

Cryogenic Solid-State Detectors

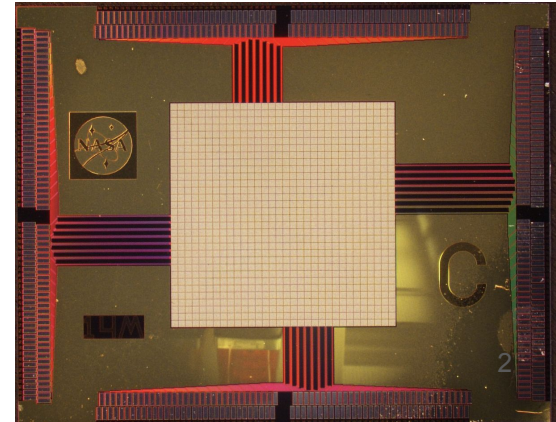
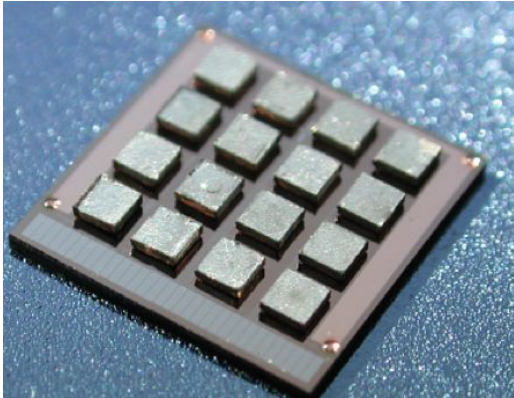
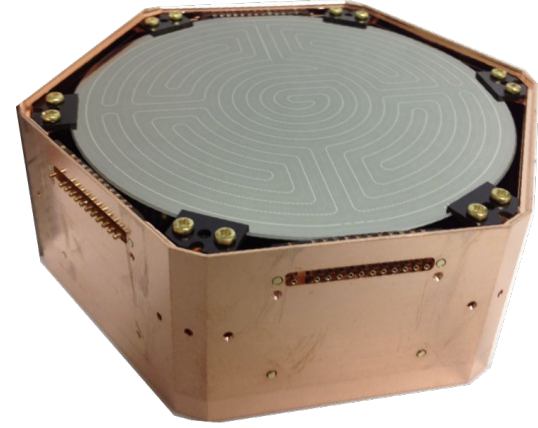
for Dark Matter Searches and beyond

Ziqing Hong, University of Toronto
33rd Rencontres de Blois on "Exploring the Dark
Universe"



Cryogenic Crystal Detectors are used in

- Particle Physics
 - Dark Matter Detectors
 - Neutrino Physics
 - Coherent Elastic Neutrino Nucleus Scattering (CEvNS) and neutrinoless double-beta decay ($0\nu\beta\beta$)
- Astrophysics
 - mm to gamma-ray energies



Disclaimer

- This is my personal biased view focusing on detectors working at < 1 K
- Lots of efforts ongoing with solid-state detectors for dark matter search efforts, a good fraction of them utilizing “cold” to battle against “noise”.
 - CCD-based detectors
 - High Purity Germanium detectors
 - ...
- Also lots of fantastic cryogenic detectors I will likely fail to include here
- Apologies for efforts that I missed in this discussion

Dark Matter Detection Channels

Hidden Sector Particles

ALPs

Axions

Sterile
 ν 's

WIMPs

feV peV neV μeV meV eV keV MeV GeV TeV PeV

Dark Matter Mass

10^{-46} 10^{-40} 10^{-34} 10^{-28} 10^{-22} 10^{-16} 10^{-10} 10^{-4} 10^2 10^5 10^5

Max Recoil Energy in Silicon [eV]

10^{26} 10^{23} 10^{20} 10^{17} 10^{14} 10^{11} 10^8 10^5 10^2 10^{-1} 10^{-4}

Dark Matter Particle Density per Liter

Dark Matter Detection Channels

Hidden Sector Particles

ALPs

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Sterile
 ν 's

WIMPs

feV peV neV μeV meV eV keV MeV GeV TeV PeV

Dark Matter Mass

10^{-41} 10^{-35} 10^{-29} 10^{-23} 10^{-17} 10^{-11} 10^{-5} 10^0 10^1 10^1 10^1

Max Electron Recoil Energy [eV]

10^{26} 10^{23} 10^{20} 10^{17} 10^{14} 10^{11} 10^8 10^5 10^2 10^{-1} 10^{-4}

Dark Matter Particle Density per Liter

Electron
Recoils

Nuclear
Recoils

Dark Matter Detection Channels

Hidden Sector Particles

ALPs

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Sterile
 ν 's

WIMPs

feV peV neV μeV meV eV keV MeV GeV TeV PeV

Dark Matter Mass

10^{-41} 10^{-35} 10^{-29} 10^{-23} 10^{-17} 10^{-11} 10^{-5} 10^0 10^1 10^1 10^1

Max Electron Recoil Energy [eV]

10^{-10} 10^{-9} 10^{-8} 10^{-7} 10^{-6} 10^{-5} 10^{-4} 10^{-3} 10^{-2} 10^{-1} 10^0

Mean Distance Between Particles [m]

10^{12} 10^9 10^6 10^3 10^0 10^{-3} 10^{-6} 10^{-9} 10^{-12} 10^{-15} 10^{-18}

Dark Matter Particle Wavelength [m]

Coherent/Resonant
Detection

Electron
Recoils

Nuclear
Recoils

Dark Matter Detection Channels

Hidden Sector Particles

ALPs

Axions

Sterile
 ν 's

WIMPs

feV peV neV μeV meV eV keV MeV GeV TeV PeV

Dark Matter Mass

10^{-41} 10^{-35} 10^{-29} 10^{-23} 10^{-17} 10^{-11} 10^{-5} 10^0 10^1 10^1 10^1

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Mean Distance Between Particles [m]

10^{12} 10^9 10^6 10^3 10^0 10^{-3} 10^{-6} 10^{-9} 10^{-12} 10^{-15} 10^{-18}

Dark Matter Particle Wavelength [m]

