
e.g. needs for large boost factor or a TeV DM mass

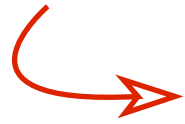
But what about a  $\gamma$ -ray line from DM decay?????

 has been considered from gravitino decay through R-parity violation

Buchmuller, Covi, Hamagushi, Ibarra, Tran 07';  
Ibarra, Tran 07'; Ishiwata, Matsumoto, Moroi 08';  
Buchmuller, Ibarra, Shindou, Takayama, Tran 09';  
Choi, Lopez-Fogliani, Munoz, de Austri 09'

# Dimension-6 operators breaking the custodial symmetry

C.Arina, T.H., A. Ibarra, C. Weniger 09'



$$(A) \quad \frac{1}{\Lambda^2} \mathcal{D}_\mu \phi^\dagger \phi \mathcal{D}_\mu H^\dagger H$$

$$(B) \quad \frac{1}{\Lambda^2} \mathcal{D}_\mu \phi^\dagger \phi H^\dagger \mathcal{D}_\mu H$$

$$(C) \quad \frac{1}{\Lambda^2} \mathcal{D}_\mu \phi^\dagger \mathcal{D}_\nu \phi F^{\mu\nu Y}$$

$$(D) \quad \frac{1}{\Lambda^2} \phi^\dagger F_{\mu\nu}^a \frac{\tau^a}{2} \phi F^{\mu\nu Y}$$

← all give 2-body decay to  $\gamma h$  or  $\gamma \eta$

examples of branching ratios:

Benchmark	$M_A$	$g_\phi$	$v_\phi$	$M_\eta$	$M_h$	$\sin \beta$
1	300 GeV	0.55	1090 GeV	30 GeV	150 GeV	$\approx 0$
2	600 GeV	0.6	2000 GeV	30 GeV	120 GeV	$\approx 0$
3	14 TeV	12	2333 GeV	500 GeV	145 GeV	$\approx 0$
4	1550 GeV	2.1	1457 GeV	1245 GeV	153 GeV	0.25

Benchmark	$\eta\eta$	$h\eta$	$hh$	$\gamma\eta$	$Z\eta$	$\gamma h$	$Zh$
1	-	0.09	-	0.04	0.02	0.65	0.20
2	-	0.04	0.62	0.002	0.003	0.15	0.18
3	-	0.04	0.80	$3 \times 10^{-6}$	0.002	0.0003	0.16

operator A & B

Benchmark	$Z\eta$	$\gamma\eta$	$Zh$	$\gamma h$
1	0.19	0.81	0	0
2	0.22	0.78	0	0
3	0.23	0.77	0	0
4	0.028	0.79	0.041	0.14

operator C

Benchmark	$Z\eta$	$Zh$	$\gamma\eta$	$W^+W^-$	$\nu\bar{\nu}$	$e^+e^-$	$u\bar{u}$	$d\bar{d}$
1	0.01	0.005	0.04	0.02	0.09	0.39	0.29	0.15
2	0.019	0.004	0.036	0.014	0.072	0.35	0.39	0.12
3	0.22	0.0002	0.73	0.0005	0.003	0.016	0.018	0.005

