

Sub-wavelength quantum imaging for astronomy

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Motivation: surpassing Rayleigh's criterion

Quantum-limited estimation for LIDARs

Quantum hypothesis testing for exoplanet detection

Experiment: quantum super-resolution imaging and hypothesis testing

Exoplanet atmospheric spectroscopy

The Rayleigh Criterion

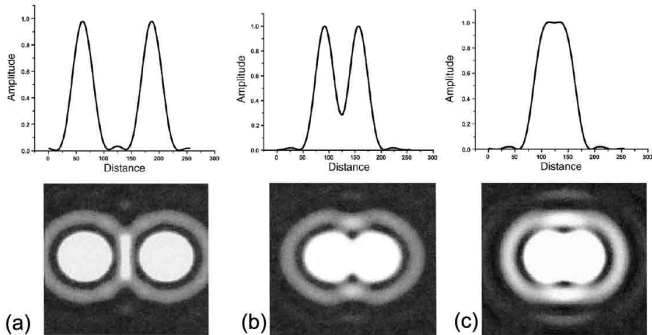


Figure: taken from www.globalsino.com

Minimum resolvable angular separation

$$\theta \approx 1.22 \frac{\lambda}{D}$$

The Rayleigh Criterion

Task: estimate θ_2

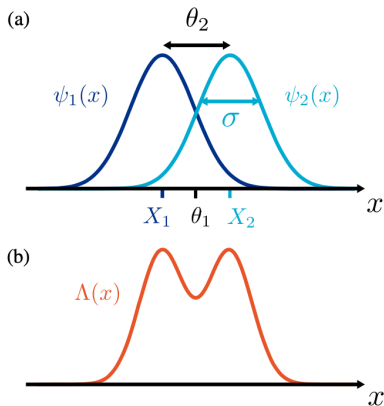


Figure: [1] Tsang et al., Phys. Rev. X 6, 031033 (2016)

Quantum metrology

Cramer-Rao bound:

$$\Delta^2 \varphi = \langle \varphi^2 \rangle - \langle \varphi \rangle^2 \geq \frac{1}{\nu I(\varphi)}$$

Fisher information:

$$I(\varphi) = \sum_i p(i|\varphi) \left(\frac{\partial \log[p(i|\varphi)]}{\partial \varphi} \right)^2$$

Quantum state:

$$\rho_\varphi = \sum_j \lambda_j |j\rangle \langle j|$$

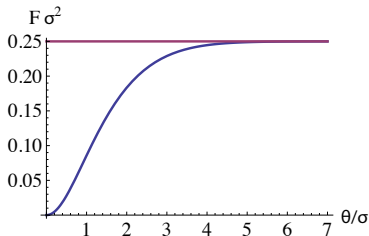
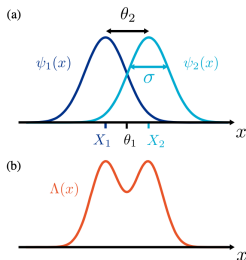
Quantum Fisher information

$$F(\rho_\varphi) = \sum_{\lambda_j + \lambda_k \neq 0} 2 \frac{|\langle j | \frac{\partial \rho}{\partial \varphi} | k \rangle|^2}{\lambda_j + \lambda_k},$$

Superresolution of two sources

Model: two incoherent, quasi-monochromatic point sources

$$\rho \approx \frac{1}{2}(|\psi_1\rangle \langle \psi_1| + |\psi_2\rangle \langle \psi_2|), \quad |\psi_i\rangle = \int_{-\infty}^{\infty} \psi_i(x) |x\rangle$$