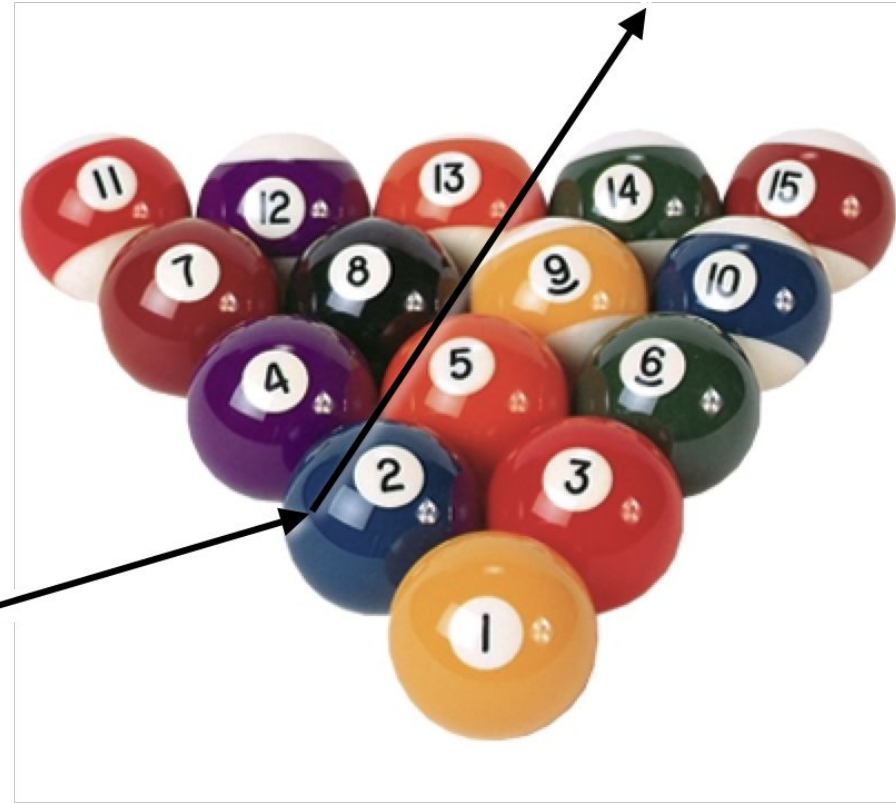
Coherent/Resonant
Detection

Cryo
Detectors

Liquid Noble
experiments

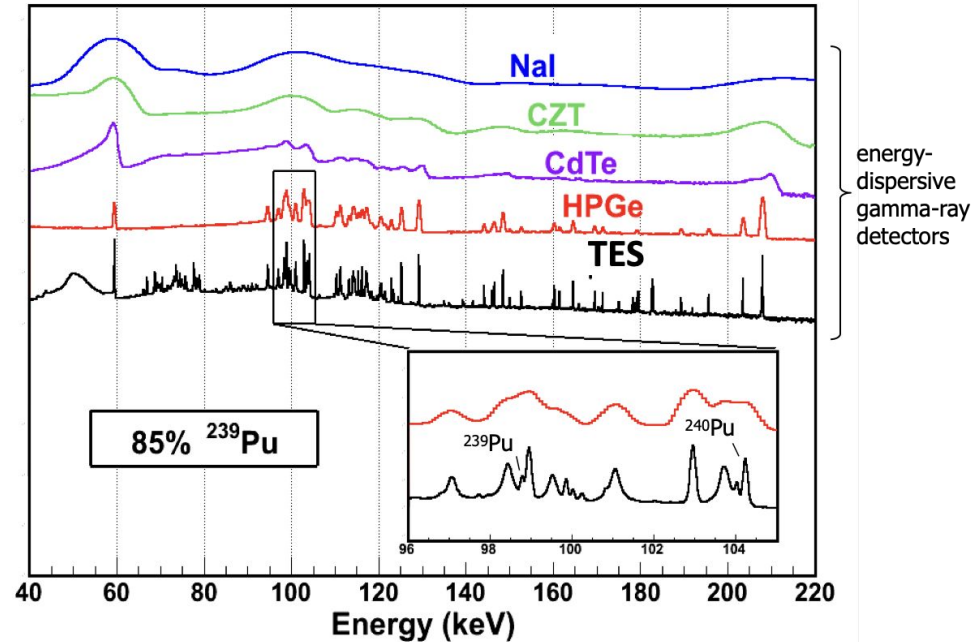
Dark matter direct detection

- Dark matter particle interaction deposits energy
- Detectors measure energy in forms of
 - Ionization → Charge
 - Scintillation → Light
 - **Heat → Phonons**
- Stealth signal calls for sensitive detectors

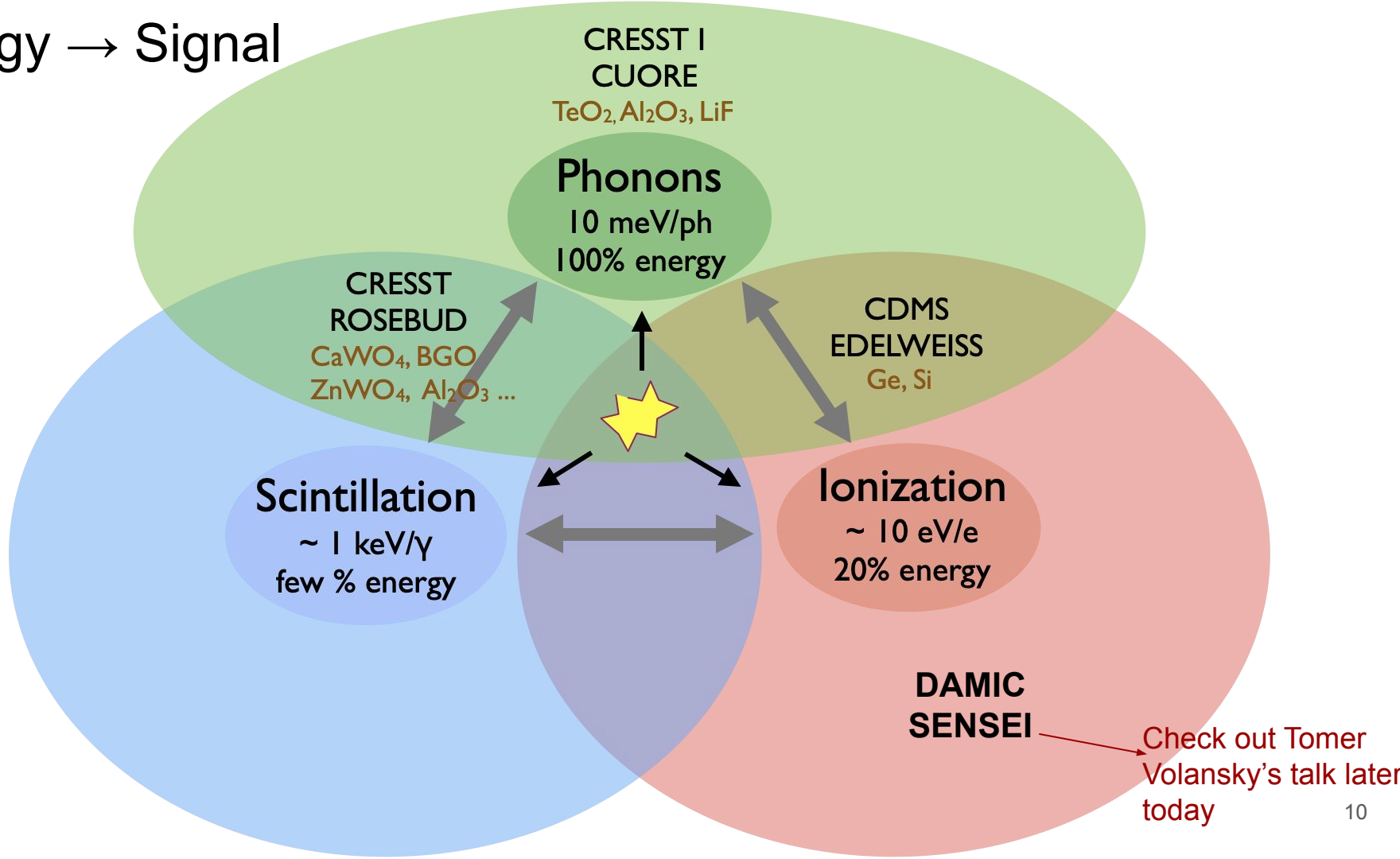


Why Use Cryogenic Crystal Detectors?

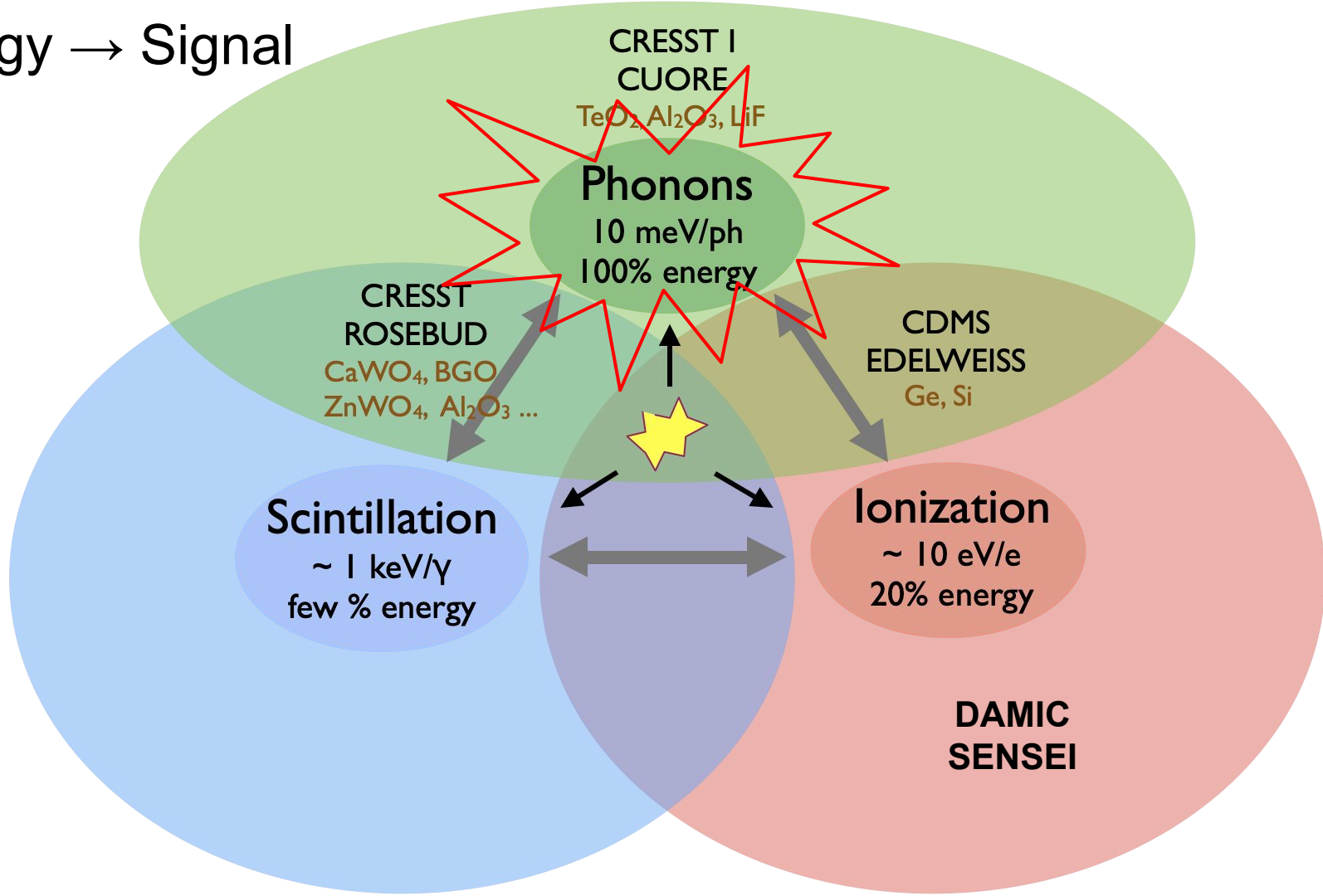
- Cryogenic detectors can provide a unique combination of energy sensitivity, low threshold and efficiency
- Exploiting the fundamental idea of lower temperature
 - Lower amount of random motions
 - Lower noise
 - Better energy sensitivity & low threshold
- Well matched to DM detection requirements
- CEvNS and $0\nu\beta\beta$ experiments share the same needs, thus cryogenic detectors are often developed by these fields jointly



