The IceCube real-time analysis framework

Marcos Santander University of Alabama

Town Hall KM3NeT meeting

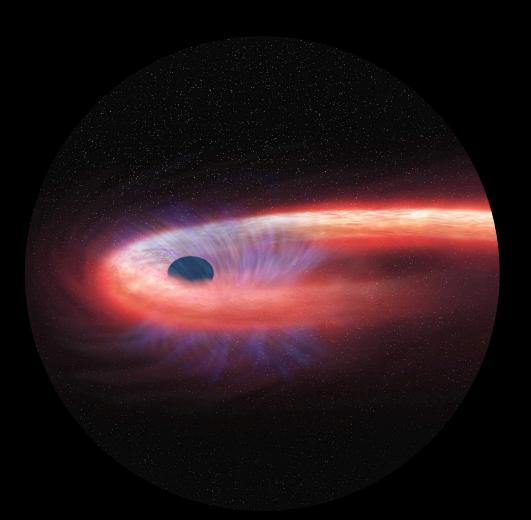




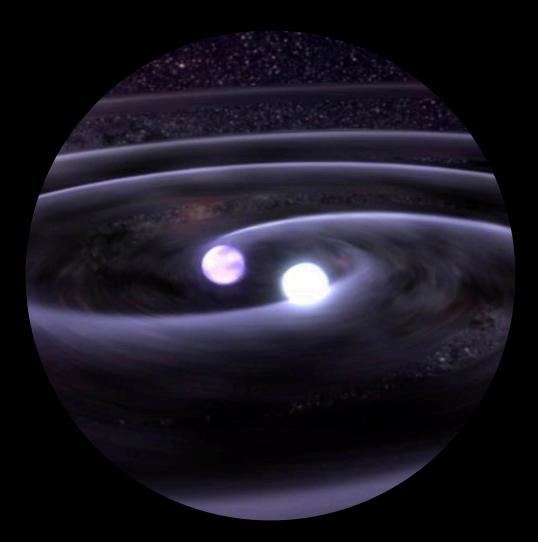


Neutrinos from transient astrophysical sources

Active galaxies

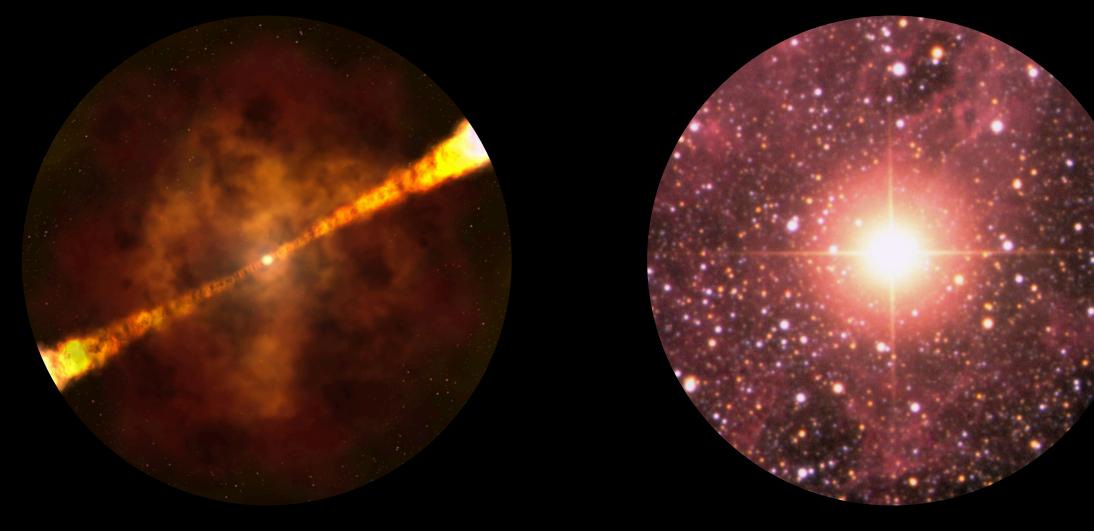


Tidal disruption events



Compact object mergers

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Gamma-ray bursts

Core-collapse supernovae

• Transient and highly-variable persistent astrophysical sources display high-energy non-thermal emission potentially from hadronic processes.

 Searching for neutrinos correlated with known MM signals provides direct insights into particle acceleration processes in these sources.





IceTop

The IceCube Neutrino Observatory

DOM

DeepCore

Inice

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- First km³-scale neutrino detector
- 4π sensitivity
- High uptime (>99%)
- Realtime program:
 - HE neutrino alerts
 - Follow-up of astrophysical events. Realtime correlations.

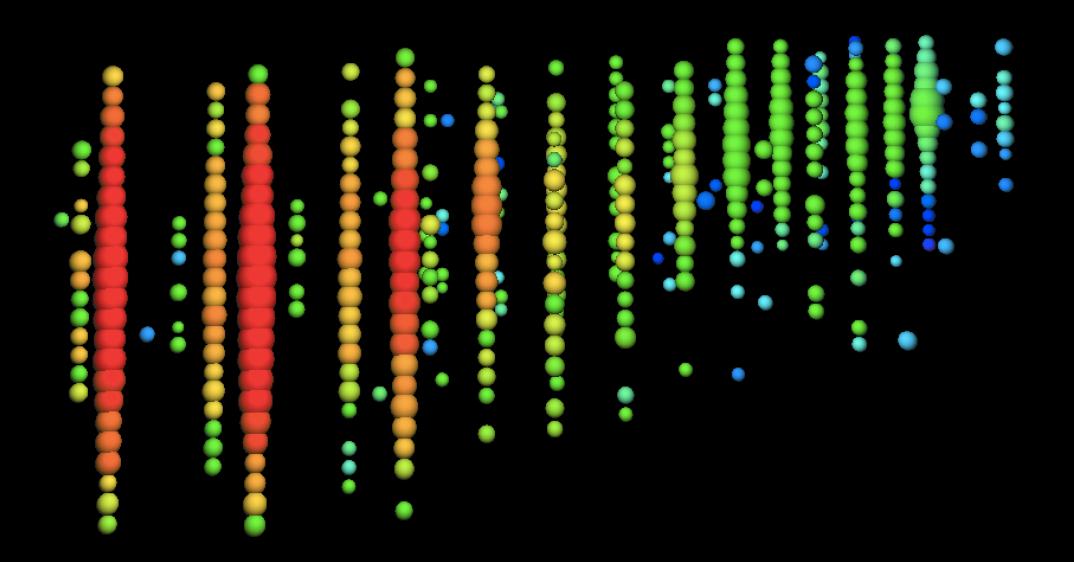
[Astropart. Phys., 92, 30, 2017]





Event topologies Muon tracks

 $CC \nu_{\mu}$ interactions



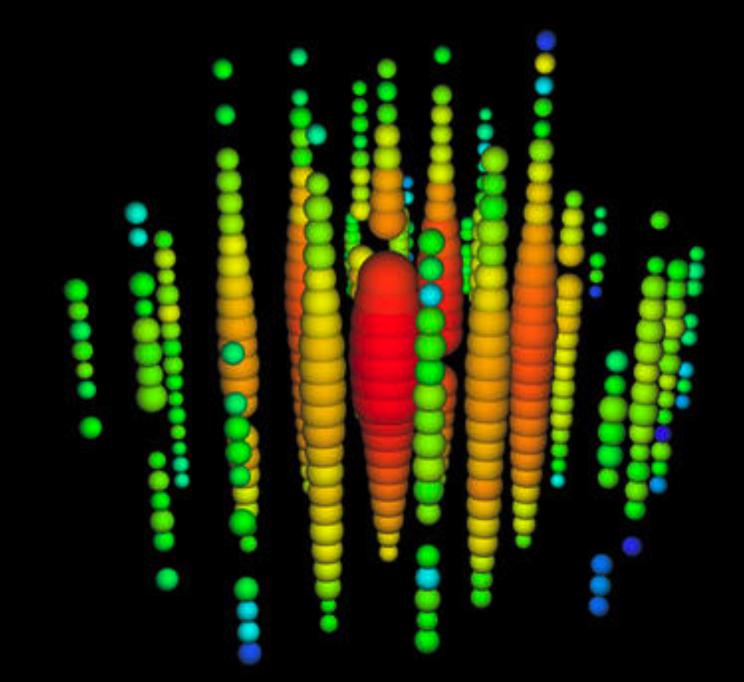
Factor of ~2 energy resolution on neutrino energy (~15% on deposited visible muon energy) Angular resolution **0.5° @ 10 TeV**, 0.3° @ 100TeV

Angular resolution better suited for correlating with pointing instruments



Cascades

NC / CC v_e , most v_T



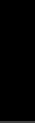
15% deposited energy resolution ~15° median angular resolution @ 10 TeV (8° @ 100 TeV)