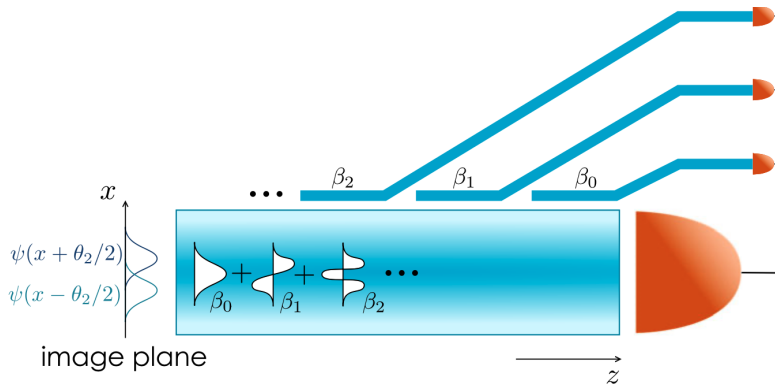


$$F(\theta_2) = \frac{1}{4\sigma^2}$$

[1] Tsang et al., Phys. Rev. X 6, 031033 (2016)

Optimal measurement

Optimal measurement: $FI = QFI$



[1] Tsang et al., Phys. Rev. X 6, 031033 (2016)

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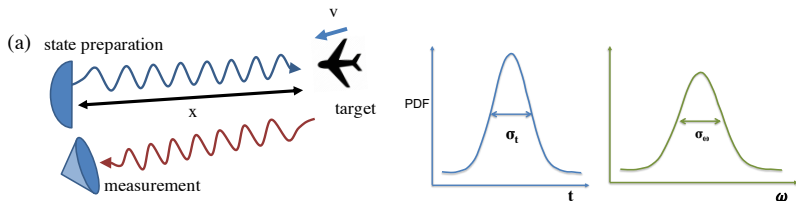
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Position and velocity estimation



Range x and velocity v :

$$\bar{t} = \frac{2x}{c}, \quad \bar{\omega} \approx \omega_0 \left(1 - 2\frac{v}{c}\right)$$

The energy-time uncertainty relation [2]

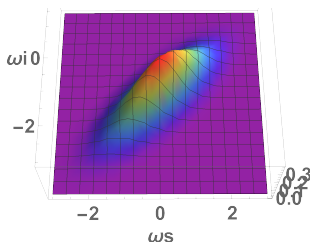
$$\sigma_t \sigma_\omega \geq 1/2.$$

[2] E. Arthurs and J. Kelly, the Bell System Technical Journal 44, 725 (1965)

State description

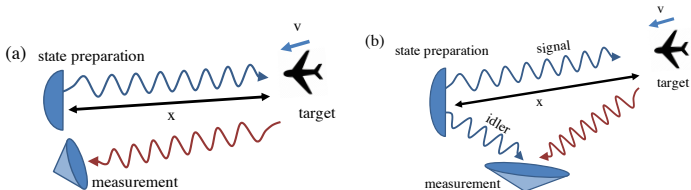
Time-frequency entanglement (SPDC) [3] :

$$|\Psi_0\rangle = \int d\omega \int d\omega_i \tilde{\Psi}_0(\omega, \omega_i, \kappa) |\omega\rangle |\omega_i\rangle, \quad \kappa \in \{0, 1\}$$



[3] MacLean et al., Phys. Rev. Lett. 120, 053601

Position and velocity estimation



$$J(\bar{t}, \bar{\omega}) = \begin{pmatrix} 4\sigma_\omega^2 & 0 \\ 0 & 1/\sigma_\omega^2 \end{pmatrix}$$

$$\sigma_t \sigma_\omega = \frac{1}{2}$$

$$J_{\text{ent}}(\bar{t}, \bar{\omega}) = \begin{pmatrix} 4\sigma_\omega^2 & 0 \\ 0 & \frac{1}{\sigma_\omega^2} \frac{1}{1-\kappa^2} \end{pmatrix}$$

