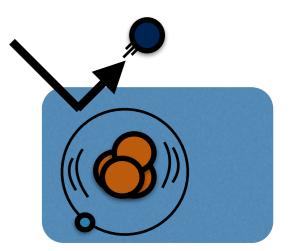
# Superfluid Effective Field Theory for Dark Matter Direct Detection

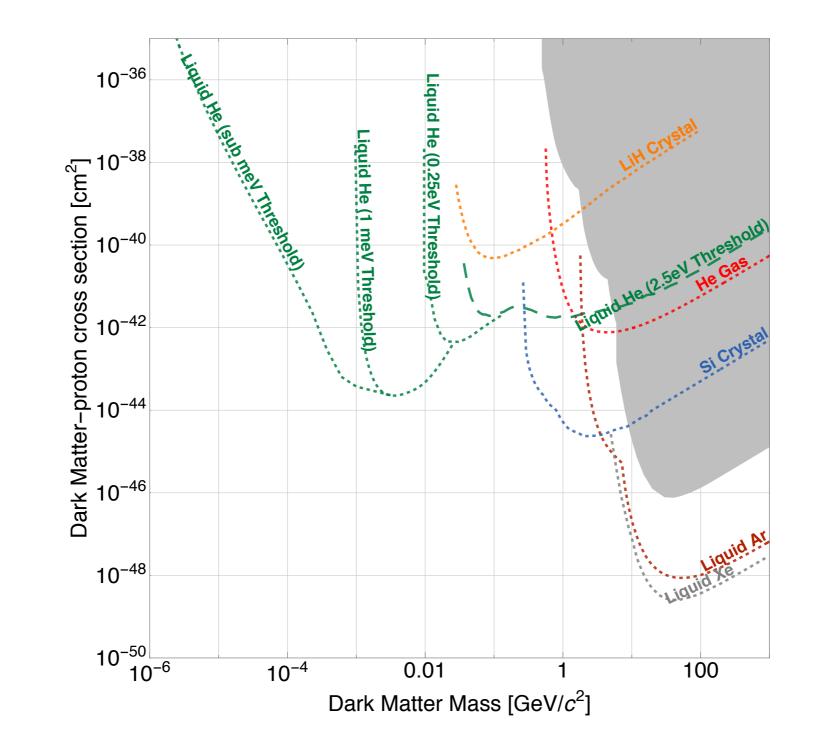
Wei Xue

June 9, 2022 CERN-CKC workshop

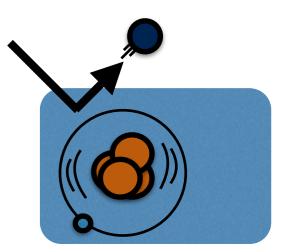


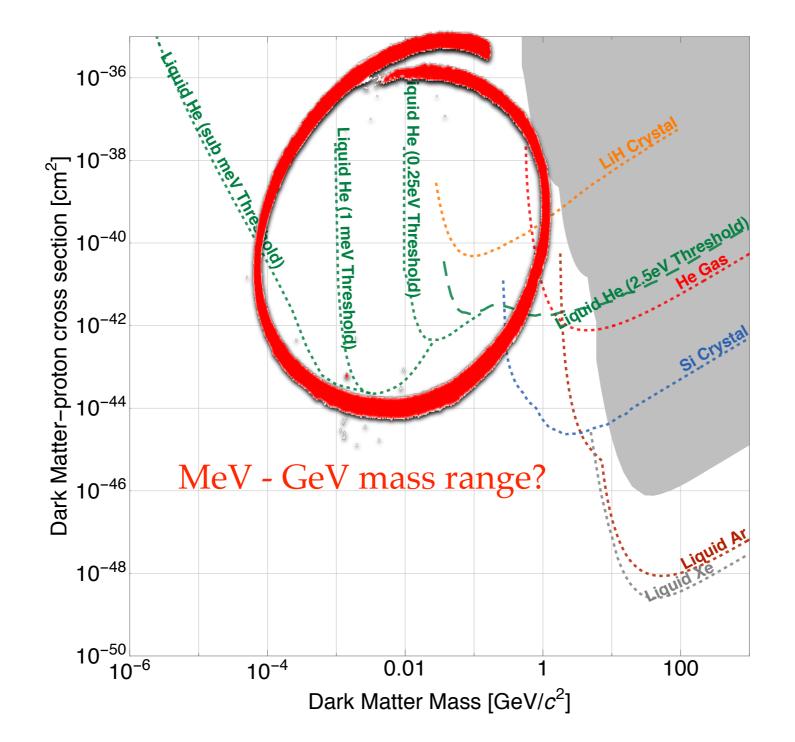
#### Dark matter direct detection





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#### Plan

• theoretical framework to understand the quasi-particle production and thermalization

arXiv: 2108.07275

- simulation to know the momentum spectrum, flux, thermalization
- a prototype experiment at University of Florida

#### Collaboration











Jordan Smolinsky

Yoonseok Lee

Tarek Saab Konstantin Matchev

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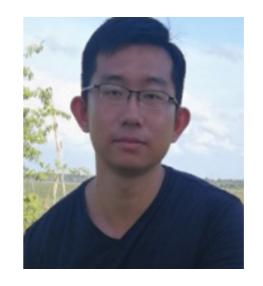
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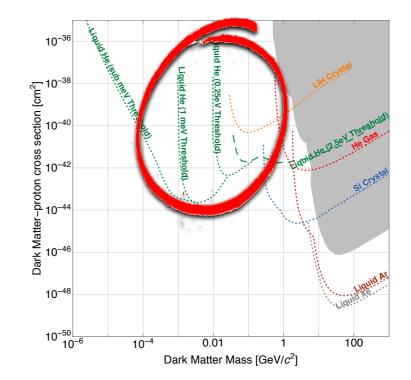
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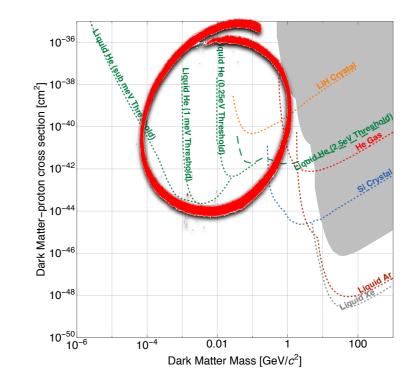
Yining You

# Why <sup>4</sup>He superfluid?



## Why <sup>4</sup>He superfluid?

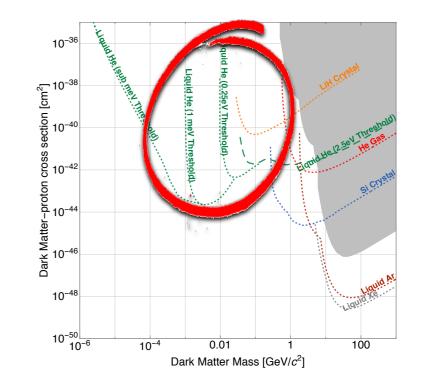
• Helium as the second lightest element an excellent target material for detecting light particles



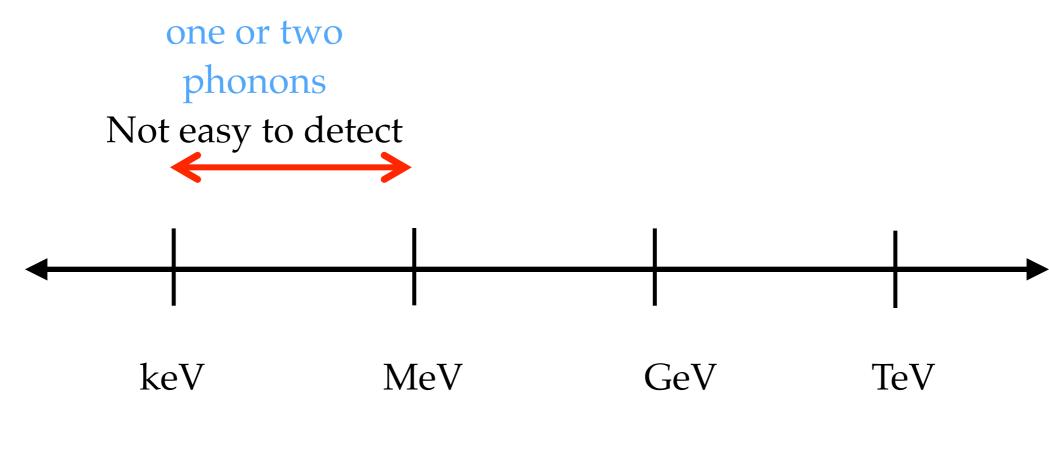
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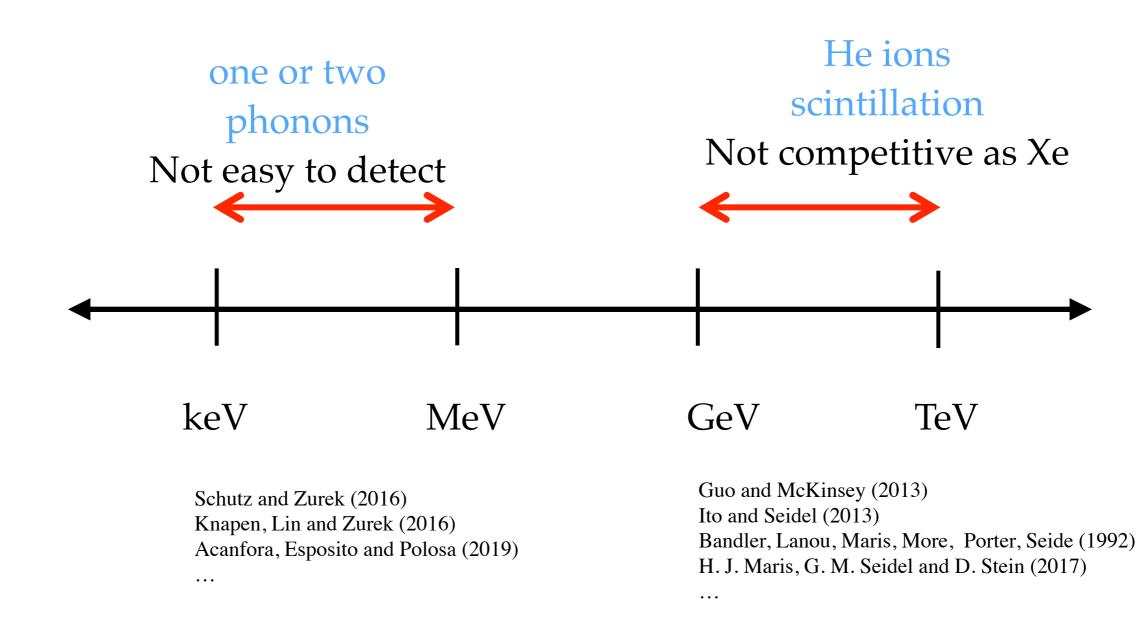
 superfluid Helium will be cooled to ~0.1 K the system behaves as a vacuum sensitive to tiny perturbations

