

Planck 2011, IST, Lisboa

May 31st 2011

# **Light Neutralino in the MSSM**

Lorenzo Calibbi

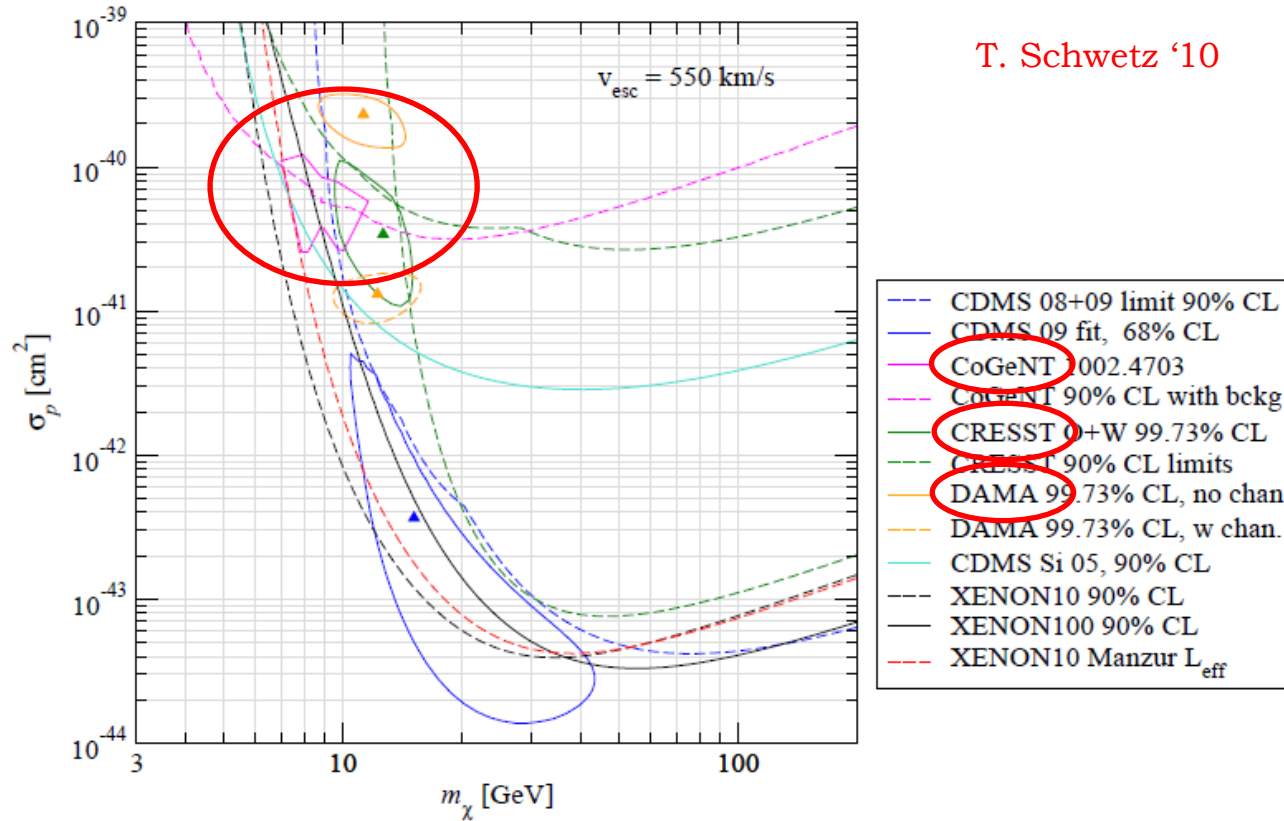
Max-Planck-Institut für Physik, Munich



MAX-PLANCK-GESELLSCHAFT

based on: L.C., T. Ota, Y. Takanishi, arXiv:1104.1134 [hep-ph]

# Motivations



T. Schwetz '10

$\Rightarrow m_\chi \sim 7 \div 10 \text{ GeV}, \quad \sigma_{\chi N}^{\text{SI}} \sim 10^{-41} \div 10^{-40} \text{ cm}^2$

e.g. Hooper et al. '10

Is it possible with MSSM Neutralino DM?

Kuflik et al. '10, Feldman et al. '10,  
Vasquez et al. '10, Fornengo et al. '10

# Light Neutralino Dark Matter

MSSM neutralinos:  $(\tilde{B}, \tilde{W}_3, \tilde{H}_d^0, \tilde{H}_u^0)$

MSSM charginos:

$$\mathbf{M}_{\tilde{N}} = \begin{pmatrix} M_1 & 0 & -c_\beta s_W m_Z & s_\beta s_W m_Z \\ 0 & M_2 & c_\beta c_W m_Z & -s_\beta c_W m_Z \\ -c_\beta s_W m_Z & c_\beta c_W m_Z & 0 & -\mu \\ s_\beta s_W m_Z & -s_\beta c_W m_Z & -\mu & 0 \end{pmatrix} \quad \mathcal{M}_\pm = \begin{pmatrix} M_2 & \sqrt{2} M_W \sin\beta \\ \sqrt{2} M_W \cos\beta & \mu \end{pmatrix}$$

Lightest eigenstate:

$$\tilde{\chi}_1^0 = a_1 \tilde{B} + a_2 \tilde{W}_3 + a_3 \tilde{H}_d^0 + a_4 \tilde{H}_u^0$$

LEP chargino searches:



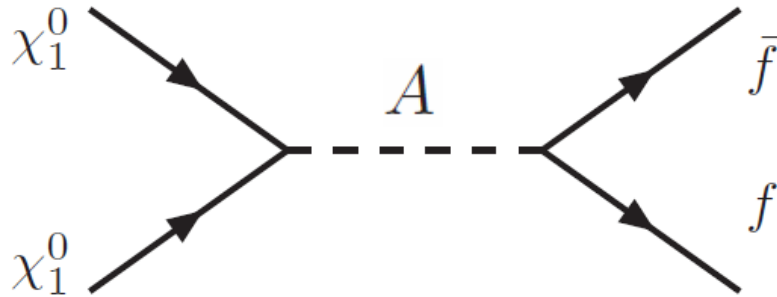
$$M_2, \mu \gtrsim 90 \text{ GeV}$$

$$M_1 \ll M_2, \mu \iff \tilde{\chi}_1^0 \approx \tilde{B}$$

Efficient annihilation required

(sfermion exchange not enough due to LEP mass limits e.g. Dreiner et al. '09)