$$\sigma_t \sigma_\omega = \frac{1}{2} \sqrt{1-\kappa^2}$$

Joint optimal estimation condition [5] not satisfied,

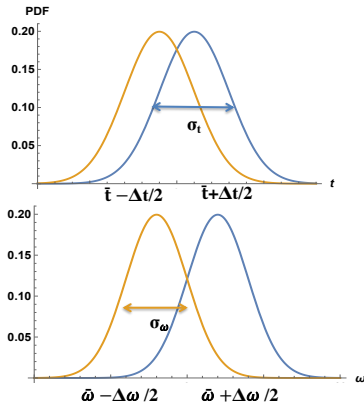
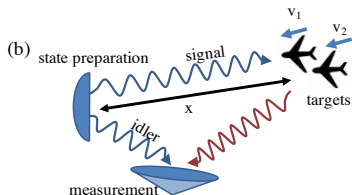
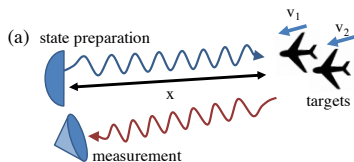
$$\text{Tr}(\rho[L_{\bar{t}}, L_{\bar{\omega}}]) = -4i$$

[5] S. Ragy, et al. Phys. Rev. A 94, 052108 (2016).

[6] Q. Zhuang, et al., Phys. Rev. A 96, 040304 (2017) .

Position and velocity *separation* estimation

$$\rho = \frac{1}{2}(|\psi_1\rangle\langle\psi_1| + |\psi_2\rangle\langle\psi_2|)$$



Quantum metrology for superresolution

PRL 117, 190802 (2016)

PHYSICAL REVIEW LETTERS

week ending
4 NOVEMBER 2016



Ultimate Precision Bound of Quantum and Subwavelength Imaging

Cosmo Lupo¹ and Stefano Pirandola^{1,2}

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²Computer Science, University of York, York YO10 5GH, United Kingdom

(Received 6 July 2016; published 4 November 2016)

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PHYSICAL REVIEW LETTERS **122**, 140505 (2019)

Towards Superresolution Surface Metrology: Quantum Estimation of Angular and Axial Separations

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PHYSICAL REVIEW LETTERS **121**, 180504 (2018)

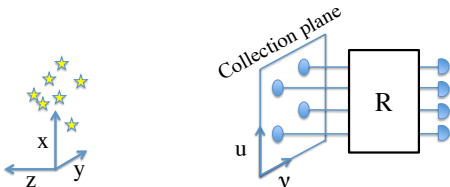
Quantum Limited Superresolution of an Incoherent Source Pair in Three Dimensions

Zhixian Yu and Sudhakar Prasad[†]

Department of Physics and Astronomy, University of New Mexico, Albuquerque, New Mexico 87131, USA

(Received 26 May 2018; published 31 October 2018)

Arbitrary number of sources



N_s incoherent sources, quasi-monochromatic, coordinate \vec{r}_s .
 N_c collectors, position $\vec{\omega}_j$. Photon impinging $\rightarrow |j\rangle$

$$|\psi(r_s)\rangle = \frac{1}{\sqrt{N_c}} \sum_j^{N_c} e^{i\phi(\vec{r}_s, \vec{\omega}_j)} |j\rangle,$$

$$\rho = \sum_s^{N_s} p(s) |\psi(r_s)\rangle \langle \psi(r_s)|$$

C Lupo, Z Huang, P Kok, Phys. Rev. Lett. 124, 080503 (2020)

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1 or 2 sources?

Two hypotheses H_0, H_1

Classical: $p_0(x), p_1(x)$, quantum: ρ_0, ρ_1

(a) H_0



(b) H_1

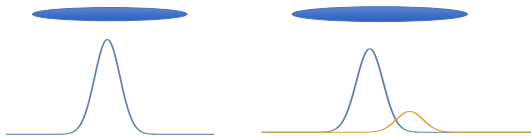
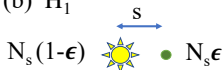


Figure: (a) H_0 : there is only 1 source. (b) H_1 : two near-by sources present.