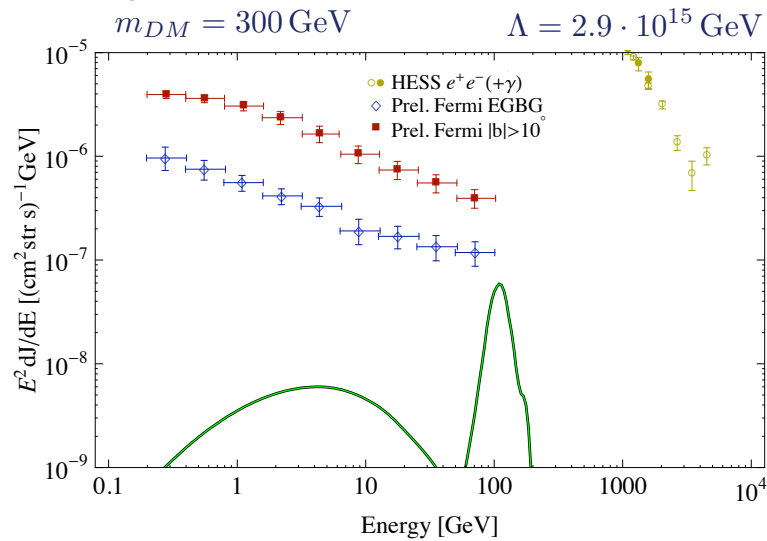
operator D



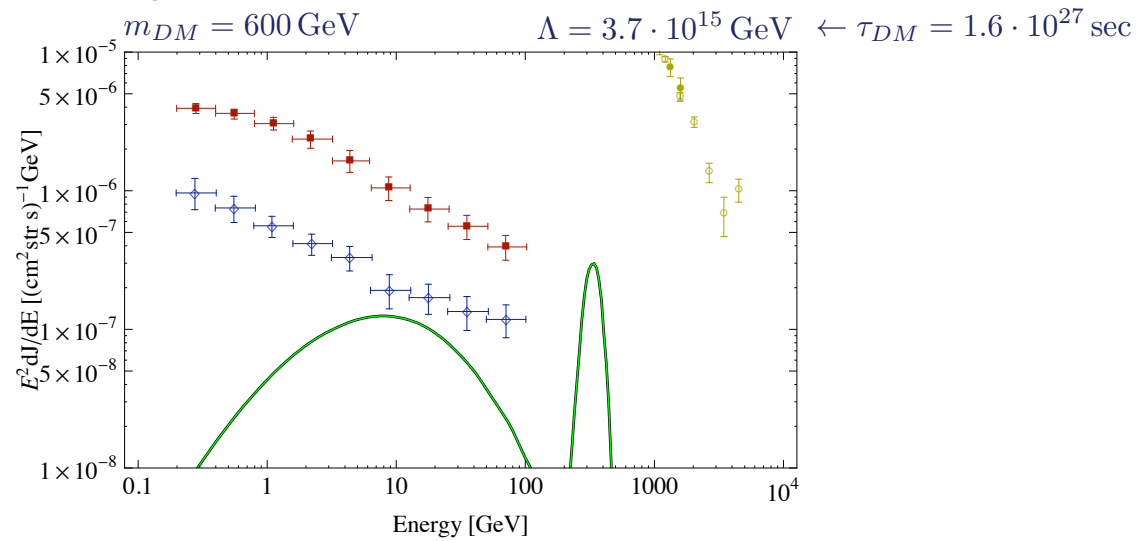
# Flux of monochromatic $\gamma$ -rays

$$0 \leq l \leq 360^\circ, 10^\circ \leq |b| \leq 90^\circ$$

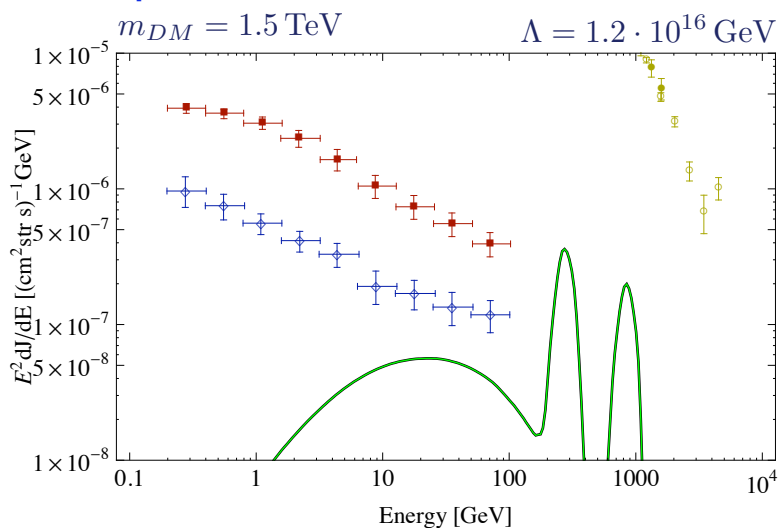
## operator A & B



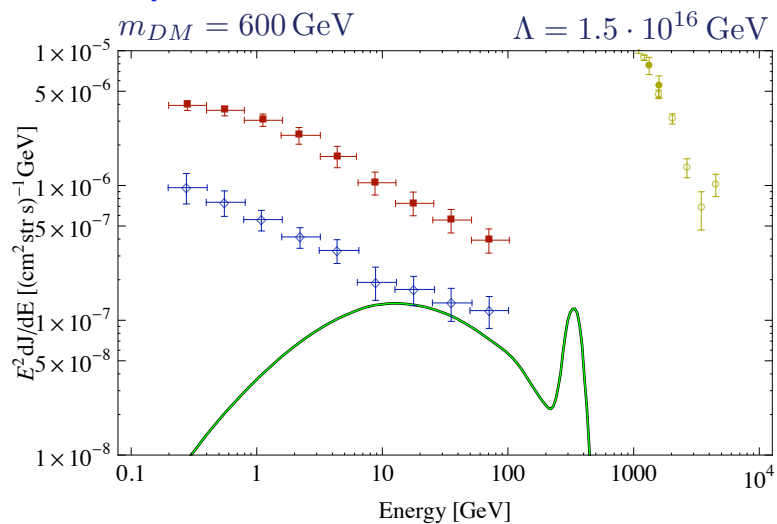
## operator A & B



## operator C



## operator D



# DM stability from accidental symmetry: Weakly interacting stable pions

Bai, Hill 10'

QCD interactions conserve G-parity:  $\rho \rightarrow \pi\pi$  but  $\rho \rightarrow \pi\pi\pi$