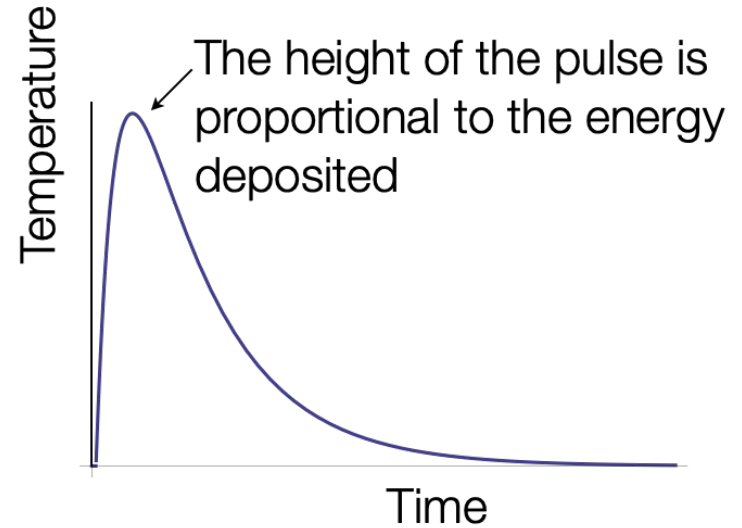
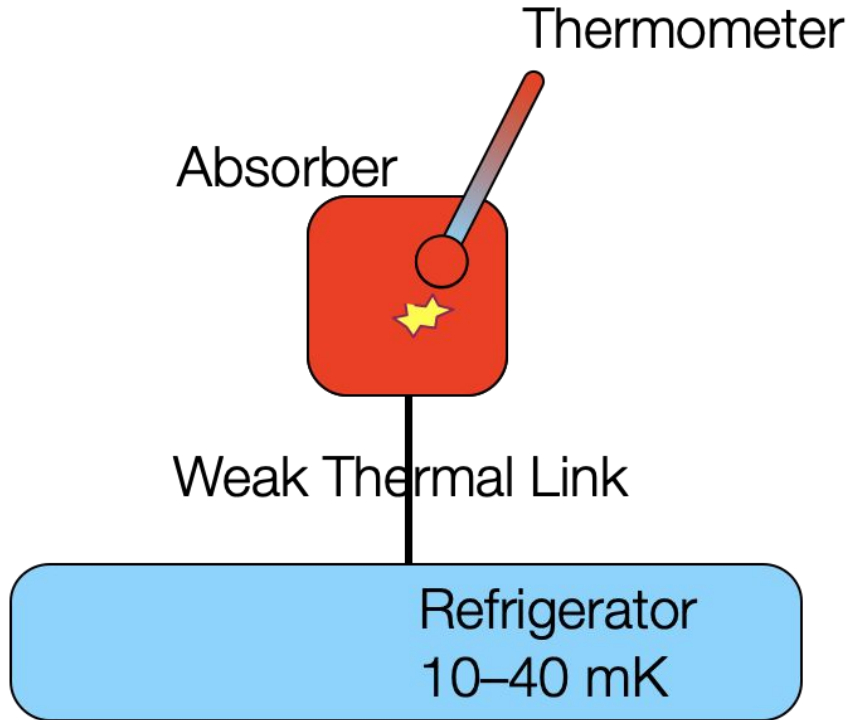
Energy \rightarrow Signal



Energy \rightarrow Signal



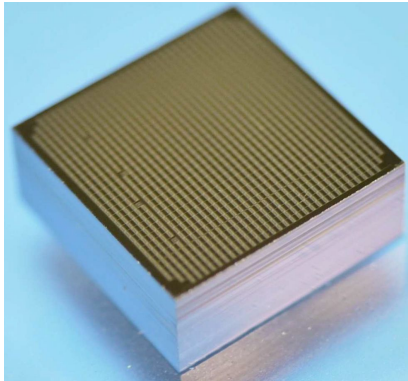
Cryogenic Phonon Detectors



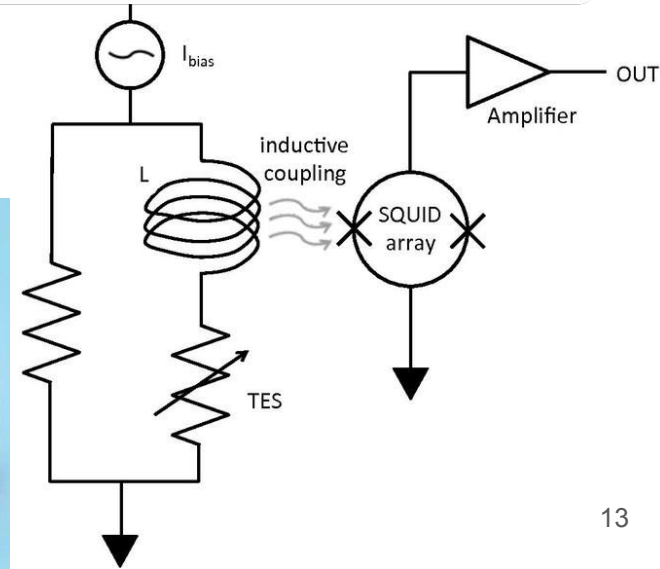
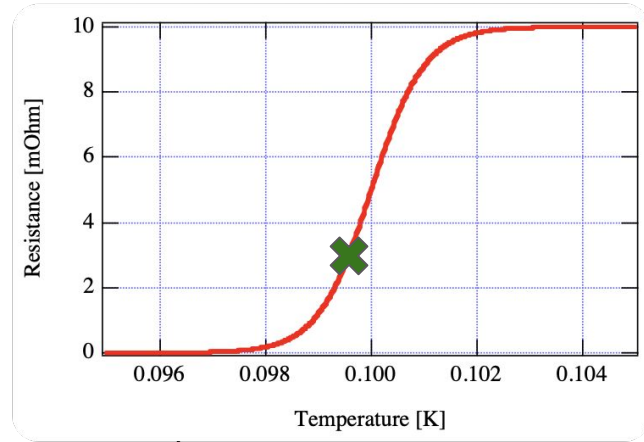
Everybody has their choice of the favourite thermometer!

Transition Edge Sensor

- Metal films, tuned to have suitable superconducting transition temperatures
- Operating in the middle of its transition
- Heat warms up the sensor
→ Increase in resistance
- Often read out by Superconducting QUantum Interference Devices (SQUIDs)
- High resistivity films can also be readout with FETs (schematics in next slide)



Transition Edge Sensor (TES)

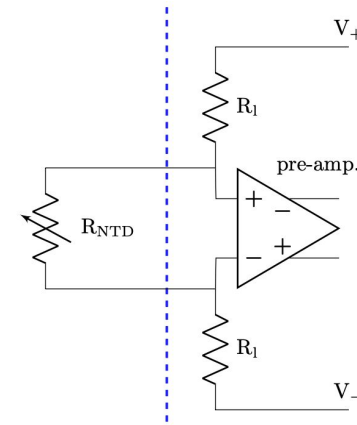
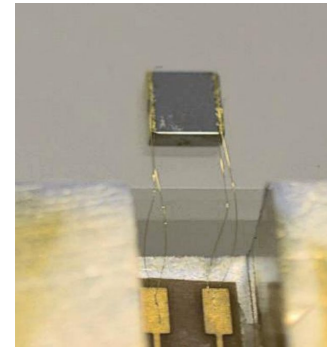
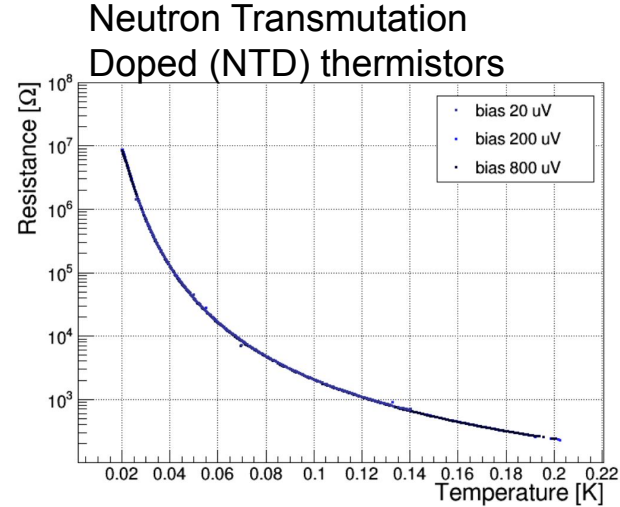