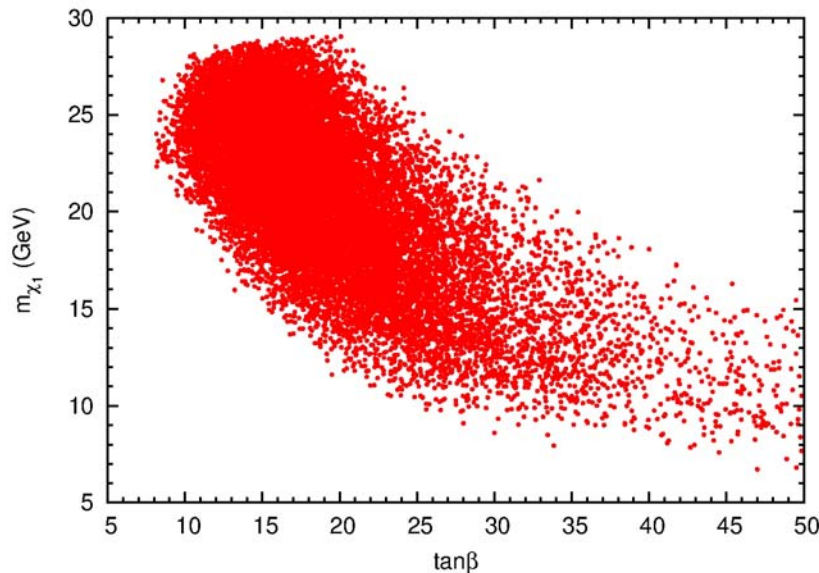
# Light Neutralino Dark Matter



$$\propto \frac{m_{\tilde{\chi}_1^0}^2}{m_A^2} \frac{m_f}{v} (a_1 a_{3,4}) \tan \beta$$

Drees Nojiri '92

Small  $m_{\tilde{\chi}_1^0} \iff$  large  $\tan \beta$ , small  $m_A$  (and  $\mu$ ) Bottino et al. '02, '04, '08, '10



Light Higgs sector, large(-ish)  $\tan \beta$   
 $\implies$  phenomenologically challenging

see e.g.: Buras et al. '02,  
Isidori Paradisi '06,  
Barenboim et al. '07,  
Eriksson et al. '08,  
Altmannshofer et al. '09,  
Altmannshofer Straub '10,

...

# Low-energy constraints

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Mild dependence on SUSY parameters

Strong dependence on SUSY parameters

$$B^+ \rightarrow \tau^+ \nu_\tau$$

$$B^+ \rightarrow D^0 \tau^+ \nu_\tau$$

$$K^+ \rightarrow \mu^+ \nu_\mu$$

$$D_s^+ \rightarrow \tau^+ \nu_\tau$$

$$b \rightarrow s \gamma$$

$$B_s \rightarrow \mu^+ \mu^-$$

# Low-energy constraints

Mild dependence on SUSY parameters

Strong dependence on SUSY parameters

$$\begin{aligned}
 B^+ &\rightarrow \tau^+ \nu_\tau \\
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 \end{aligned}$$

$$\begin{aligned}
 b &\rightarrow s \gamma \\
 B_s &\rightarrow \mu^+ \mu^-
 \end{aligned}$$

$$R_{B\tau\nu} \equiv \frac{\text{BR}(B \rightarrow \tau\nu)}{\text{BR}(B \rightarrow \tau\nu)_{\text{SM}}} \simeq \left[ 1 - \frac{m_B^2}{m_{H^\pm}^2} \frac{\tan^2 \beta}{1 + \epsilon \tan \beta} \right]^2$$

Hou '93,  
Akeroyd Recksiegel '03,  
Isidori Paradisi '06

$$R_{\ell 23} \equiv \left| \frac{V_{us}(K \rightarrow \ell\nu)}{V_{us}(K \rightarrow \pi\ell\nu)} \times \frac{V_{ud}(\beta \text{ decay})}{V_{ud}(\pi \rightarrow \ell\nu)} \right| \simeq \left| 1 - \frac{m_K^2}{m_{H^\pm}^2} \left[ 1 - \frac{m_d}{m_s} \right] \frac{\tan^2 \beta}{1 + \epsilon \tan \beta} \right|$$

Antonelli et al. '08, '10

$$R_{D\ell\nu} \equiv \frac{\text{BR}(B \rightarrow D\tau\nu)}{\text{BR}(B \rightarrow D\ell\nu)}$$

