

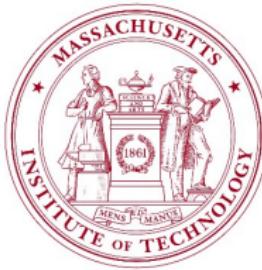
Decoupling the Gravity Multiplet from Supergravity

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C. Cheung, FDE, J. Thaler; arXiv:1104.2598 [hep-ph], arXiv:1104.2600 [hep-ph].



Motivation

SUSY symmetry of nature \Rightarrow SUGRA

Not all of the SUGRA formalism relevant
for phenomenology at colliders and in cosmology

Our Goal: Framework to Simplify Supergravity Calculations

decoupling the gravity multiplet from matter fields
and
accounting for SUGRA effects

Conformal SUGRA

Minimal SUGRA: gauge fixing of conformal SUGRA

Gates et al, NPB 147(1979), FP 58(1983); Linch et al, PRD 68(2003), [hep-th/0209060].

Gauge Fixing in Superspace: Conformal Compensator Formalism

Usual superspace of global SUSY + conformal compensator superfield Φ

$$\begin{aligned}\mathcal{L}_{\text{SUGRA}} = & -3 \int d^4\theta \Phi^\dagger \Phi e^{-K/3} + \int d^2\theta \Phi^3 W + \text{h.c.} \\ & + \frac{1}{4} \int d^2\theta f_{ab} W^{a\alpha} W_\alpha^b + \text{h.c.} + \dots\end{aligned}$$

Standard Gauge Fixing

Lowest and fermionic components
of Φ are pure gauge mode:

$$\Phi = 1 + \theta^2 F_\phi$$

Naive application of standard Φ :

neglect terms in $\dots \Rightarrow$ incorrect answers in many cases

Problematic Terms

Standard choice for Φ : mixing between gravity and matter multiplets

Graviton normalization and kinetic mixing

$$C R/6$$

E.H. action normalization
Graviton/matter mixing

Gravitino kinetic mixing

$$i\xi\sigma^{\mu\nu}\partial_\mu\psi_\nu + \text{h.c.}$$

Gravitino/matter mixing
Non-canonical gravitino

Gravitino mass phase

$$-z^\dagger\psi_\mu\sigma^{\mu\nu}\psi_\nu + \text{h.c.}$$

Gravitino mass real

C , ξ and z functions of matter fields

Additional terms must be taken into account in the calculations

Standard gauge fixing:
need to work with component fields
no simple interpretations in terms of superfields

The Kugo-Uehara gauge

Gauge Freedoms

Lowest and fermionic components of Φ pure gauge modes

They can be fixed to **any** values by an appropriate gauge choice

Kugo-Uehara gauge

Enough freedom for $C = -3$, $\xi_\alpha = 0$, $\text{Arg}[z] = 0$ to all orders in fields

$$\Phi = \exp \left[\frac{1}{3} (K/2 - i \text{Arg } W) \right] \times \left\{ 1, \frac{K_i \chi^i}{3}, F_\phi \right\}$$

Kugo and Uehara, NPB B222(1983).

Hidden problem of KU gauge

Mixing with vector auxiliary field b_μ : $b_\mu \partial^\mu \phi$

Integrating out b_μ \Rightarrow gravity/matter mixing

Need to work again with component fields

A Novel Gauge Fixing

A less stringent gauge choice

$C = -3$, $\xi_\alpha = 0$, $\text{Arg}[z] = 0$ to linear order in field fluctuations

Cheung, FDE, Thaler, arXiv:1104.2598 [hep-ph].

$$\Phi = e^{Z/3}(1 + \theta^2 F_\Phi), \quad Z = \langle K/2 - i \text{Arg } W \rangle + \langle K_i \rangle X^i$$

Coupling with the vector auxiliary field?

Gauge fixing $\Rightarrow b_\mu = 0 + \mathcal{O}(1/M_{\text{Pl}})$ $\Rightarrow 1/M_{\text{Pl}}^2$ suppressed operators

No need to perform component manipulations

Gravity multiplet **decoupled** from matter fields calculations

Phenomenological SUGRA Lagrangian

$$\begin{aligned}\mathcal{L}_{\text{SUGRA}} = & -3 \int d^4\theta \Phi^\dagger \Phi e^{-K/3} + \int d^2\theta \Phi^3 W + \text{h.c.} \\ & + \frac{1}{4} \int d^2\theta f_{ab} W^{a\alpha} W_\alpha^b + \text{h.c.} + \mathcal{O}(1/M_{\text{Pl}})\end{aligned}$$

$$\Phi = e^{Z/3}(1 + \theta^2 F_\Phi), \quad Z = \langle K/2 - i \operatorname{Arg} W \rangle + \langle K_i \rangle X^i$$

Applications: Fermionic Spectra

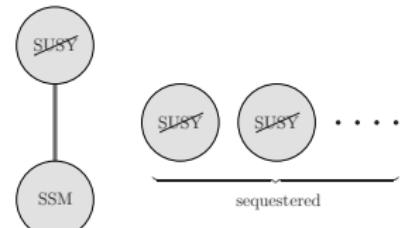
N sequestered SUSY sectors: $\text{SUSY}^N \equiv \otimes \prod_{i=1}^N \text{SUSY}_i$

Gravity explicitly breaks to diagonal SUSY

N goldstini η_i in the theory:

- linear combination η_{long} eaten by the gravitino
- $N - 1$ goldstini in the spectrum

Cheung, Nomura, Thaler, JHEP 1003 (2010) [arXiv:1002.1967 [hep-ph]].



Minimal scenario: N sequestered sectors, F and D breaking, $\langle K_i \rangle = 0$

$$m_\eta = 2 m_{3/2}$$

Modulus field with shift symmetry (almost "no-scale" SUSY breaking)

massless modulino

Cheung, FDE, Thaler, arXiv:1104.2600 [hep-ph].

Summary

Supergravity Computations without Gravity Complications

C. Cheung, FDE, J. Thaler, arXiv:1104.2598 [hep-ph].

$$\begin{aligned}\mathcal{L}_{\text{SUGRA}} = & -3 \int d^4\theta \Phi^\dagger \Phi e^{-K/3} + \int d^2\theta \Phi^3 W + \text{h.c.} \\ & + \frac{1}{4} \int d^2\theta f_{ab} W^{a\alpha} W_\alpha^b + \text{h.c.} + \mathcal{O}(1/M_{\text{Pl}})\end{aligned}$$

$$\Phi = e^{Z/3}(1 + \theta^2 F_\Phi), \quad Z = \langle K/2 - i \arg W \rangle + \langle K_i \rangle X^i$$

The Spectrum of Goldstini and Modulini

C. Cheung, FDE, J. Thaler, arXiv:1104.2600 [hep-ph].

Fermionic spectra calculation directly in superspace