$\begin{matrix} \uparrow & & \uparrow\uparrow \\ + & & -- \end{matrix}$

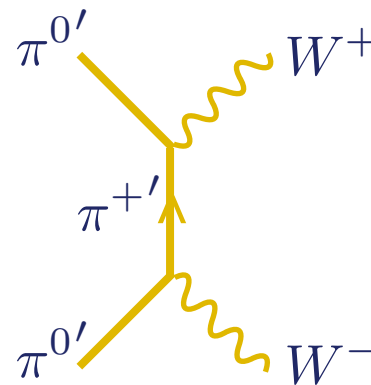
pions would be stable if there were only QCD

but G-parity is not conserved by  $\begin{matrix} \rightarrow \\ \rightarrow \end{matrix}$  weak interactions ( $V - A$  structure)  
hypercharge interactions

$\Rightarrow$  assume a new QCD structure with new quarks  $\psi$  with  $\begin{matrix} \rightarrow \\ \rightarrow \end{matrix}$  no hypercharge  
vector weak interact.

$\Downarrow$   
G-parity is exactly conserved

$\Downarrow$   
the  $\pi^{0'}$  is stable and is a WIMP



$\Downarrow$   
 $\psi_L = \text{doublet}$   
 $\psi_R = \text{doublet}$

$m_{DM} = 100 \text{ GeV} - 2 \text{ TeV}$

$\sigma_N^{SI} \sim 10^{-45} \text{ cm}^2$

but dim-5 operators breaking G-parity are allowed  $\Rightarrow$  DM decay would be too fast

# Other naturally stable DM candidates

---

- DM stability from small couplings (suppressed by heavy scale) and/or small DM mass:

- axion
- KeV right-handed neutrino
- gravitino
- ...

P. Sikivie's talk

F. Bezrukov's talk

Servant, Tait 06', ...

- KK parity in Universal Extra Dimension models (assuming orbifold,...)

- $U(1)_{B'}$  in technicolor models

Gudnason, Kouvaris, Sannino 06'

- DM stability from a flavour symmetry


Hirsch, Morisi, Peinado, Valle 10'

- .....

More Backup

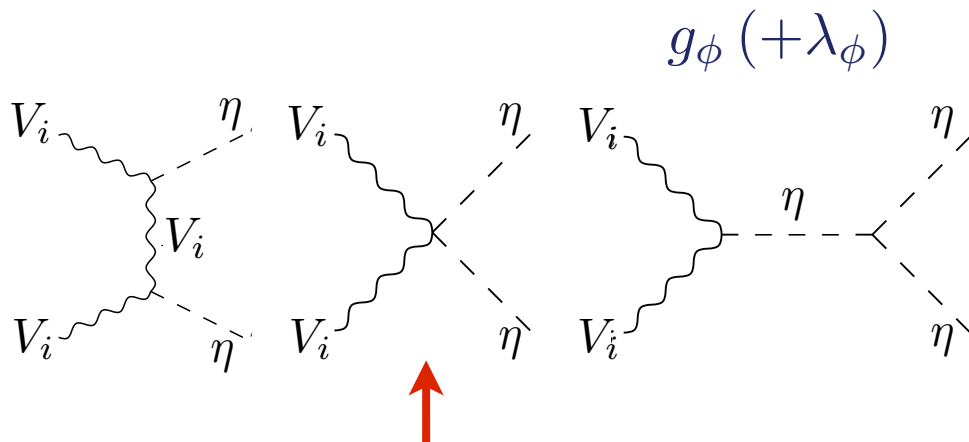
# Relic density

- $T \gtrsim m_V : V_{1,2,3}^\mu$  in thermal equilibrium with SM thermal bath


 $\eta$  with  $h$  : due to  $\lambda_m$  coupling  
 $V_i$  with  $\eta$  : due to  $g_\phi$  coupling

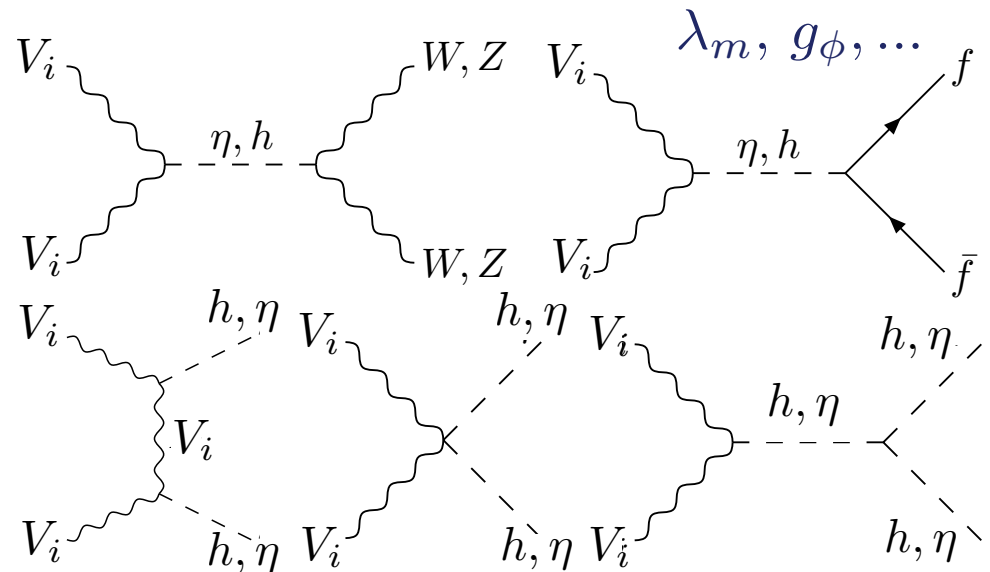
- $T < m_V : n_V^{eq.} \sim e^{-m_V/T} \Rightarrow$  annihilation freeze out (WIMP)

to two real  $\eta$ :



with subsequent decay of  $\eta$  to SM particles via  $h - \eta$  mixing

with at least one SM part. in final state:



# Relic density: additional new type of contribution

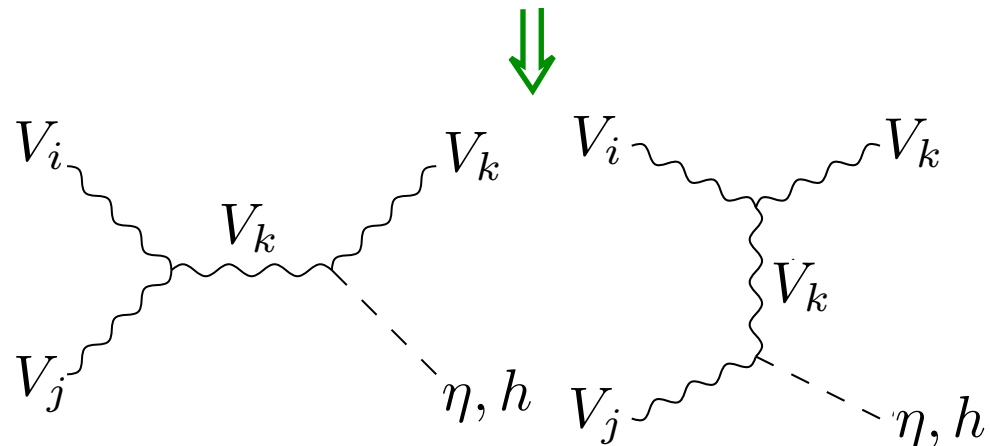
→ non abelian trilinear gauge couplings:

$$F_{\mu\nu}^a F^{\mu\nu a} \ni \varepsilon_{ijk} \partial_\mu A_{i\nu} (A_j^\mu A_k^\nu - A_j^\nu A_k^\mu)$$

do not lead to any  $V_i$  decay even if trilinear (carries 3  $\neq$  indices)

but induces two DM to one DM particle annihilation

$\neq$  from the  $Z_2$  case



⇒ no dramatic effect for the freeze out (same order as other diagrams)

# Small Higgs portal regime

→  $\lambda_m \lesssim 10^{-3}$  ← (but larger than  $\sim 10^{-7}$  to have thermalization with the SM bath)

→  $V_i V_i \rightarrow \eta\eta, V_i V_j \rightarrow V_k \eta$  dominant

→ depend only on  $g_\phi, v_\phi, \lambda_\phi$  with  $m_V = \frac{g_\phi v_\phi}{2}, m_\eta \simeq \sqrt{2\lambda_\phi} v_\phi$

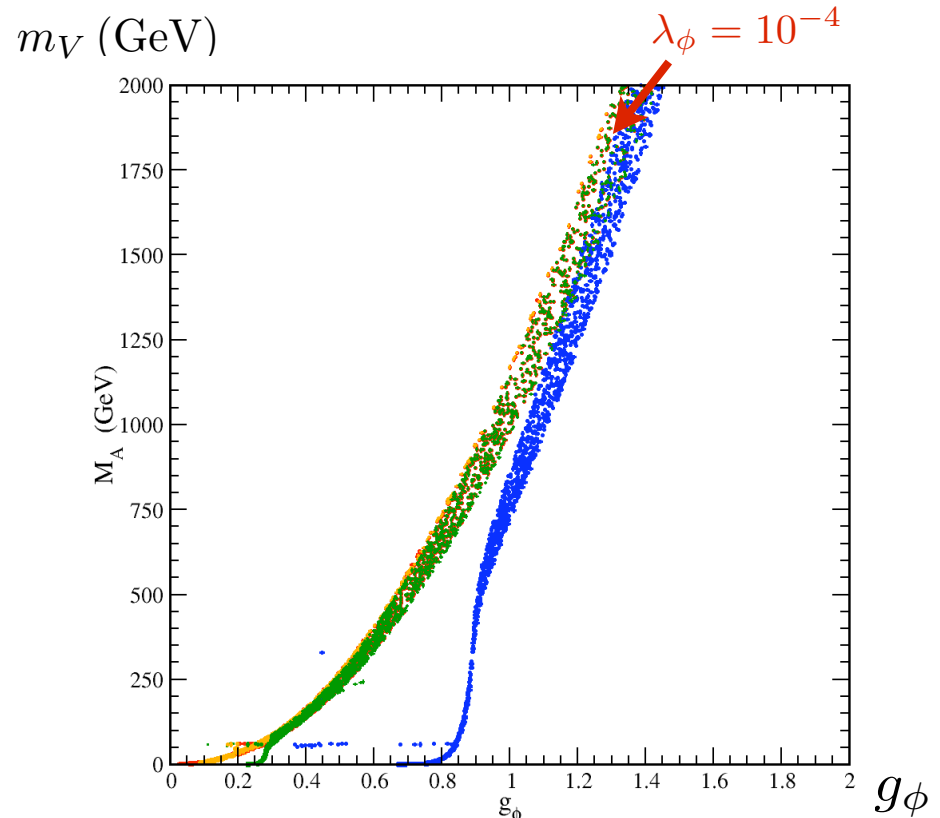
⇒ if  $\lambda_\phi$  also small:

$$\sigma_{\text{annih.}} \sim \frac{g_\phi^4}{m_V^2} \sim \frac{g_\phi^2}{v_\phi^2}$$



$$m_V \propto g_\phi^2 \quad (\propto v_\phi^2)$$

⇒  $1 \text{ MeV} \lesssim m_{DM} \lesssim 25 \text{ TeV}$



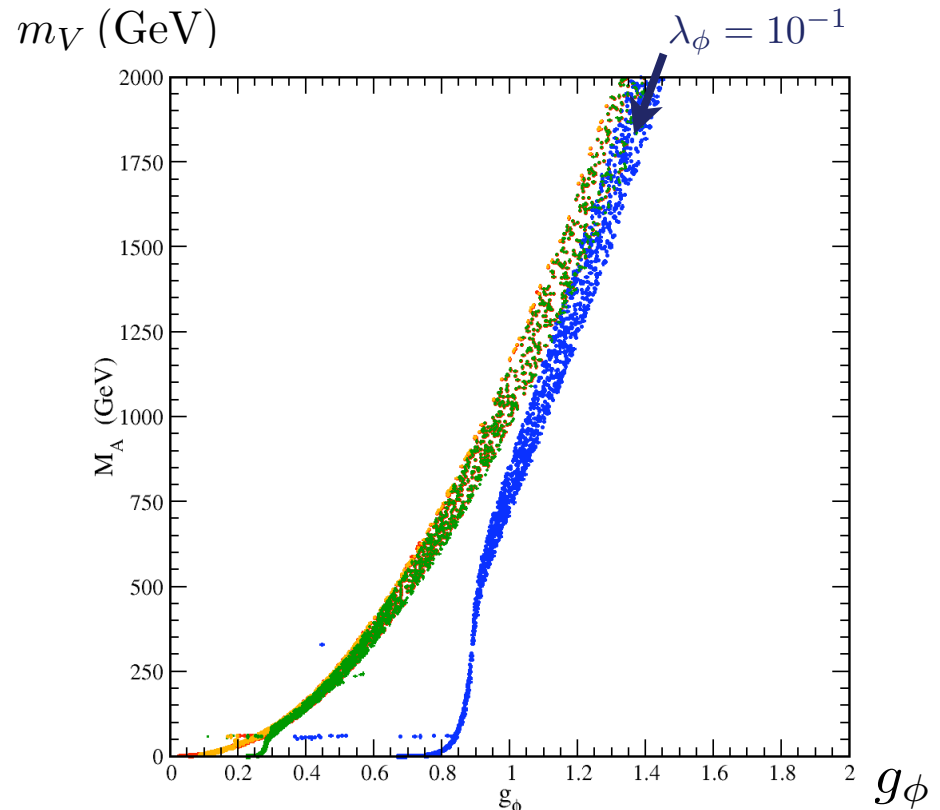
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→ depend only on  $g_\phi, v_\phi, \lambda_\phi$  with  $m_V = \frac{g_\phi v_\phi}{2}, m_\eta \simeq \sqrt{2\lambda_\phi} v_\phi$

⇒ if  $\lambda_\phi$  large:

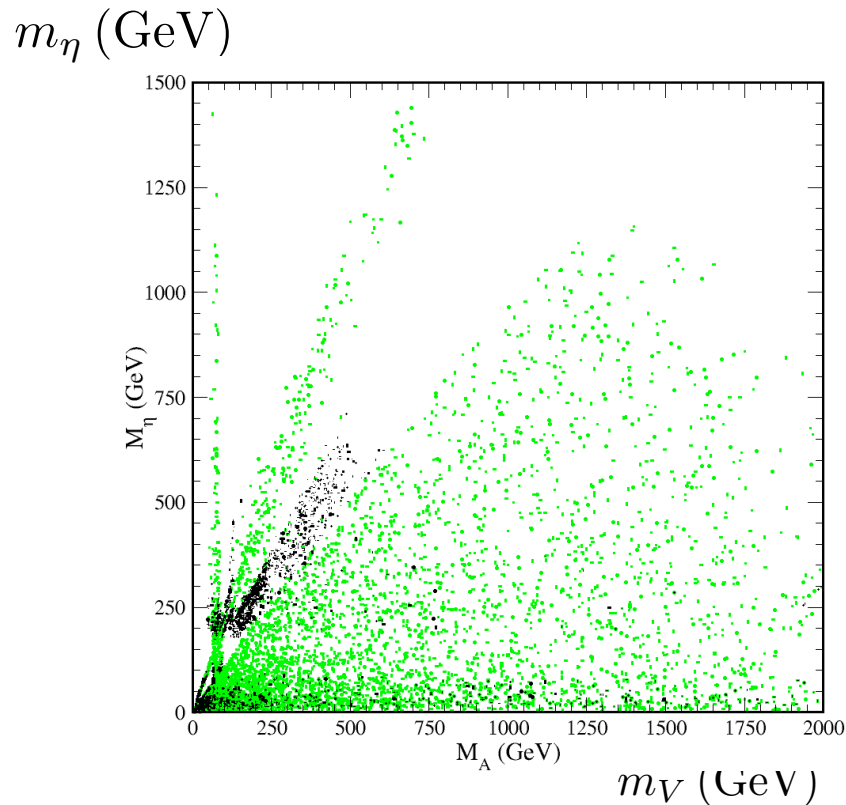




# Large Higgs portal regime

$\lambda_m \gtrsim 10^{-3} \Rightarrow$  large  $\eta - h$  mixing  $\Rightarrow$  large hidden sector - SM mixing

$\Rightarrow$  can lead to the right  $\Omega_{DM}$  even for maximal mixing



production at LHC of  $\eta$  just as for the Higgs in the SM but with possibly a larger mass



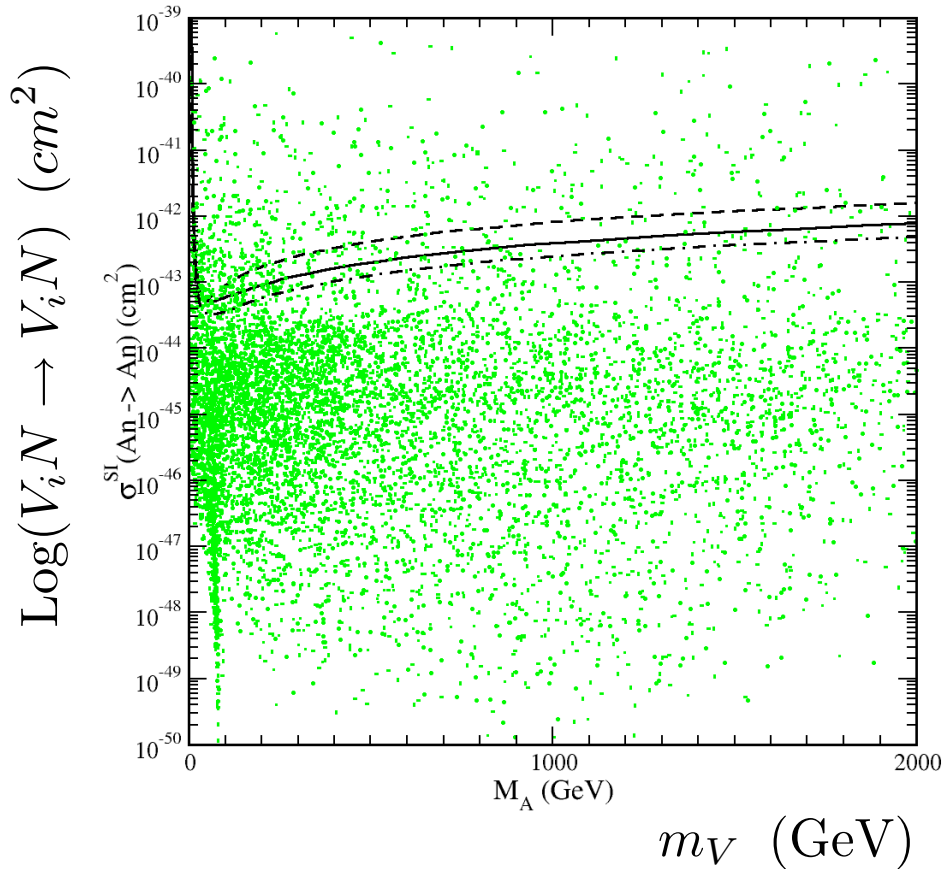
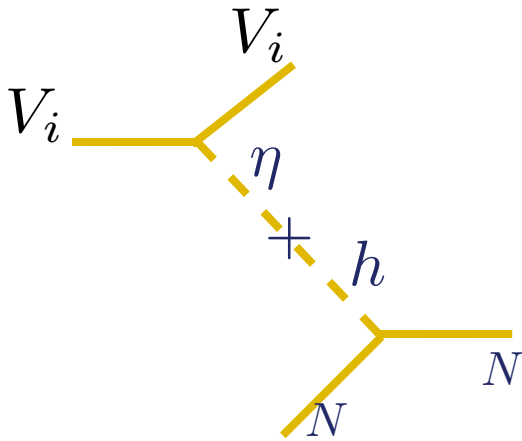
T parameter constraint:

if  $m_h = 120$  GeV  $\Rightarrow m_\eta < \sim 240$  GeV ( $3\sigma$ )