Added statistics at high energies



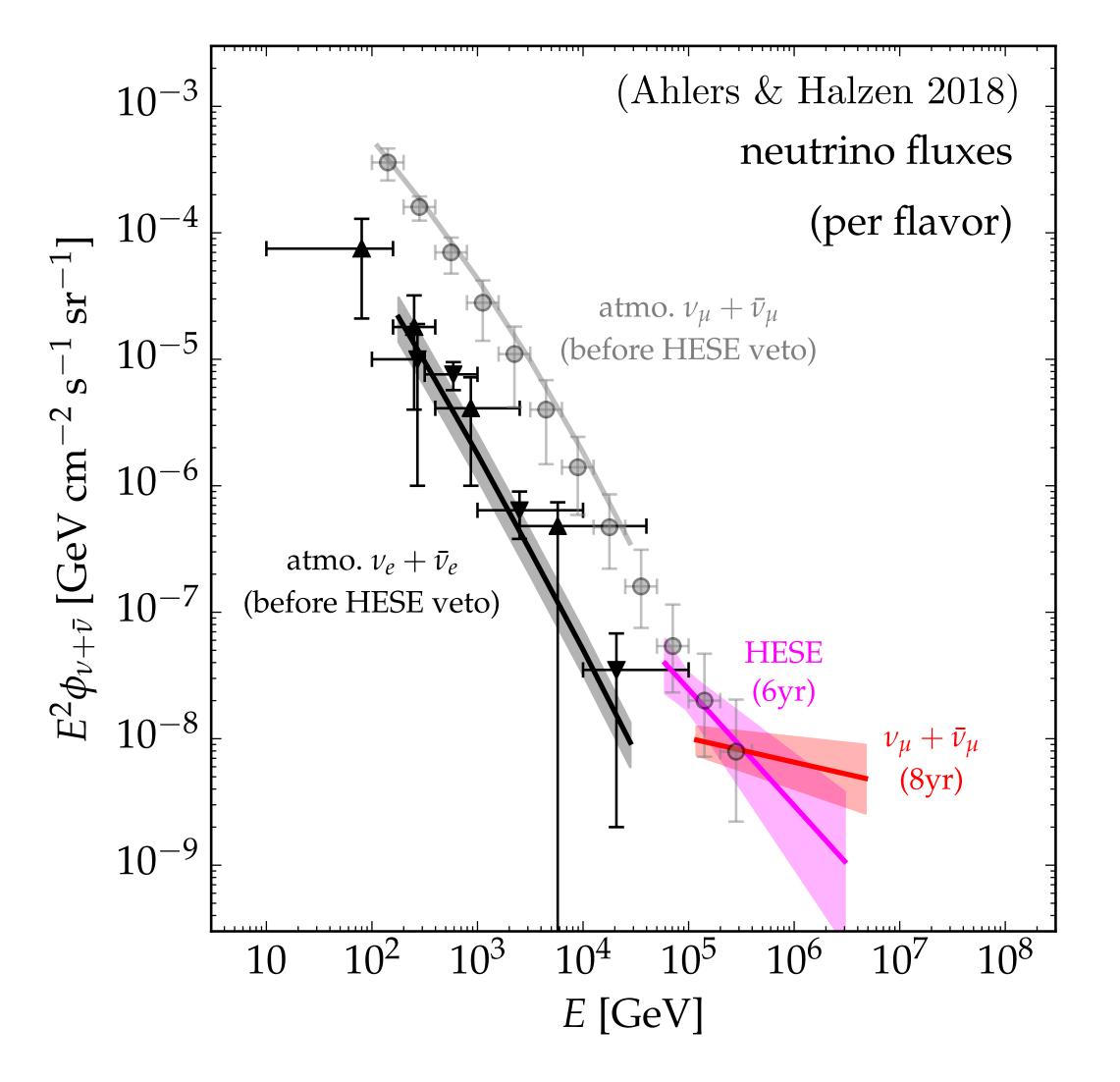




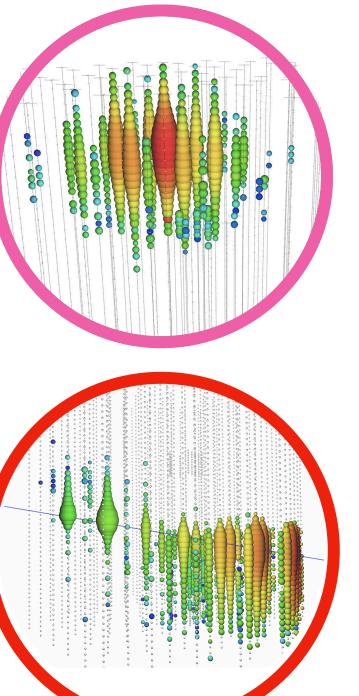


Astrophysical neutrino flux









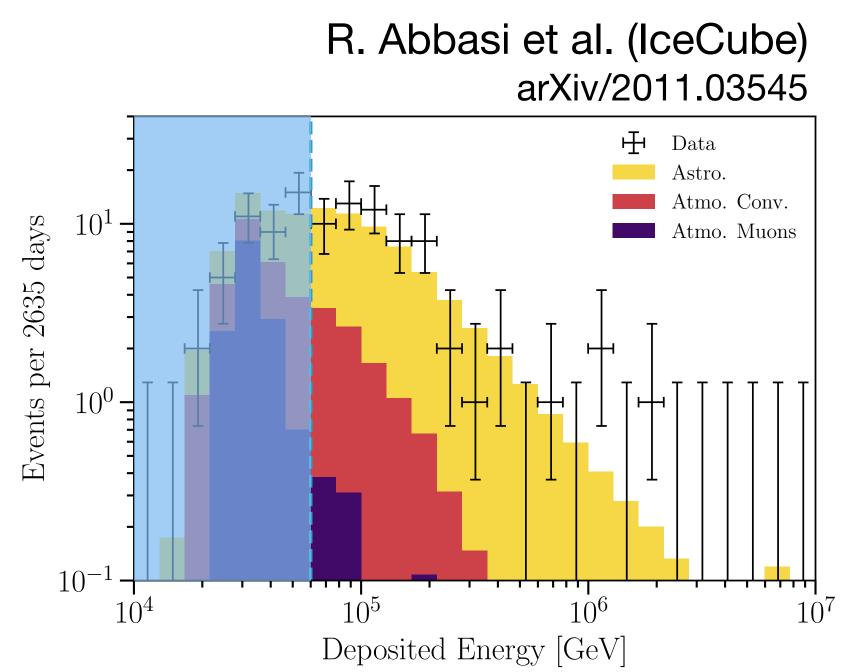
Up-going tracks

Northern sky

Muon-dominated

High-energy starting events (HESE)

Interaction vertex within the detector All flavor, all sky

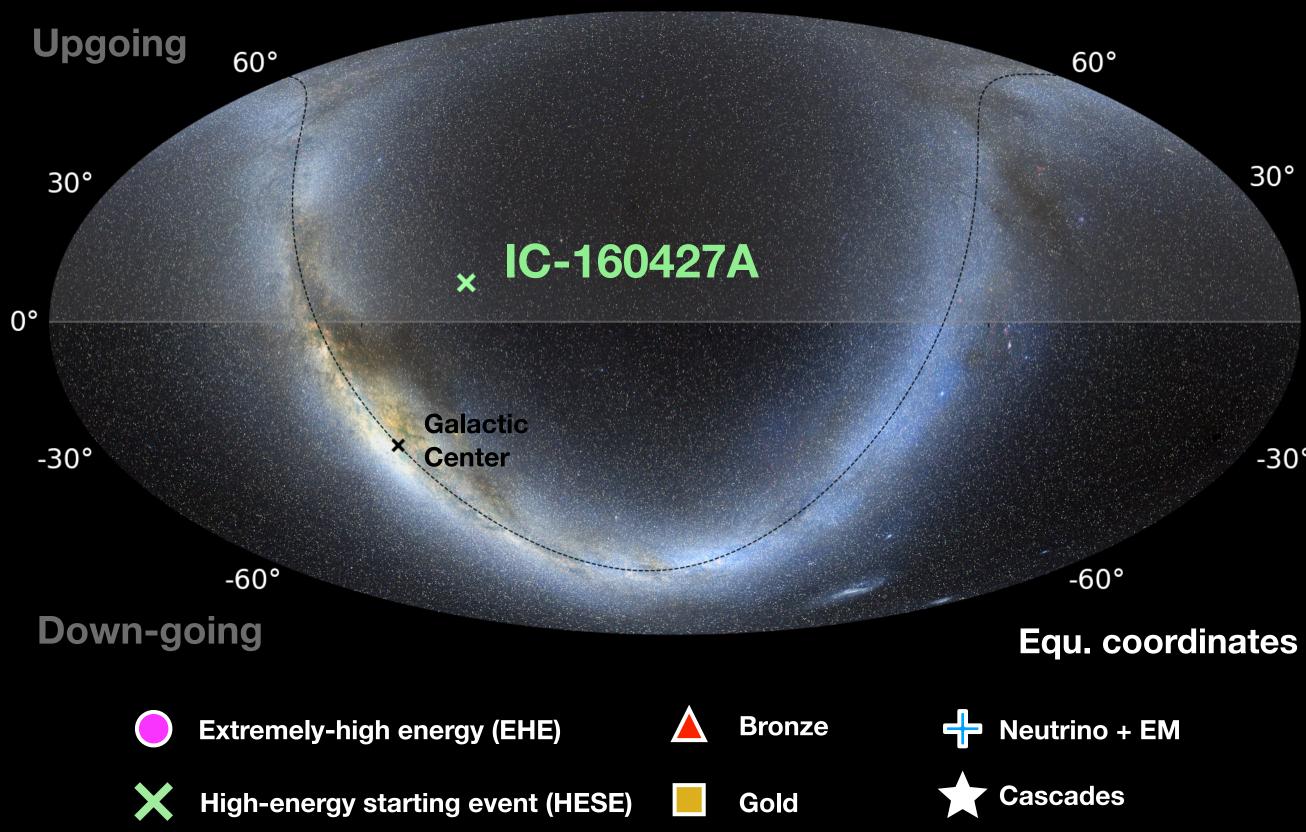


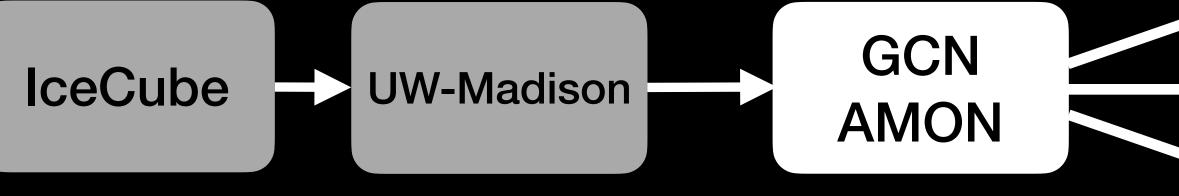
- Astrophysical flux in the 20 TeV 9PeV range, dominant over atmospheric background at \geq 100 TeV
- Detected in multiple analyses.