Kamenik Mescia '08

$$\text{BR}(D_s \rightarrow \tau\nu)$$

Akeroyd Mahmoudi '09

# Low-energy constraints

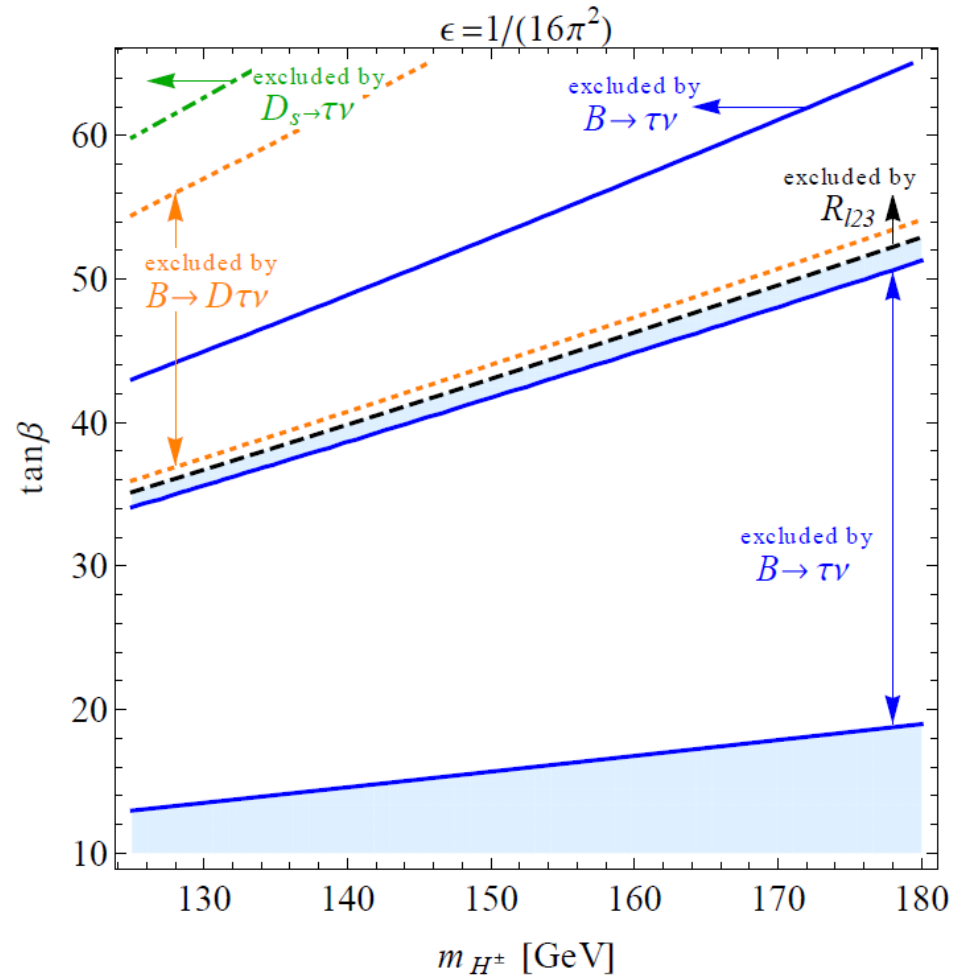
Experimental ranges:

$$0.52 < R_{B\tau\nu} < 2.61$$

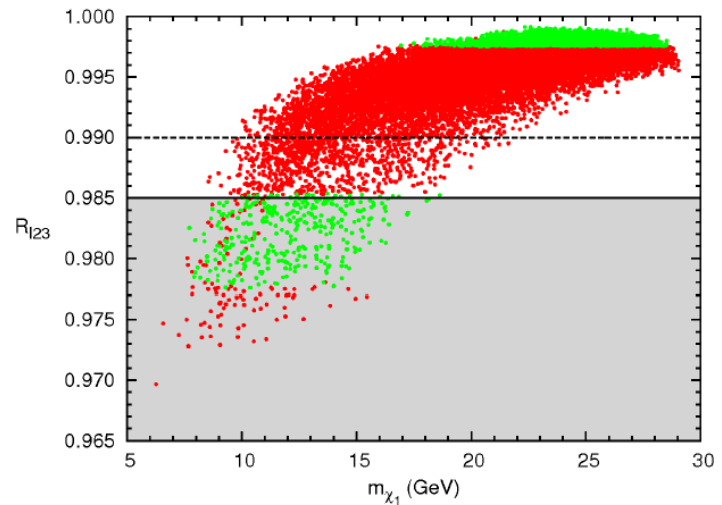
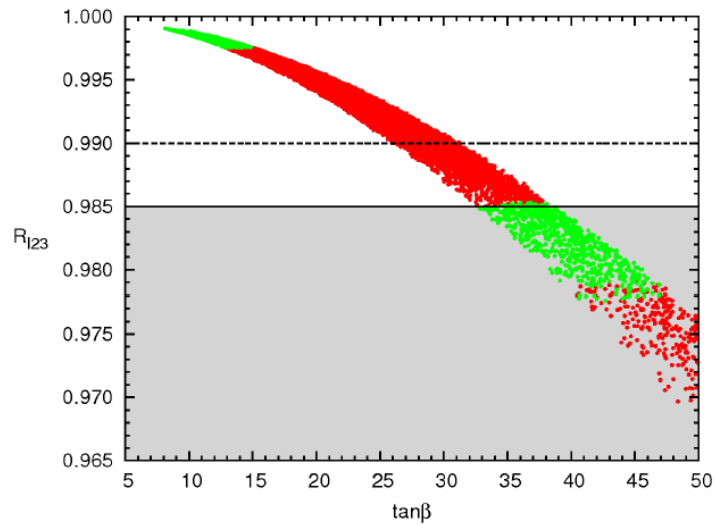
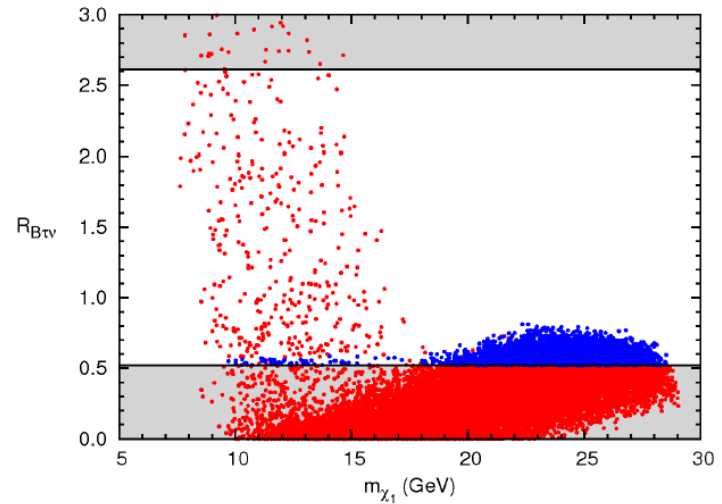
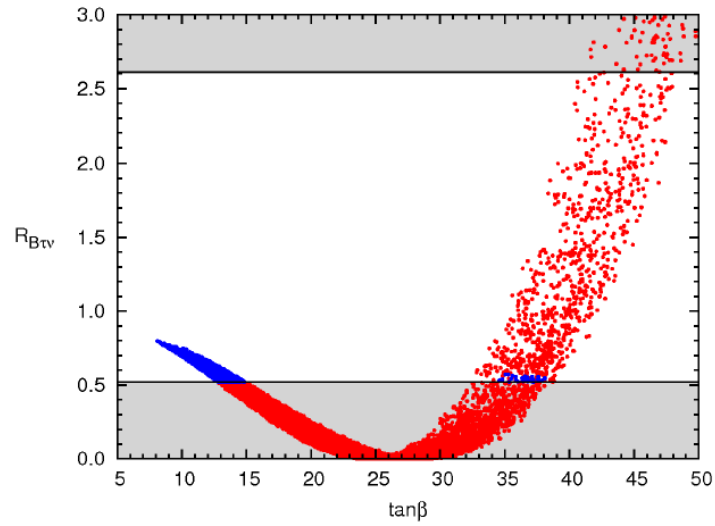
$$0.985 < R_{\ell 23}(K \rightarrow \mu\nu) < 1.013$$