Neutron Transmutation Doped (NTD) thermistors

- Doped germanium/silicon chips
- Resistance follows Efros-Shklovskii law:

$$R = R_0 e^{\sqrt{T_0/T}}$$

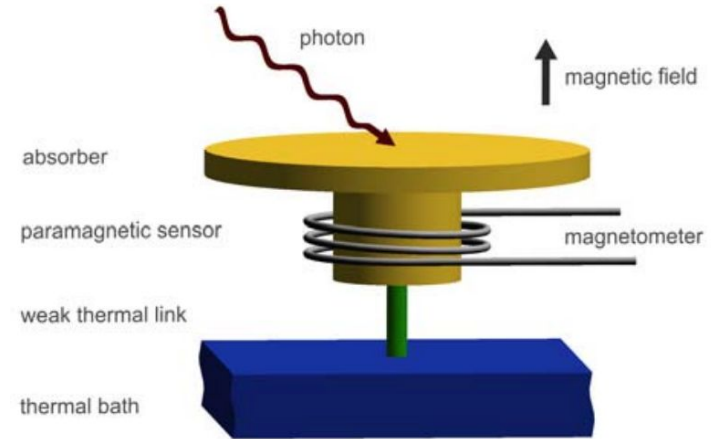
- Taking advantage of the steep slope at low temperature
- Also comes with high dynamic range
- Readout with FETs, operating at room temperature or in cold



Metallic magnetic calorimeter (MMC)

- Paramagnetic sensor positioned in weak magnetic field
- Heat changes its induced magnetic field
- Readout by SQUIDs as magnetometer

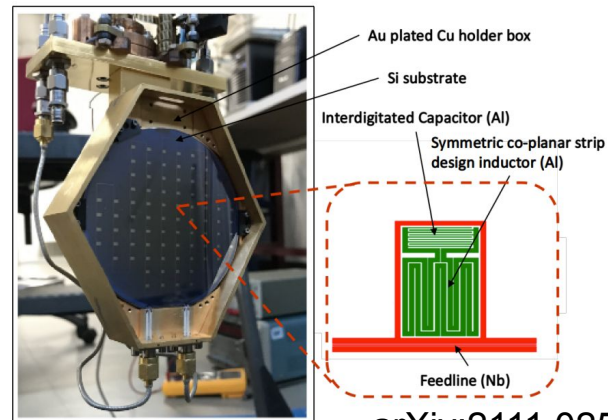
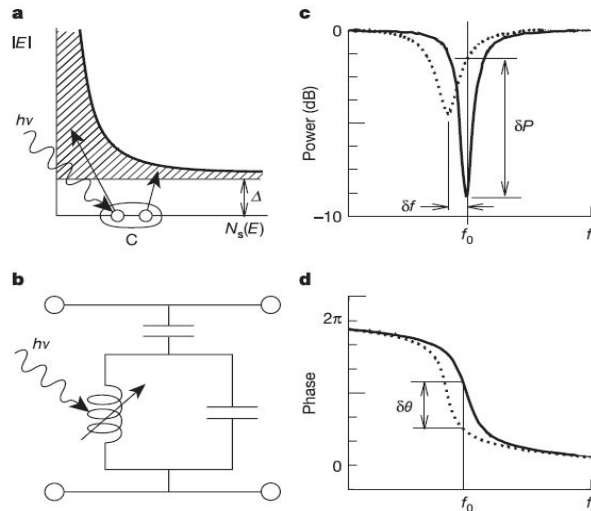
Metallic magnetic calorimeter (MMC)



TASC.2009.2012724

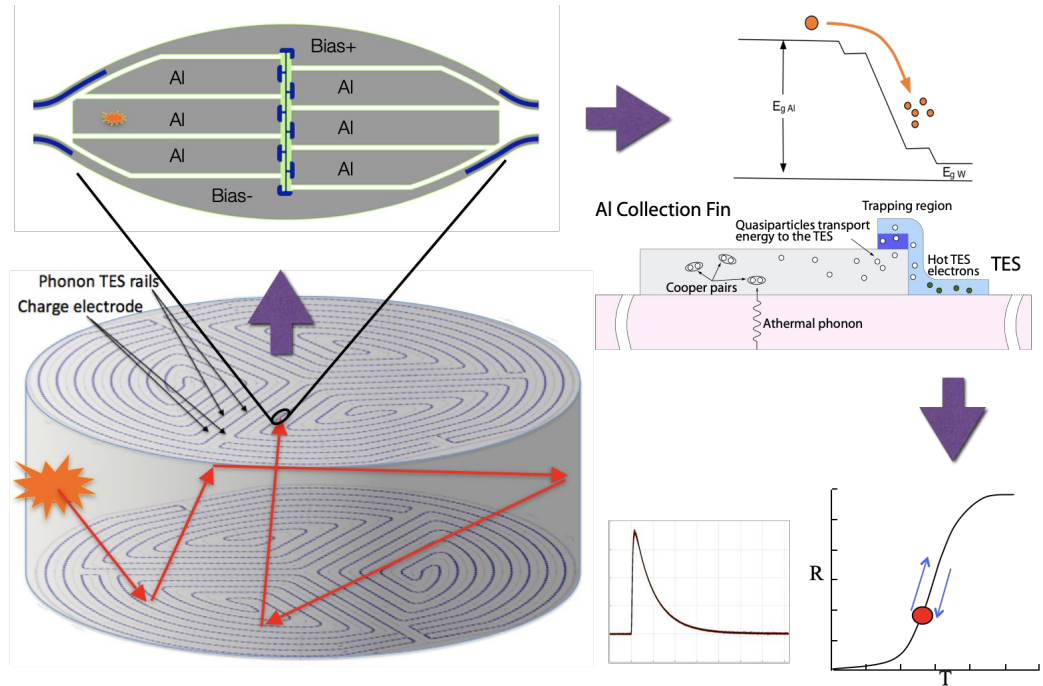
Microwave Kinetic Inductance Detectors (MKIDs)

- Resonators made of superconducting metal films
- Resonance frequency and phase response depending on its temperature
- Radio-Frequency (RF) Readout system
- Intrinsic capability for multiplexing



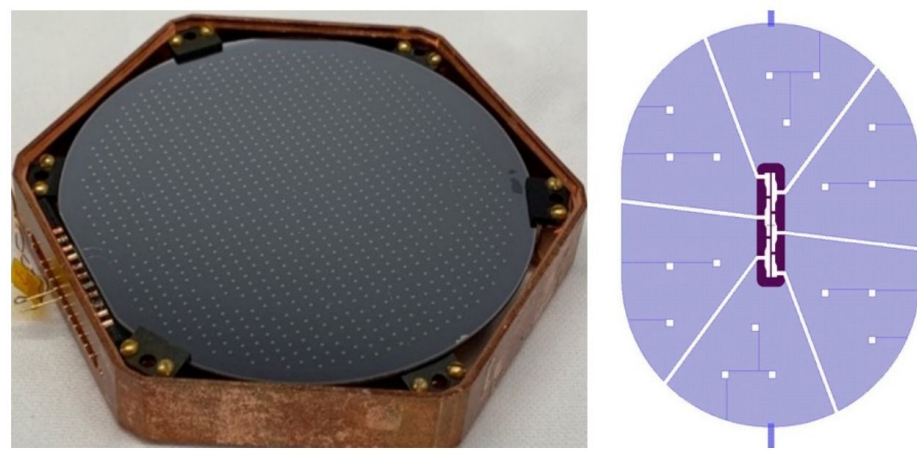
Funneling Energy to the Phonon Sensors (QETs)

- Quasiparticle-trap-assisted electrothermal-feedback TES (QETs)
- Targeting at high energy **athermal phonons** before they down-convert
- Utilizing superconducting “fins” to trap phonons and funnel them to TES
- Fast ($O(10\text{ us})$) detector response (in cryogenic detector sense)

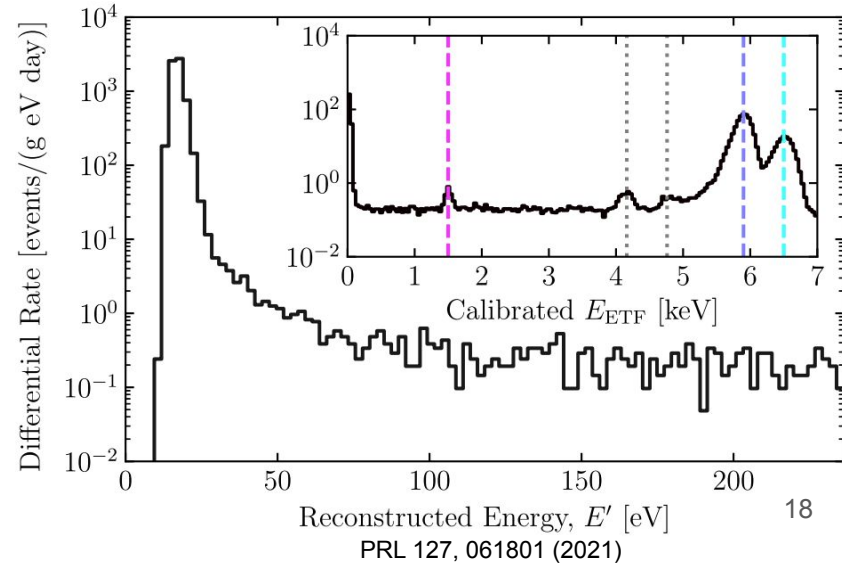


Cryogenic Photon Detector (CPD)

- 10 gram, silicon, QET-based
- One channel construction, easy to operate
- 3.9 eV phonon resolution
- Works great in both sensing photons and DM direct detection
- Future development by TESSERACT collaboration
 - See D. Mckinsey's presentation on Wed.