$\Rightarrow$  or larger if non maximal mixing

if  $m_\eta = m_h \Rightarrow m_h = m_\eta < 154$  GeV ( $3\sigma$ )

# Hidden vector: direct detection



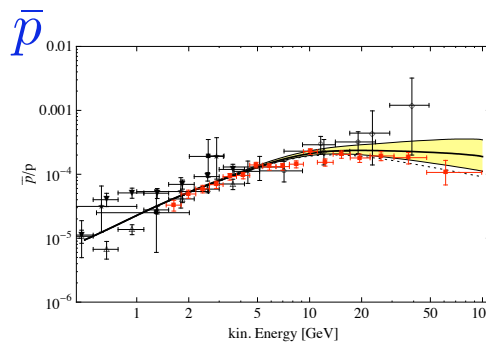
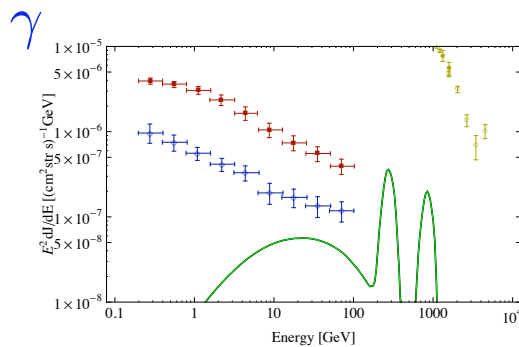
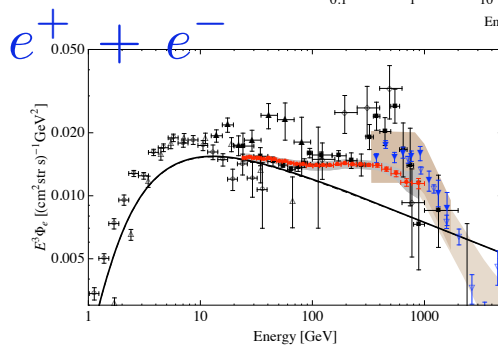
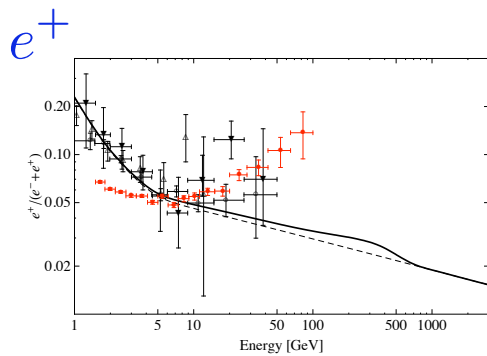
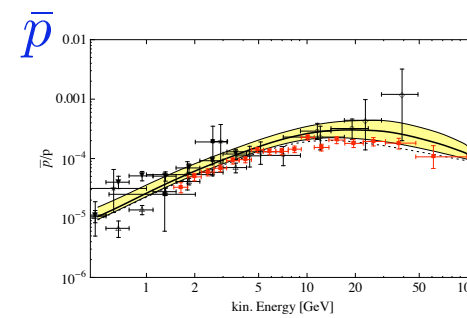
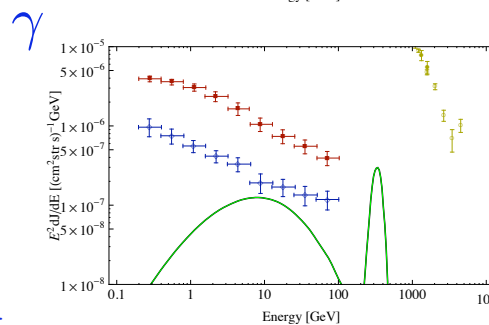
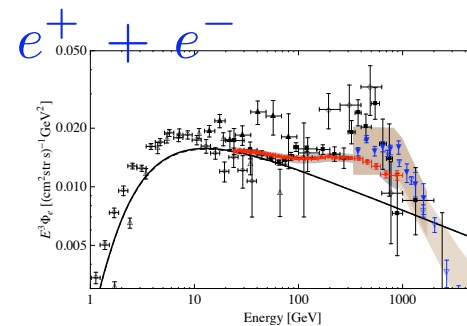
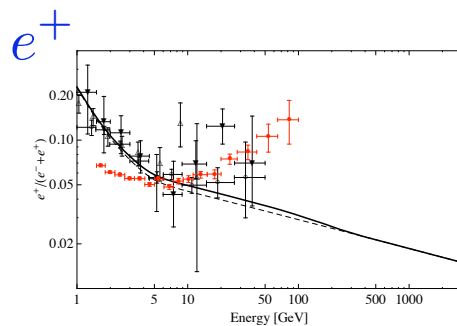
$\Rightarrow$  can saturate the experimental bound easily