$$0.151 < R_{D\ell\nu} < 0.681$$

$$4.7 \times 10^{-2} < \text{BR}(D_s \rightarrow \tau\nu) < 6.1 \times 10^{-2}$$



# Low-energy constraints





# Numerical analysis

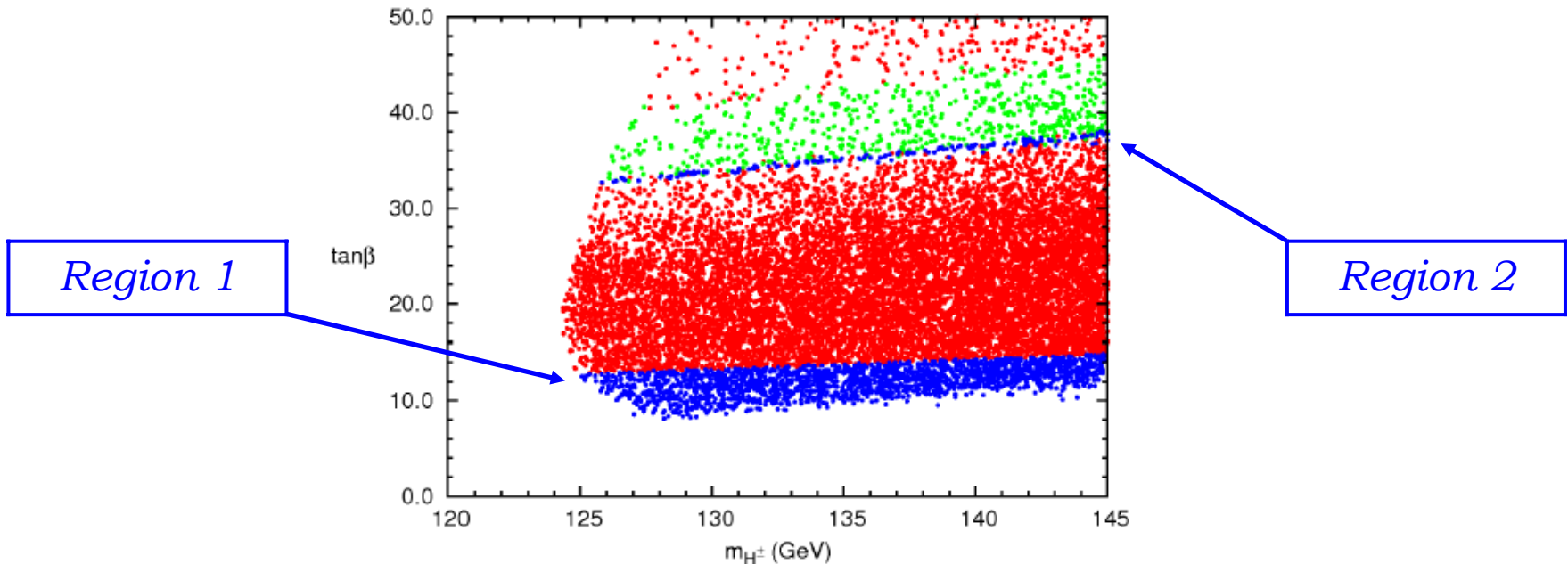
MSSM parameters to scan: (SuSpect, micrOMEGAs, SuperIso)

Djouadi et al. '02  
Belanger et al. '06  
Mahmoudi '08

Low-E values  $\tan\beta$ ,  $M_1$ ,  $M_2$ ,  $M_3$ ,  $a_0$ ,  $\mu$ ,  $m_A$ ,  $m_{\tilde{q}}$ ,  $m_{\tilde{\ell}}$ , ( $A_{u,d} = a_0 Y_{u,d} m_{\tilde{q}}$ )

Ranges:

$$\begin{aligned} M_1 &\in [7, 30] \text{ GeV}, & M_2 &\in [100, 600] \text{ GeV}, & M_3 &\in [400, 1200] \text{ GeV}, \\ m_A &\in [90, 120] \text{ GeV}, & \mu &\in [100, 200] \text{ GeV}, & a_0 &\in [-2, 2], \\ m_{\tilde{q}} &\in [400, 1200] \text{ GeV}, & m_{\tilde{\ell}} &\in [100, 1200] \text{ GeV}, & \tan\beta &\in [5, 50]. \end{aligned}$$