Median alert latency: 33 seconds

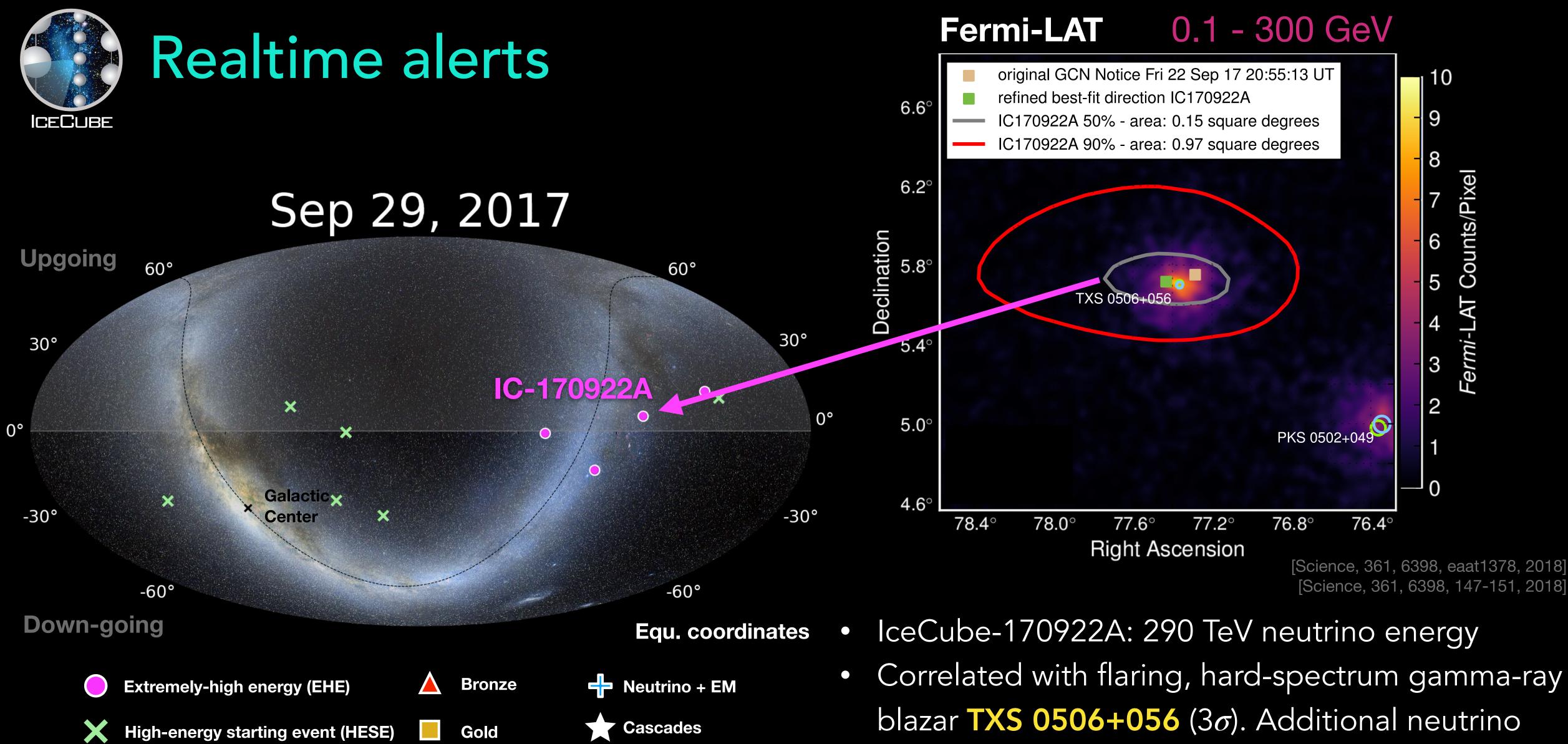
https://gcn.gsfc.nasa.gov/amon.html

- 30° 0° -30°

- Original alert streams (2016-2019)
 - **HESE**: HE starting muon events. Median ang. resolution ~1.5°. 3-4 / year. ~25% astrophysical fraction.
 - EHE: HE through-going muons. Median ulletangular resolution $< 0.5^{\circ}$. 4-6 / year. ~50% astrophysical fraction.
- First alert on April 2016.
- Alerts issued via GCN: 18 HESE, 9 EHE.

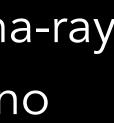


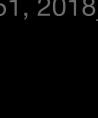




IceCube-170922A: 290 TeV neutrino energy

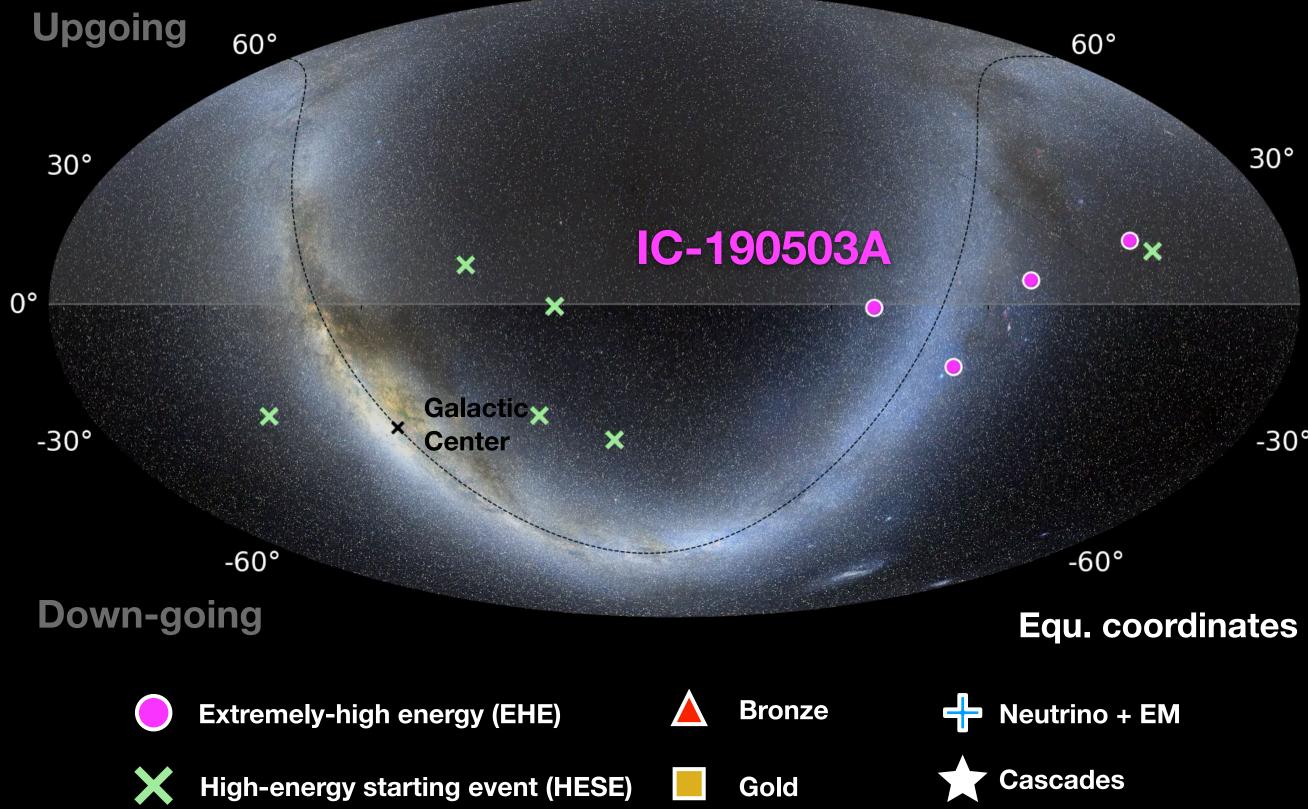
- Correlated with flaring, hard-spectrum gamma-ray blazar TXS 0506+056 (3 σ). Additional neutrino emission in 2014-2015.
- Detected in VHE gamma-rays.



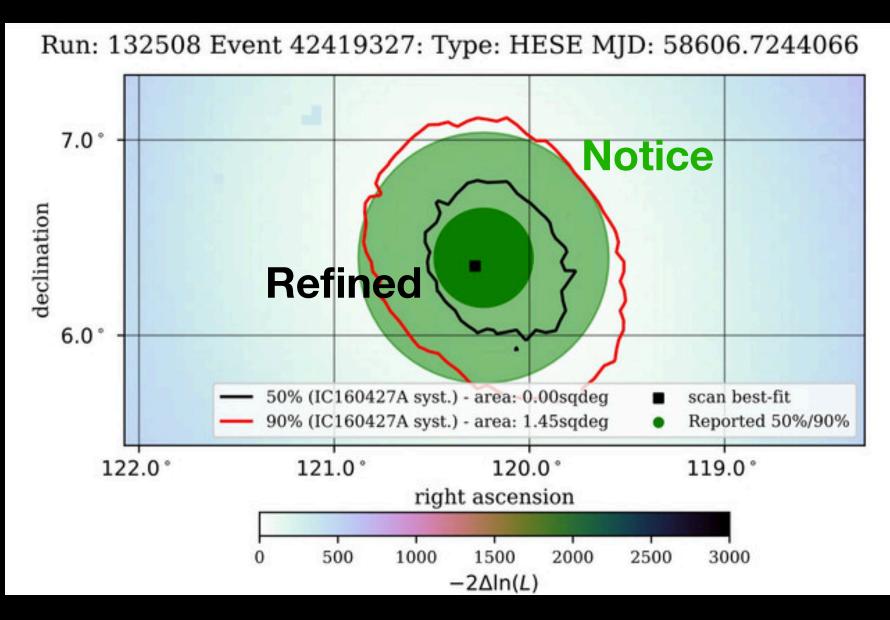




Sep 25, 2017



Initial GCN Notice followed by GCN circular with refined ightarrowposition and error estimates (within couple of hours)