# Hidden vector: cosmic ray fluxes

operator A & B

$$m_{DM} = 300 \text{ GeV}$$

$$\Lambda = 2.9 \cdot 10^{15} \text{ GeV}$$



operator C

$$m_{DM} = 1.5 \text{ TeV}$$

$$\Lambda = 1.2 \cdot 10^{16} \text{ GeV}$$

# What about the non-perturbative regime of this model?

T.H., M. Tytgat, arXiv:0907.1007

→  $SU(2)_{\text{Hidden Sect.}}$  confines automatically if  $\Lambda_{SU(2)} \gg v_\phi$

↓                      ↓  
dynamical          perturbative  
scale                breaking scale

→ but the custodial symmetry remains exact in this case too

't Hooft '98

⇒  $\phi$  confines: boundstates are eigenstates of the custodial sym.:

- scalar state:  $S \equiv \phi^\dagger \phi$  singlet of  $SO(3)$  expected the lightest

- “charged” vector state:  $V_\mu^+ \equiv \phi^\dagger D_\mu \tilde{\phi}$

$$V_\mu^- \equiv \tilde{\phi}^\dagger D_\mu \phi$$

- “neutral” vector state:  $V_\mu^0 \equiv \frac{\phi^\dagger D_\mu \phi - \tilde{\phi}^\dagger D_\mu \tilde{\phi}}{\sqrt{2}}$

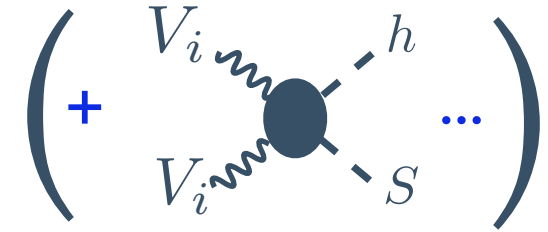
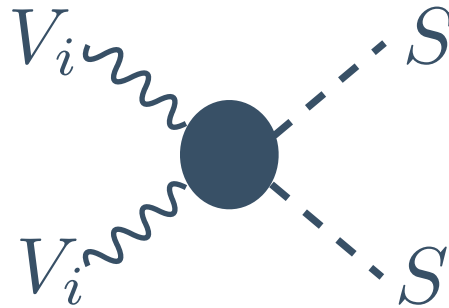
}  $SO(3)$  triplet  
↓ ↓  
stable DM candidates!

# Relic density in the confined regime

strongly interactive massive particle (SIMP)

annihilation cross section cannot be calculated perturbatively

expected dominant channel:



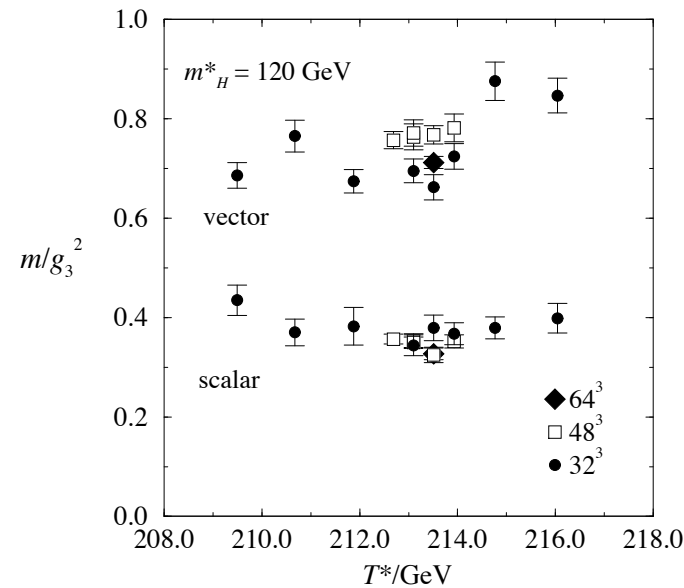
if  $S - h$  mixing is large (for large  $\lambda_m$ )

$$\sigma_{annih.} \sim \frac{A}{\Lambda_{SU(2)}^2} \xrightarrow{A = 10 - 50} \underline{m_{DM} \simeq 20 - 120 \text{ TeV}}$$

confining non-abelian hidden sector coupled to the SM through the Higgs portal: perfectly viable DM candidate

# Expected spectrum (in a similar case)

vector states e.g. expected heavier than scalar ones:



Kajantie, Laine, Rummukainen, Shaposhnikov '96

# Possible effects on Electroweak Symmetry Breaking

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→ contribution of the vev of the hidden scalar to the Higgs mass term:

$$\mathcal{L}_{Higgs\ portal} = -\lambda_m \phi^\dagger \phi H^\dagger H$$

→  $\ni -\lambda_m v_\phi^2 H^\dagger H$



gives a contribution to the Higgs vev:  $v^2 \propto \frac{\lambda_m}{\lambda_H} v_\phi^2 \propto m_{DM}^2$



gives a hint for the  $m_{DM}$  versus  $v$  WIMP coincidence

see also T.H, M.Tytgat, arXiv 0707.0633, (PLB 659)