Huang & Lupo, Phys. Rev. Lett. 127, 130502 (2021);
Editor's Selection and Featured in Physics

Quantum state discrimination

1. Symmetric discrimination: trace distance - quantum Chernoff bound [6], $P_e \sim \exp[-n f(T)]$

$$T_c(\rho_0, \rho_1) = 1/2 \int dx |\rho_0(x) - \rho_1(x)|$$

$$T_Q(\rho_0, \rho_1) = 1/2 \|\rho_0 - \rho_1\|_1$$

2. Asymmetric: relative entropy - quantum Stein lemma [7],

$$P_e \sim \exp[-nD(\rho_0||\rho_0) + O(\alpha^{-1}, \ln n)]$$

$$D_c(\rho_0||\rho_1) = \int dx \rho_0(x)(\ln \rho_0(x) - \ln \rho_1(x))$$

$$D_Q(\rho_0||\rho_1) = \text{Tr}[\rho_0(\ln \rho_1 - \ln \rho_0)]$$

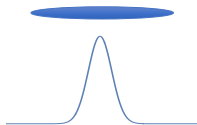
[6] Audenaert et al., Phys. Rev. Lett. 98, 160501 (2007)

[7] F. Hiai, D. Petz, Commun. Math. Phys. 143, 99 (1991)



Density matrices

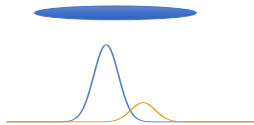
(a) H_0

N_s 



(b) H_1

$N_s(1-\epsilon)$  \longleftrightarrow $N_s\epsilon$ 



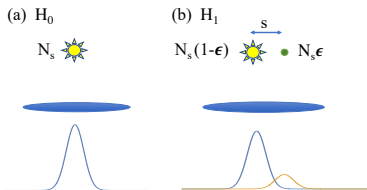
Two hypotheses associated with the density matrices

$$\rho_0 = |\psi_{x_0}\rangle \langle \psi_{x_0}| ,$$

$$\rho_1 = (1 - \epsilon) |\psi_{x_0}\rangle \langle \psi_{x_0}| + \epsilon |\psi_{x_0+s}\rangle \langle \psi_{x_0+s}| .$$

Huang & Lupo, Phys. Rev. Lett. 127, 130502

Classical probability distributions



On-screen probability distributions:

$$p_0(x) = |\psi(x - x_0)|^2,$$

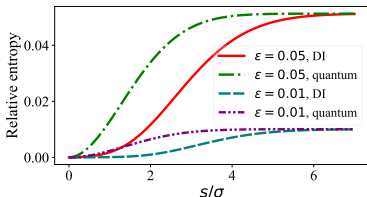
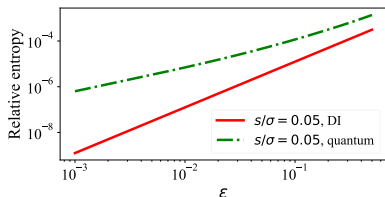
$$p_1(x) = (1 - \epsilon)|\psi(x - x_0)|^2 + \epsilon|\psi(x - (x_0 + s))|^2.$$

Classical relative entropy:

$$D(p_0||p_1) \approx \frac{s^2\epsilon^2}{2\sigma^2} + O(\epsilon^3)$$

Huang & Lupo, Phys. Rev. Lett. 127, 130502

Classical vs Quantum



Classical relative entropy:

$$D(\rho_0 || \rho_1) \approx \frac{s^2 \epsilon^2}{2\sigma^2}$$

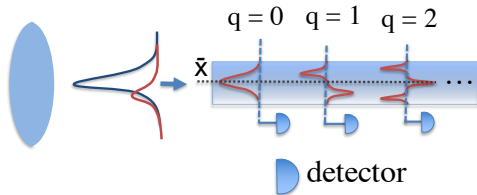
Quantum relative entropy;

$$D(\rho_a || \rho_b) \approx \frac{s^2 \epsilon}{4\sigma^2} + O(\epsilon^2 s^2).$$

Huang & Lupo, Phys. Rev. Lett. 127, 130502

Optimal measurement - SPADE

Near-optimal measurement: SPADE



For our case, we point the optical imaging system towards the optical “center of mass”.