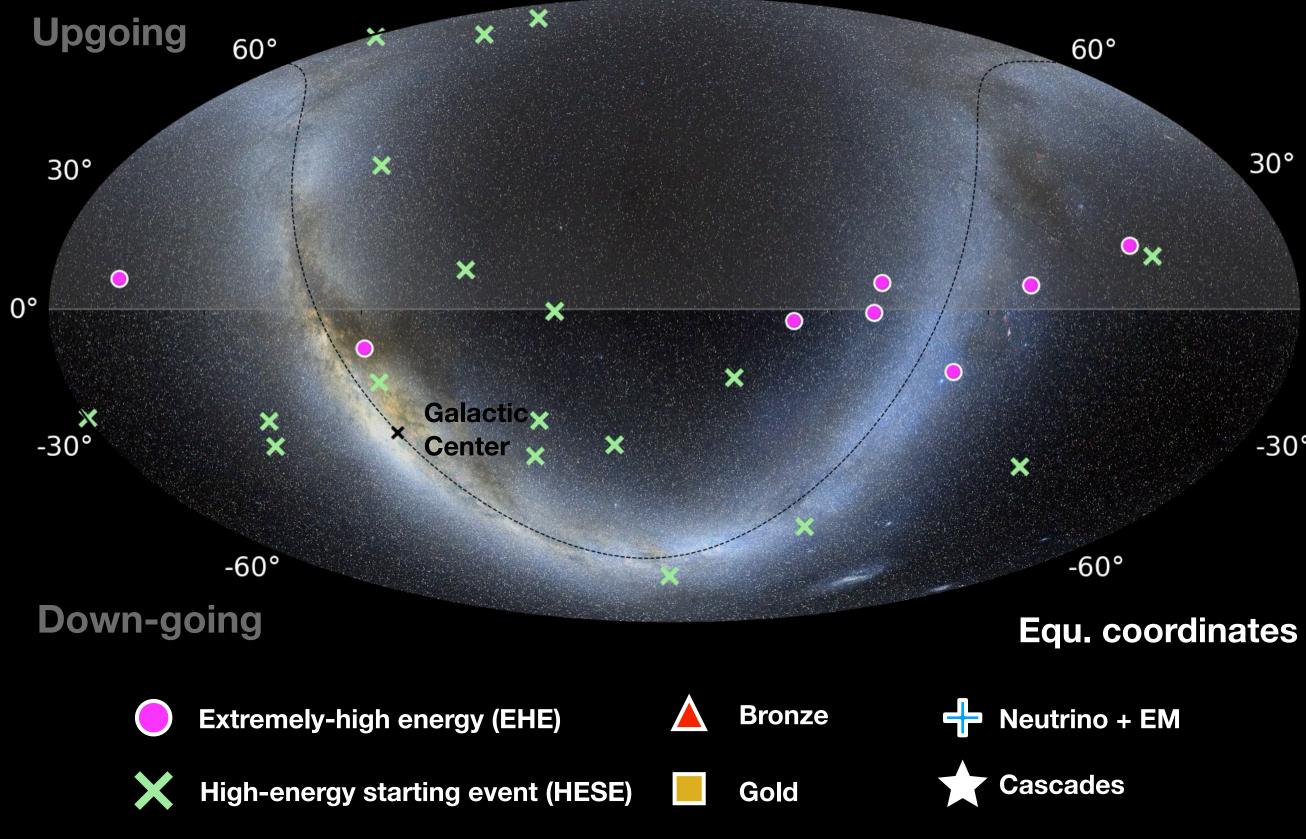


- 30° 0° -30°
 - Typically followed up by multiple multimessenger/ ulletmultiwavelength facilities.
- Example: IC-190503A event
 - ~145 TeV EHE event
 - Follows up by ZTF, ASAS-SN, Kanata, INTEGRAL, IceCube, Fermi-GBM, ANTARES, Fermi-LAT, Lick/ KAIT, Swift-XRT, Insight-HXMT (9 GCNs, 3 ATels)





Jun 15, 2019



[PoS-ICRC2019-1021]

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Unified track alert streams

- Signalness = N_{signal} / (N_{signal} + N_{background})
- Improved selection based on signalness combines through-going and starting tracks.
- Doubled effective area at 0.1 1 PeV

30°

0°

-30°

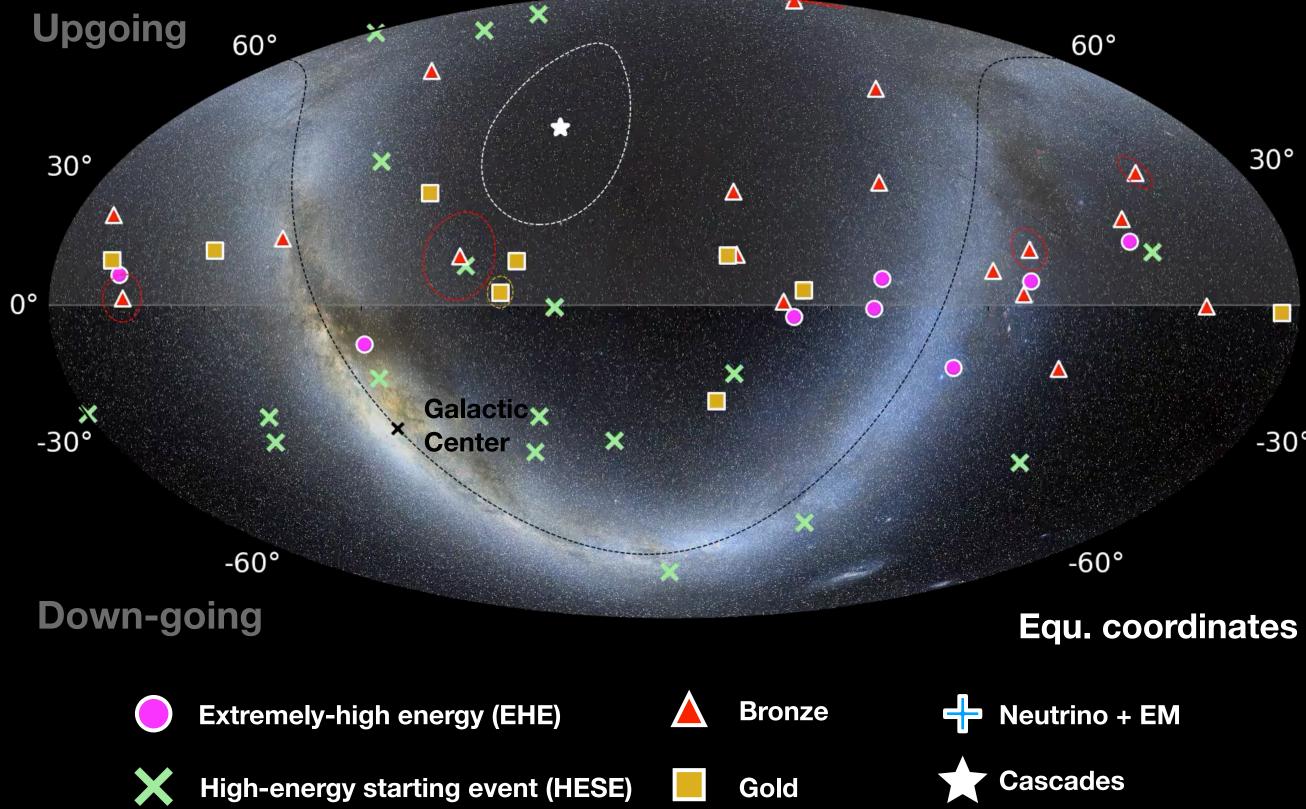
- Gold stream: ~50% signalness (16 issued)
- Bronze stream: ~30% signalness (26 issued)

	Gold Events	Bronze Events
Signal $(E^{-2.19})$	6.6 (Total)	2.8 (Total)
	5.1 (GFU)	2.5 (GFU)
	0.5 (HESE)	0.3 (HESE)
	2.1 (EHE)	
Atmospheric Backgrounds	6.1 (Total)	14.7 (Total)
	4.7 (GFU)	13.8 (GFU)
	0.4 (HESE)	0.9 (HESE)
	1.9 (EHE)	
Observed historical rate	9.9 (Total)	19.5 (Total)
	7.8 (GFU)	18.4 (GFU)
	1.1 (HESE)	0.9 (HESE)
	4.3 (EHE)	

As of Dec 2nd, 2020: 16 gold and 26 bronze alerts issued https://gcn.gsfc.nasa.gov/amon_icecube_gold_bronze_events.html



Jul 03, 2020



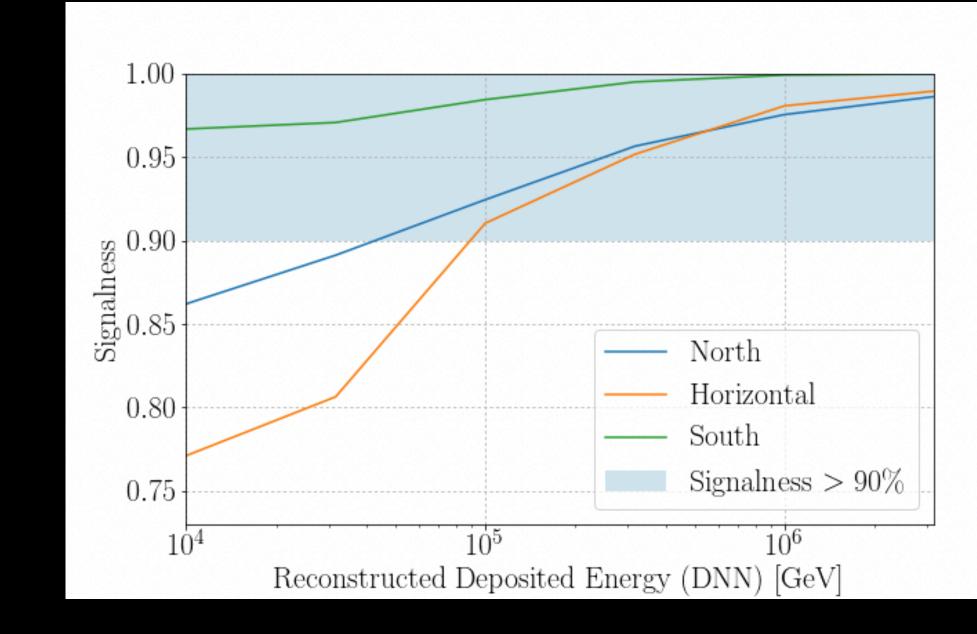
Cascade alerts

30°

-30°

0°

- HESE events are selected using a deep neural network (DNN) classifier.
- 50% have an uncertainty $< 7^{\circ}$, 68% is $< 9^{\circ}$.
- Signalness > 0.9 at energies above 100 TeV
- Online July 2020, two alerts as of Dec 2020.