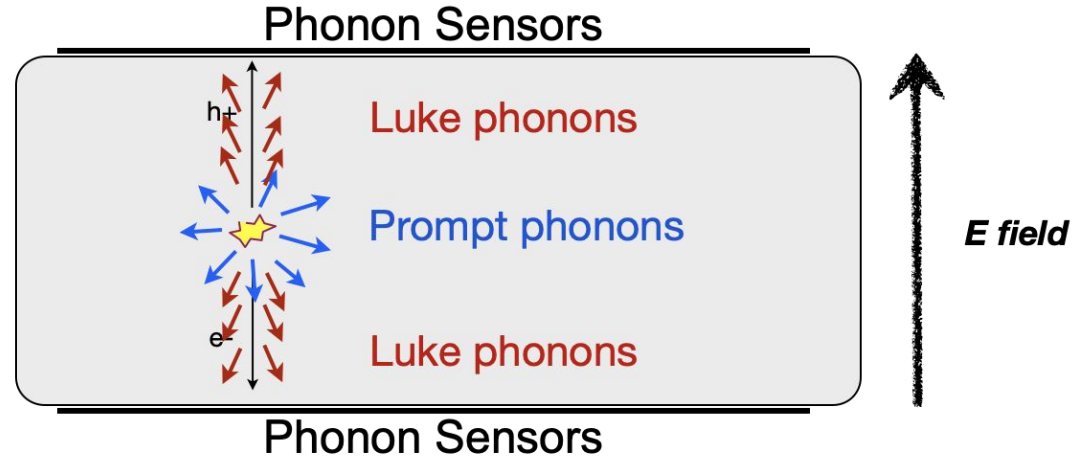


APL 118, 022601 (2021)



Internal amplification - the NTL effect

Phonon sensors measure amount of charge produced:
Phonon-based charge amplification!

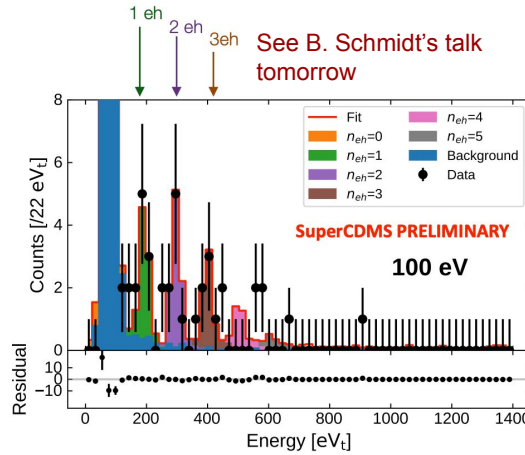
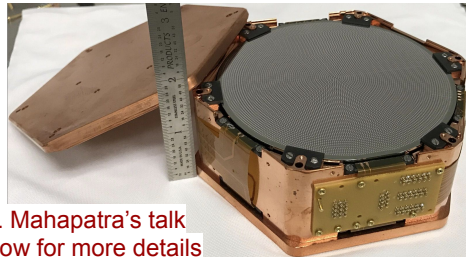
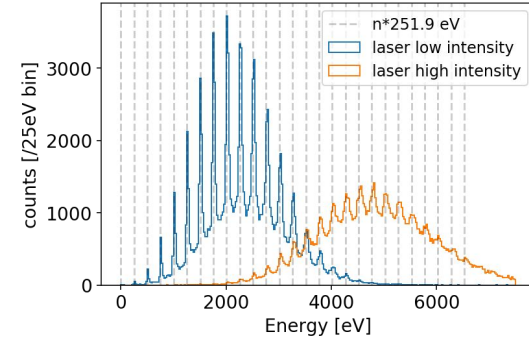
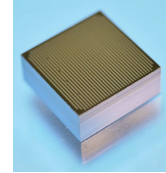


$$\begin{aligned}\text{Phonon energy} &= E_{\text{recoil}} + E_{\text{Luke}} \\ &= E_{\text{recoil}} + n_{\text{eh}} e^- \Delta V\end{aligned}$$

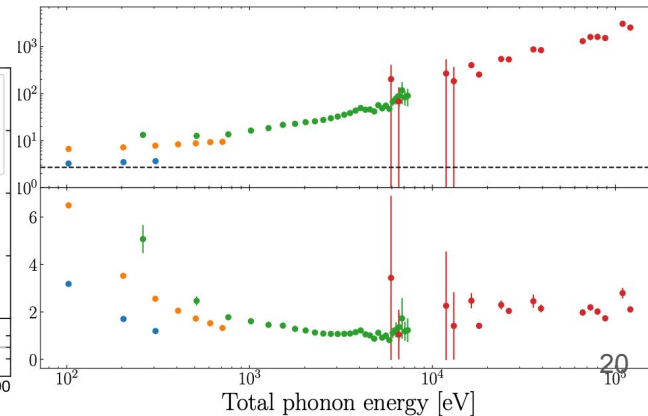
SuperCDMS High Voltage eV sensitivity Detector (HVeV)

PRD 104, 032010 (2021)

- 1 gram, silicon, QET-based, 2.7 eV resolution
- Can apply O(100 V) for NTL boost or operate at 0 V as pure phonon sensors
 - Particle identification by statistics
- Quantized electron-hole pair sensitivity for both ER and NR
- Scaling it up to 1 kg SuperCDMS HV detector



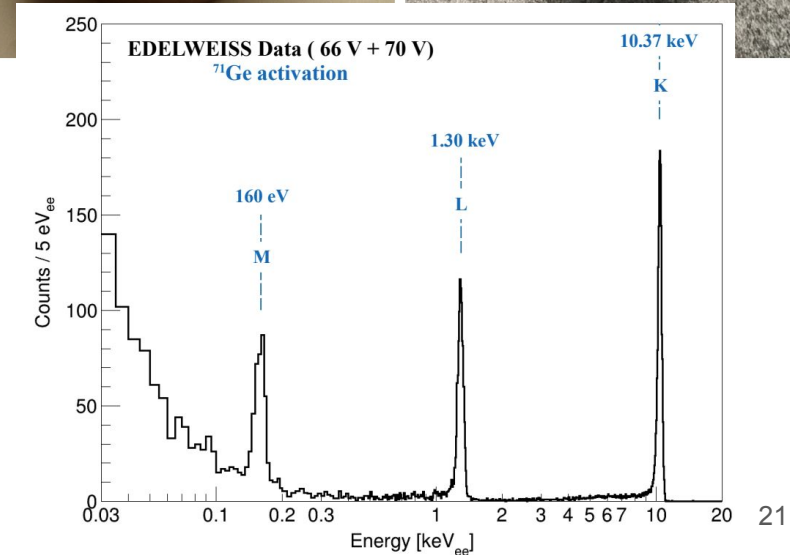
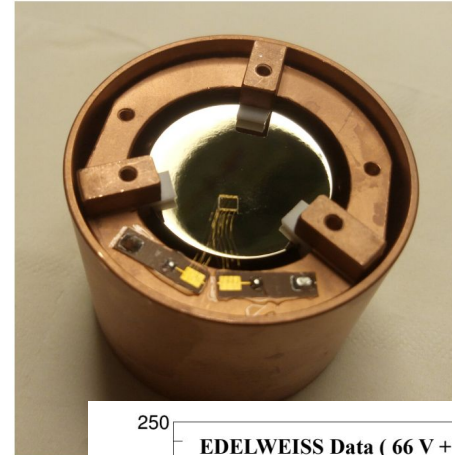
- baseline - OF
- Laser 100 V - OF
- Laser 100 V - MF
- Laser 250 V - MF
- ⁵⁵Fe 0 V - 70 V - MF