$$\bar{x} = (1 - \epsilon)x_0 + \epsilon(x_0 + s). \quad (1)$$

Huang & Lupo, Phys. Rev. Lett. 127, 130502 (2021)

Motivation: surpassing Rayleigh's criterion

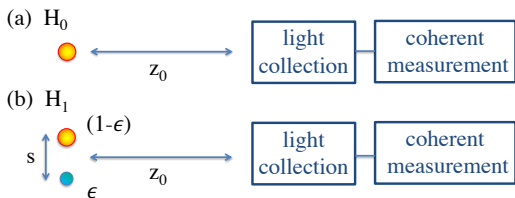
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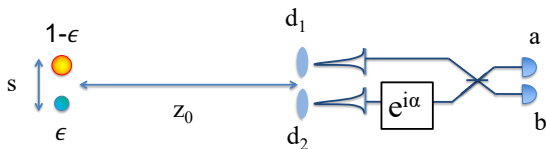
The experiment



- ▶ Sub-Rayleigh quantum state discrimination
- ▶ Optimal angular separation estimation (equally bright sources)
- ▶ Two tasks achieved with *a single measurement*
- ▶ Binary stars, microscopy

Experimental scheme

Instead of a lens, place two collectors at d_1 , d_2 :



$$|\psi_{\text{star}}\rangle = \frac{1}{\sqrt{2}}(|d_1\rangle + e^{i\phi} |d_2\rangle), \quad |\psi_{\text{planet}}\rangle = \frac{1}{\sqrt{2}}(|d_1\rangle + e^{-i\phi} |d_2\rangle)$$

In the limit of $\epsilon \ll 1$, $\theta = s/z_0$ (angular)

$$D(\rho_0 || \rho_1) \approx \frac{\epsilon \theta^2 k^2 d^2}{4}$$

Zanforlin, et al., Nat. Commun. 13, 5373 (2022)

Simplifying the experiment

H_0 : the probabilities of photon arriving at detectors a and b are

$$p_{H_0}(a) = \frac{1}{2}(1 + \nu \cos(\phi + \alpha))$$

$$p_{H_0}(b) = \frac{1}{2}(1 - \nu \cos(\phi + \alpha))$$

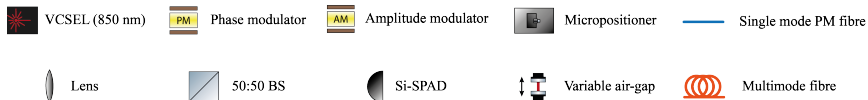
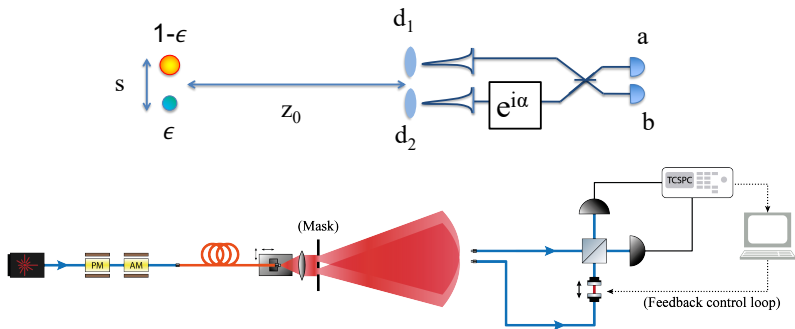
H_1 :

$$p_{H_1}(a) = \frac{1}{2}[(1 - \epsilon)(1 + \nu \cos(\phi + \alpha)) + \epsilon(1 + \nu \cos(-\phi + \alpha))]$$