https://gcn.gsfc.nasa.gov/amon_icecube_cascade_events.html

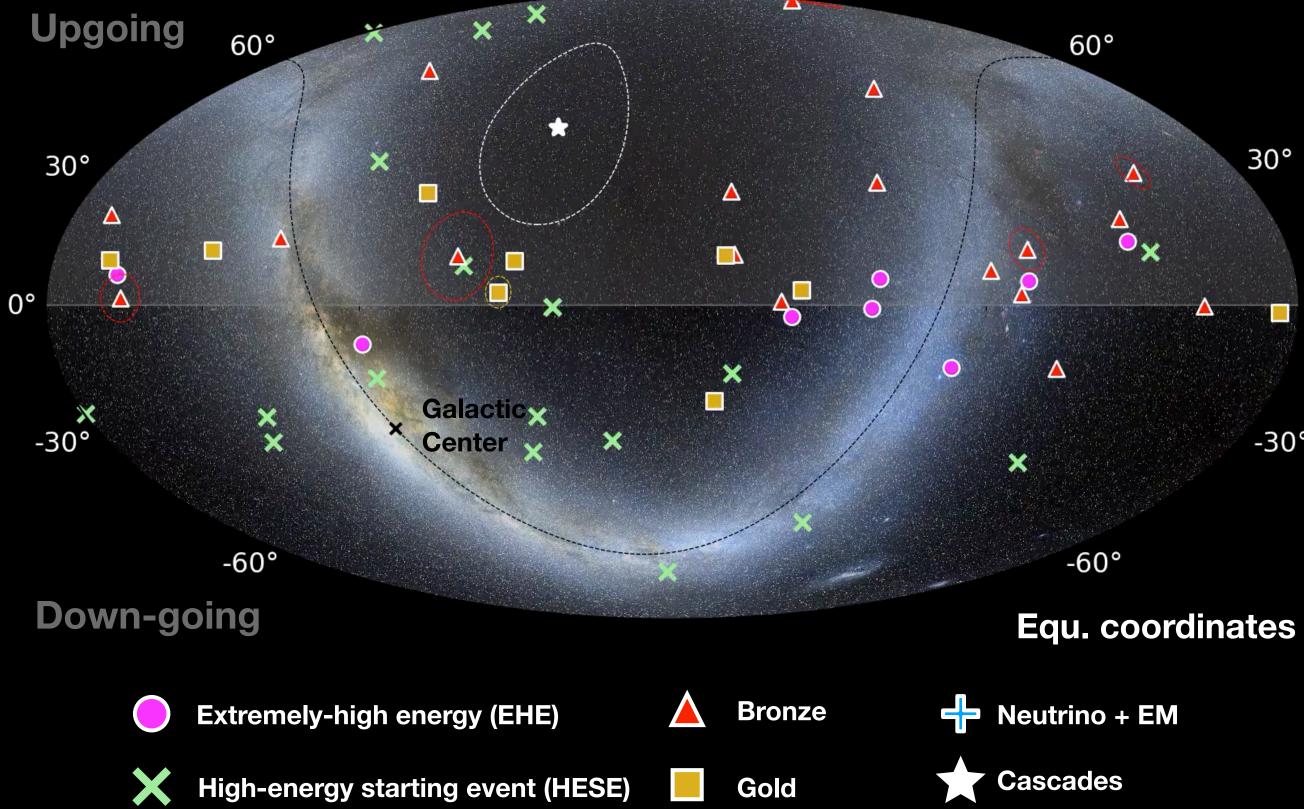








Jul 03, 2020

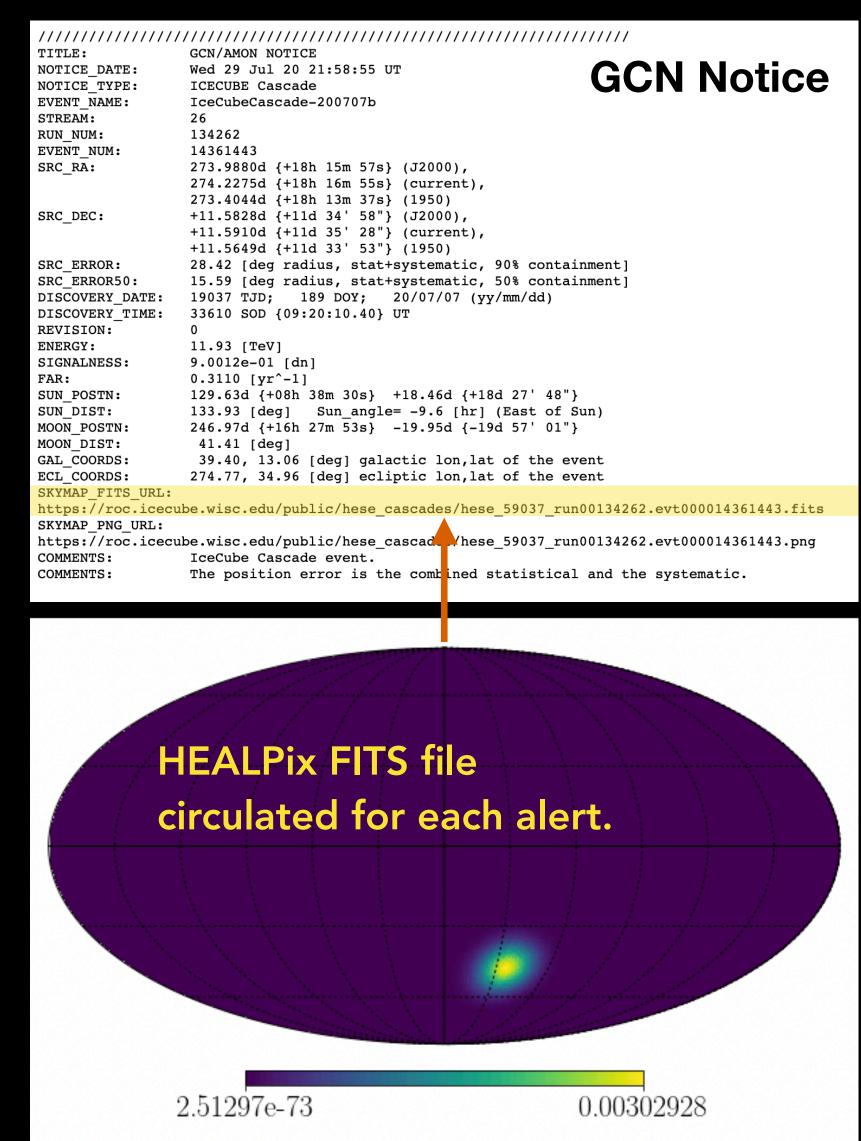


Cascade alerts

30°

-30°

0°



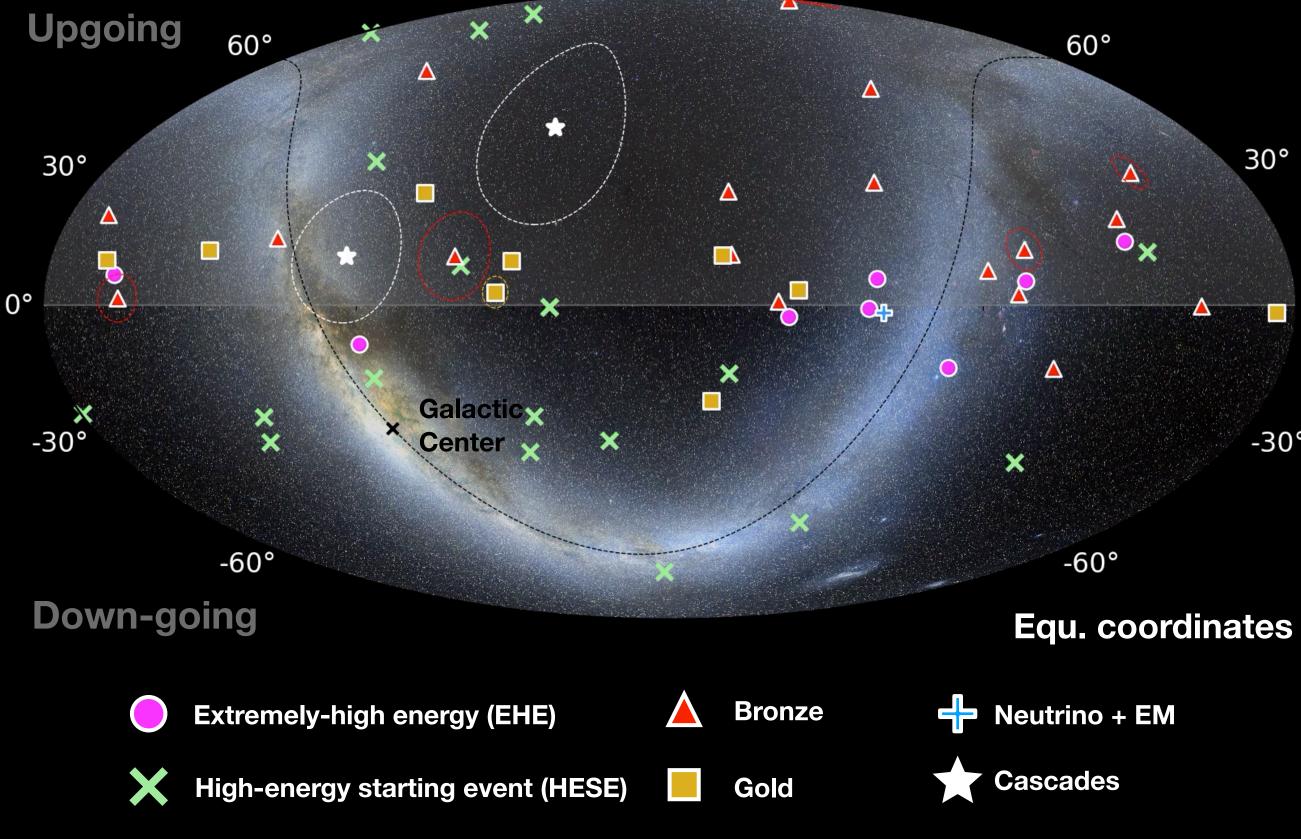
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Jul 20, 2020



[PoS-ICRC2019-841]

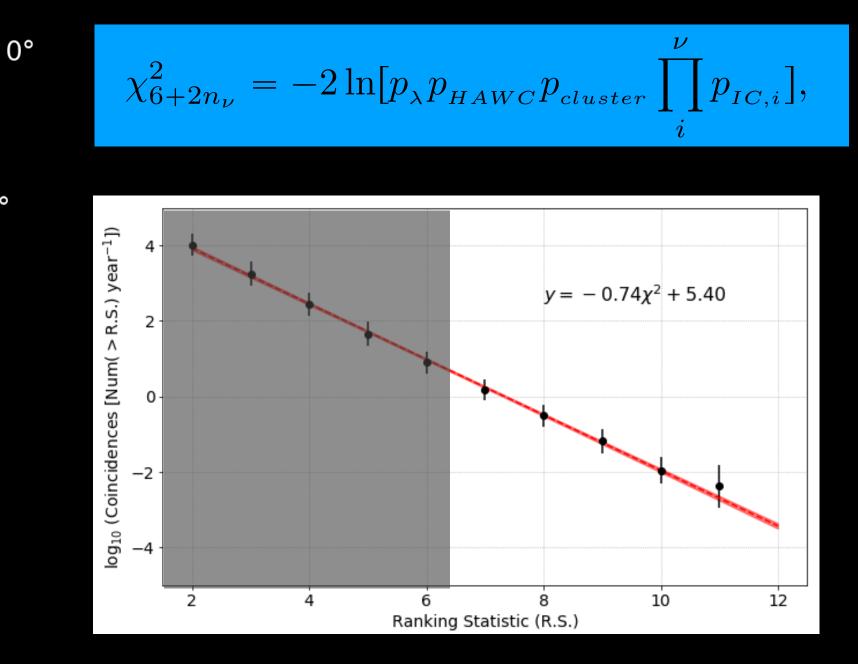
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Neutrino-gamma coincidences