See R. Mahapatra's talk tomorrow for more details

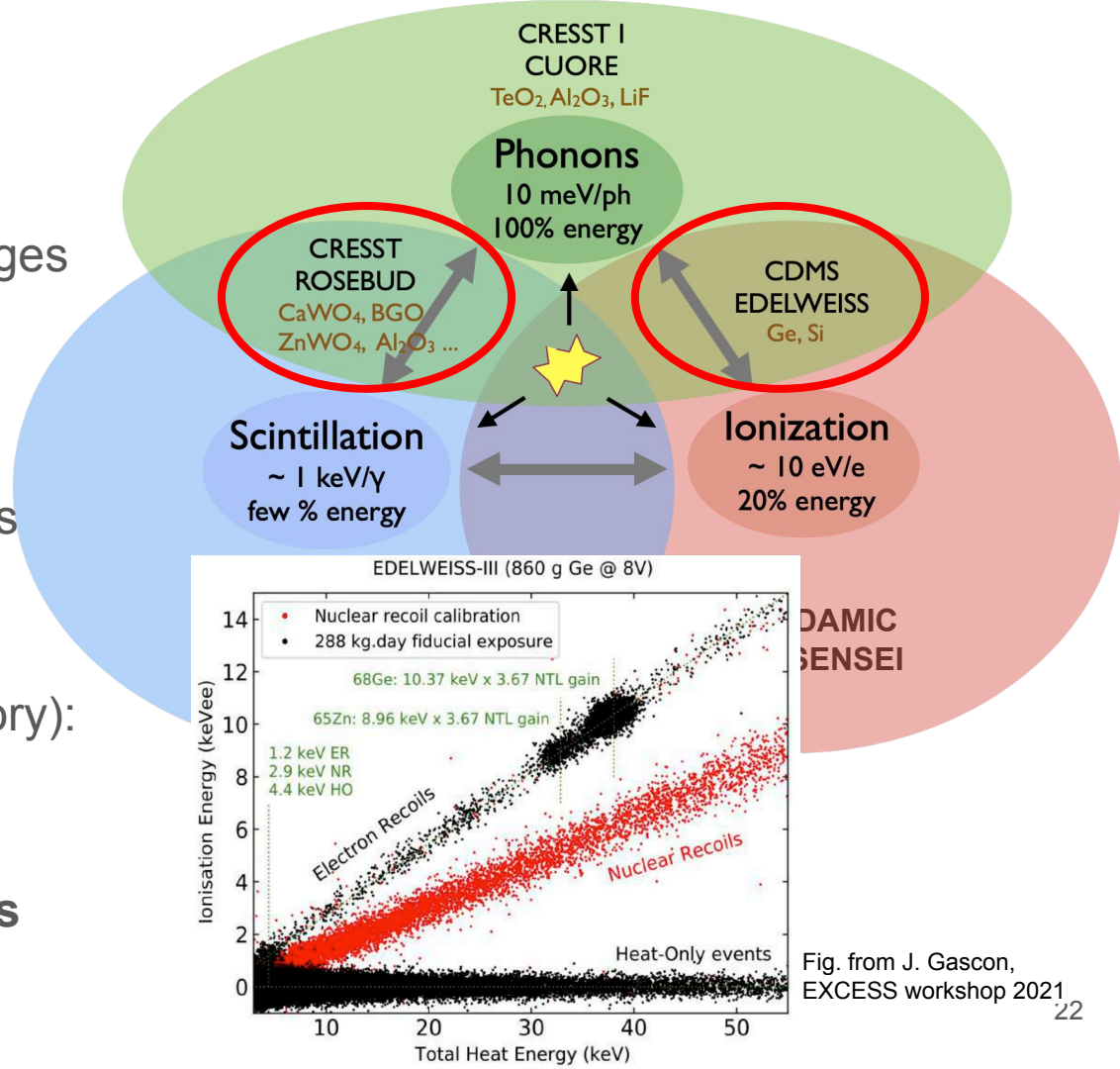
Edelweiss: RED20 and RED30

- 33 gram, germanium, NTD-based
- 18 eV phonon resolution
- RED20 operated with no E-field
- RED30 employs NTL gain to boost signals
 - With a planar electrode design
 - 8 eVee at 160 eV, consistent with Fano fluctuations
- See E. Guy's presentation on Wed.



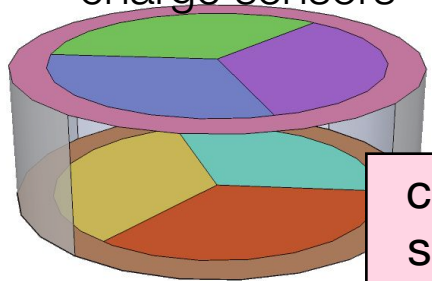
Adding information for particle identification

- Electron recoil (ER): lots of charges and scintillation lights
 - Source: photons, electrons alphas, ER DM particles
- Nuclear recoil (NR): less charges and scintillation lights
 - Source: neutrons, WIMPs
- Heat only (newly realized category): well... heat (phonon) only...
 - Source: unknown
- **Adding information sometimes degrades phonon information**
 - Careful trade-off needed



SuperCDMS: TES+charge

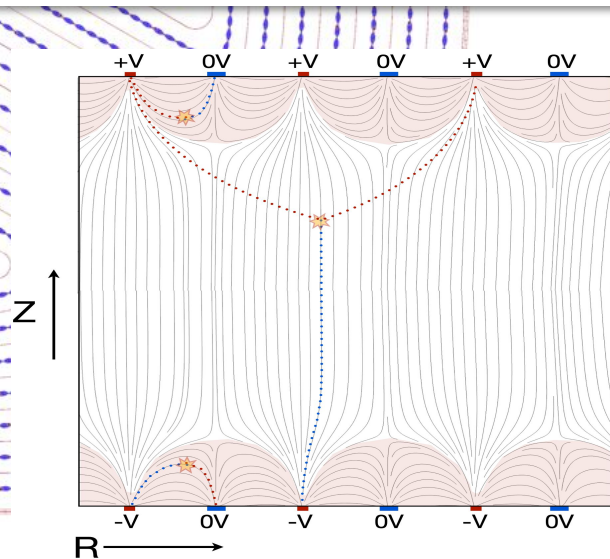
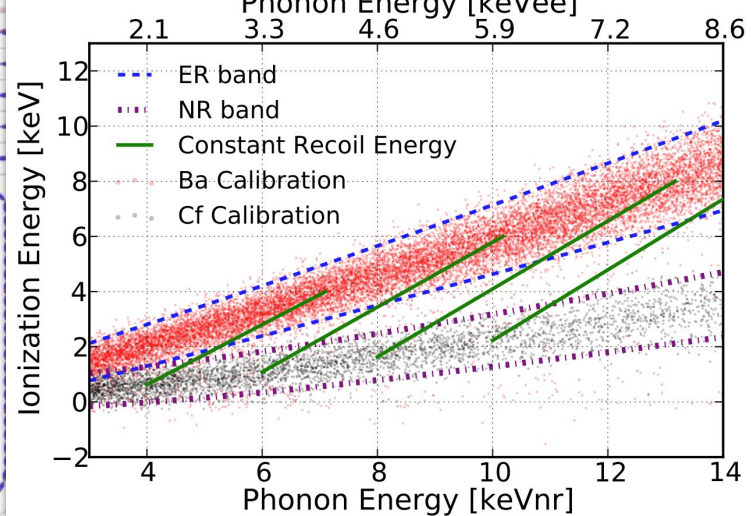
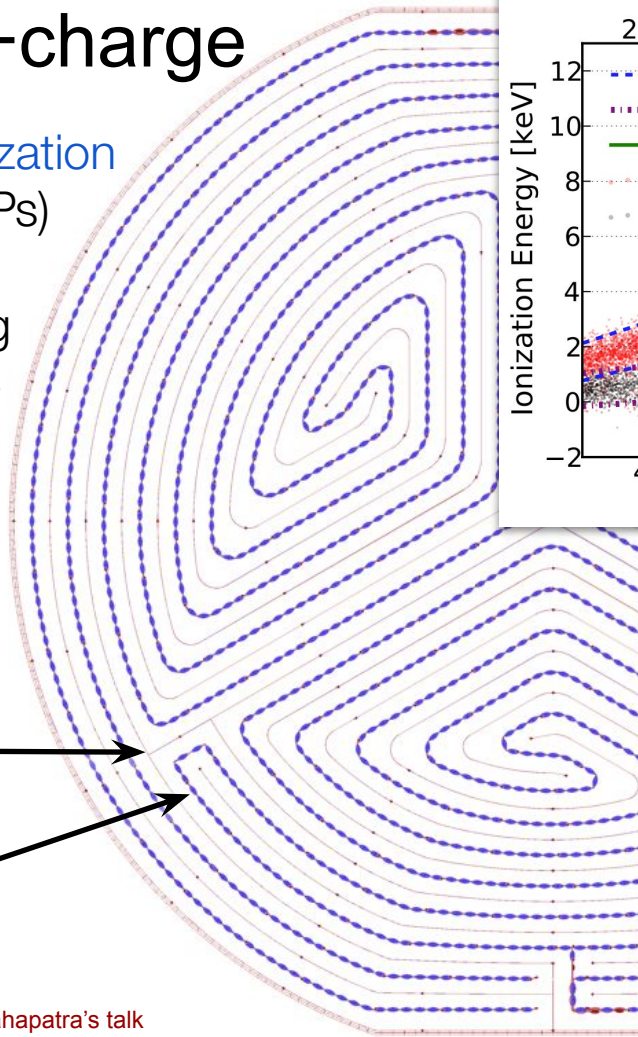
- Interleaved **z-sensitive ionization** and **phonon** detectors (iZIPs)
- Silicon and Germanium
 - Next generation ~ 1 kg
- 8-12 phonon channels + 4 charge sensors



charge sensor
(biased)

phonon sensor
(grounded)

See R. Mahapatra's talk tomorrow for more details



EDELWEISS: NTD/TES + charge

- Fully Inter-Digitized (FID) detector

- 800 g germanium
- 2 NTD + 4 charge channels

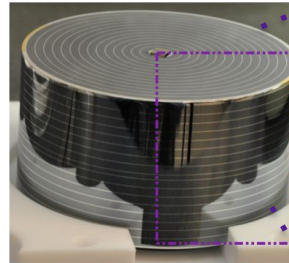
- NbSi₂₀₉

- 200 g germanium
- 2 TES channel + 2 charge planar electrodes
- TES vs NTD tests origin of the HO events
- Charge channels help rejecting HO events