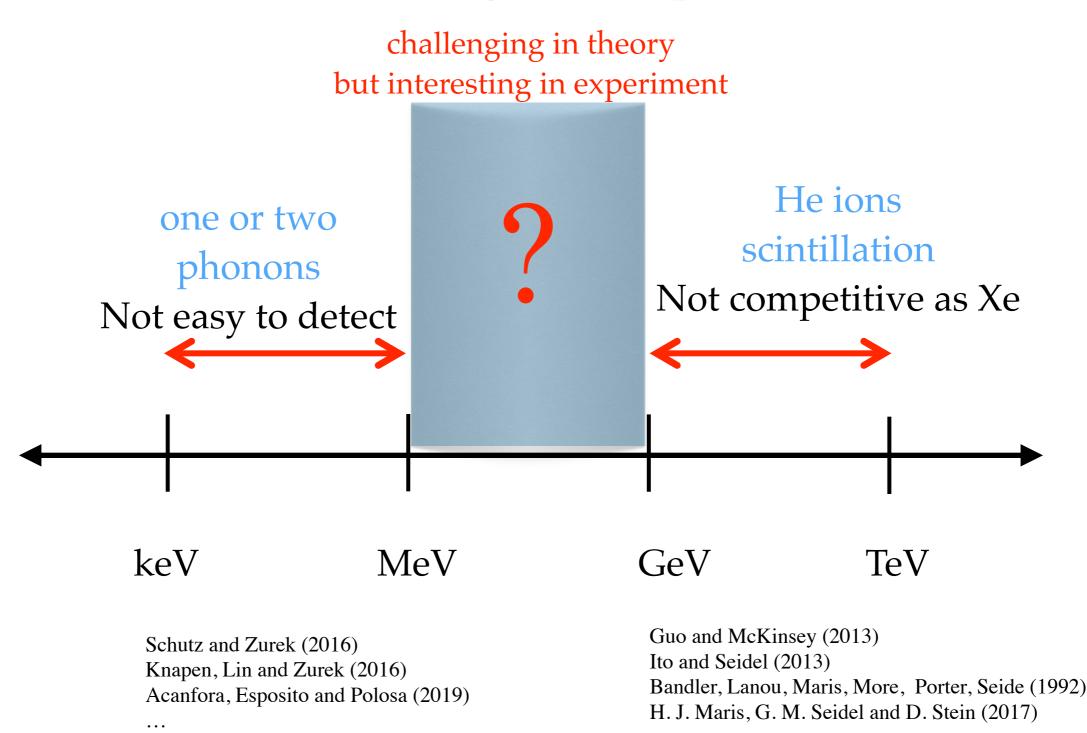






Schutz and Zurek (2016) Knapen, Lin and Zurek (2016) Acanfora, Esposito and Polosa (2019) ...





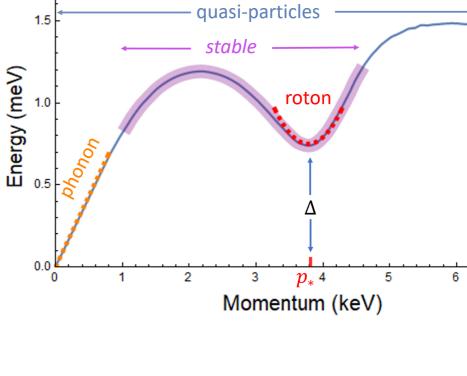
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. . .

# Challenges and motivations

• What happens when a test particle (dark matter or neutron) scatters with the helium superfluid?

 The perturbative theory of superfluid break down ~ keV ?
(inverse of the helium spacing)



1

0

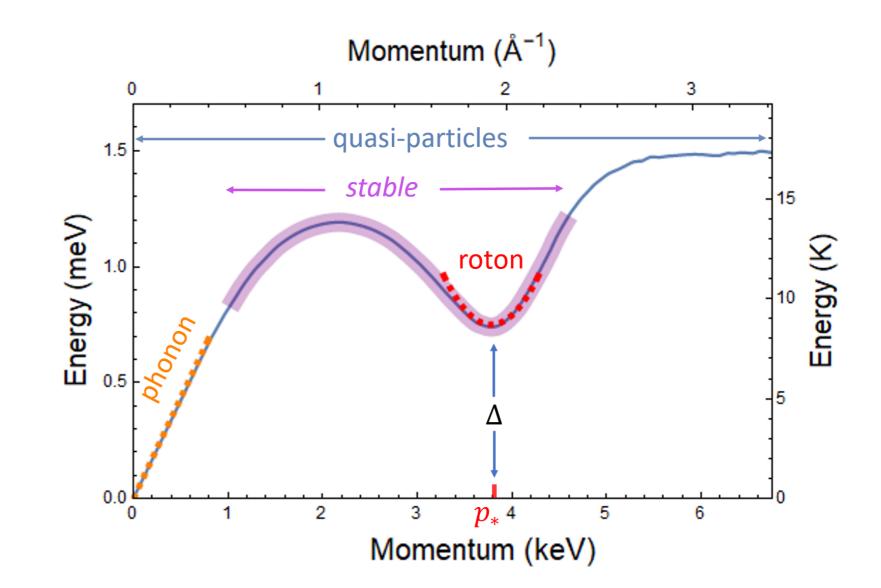
Momentum (Å<sup>-1</sup>)

2

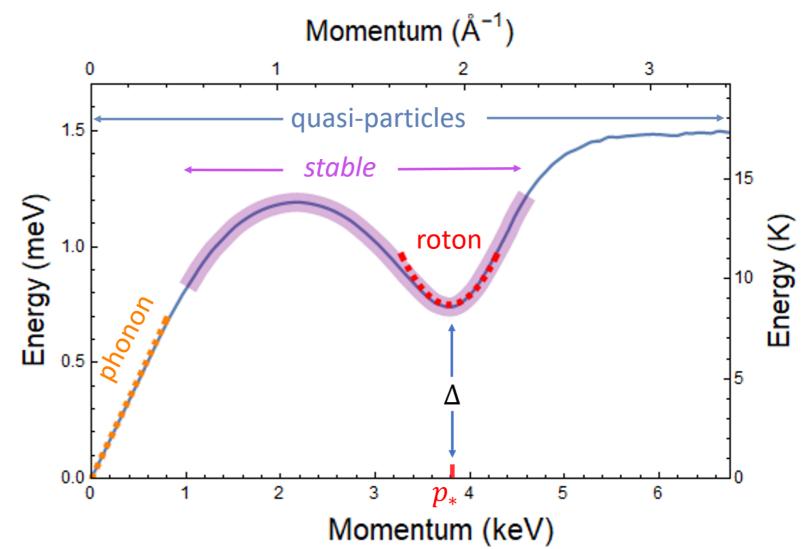
3

15

Energy (K)



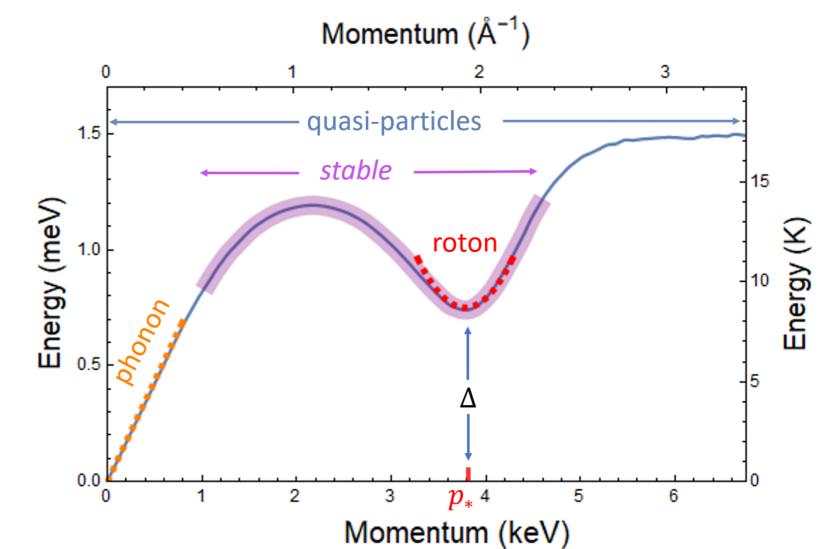


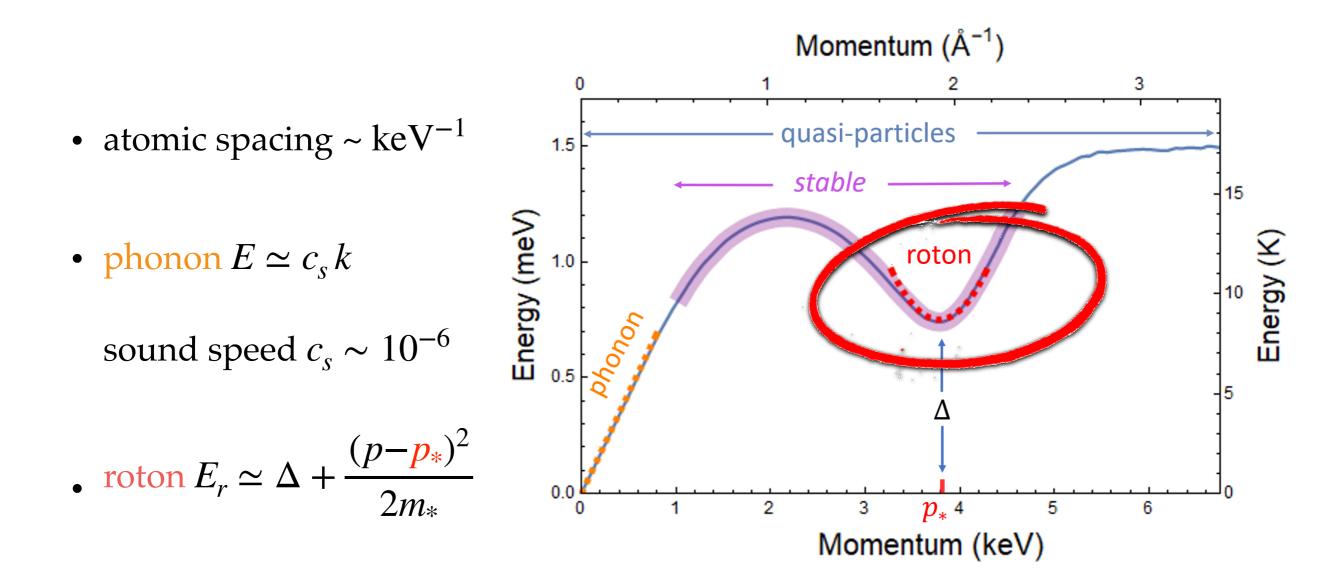




• phonon  $E \simeq c_s k$ 

sound speed  $c_s \sim 10^{-6}$ 



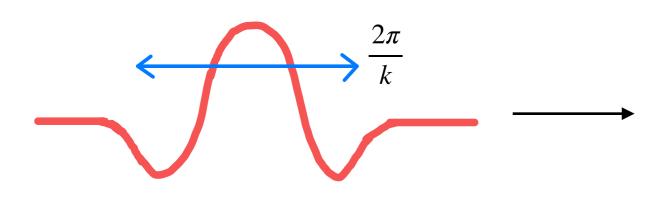


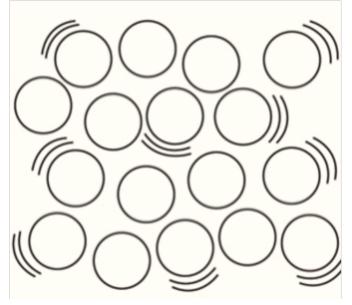
## de Broglie wavelength

• atomic spacing  $\lambda_a \sim \text{keV}^{-1}$ 

## de Broglie wavelength

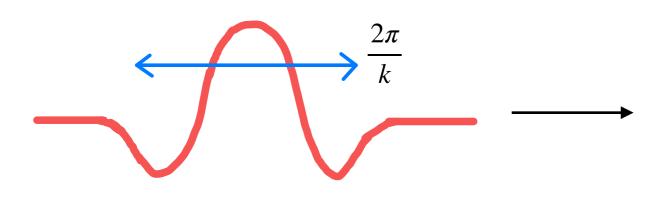
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- incoming particle de Broglie wavelength  $\gtrsim \lambda_a$  e.g. sub-MeV dark matter, v~ 10<sup>-3</sup> c