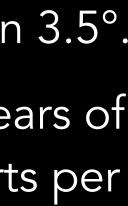
- HAWC + IceCube: HAWC daily transit hotspot correlated by AMON with IceCube neutrinos within 3.5°.
- Ranking statistic (RS) distribution derived from 2 years of scrambled data. Cuts on RS defined to send 4 alerts per year to GCN.
- Started April 2020, **3 alerts sent so far** ightarrow

30°

-30°



https://gcn.gsfc.nasa.gov/amon_nu_em_coinc_events.html

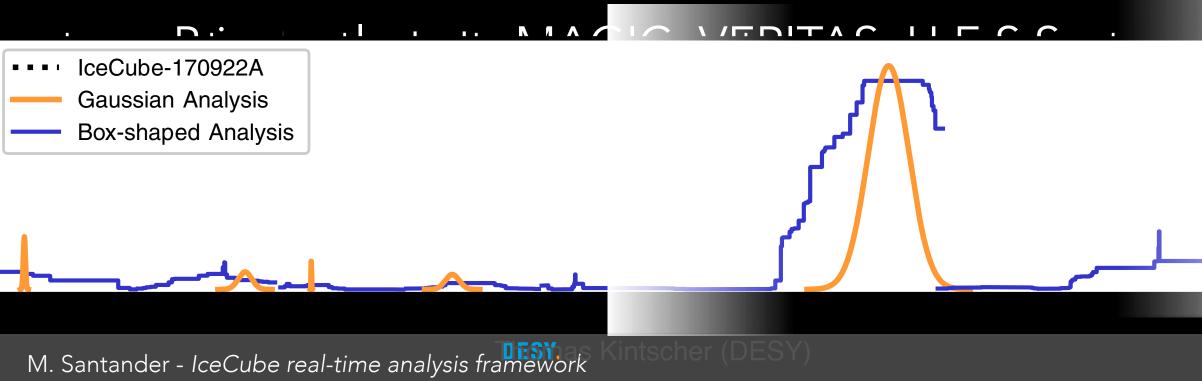






Clustering searches

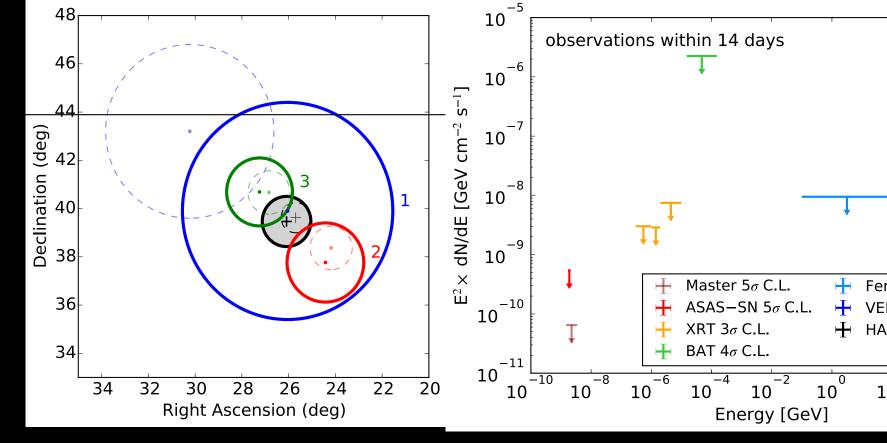
- Spatial correlations can reveal an astrophysical signal buried in the atmospheric neutrino background.
- Optical follow-up (OFU): GRB/SN
 - 2+ events in 100 s, within 3.5°
 - Private alerts to ZTF and Swift ightarrow
- Gamma-ray follow-up (GFU): Blazar flares
 - Likelihood analysis on variable time-scales correlated with ulletknown or likely VHE gamma emitters.

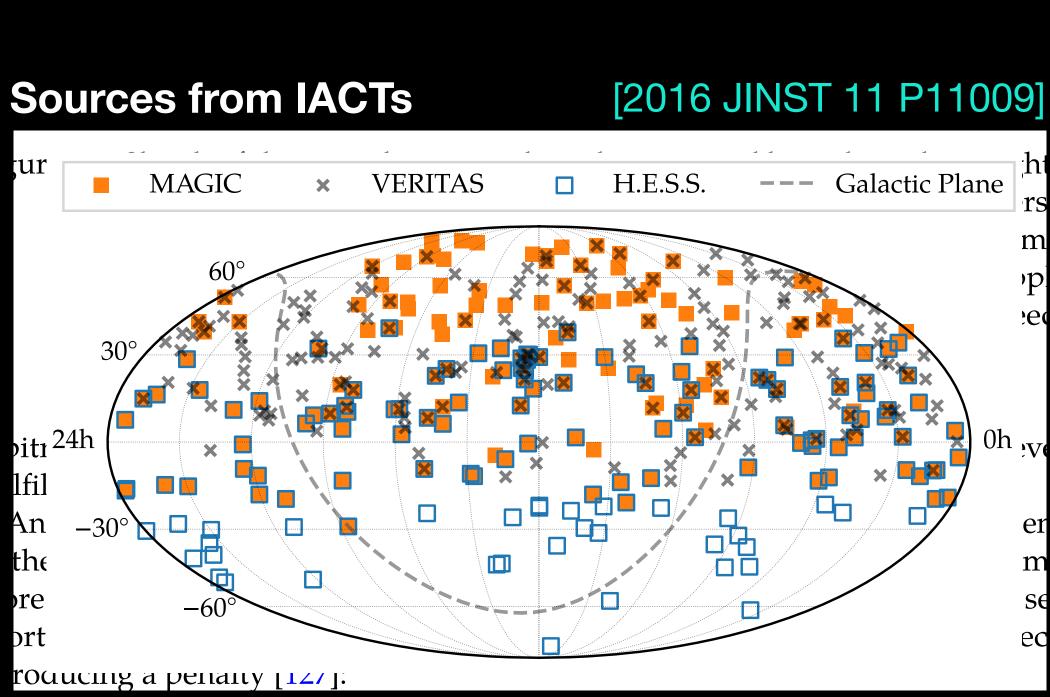


Neutrino triplet (2016)

[A&A 607, A115 (2017)]

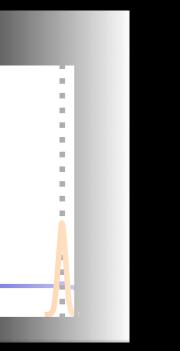




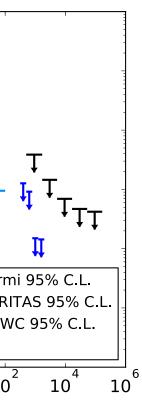


 $\mathcal{L}(\hat{n}_s, \hat{\gamma}) \mathcal{U}(t_i, t_k)$

 $\Lambda = 2 \log$





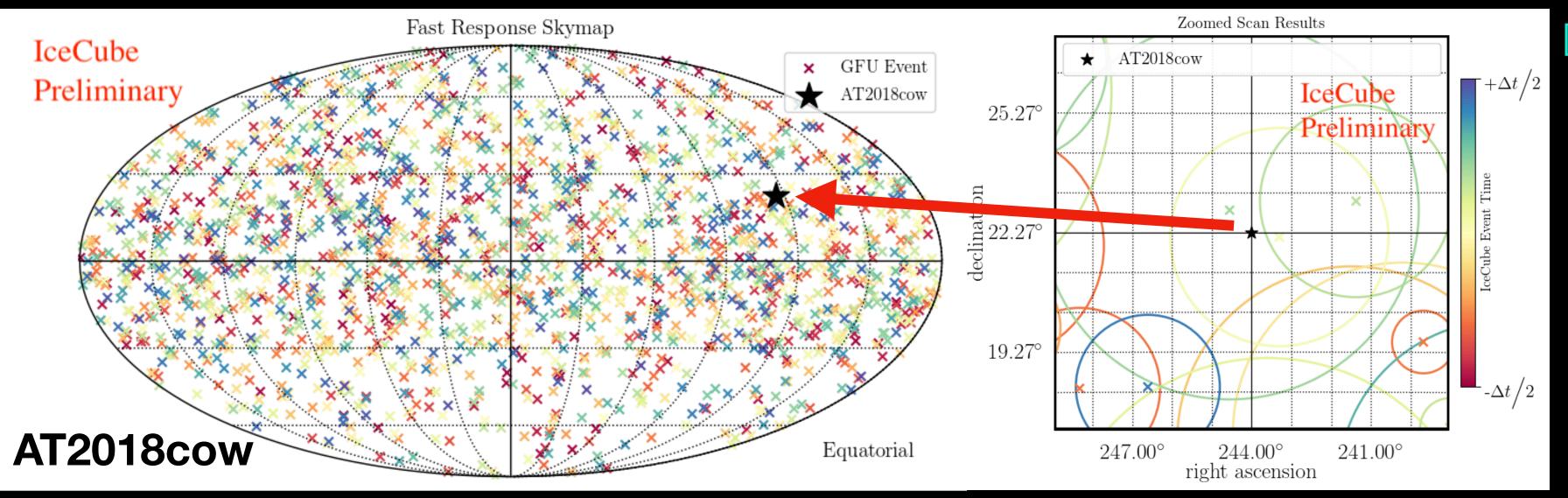




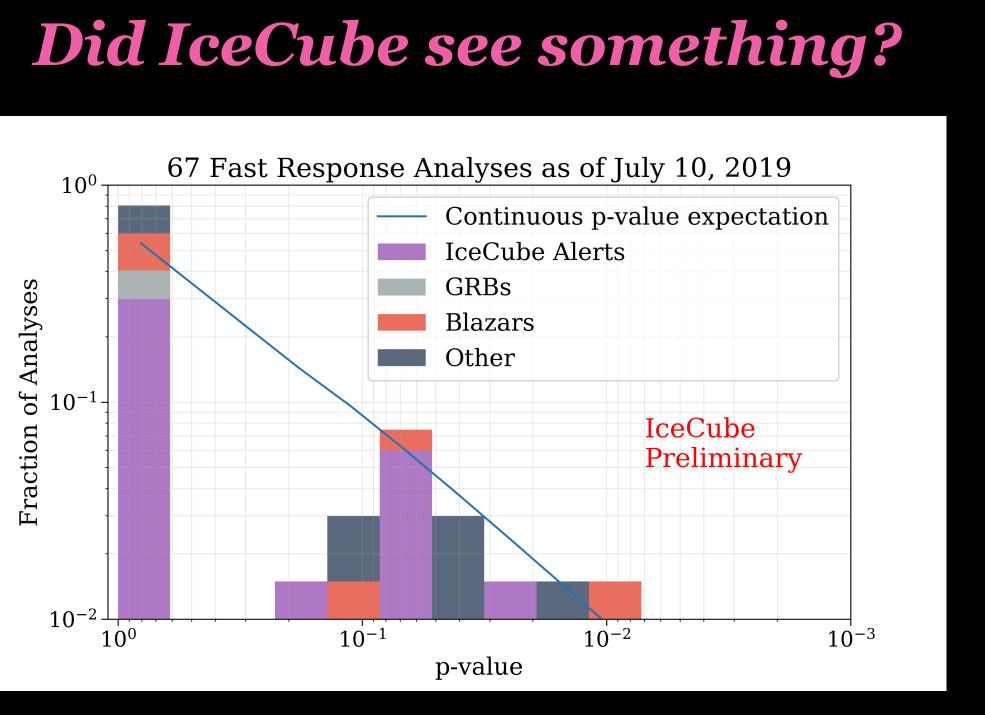


Fast response analysis

- Fast response analysis following:
 - IceCube HE alerts (search for additional, LE nus). 102 up to Dec 2020.
 - HE astrophysical events with potential neutrino emission: ATels, GCN, etc. 60 up to Dec 2020.