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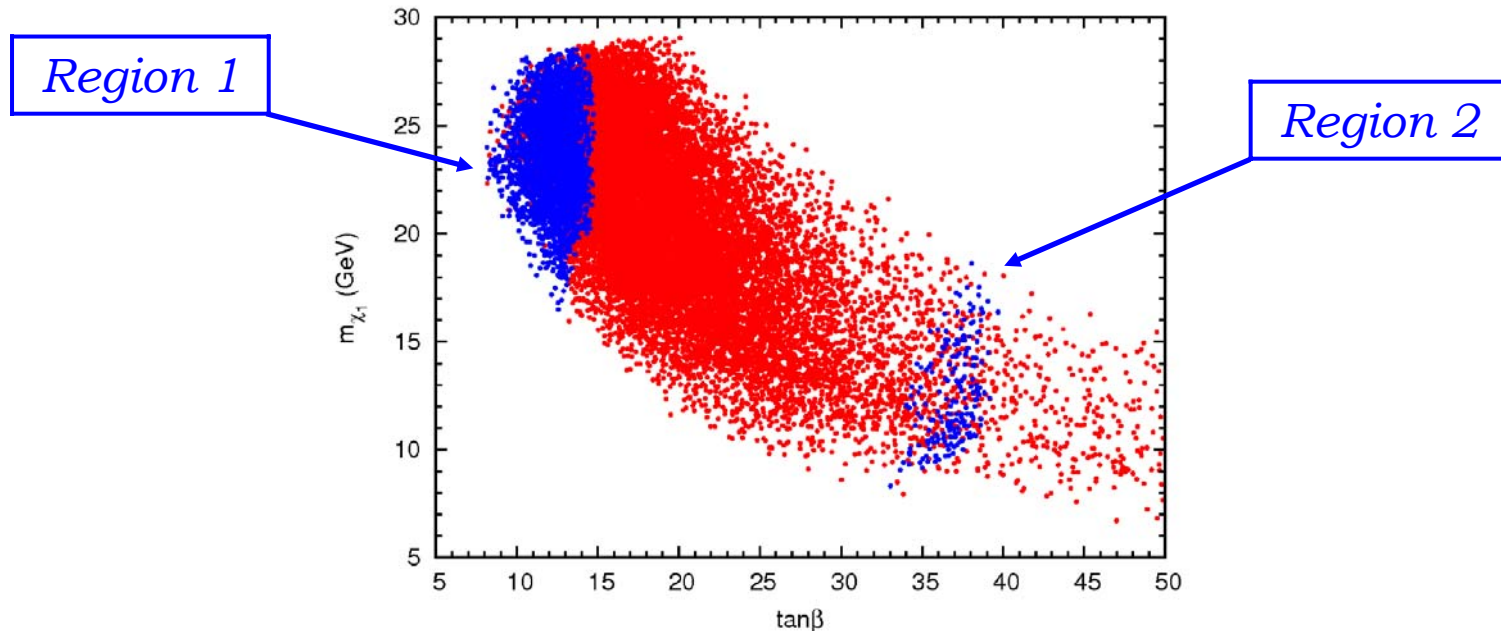
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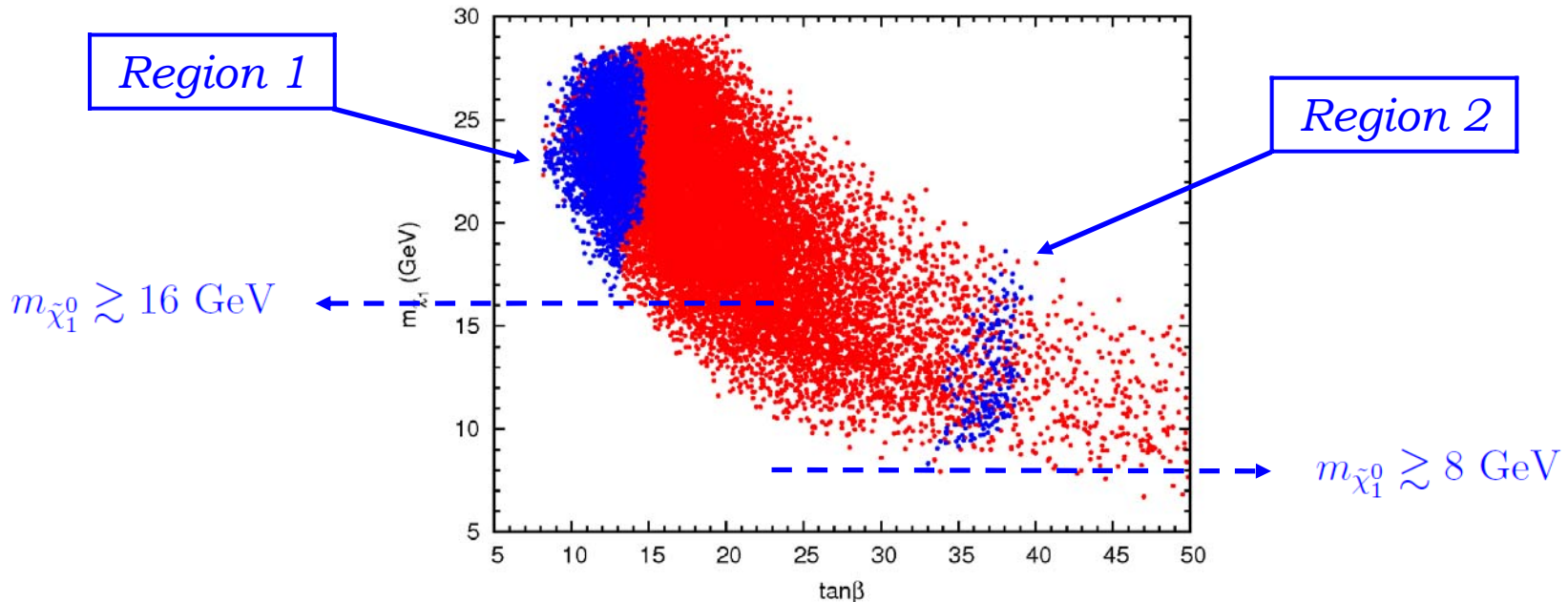
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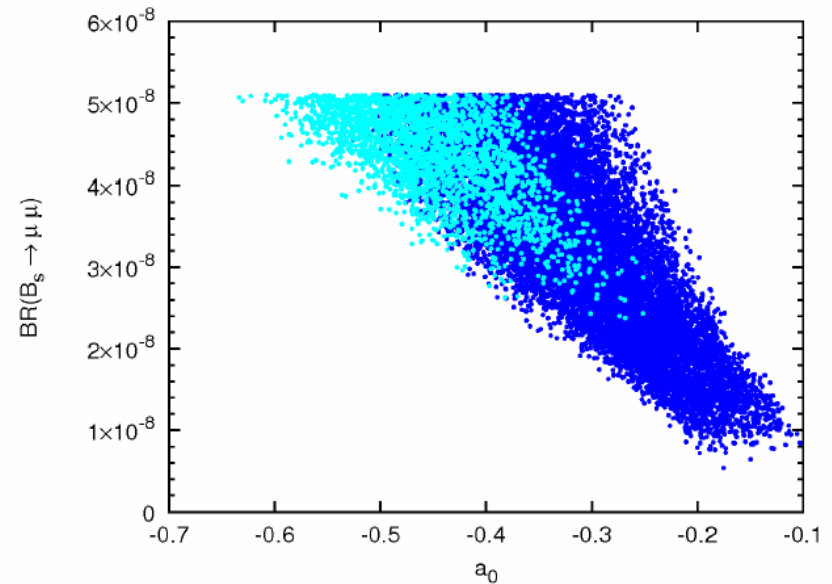
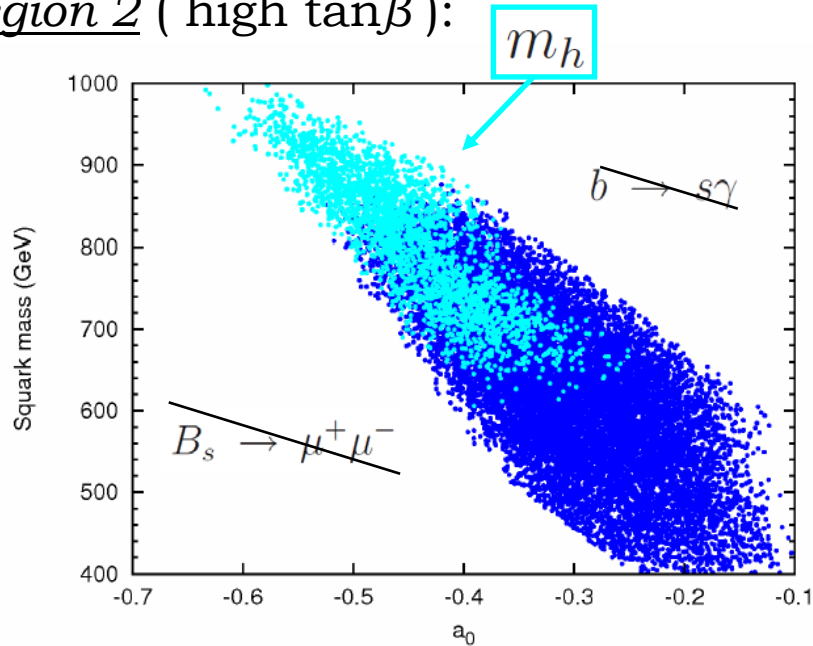