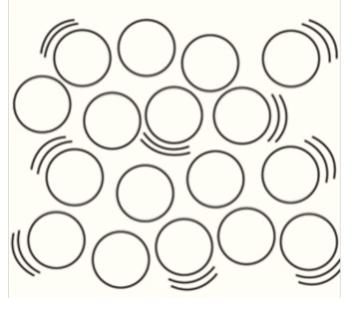




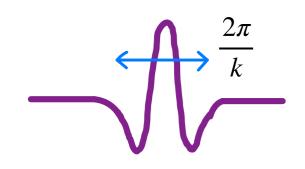
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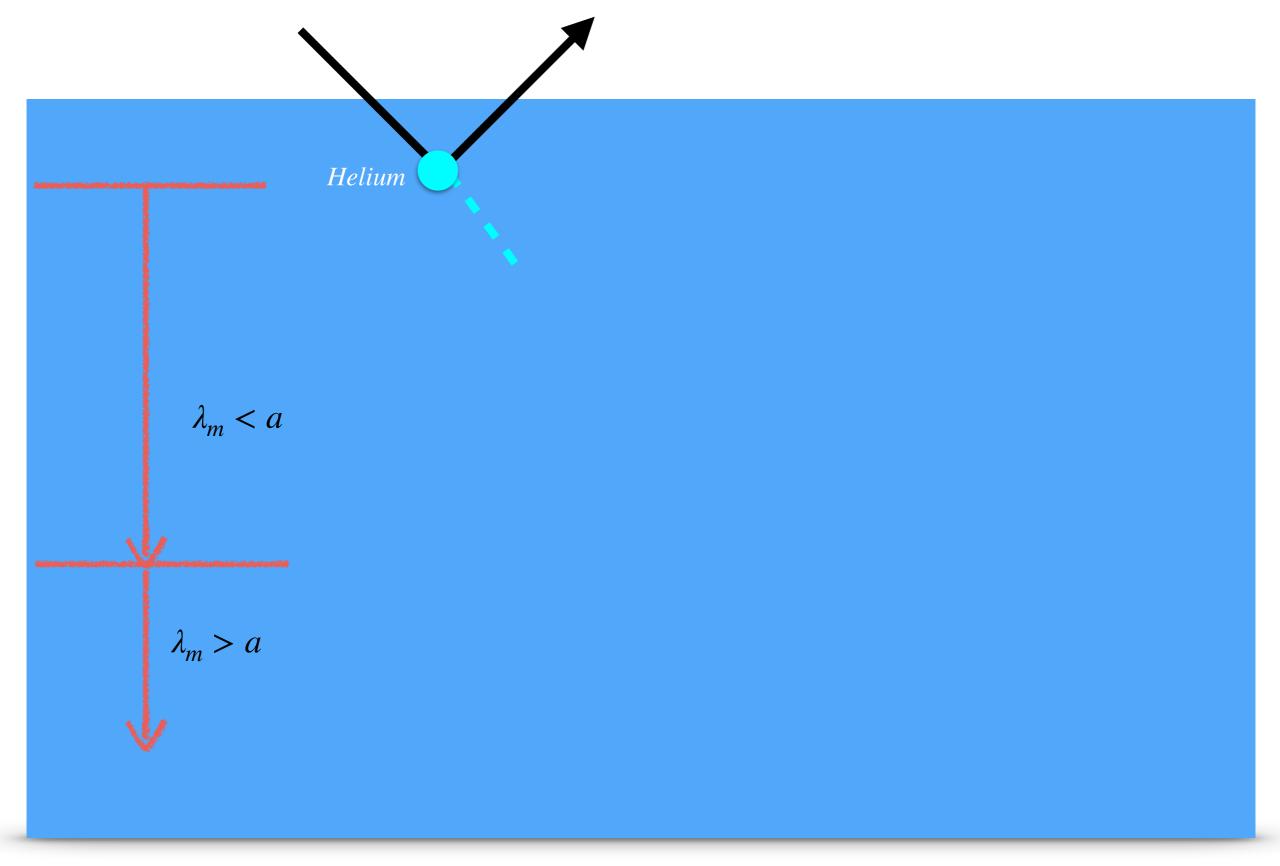