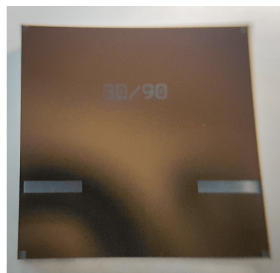
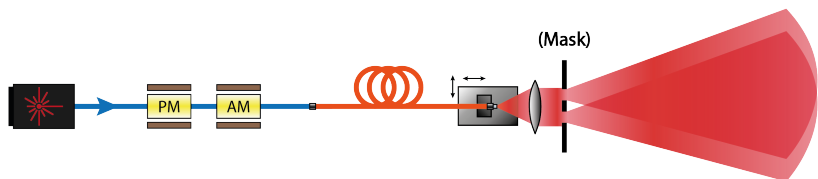
$$p_{H_1}(b) = \frac{1}{2}[(1 - \epsilon)(1 - \nu \cos(\phi + \alpha)) + \epsilon(1 - \nu \cos(-\phi + \alpha))]$$

$$D_c(p_0 || p_1) = \sum_x p_0(x) (\log p_0(x) - \log p_1(x)), \quad x = a, b$$

The experimental setup



The experimental setup



Zanforlin, et al., Nat. Commun. 13, 5373 (2022)

Results - relative entropy

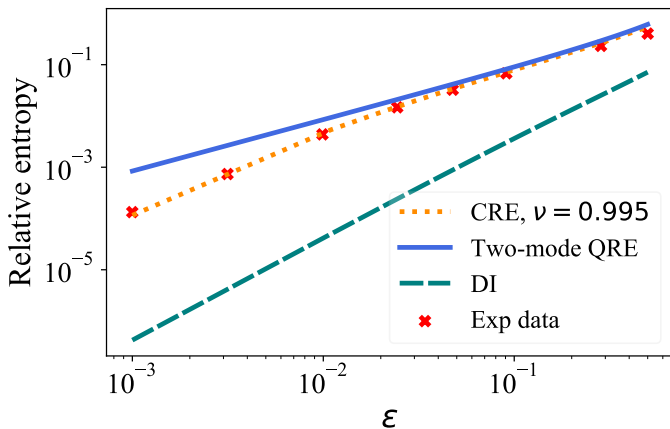


Figure: Relative entropy of the two hypotheses

Angular separation measurement

When the source intensities are equal, the QFI for the angular separation is

$$\text{QFI} = \frac{k^2}{4}(d_1 - d_2)^2$$

The probabilities of detecting the photon at either detector are

$$p_1 = \frac{1}{2}(1 + \nu \cos(\alpha) \cos[\phi]), \quad \phi = kd\theta/2$$
$$p_2 = \frac{1}{2}(1 - \nu \cos(\alpha) \cos[\phi])$$

We use Baye's theorem to update the probability distribution.

Sample distribution

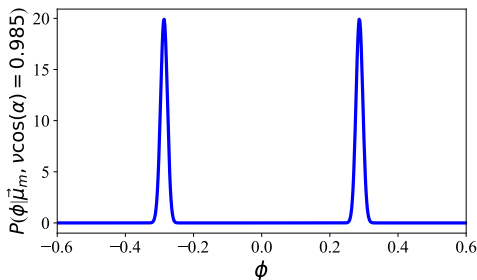
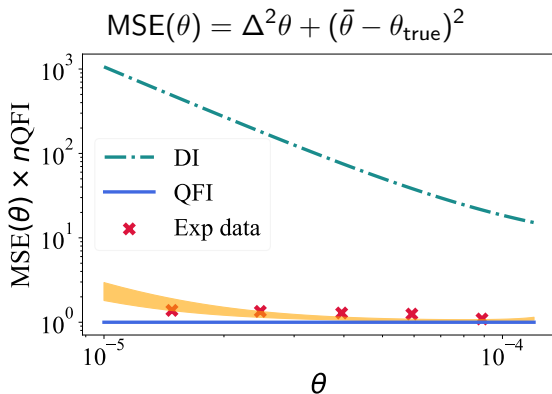


Figure: Sample distribution: after 10,000 detection events, 204 arrived at detector b .

$$\nu \cos(\alpha) \in [-1, 1]$$

Estimator:

$$\hat{\theta}_{\text{est}} = 2|\hat{\phi}_{\text{est}}|/(kd) \quad (2)$$



Two sources 15 μ rad apart, we resolve θ to 1.7% accuracy.
Factor 2 within QCR bound.