# Phenomenological consequences

## Region 1 ( low $\tan\beta$ ):

- $b \rightarrow s\gamma \implies m_{\tilde{q}} \lesssim 800 \text{ GeV}$  and  $a_0 \lesssim -0.4$
- LEP bound on  $m_h \implies m_{\tilde{q}} \gtrsim 500 \text{ GeV}$  and  $a_0 \lesssim -0.8$
- $\text{BR}(B_s \rightarrow \mu\mu)$  max. twice SM prediction

## Region 2 ( high $\tan\beta$ ):

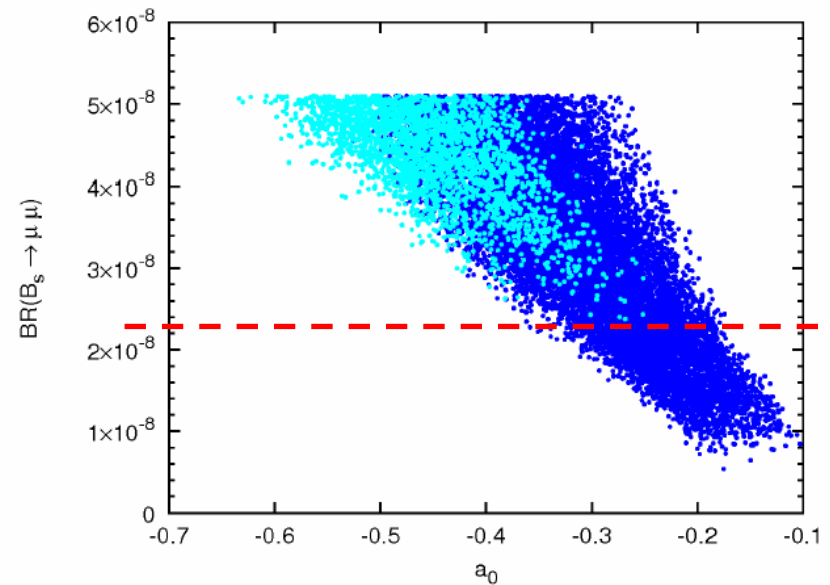
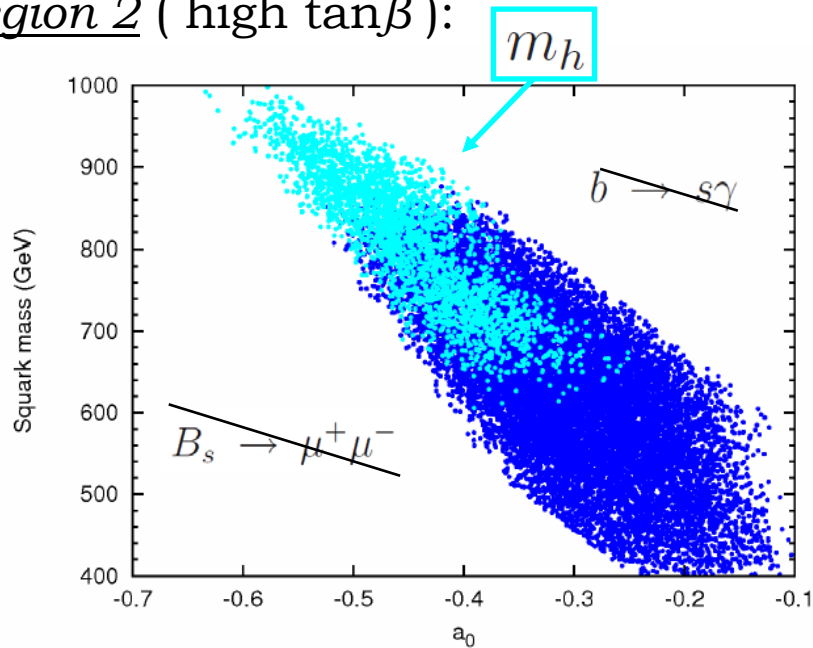