incoming particle de Broglie wavelength ≪ λ<sub>a</sub>
e.g. MeV-GeV dark matter

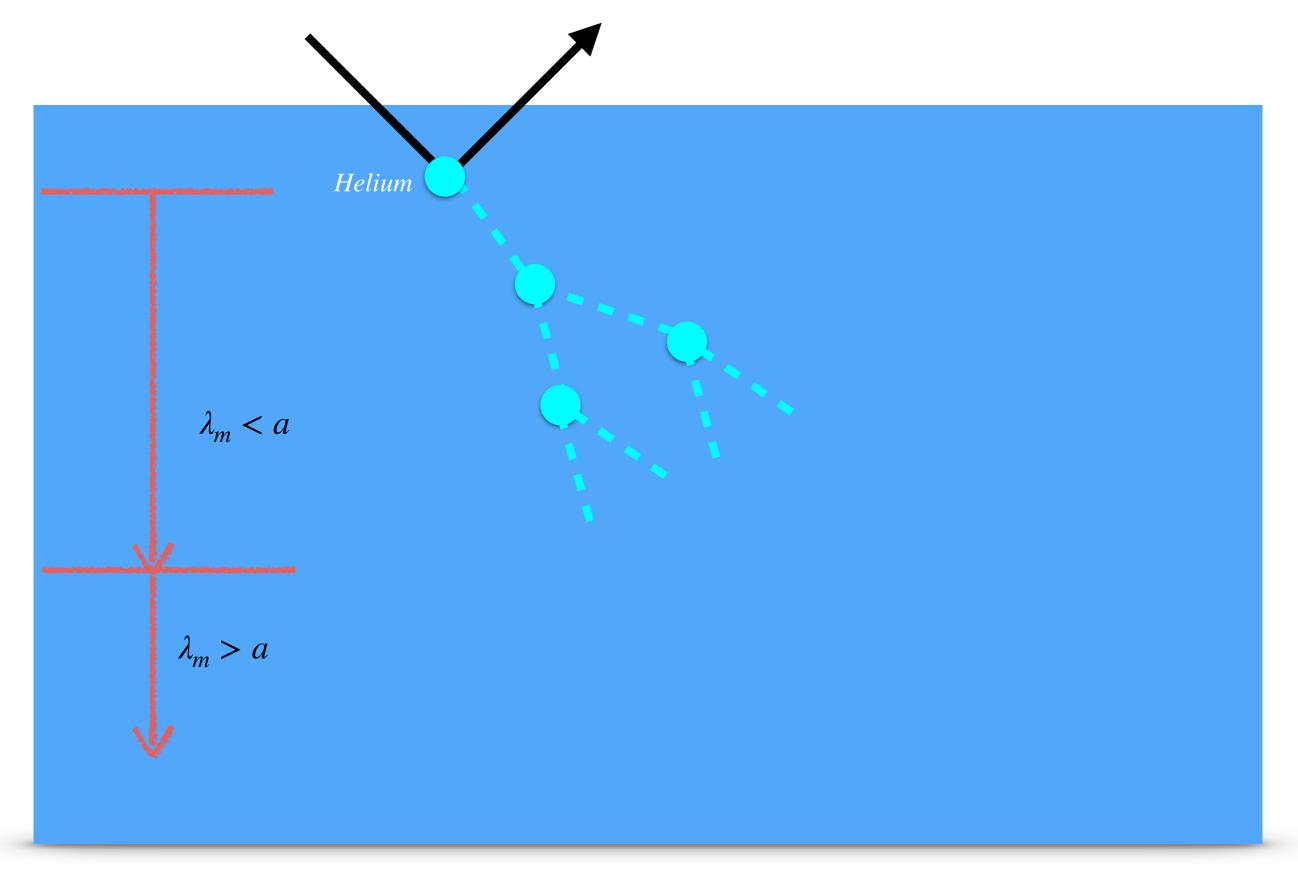




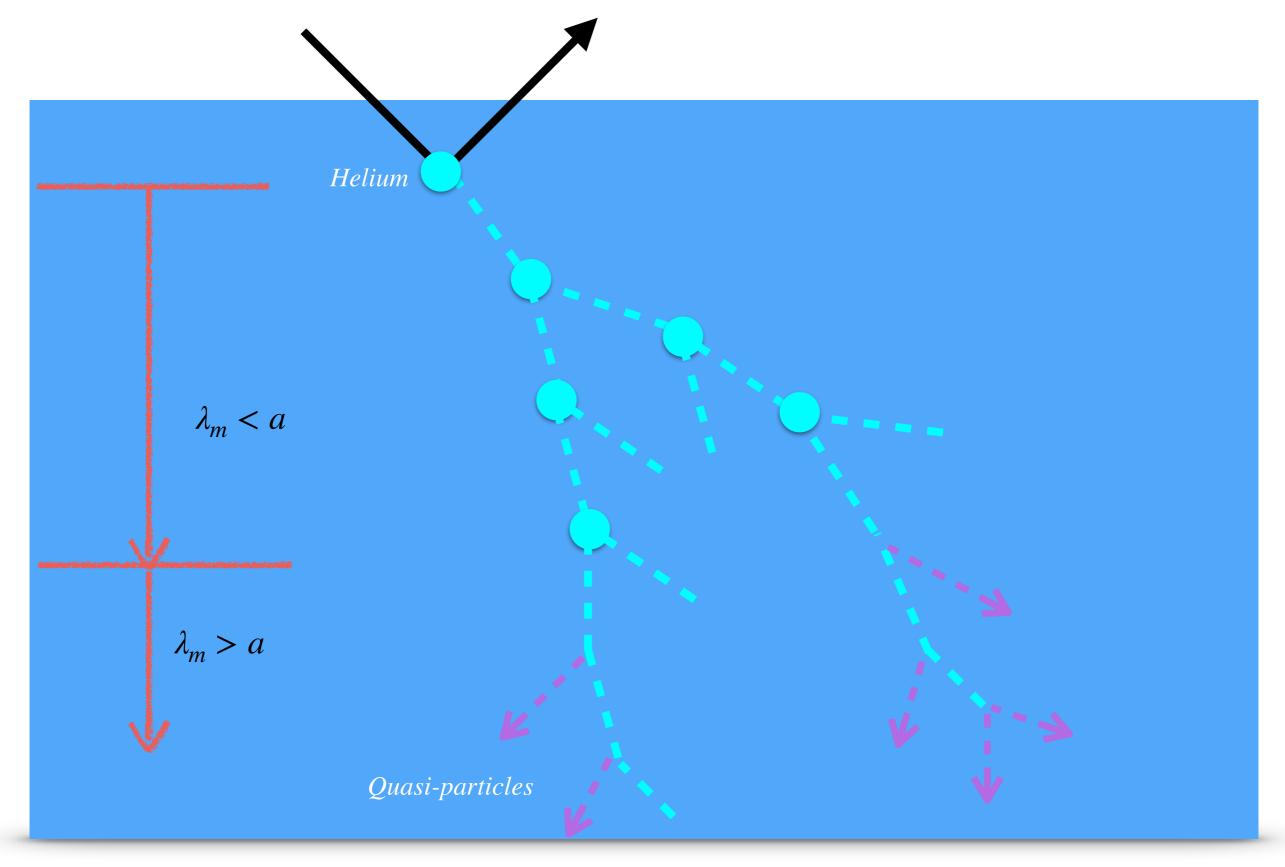
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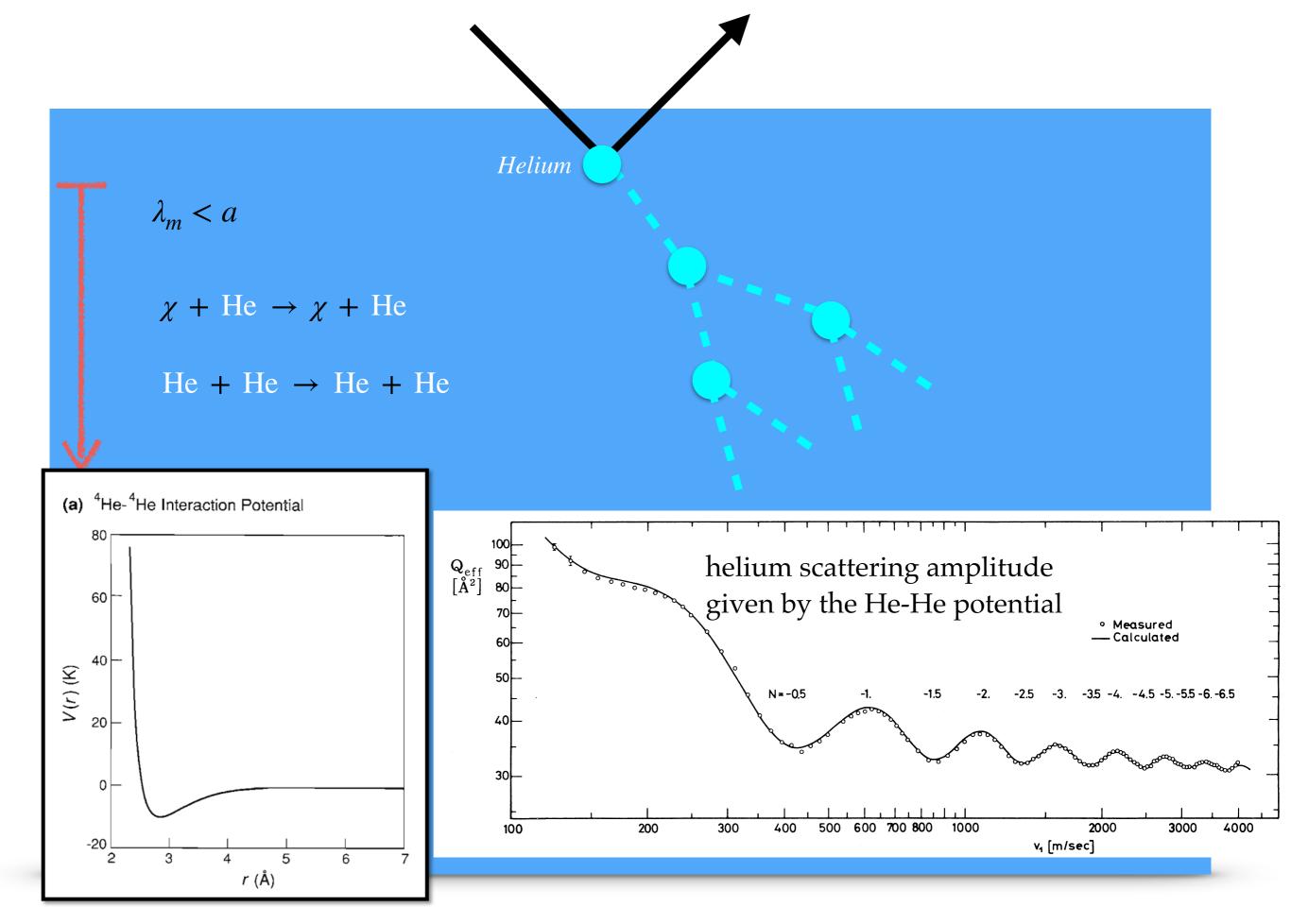


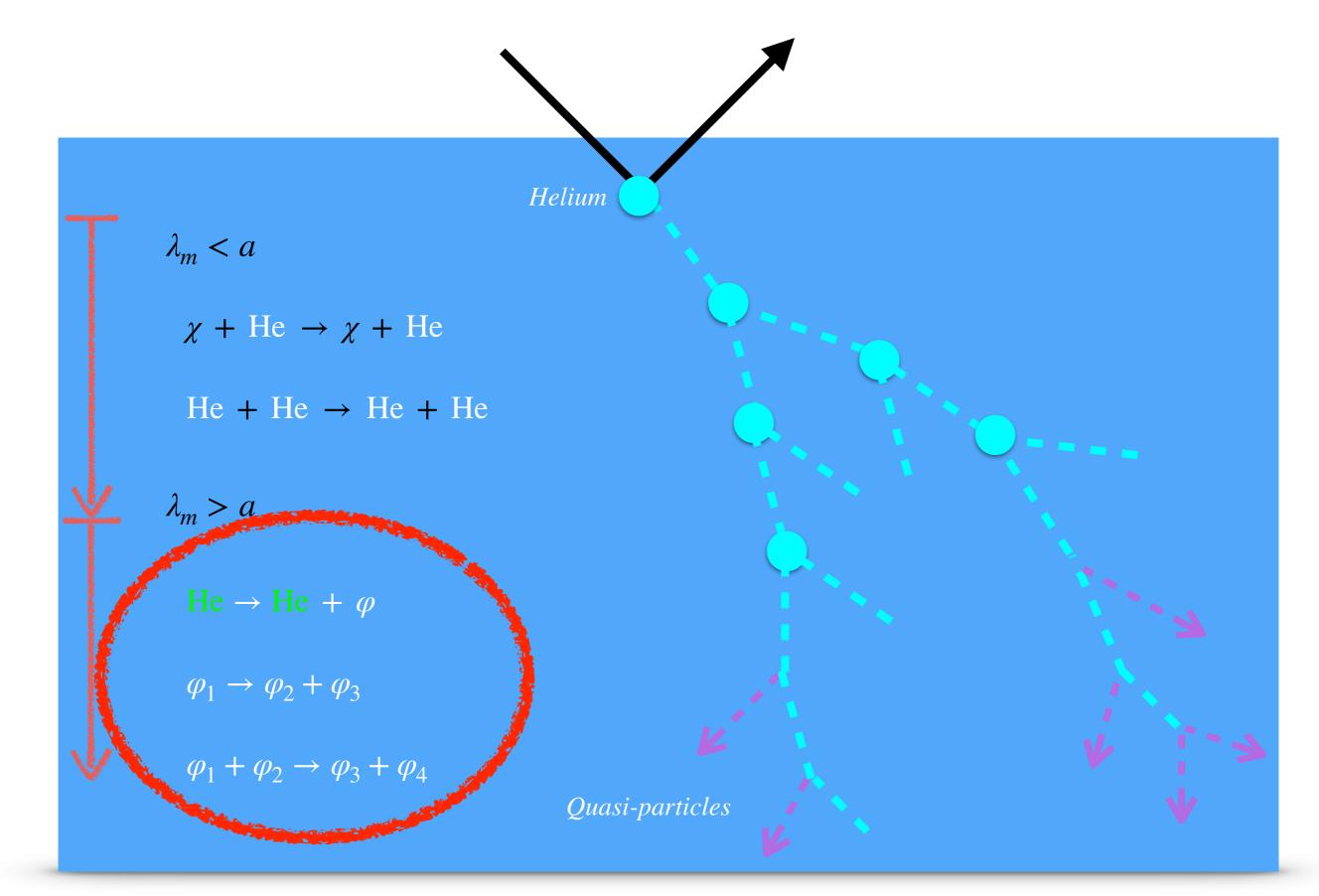
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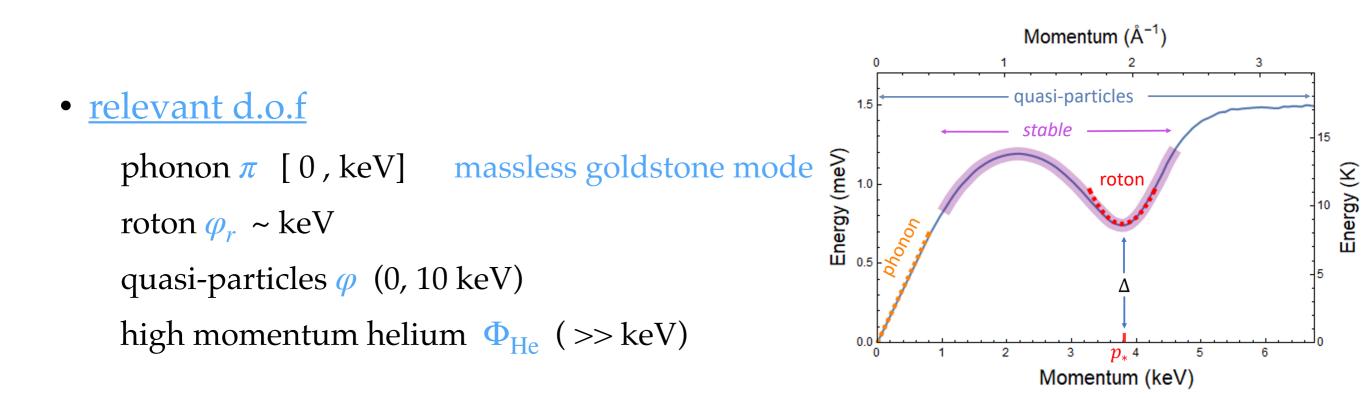