- FID38 & PL38

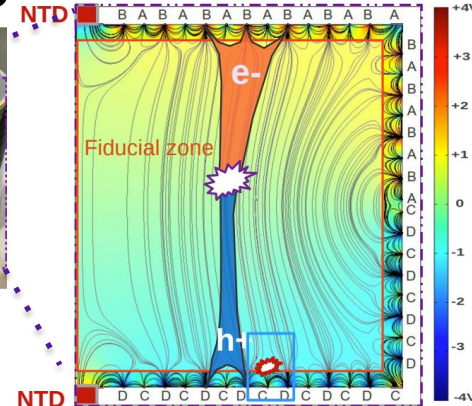
- Next-stage 38 gram germanium detectors, with NTD+charge
- Expect better phonon and charge resolutions with cryogenic HEMTs

Height : 4cm

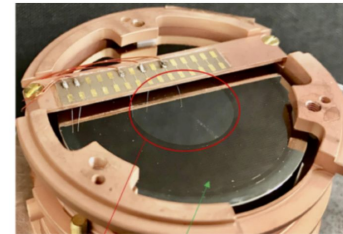
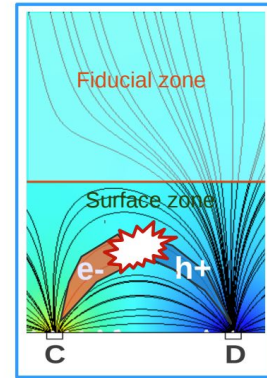


Width : 7cm

See E. Guy's presentation on Wed. for more details

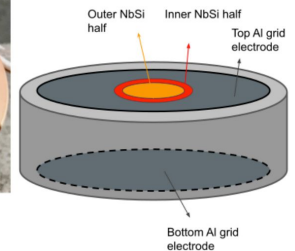


PRD 97 (2018) 022003



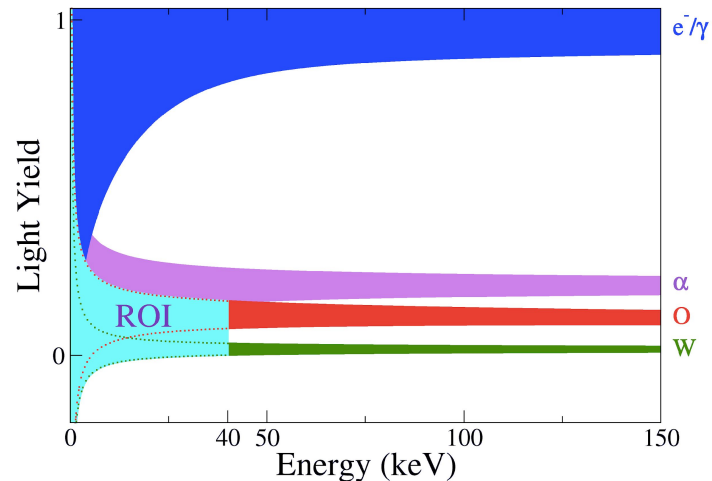
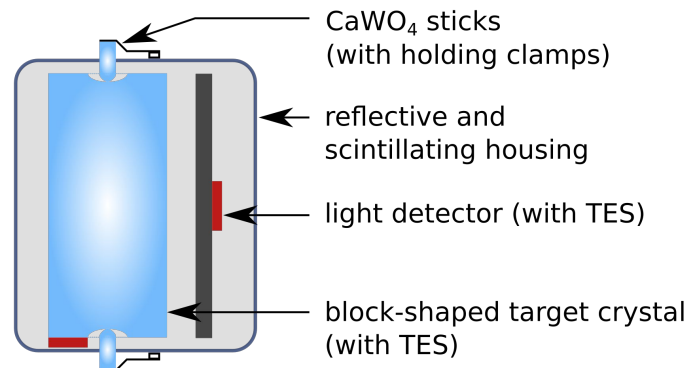
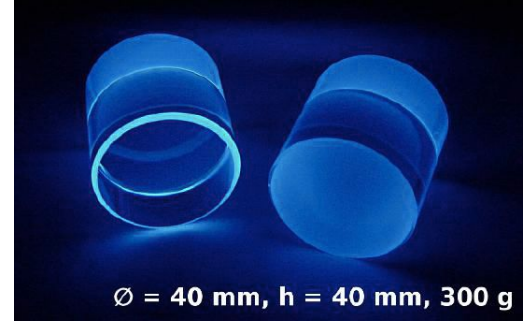
Nb_xSi_{1-x} spiral Al grid

arXiv: 2203.03993



CRESST: TES + Light + Active Veto

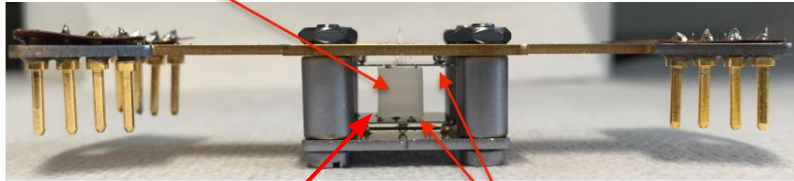
- CaWO_4 scintillating crystals, with phonon + light readout
- Mechanical structure instrumented with active sensors as well
- 300 g in CREST II \rightarrow 24g crystal in CRESST III
 - 4.6 eV resolution



NUCLEUS: TES + Active Veto

- 1 gram crystal, read out with TES
 - 3.7 eV resolution
- Outer detector provides active veto

Target (Al_2O_3 , CaWO_4 , Si, Ge, ...)



Inner veto (Si)

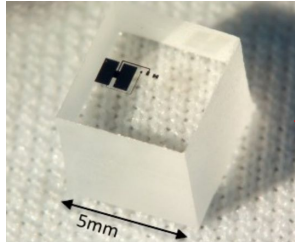
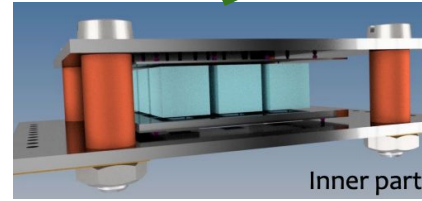
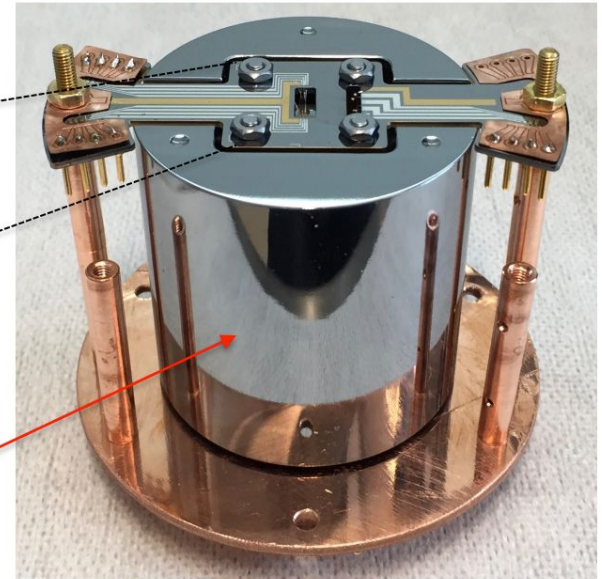


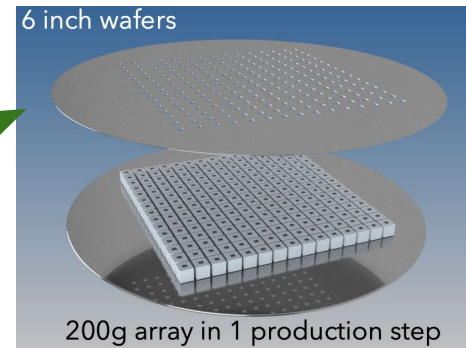
Fig. from R. Strauss, m7s 2019
PRD 96, 022009 (2017)

Outer veto (Ge , LiWO_4)



Inner part

Further up



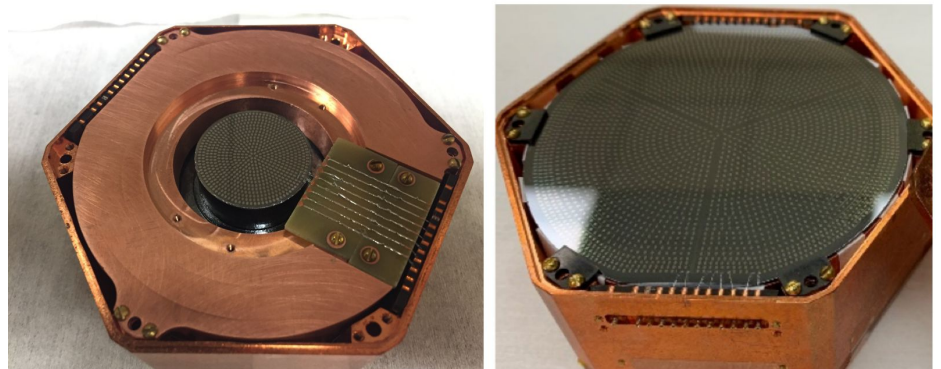
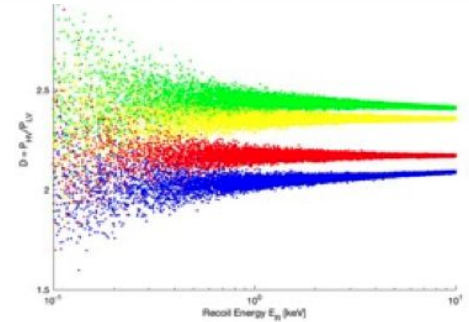
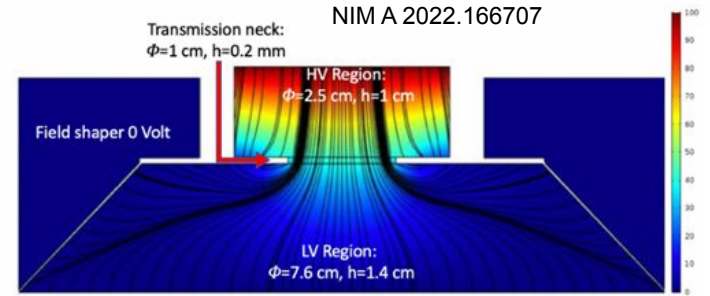
6 inch wafers