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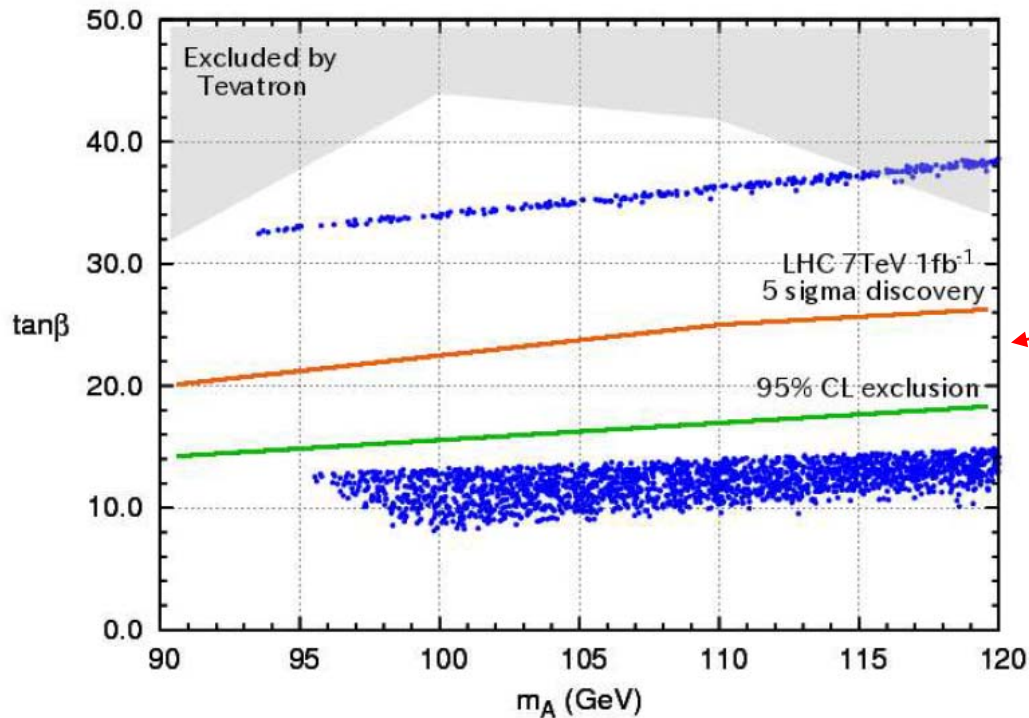


$\implies \text{BR}(B_s \rightarrow \mu\mu) \gtrsim 2 \times 10^{-8}$

Tested within 2011 by LHCb!

# Phenomenological consequences

Higgs searches (e.g.  $pp \rightarrow X \Phi \rightarrow \tau\tau$  at LHC):

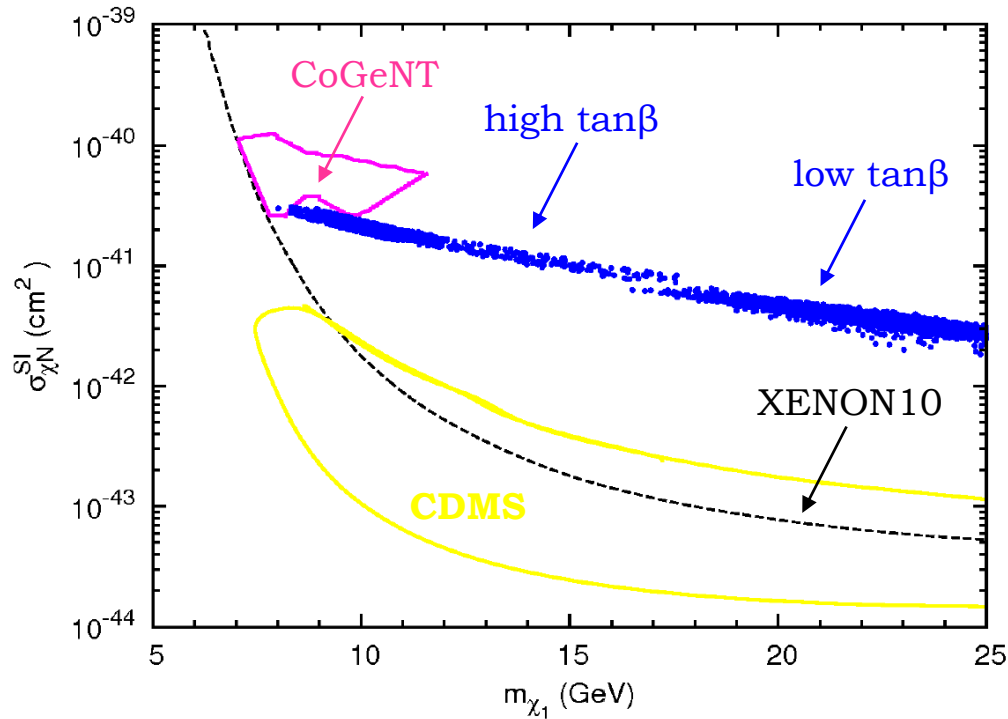


*Region 2 tested within 2011 by LHC Higgs searches!*

Updates: first CMS searches, Region 1 already excluded? [CMS arXiv:1104.1619](#)  
(see however [J. Baglio arXiv: 1105.1085](#) and [Djouadi's talk](#))