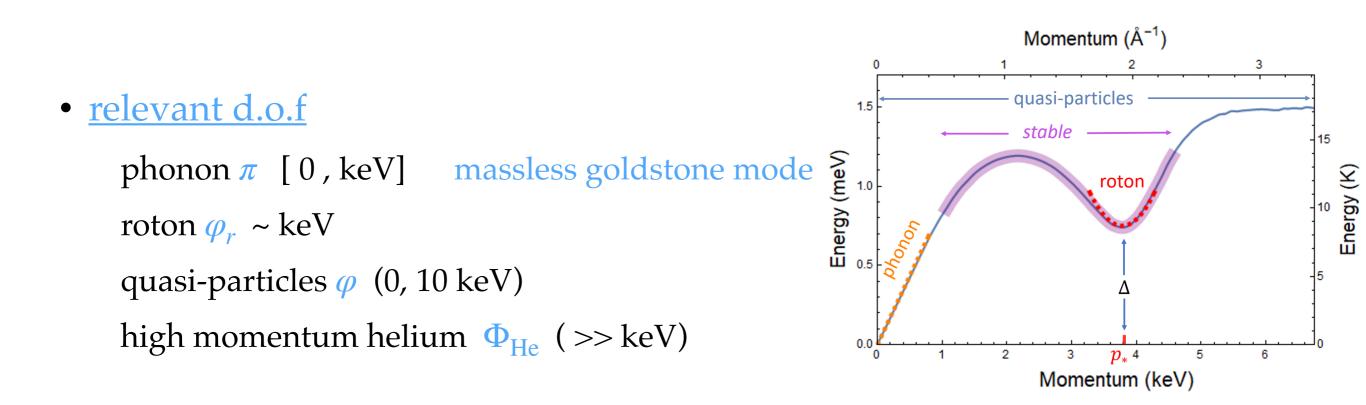
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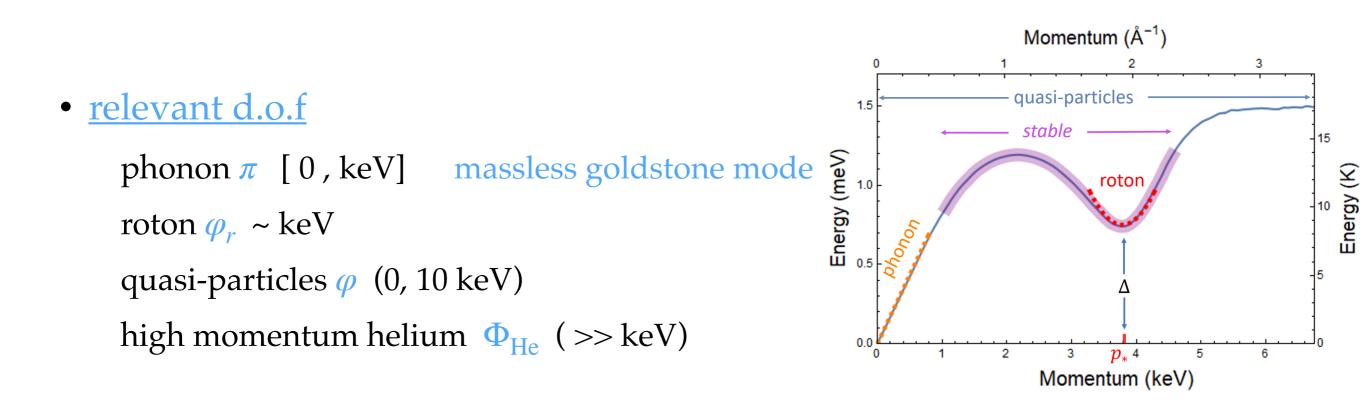








- <u>cutoff  $\Lambda \sim \text{keV}$ </u>
  - atomic spacing ~  $(\text{keV})^{-1}$



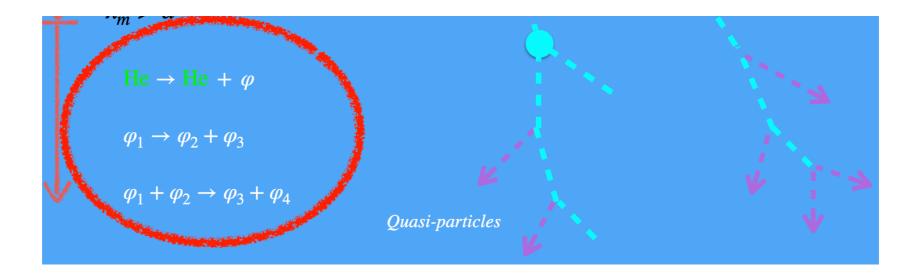
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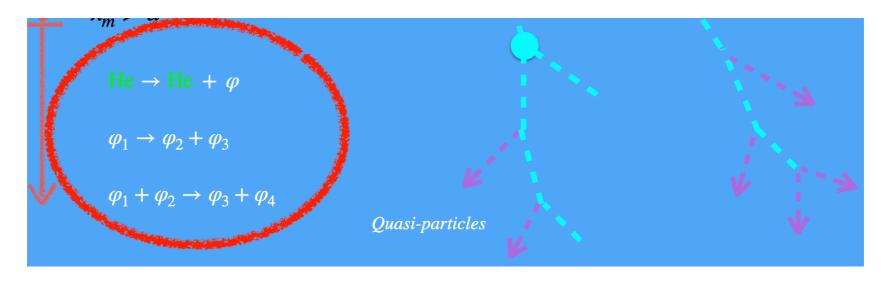
• <u>processes</u>

quasi-particle Helium interactions to study quasi-particle production quasi-particle self interactions to understand thermalization

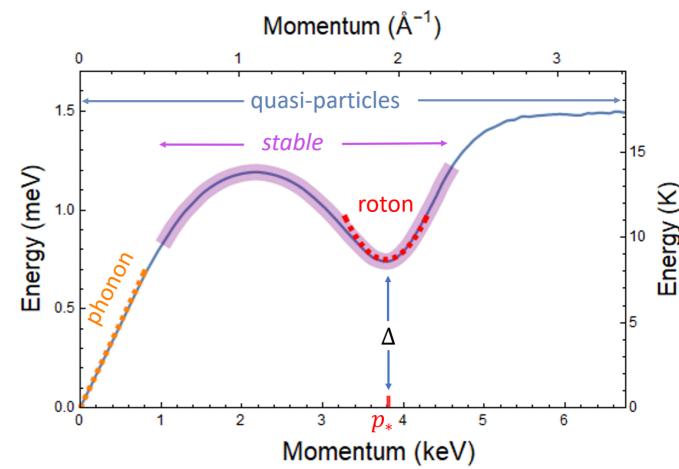
# Why it is challenging?



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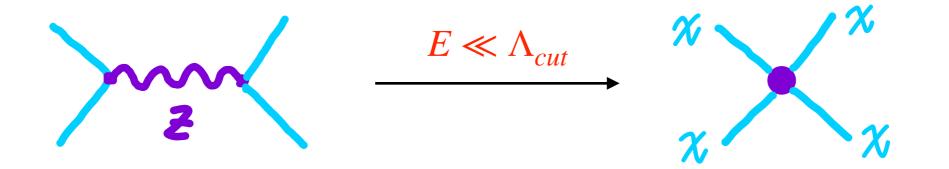
- inverse atomic spacing  $\Lambda_{cut} \sim \text{keV}$
- rotons, quasi-particles and heliums momentum  $p \gtrsim \Lambda_{cut}$



# Effective field theory

## Effective field theory

- find the relevant degrees of freedom and symmetry
- four-fermion interactions



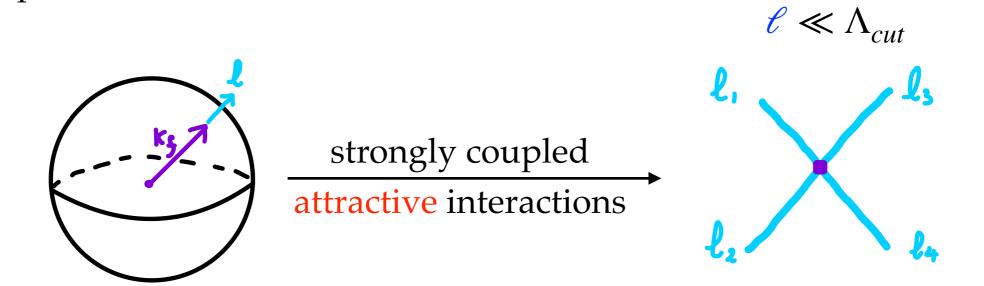
## Effective field theory

- find the relevant degrees of freedom and symmetry
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$$\underbrace{E \ll \Lambda_{cut}}_{Z}$$

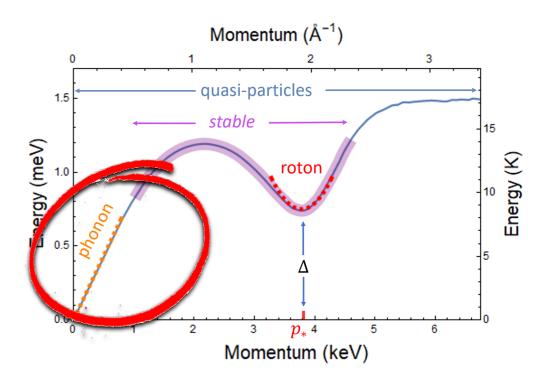
• BCS theory

perturbation around Fermi surface [J. Polchinski, 1999]