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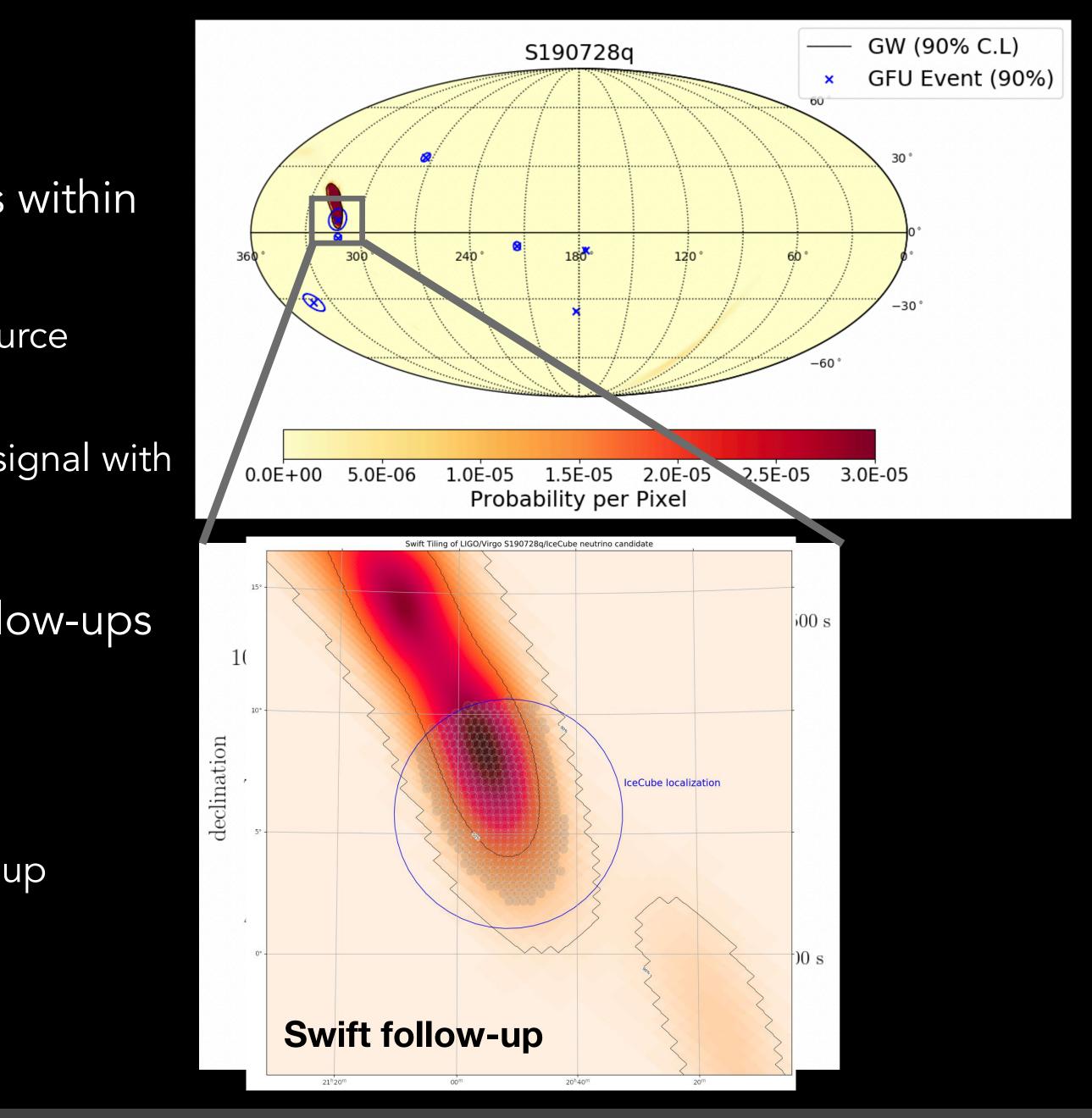




Gravitational waves

- Two independent analysis of neutrino candidates within 500 s of the GW trigger.
 - Unbinned maximum likelihood search: test for point source consistent with GW localization.
 - Bayesian approach: probability of a joint GW+nu joint signal with astrophysical priors
- Results are reported in GCN circulars. 56 GW follow-ups during O3 run of LIGO/Virgo.
- Example: GW190728, BBH merger
 - p-value ~ 0.01 in both analyses, triggered MWL follow-up
 - https://gcn.gsfc.nasa.gov/gcn3/25210.gcn3

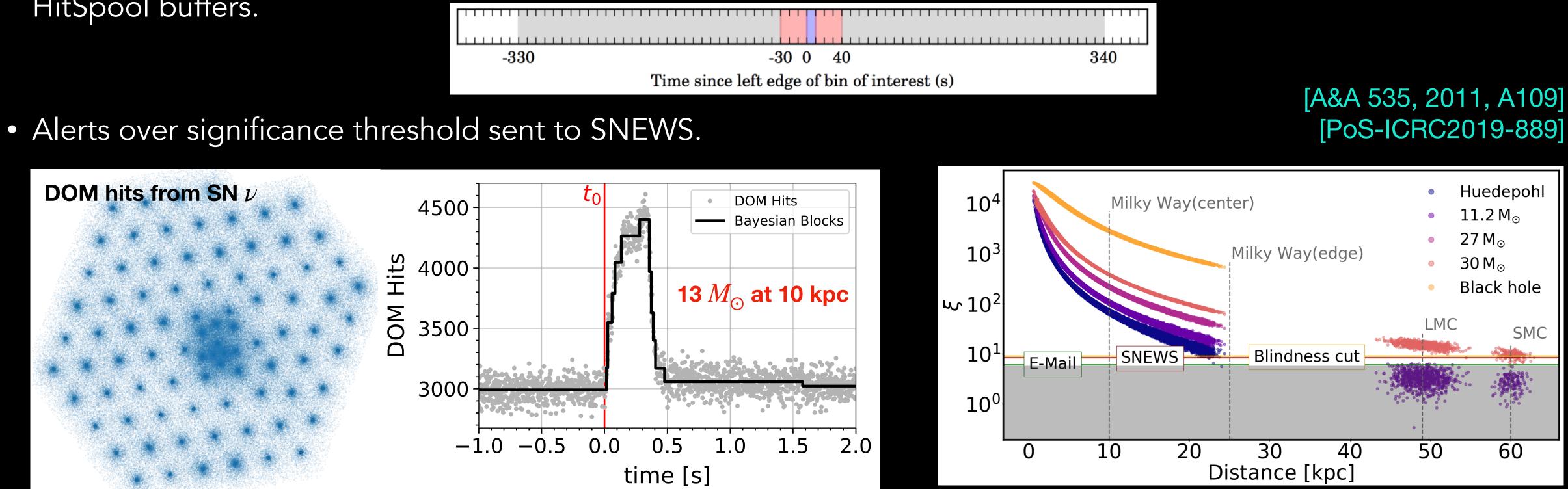
[Astrophys. J. Lett. 898 (2020) L10]

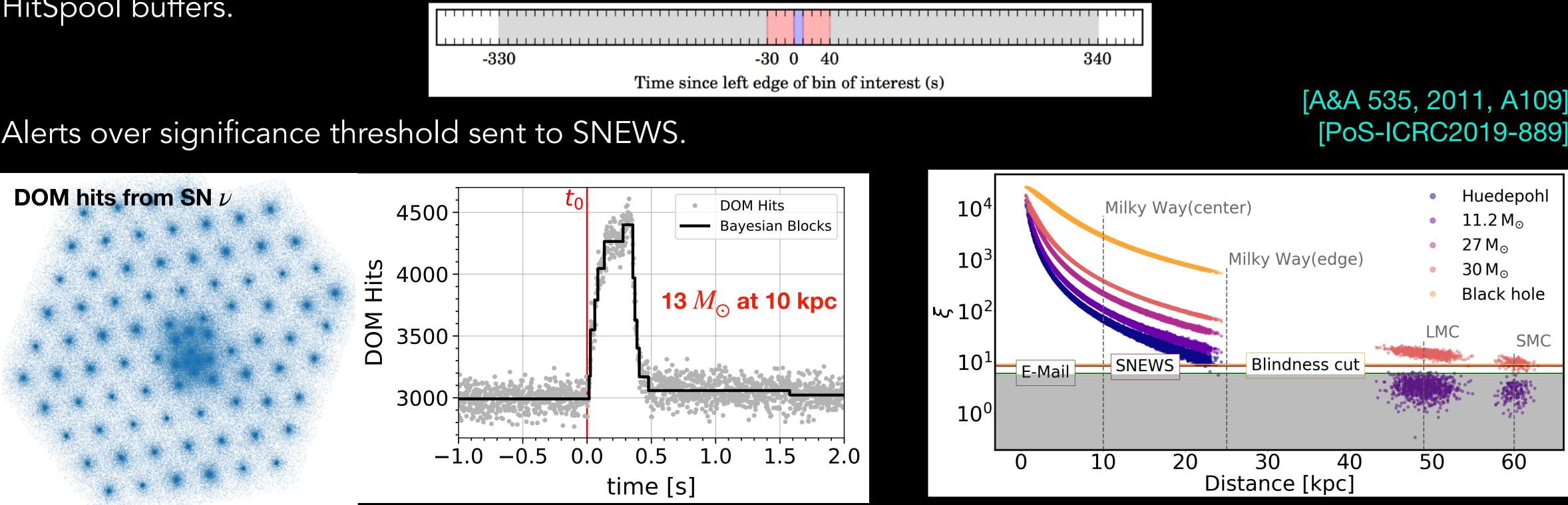




Galactic supernovae

- produced by Cherenkov photons from 10 MeV neutrino interactions.
- HitSpool buffers.