200g array in 1 production step



MINER: Hybrid Phonon detector

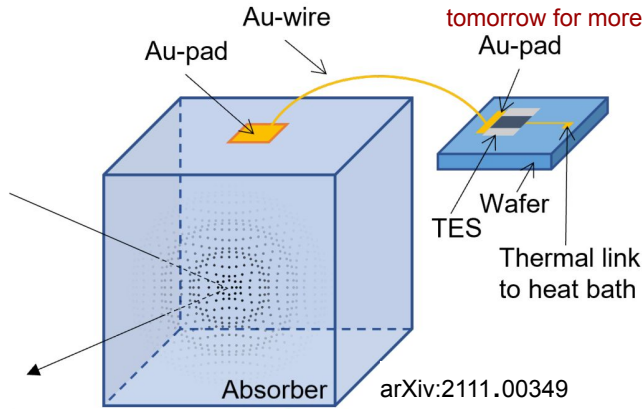
- Separate crystal into a low-voltage (LV) and a high-voltage (HV) region
- Phonon sensors on both sides
- Use the LV region as the fiducial volume
- Shape E-field to guide charges through the “neck”
- NTL phonons from the charges dominates in the HV side, whereas recoil phonons dominate the LV side
- $E_{HV} \sim$ charge measurement
 $E_{LV} \sim$ recoil phonon measurement



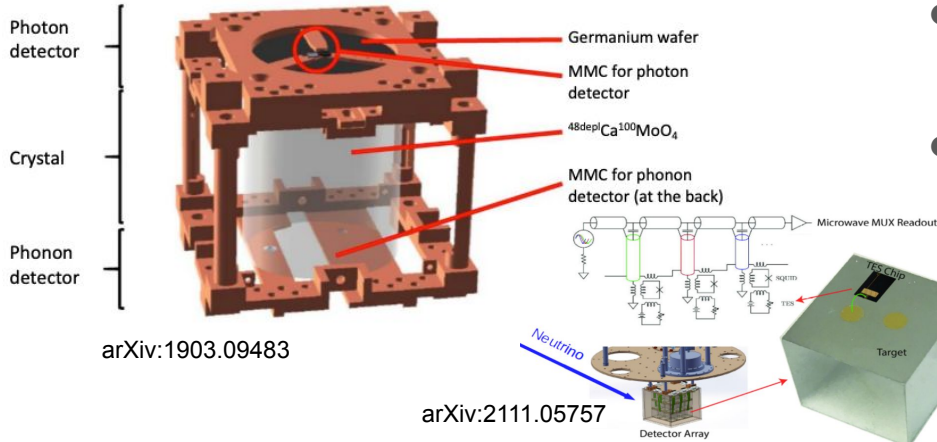
See R. Mahapatra's talk tomorrow for more details

Modular design: detaching thermometer from absorber

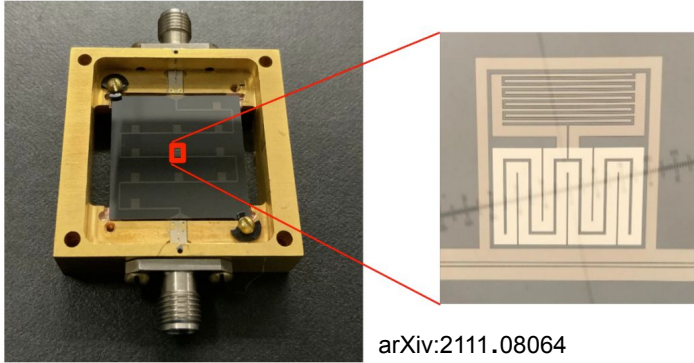
See M. Kellermann's talk tomorrow for more details



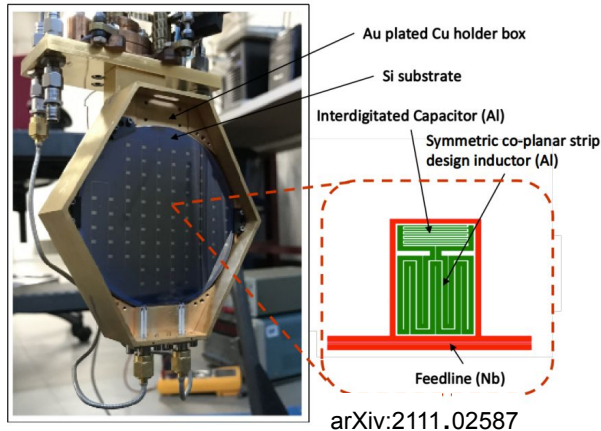
- Decoupling thermometer from crystals for ease of fabrication
- Thermal conduction facilitated by a gold wirebond
- Can be coupled to **a wide variety of target materials**
- RemoTES from COSINUS achieved <100 eV resolution with a 2-gram target
- Similar designs include
 - MMC based AMORE detector for $0\nu\beta\beta$
 - TES-based Ricochet Qarray for CEvNS



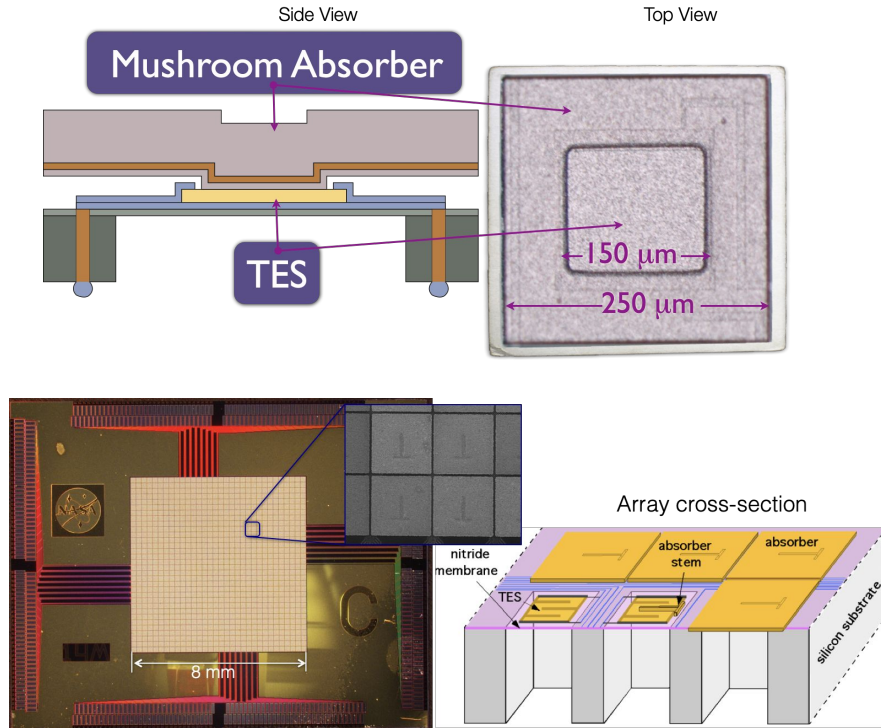
Multiplexing with RF resonators



- Significant improvement has been made on mKIDs recently
- 6 eV resolution for energy deposited in resonator demonstrated
 - Translates to a few tens of eV of resolution for energy deposited in the crystal
- Also with intrinsic capability of multiplexing
 - Promising candidates for next generation rare-event experiments



Cryogenic detector in Astrophysics and Indirect searches



- Cryogenic microcalorimeters also contribute in DM studies in astrophysics and indirect searches
- Arrays of eV-resolution sensors make perfect X-ray detectors
 - Like an ultra-sensitive camera
- Widely used in earth-based X-ray telescopes as well as rocket and satellite-based detectors

Fig. from E. Figueroa, COFI PIRE 20017

Conclusions

- Cryogenic detectors play an important role in dark matter search
 - Both in direct search and in indirect and astro approaches
- Resistance-based phonon detectors (TES and NTD) are approaching an eV-resolution regime
 - Also exploiting techniques including quasi-particle traps and internal amplifications via NTL effect
- Alternative sensing mechanisms (MMC and mKIDs) are advancing as well
 - Promising candidates for next-generation DM search detectors
- Information with phonon + charge, phonon + light, or with layered detector structure can help with particle identifications
- Stay tuned for more results -- the discovery might be around the corner