#### **Phonons** $\pi$ (p $\ll \Lambda$ ( ~ keV))

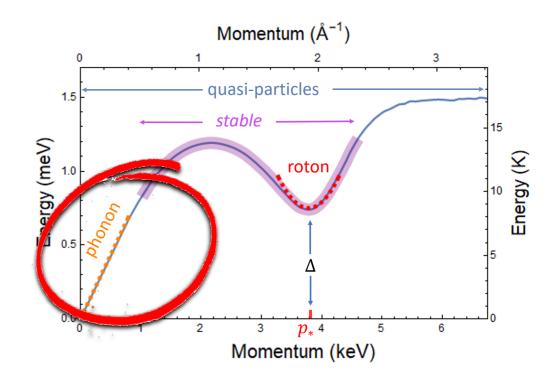
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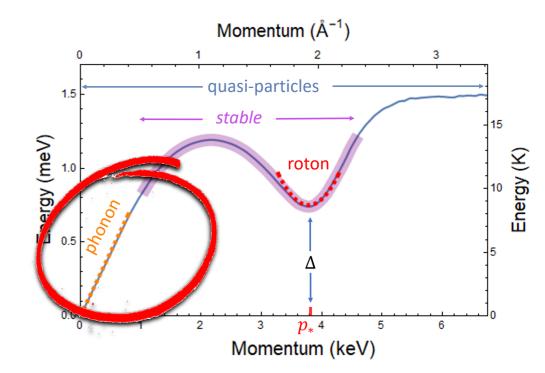
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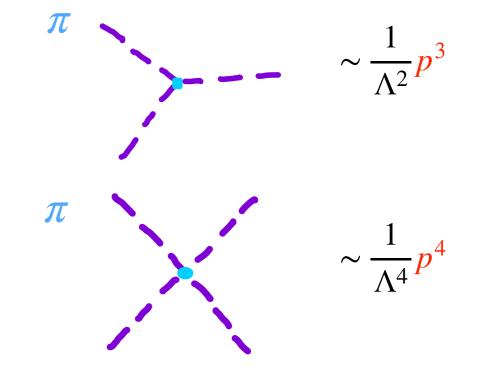


$$[p] = 1, \quad [t] = -1, [x] = -1, \quad [\pi] = 1$$

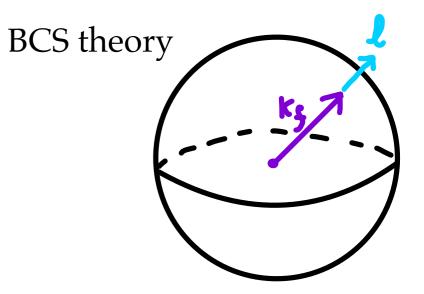
phonon interactions

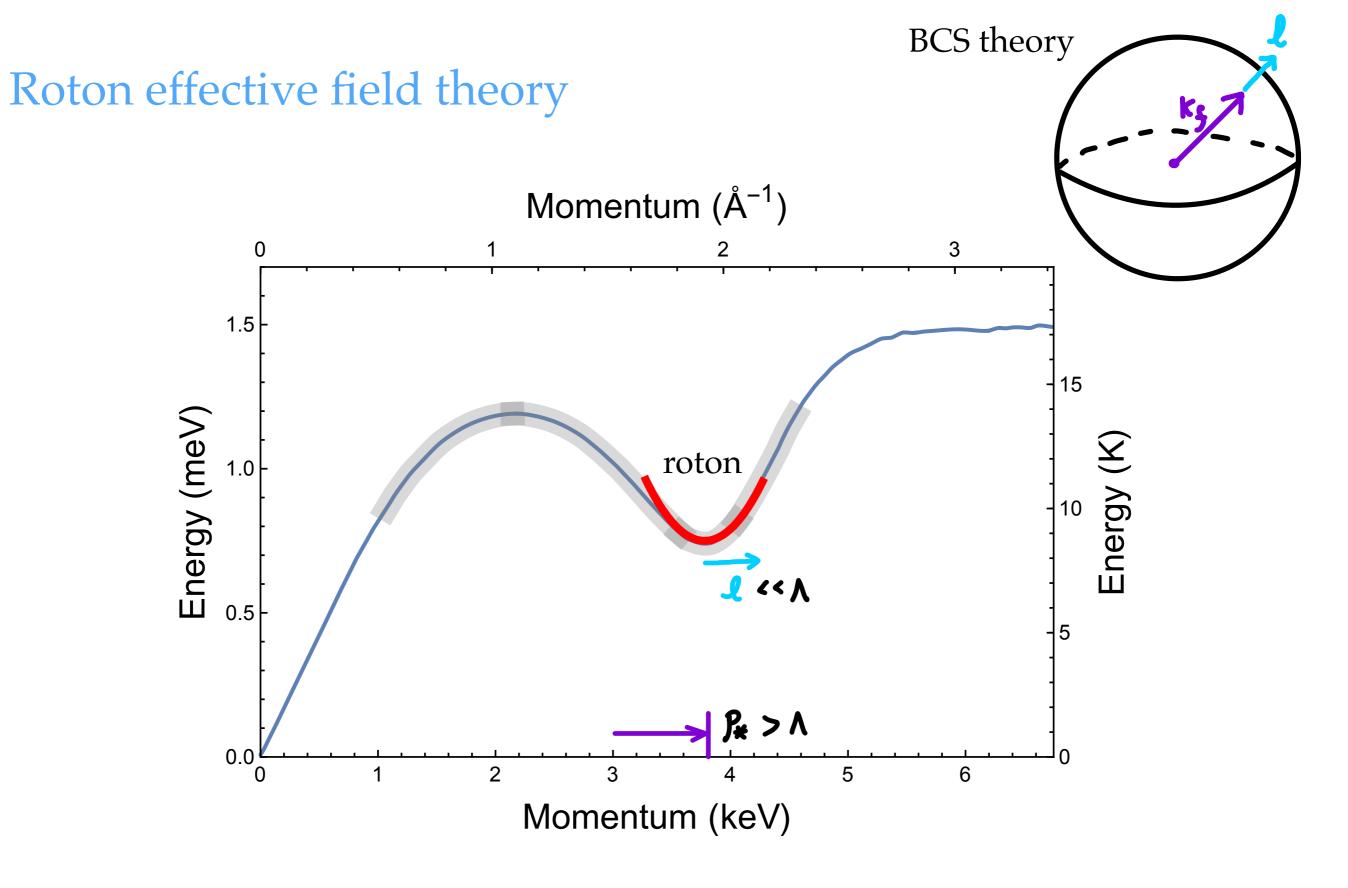
$$\mathscr{L}_{\rm ph} = -\frac{c_s^{3/2}}{2\Lambda^2} \dot{\pi} \partial_i \pi \partial_i \pi + \frac{g_3 c_s^{-1/2}}{6\Lambda^2} \dot{\pi}^3 + \mathcal{O}(\pi^4)$$