• A Galactic core-collapse supernova would be seen by IceCube as an overall increase in the DOM noise rate

• SNDAQ searches for correlated noise rate increases in a 0.5 s time bin with respect to a moving average calculated over a ±5 min window. Can be triggered by SNEWS and LIGO GW alerts. SNDAQ retrieves waveforms from





Future directions

[PoS-ICRC2019-1177]



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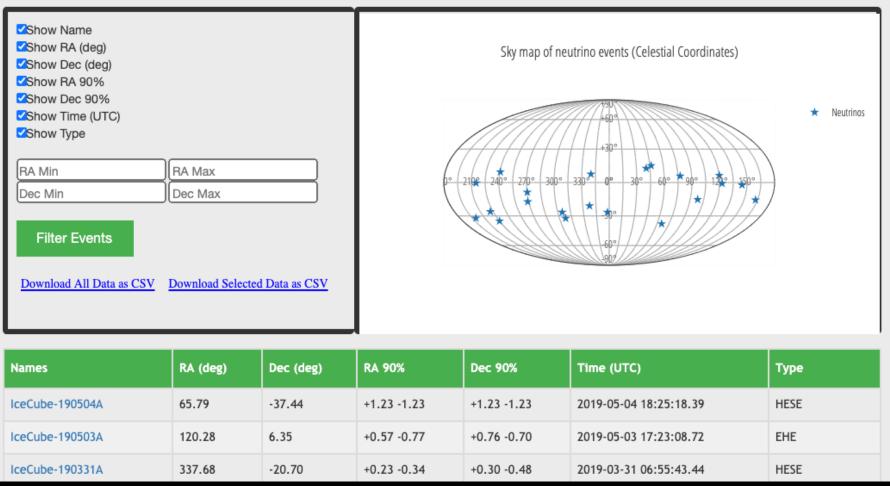
https://neutrino-catalog.icecube.aq/main

IceCube Catalogue of Astrophysical Neutrino Candidates

Welcome to this curated online catalogue of astrophysical neutrino candidates. Information on high-energy neutrinos if often published multiple times using various analyses. This catalog records all these publications and highlights the information recommended to use in further studies.

(RA & Dec): J2000

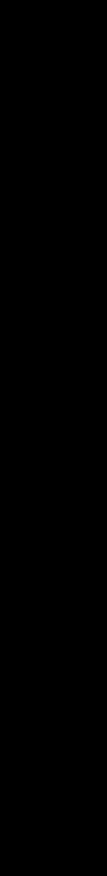
Events



Catalog of astrophysical neutrino events (both archival and realtime) in preparation.

Angular error estimation improvements for realtime alerts.

Questions? Requests? Contact the IceCube Realtime Oversight Committee (ROC) at roc@icecube.wisc.edu







Conclusions

- IceCube operates a realtime alert program from MeV to PeV energies
 - Realtime high-energy neutrino events of potential astrophysical origin
 - Monitoring of known gamma-ray emitters
 - All-sky monitoring for neutrino clusters
 - Fast-response analysis to external triggers
 - Sensitivity to Galactic supernovae



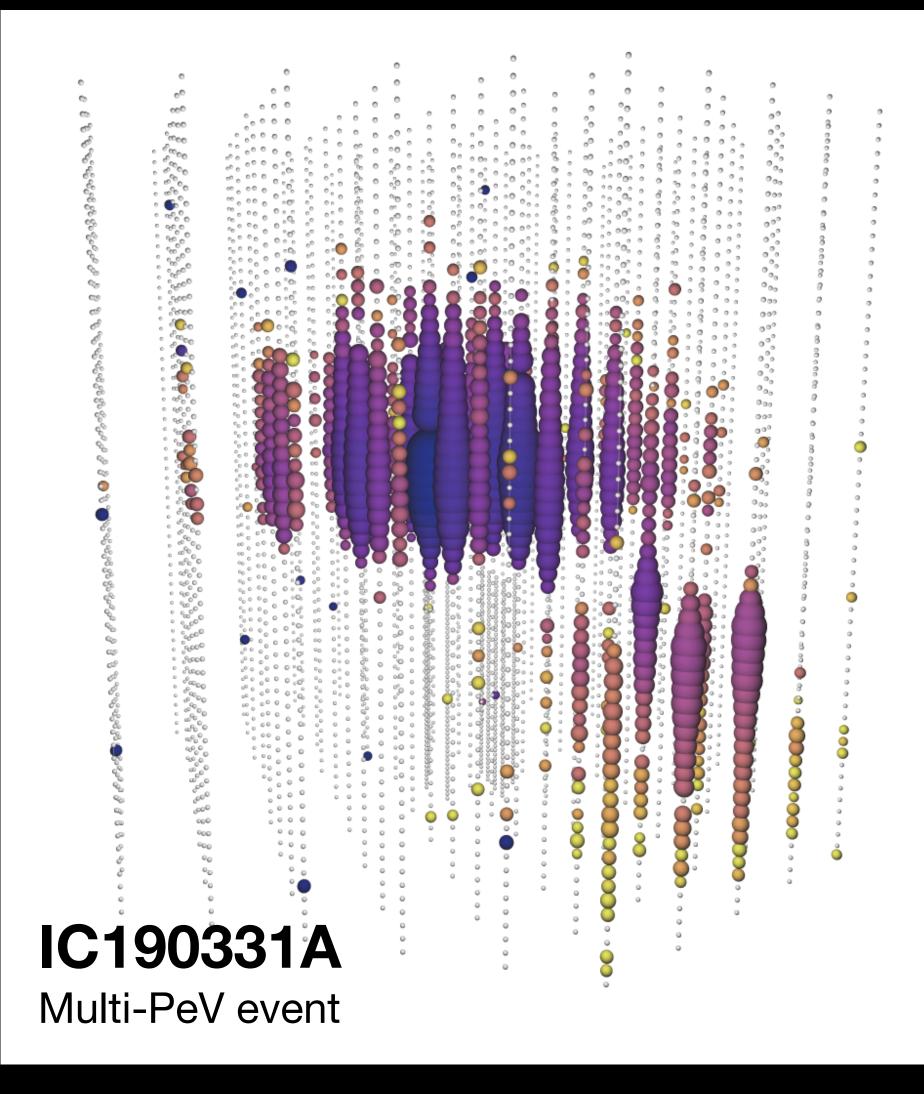
A lot of interest in collaborating and coordinating with KM3NeT on realtime activities!



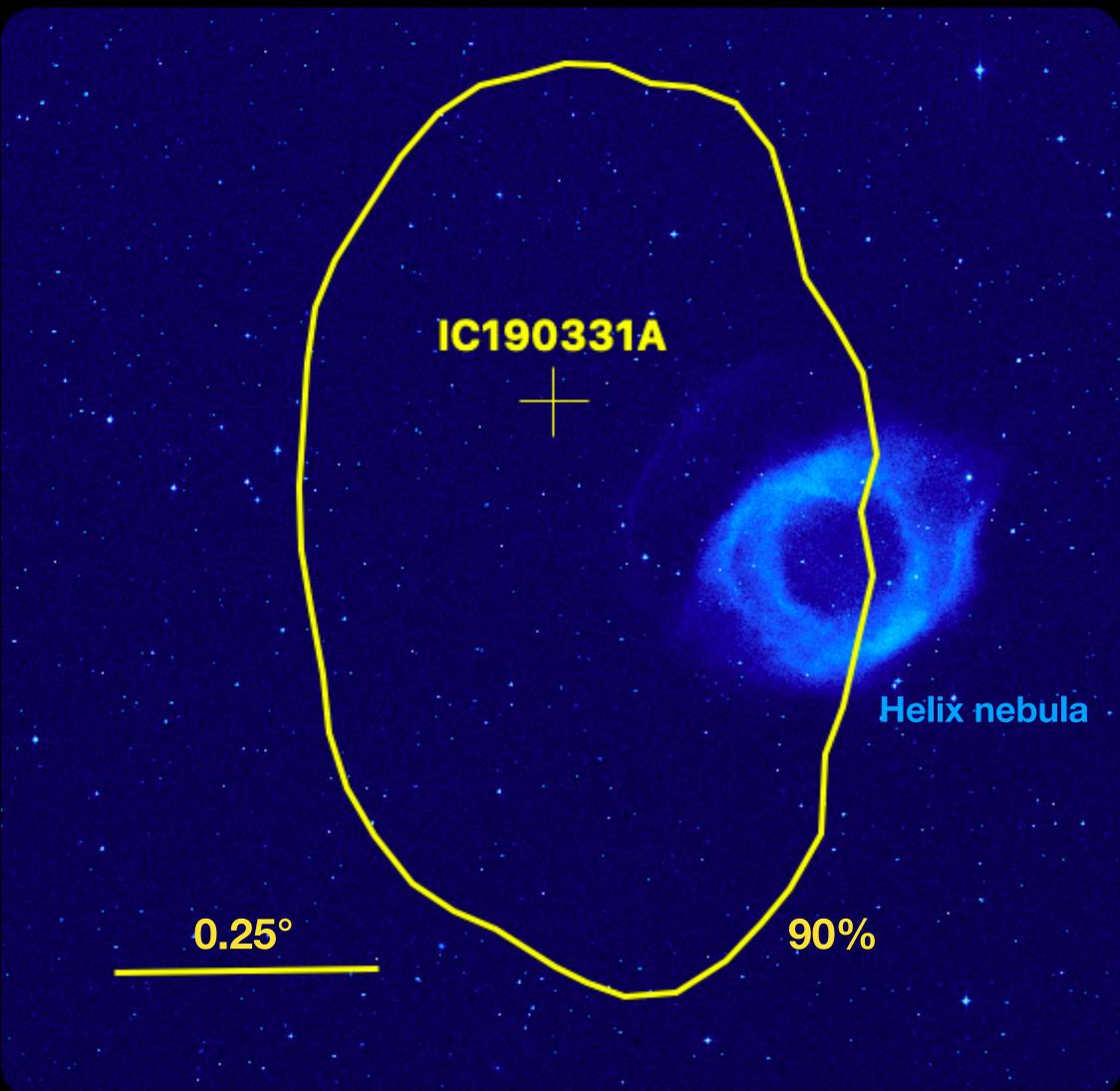




More recent alerts



Digitized Sky Survey (optical)



No obvious high-energy EM counterparts