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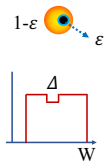
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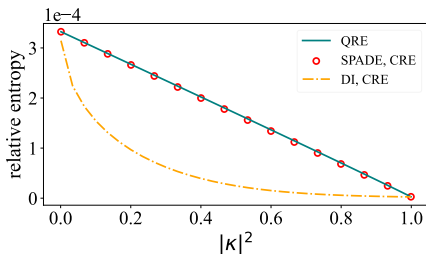
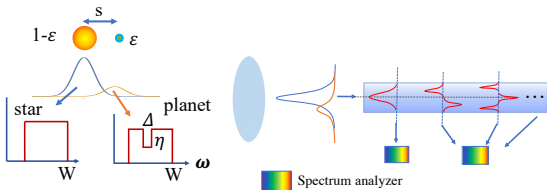
Exoplanet atmospheric spectroscopy

Application to spectroscopy

(a) H_1 : transit



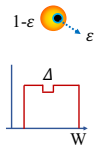
(b) H_1 : spatially separated



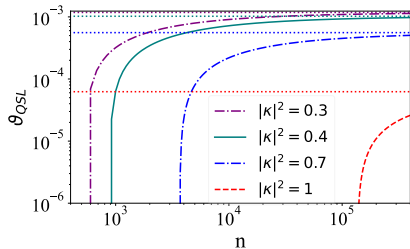
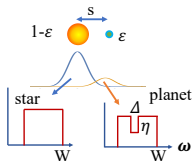
Huang, Schwab, Lupo, arXiv:2211.06050 (2022)

Application to spectroscopy

(a) H_1 : transit



(b) H_1 : spatially separated



$$\vartheta_{\text{QSL}} := D(\rho_0 \parallel \rho_1) + \sqrt{\frac{V(\rho_0 \parallel \rho_1)}{n}} \Phi^{-1}(\alpha) + O\left(\frac{\log n}{n}\right)$$

Huang, Schwab, Lupu, arXiv:2211.06050 (2022)

Conclusions

- ▶ We compute the type-II error probability exponent of discriminating between 1 or two sources with arbitrary intensity.
- ▶ in the limit that $\epsilon \ll 1$, the quantum relative entropy is larger than that of direct imaging by a factor of $1/\epsilon$.
- ▶ For a lens: two measurement methods that are optimal in this regime
- ▶ We significantly simplified the scheme to a two-mode interferometer, demonstrates sub-Rayleigh scaling and approach the QCRB.

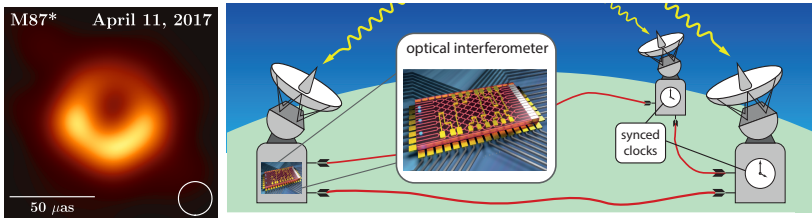
Huang & Lupo, Phys. Rev. Lett. 127, 130502

Zanforlin, et al., Nat. Commun. 13, 5373 (2022)

To the future

£360K (\approx AU \$650K) grant from the EPSRC
DECRA Fellowship (2023) from the ARC

This work is funded by the Sydney Quantum Academy



- ▶ Long-distance optical coherence, entanglement-assisted network
- ▶ Quantum error correction to combat to loss to decoherence
- ▶ Current collaborations: Bristol, Heriot-Watt, Erlangen, UWA

PhD scholarship available

Questions?

Thank you for your attention.



Figure: (Left) my hamster in the UK; (right) my jenday conure

PhD scholarship available