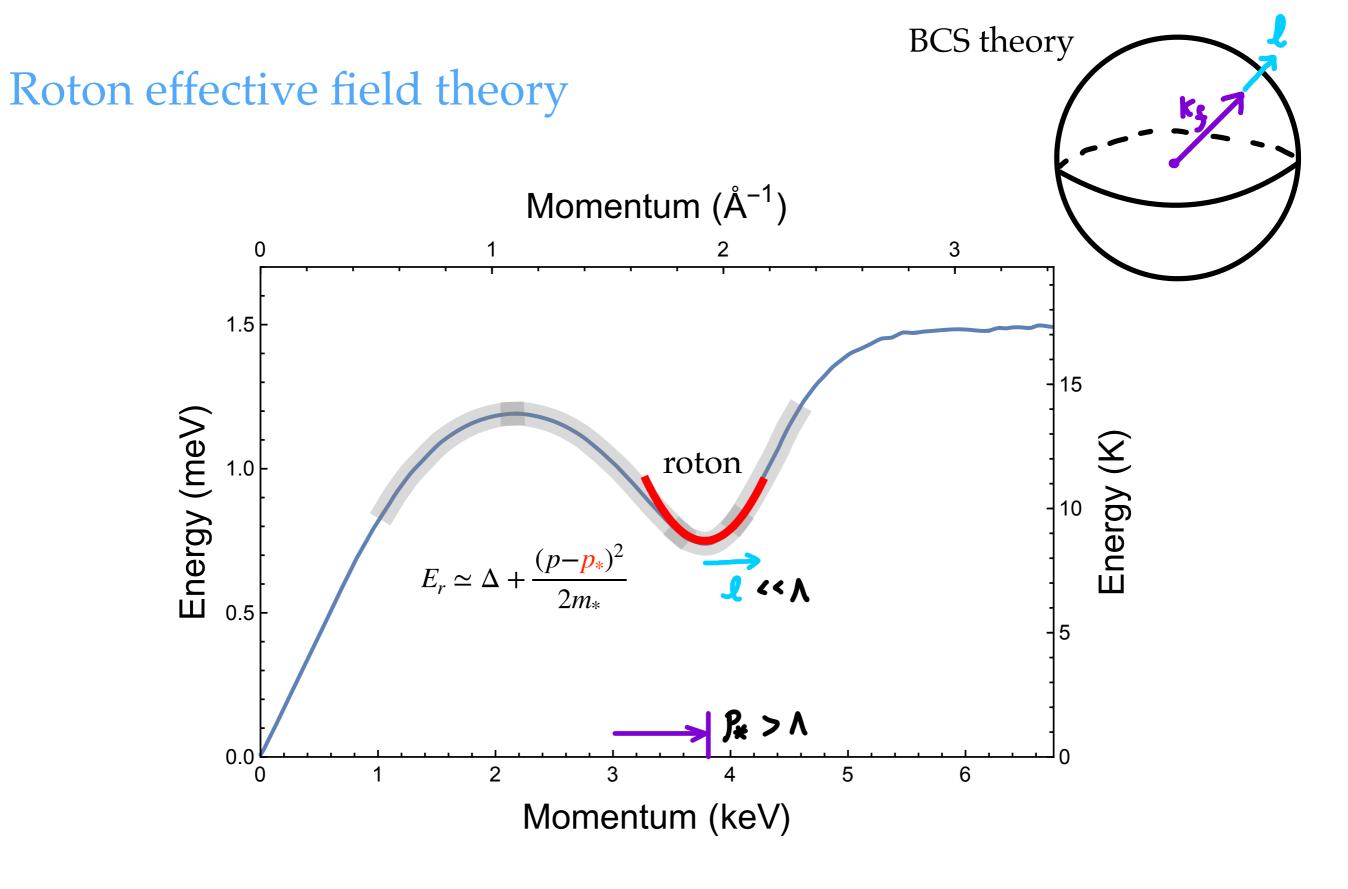


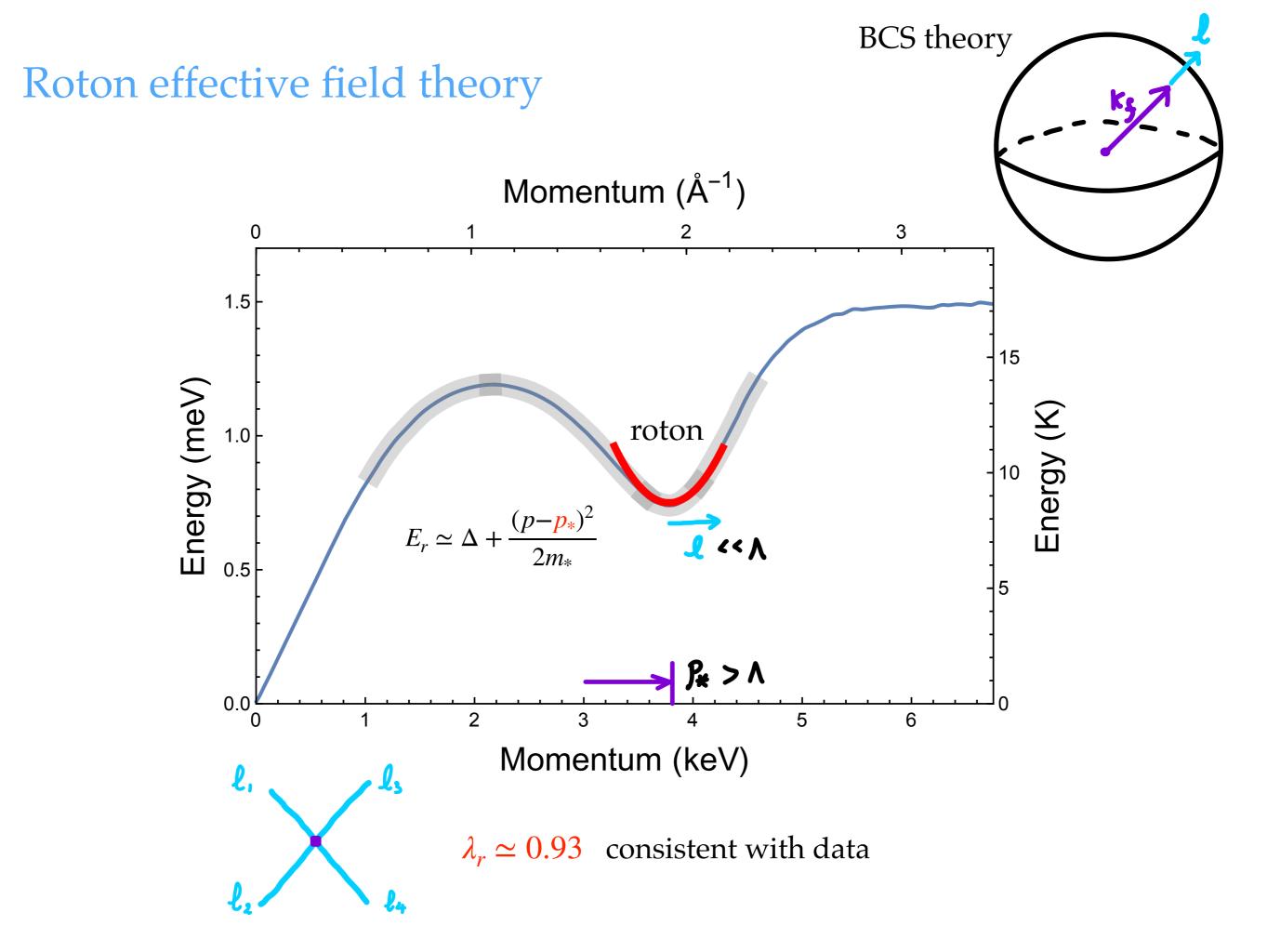


# Roton effective field theory

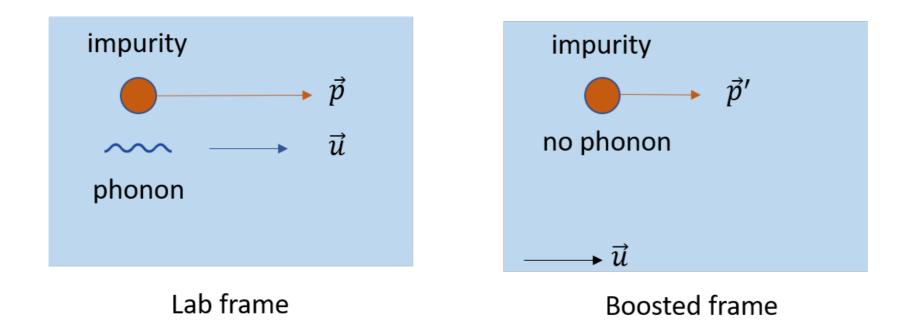




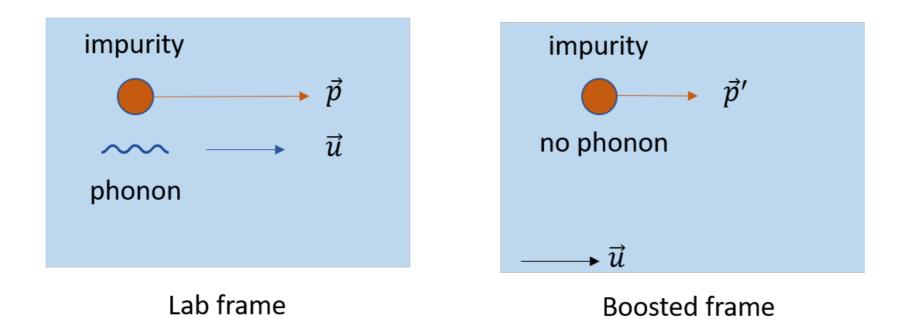




## Roton and phonon interactions



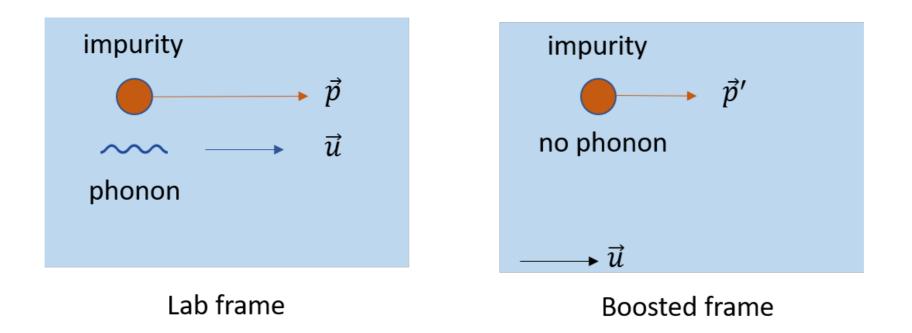
### Roton and phonon interactions



• roton (as impurity) and phonon interactions

$$V_{\text{ph-r}} = \epsilon_{boost} - \epsilon_{lab} + \boldsymbol{p} \cdot \boldsymbol{u}$$

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• This method can be applied to dark matter, quasiparticles, etc.

### Helium atom $\Phi_{\text{He}}(p \gg \text{keV})$

• U(1) symmetry  $\Phi_{\text{He}} \rightarrow e^{i\alpha} \Phi_{\text{He}}$ 

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 $\mathbf{p} = m_Q \mathbf{v} + \mathbf{k}$ ,  $\Lambda/m_Q \ll 1$ 

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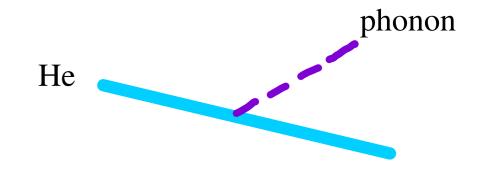
 $\mathbf{p} = m_Q \mathbf{v} + \mathbf{k}$ ,  $\Lambda/m_Q \ll 1$ 

• Helium currents  $J_{\text{He}}^0$  ,  $J_{\text{He}}^i$ 

 $J_{\text{He}}^{0} = \Phi^{\dagger}\Phi, \quad J_{\text{He}}^{i} = v^{i}\Phi^{\dagger}\Phi$ Phonon currents  $J^{0}, J^{i}$ 

$$J^0 = (\sqrt{\rho}/m_{\rm He}c_s)\,\dot{\pi} + \cdots,$$

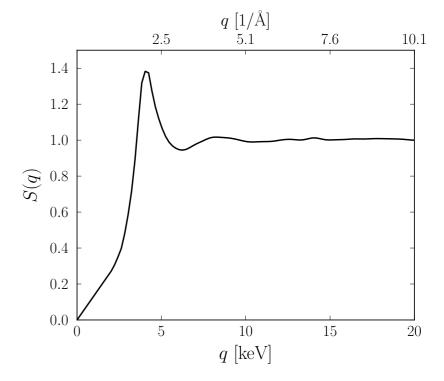
$$\mathscr{L}_{JJ} = \lambda_1 \frac{1}{m_{\text{He}}\Lambda} J^0 J^0_{\text{He}} + \lambda_2 \frac{m_{\text{He}}}{\Lambda^3} J^i J^i_{\text{He}}$$



#### $He \rightarrow He + quasi-particle from measurement$

• dynamical (static) structure function  $S(\mathbf{q}, \omega)$  is

directly measured in neutron scattering experiments



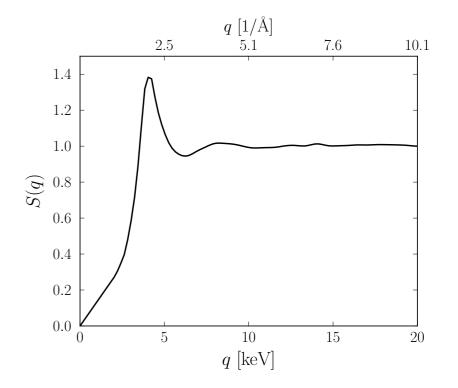
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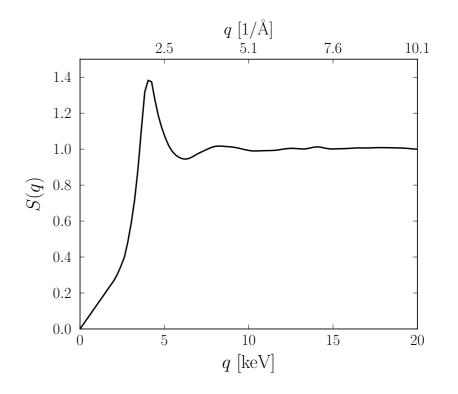
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• Helium decay to a quasi-particle,