# Direct searches



Uncertainties: hadronic matrix elements

e.g. Fornengo et al. '10

local DM density and velocity

e.g. Belanger et al. '10

Updates: CoGeNT reported annual modulation

XENON100 released new exclusion see however Arina et al. '11

# Conclusions

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- WMAP bound combined with constraints from low-energy observables selects two regions of the parameter space:

Low  $\tan\beta$  region with:

$$m_{\tilde{\chi}_1^0} \gtrsim 16 \text{ GeV}$$

$$\sigma_{\chi N}^{\text{SI}} \sim 10^{-42} \div 10^{-41} \text{ cm}^2$$

High  $\tan\beta$  strip with:

$$m_{\tilde{\chi}_1^0} \gtrsim 8 \text{ GeV}$$

$$\sigma_{\chi N}^{\text{SI}} \gtrsim \mathcal{O}(10^{-41}) \text{ cm}^2$$

- High  $\tan\beta$  strip will be tested in the upcoming months by means of LHC Higgs bosons and  $B_s \rightarrow \mu\mu$  searches.
- High  $\tan\beta$  strip can account for CoGeNT, CRESST, DAMA signals. Direct searches are the most promising way of testing the low  $\tan\beta$  region (which indeed seems to be already strongly disfavoured by XENON100).



# Conclusions

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Obrigado!

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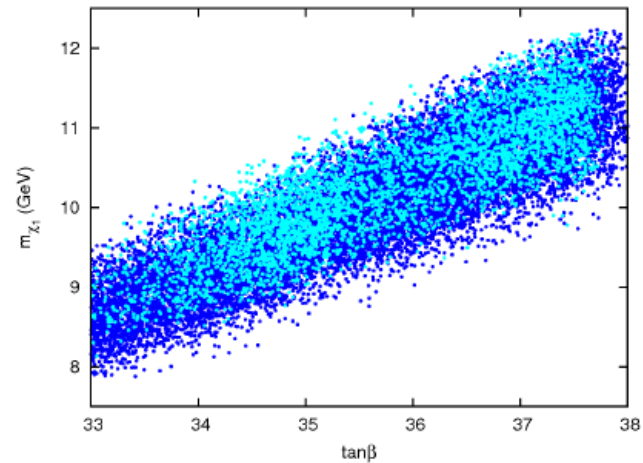
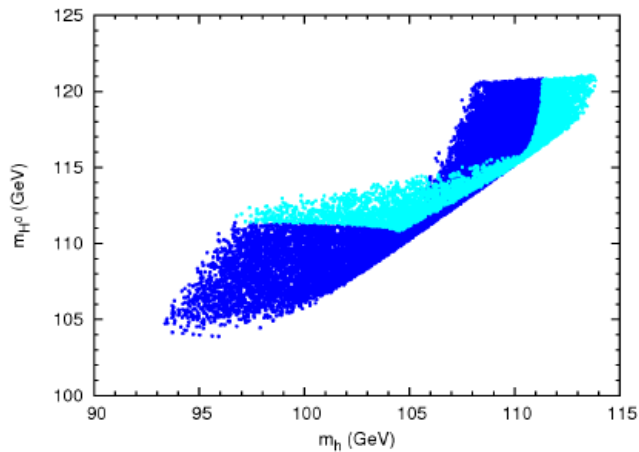
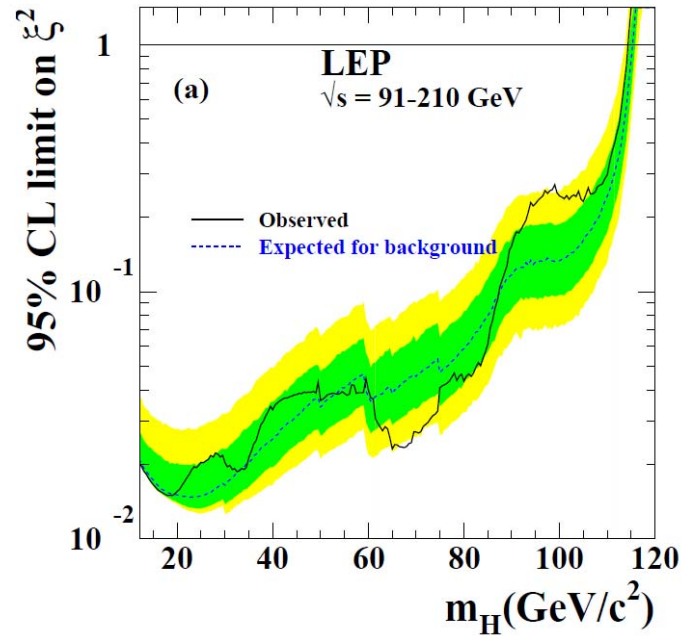
## Additional slides

# Constraints

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	Observable	Allowed range
WMAP	$\Omega_{\text{DM}}h^2$	[0.101, 0.123]
LEP	$m_h$	$> 92.8$ GeV
	$m_A$	$> 93.4$ GeV
	$M_{\tilde{\chi}_1^+}$	$> 94$ GeV
	$\Gamma(Z \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_1^0)$	$< 3$ MeV
	$\sigma(e^+e^- \rightarrow \tilde{\chi}_1^0 \tilde{\chi}_{2,3}^0)$	$< 0.1$ pb
Group (i)	$R_{B\tau\nu}$	[0.52, 2.61]
	$R_{\ell 23}$	[0.985, 1.013]
	$R_{D\ell\nu}$	[0.151, 0.681]
	$\text{BR}(D_s \rightarrow \tau\nu)$	[0.047, 0.061]
Group (ii)	$\text{BR}(b \rightarrow s\gamma)$	$[2.89, 4.21] \times 10^{-4}$
	$\text{BR}(B_s \rightarrow \mu^+\mu^-)$	$< 5.1 \times 10^{-8}$

# LEP Higgs mass constraints



# Low-energy observables within *Region 2*

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