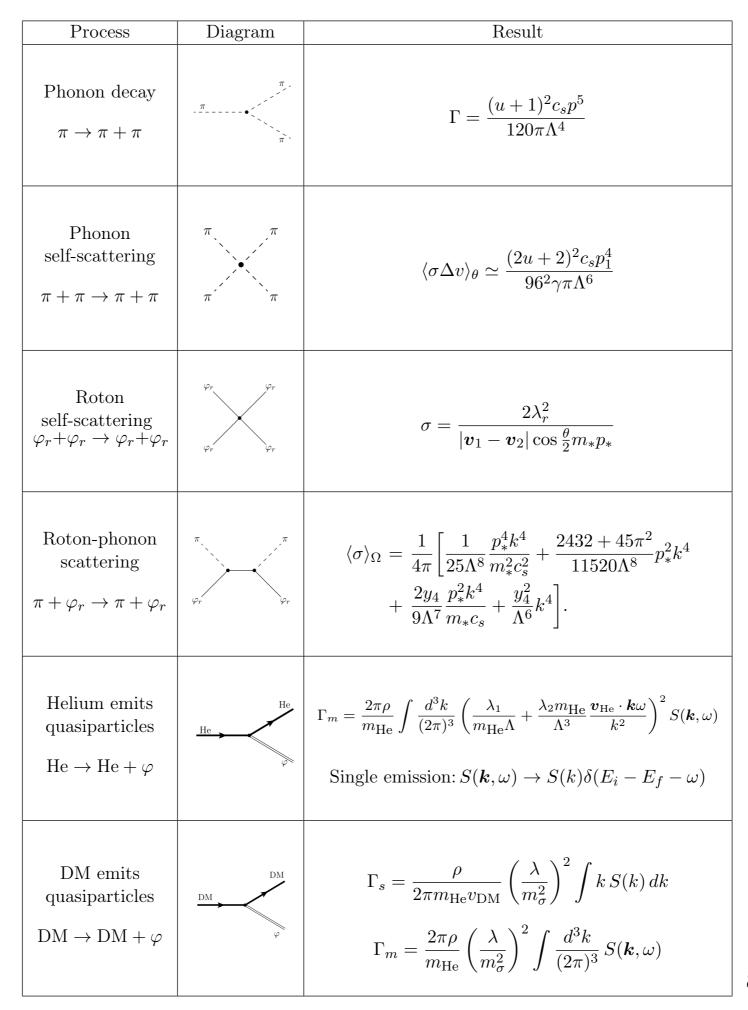
$$\mathscr{L}_{JJ} = \lambda_1 \frac{1}{m_{\text{He}}\Lambda} J^0 J^0_{\text{He}} + \lambda_2 \frac{m_{\text{He}}}{\Lambda^3} J^i J^i_{\text{He}}$$

amplitude  $\mathcal{M} \propto \lambda_1 J_{\text{He}}^0 \langle vac | J^0 | \varphi \rangle + \lambda_1 J_{\text{He}}^i \langle vac | J^i | \varphi \rangle$ 



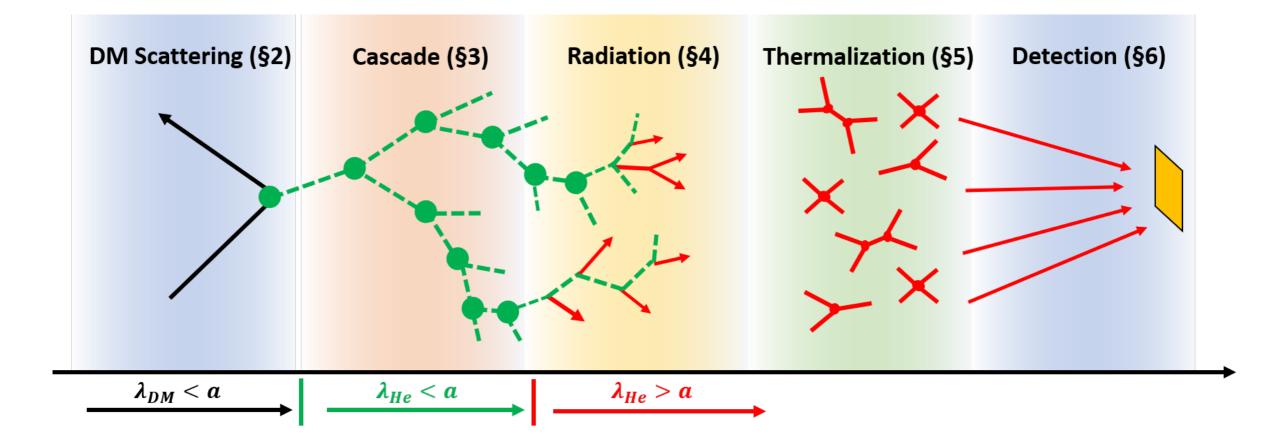
quasi-particle He

#### Rate summary

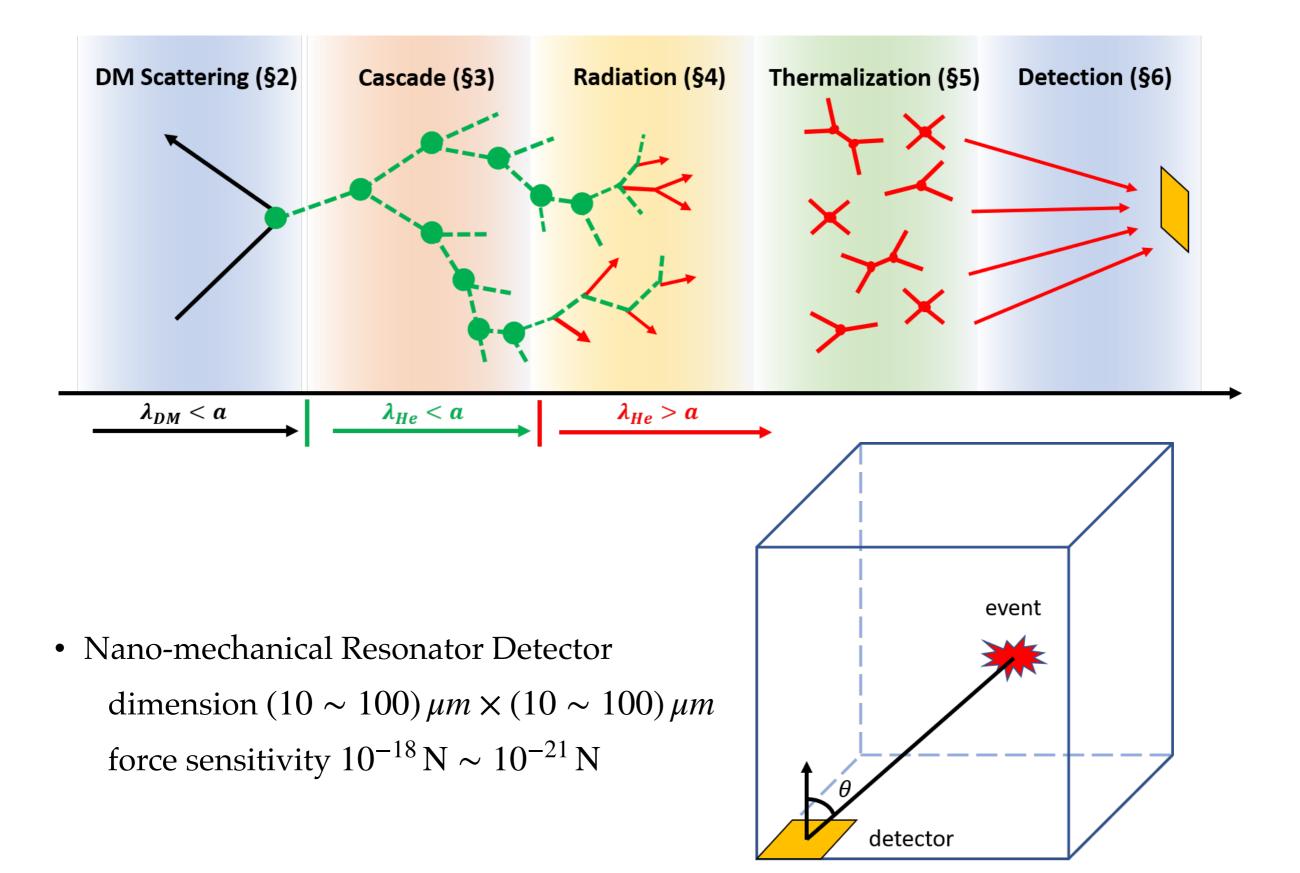


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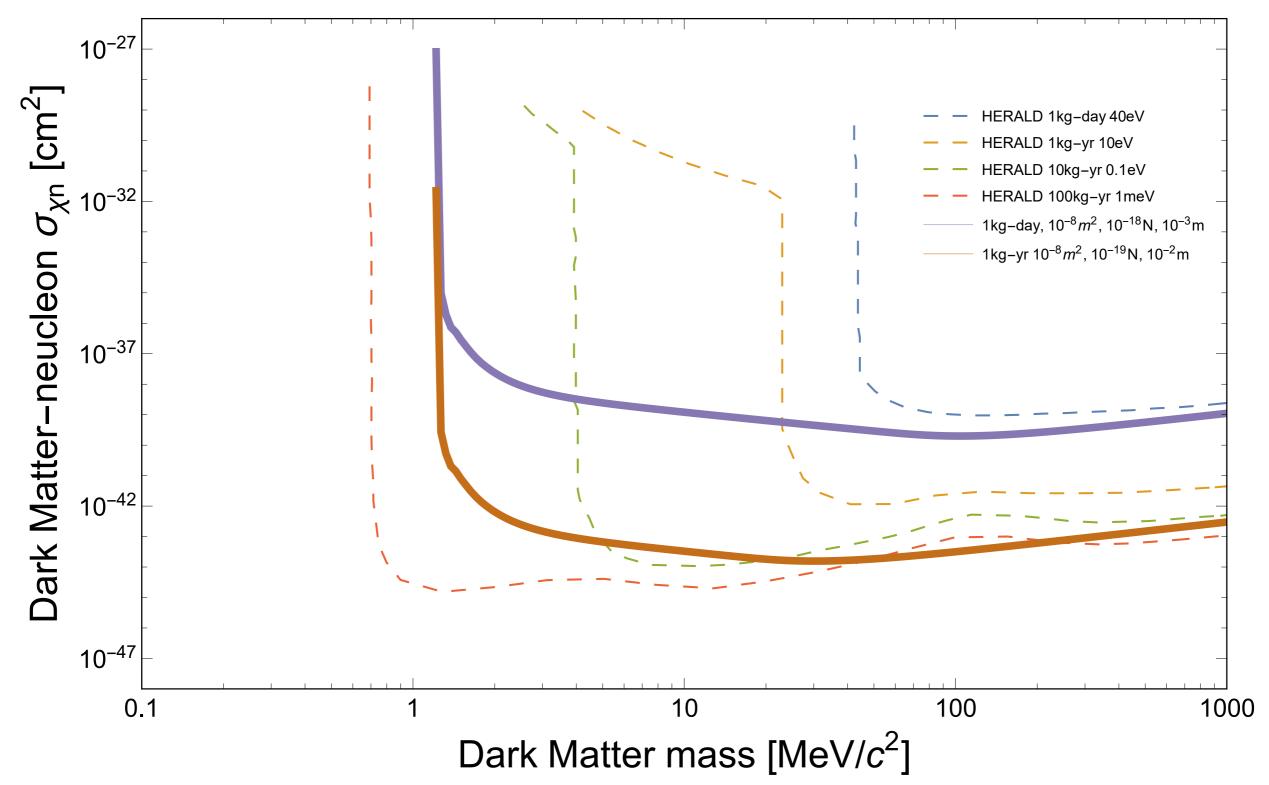
### Thermalization and Detection



### **Thermalization and Detection**



## Reach Plots



## Summary and future works

MeV - GeV dark matter theoretically and experimentally interesting *dark matter Helium*effective field theory quasi-particle interactions

Quasi-particles

## Summary and future works

