# EXPLORING THE DARK UNIVERSE WITH THE SOUTH POLE TELESCOPE

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Photo: Aman Chokshi

#### **SPT3G COLLABORATION**





### THE SOUTH POLE TELESCOPE

10-m submm-quality wavelength telescope

- ▶ 90, 150, 220 GHz
- ▶ 1.6, 1.2, 1.0 arcmin resolution

#### ▶ 2007: SPT-SZ

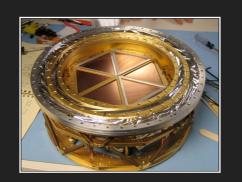
- 960 detectors
- > 90, 150, 220 GHz

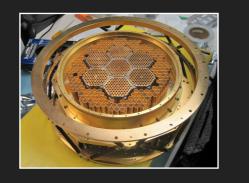
#### ▶ 2012: SPTpol

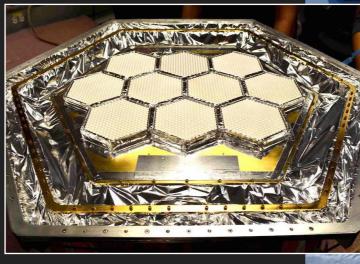
- 1600 detectors
- ▶ 90, 150 GHz
- +polarization

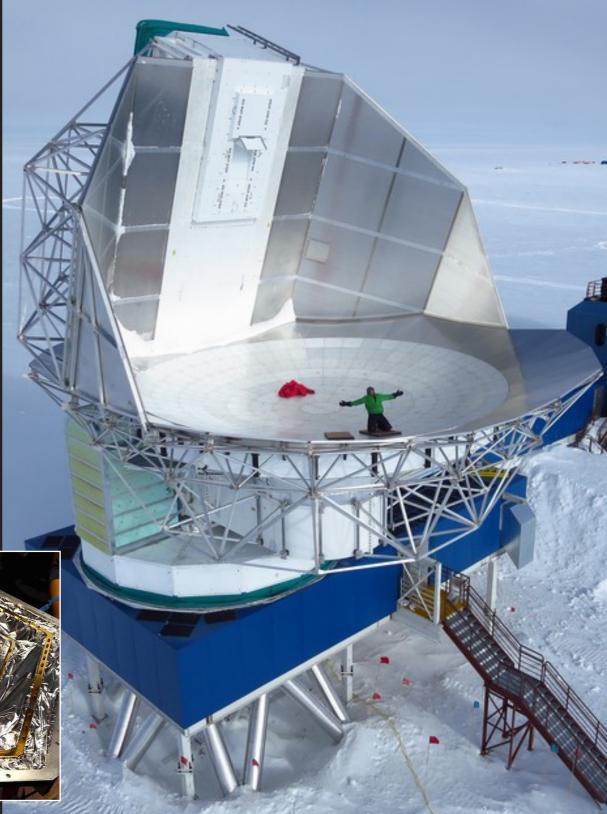
#### ▶ 2017: SPT-3G

- ~16,200 detectors
- > 90, 150, 220 GHz
- +polarization





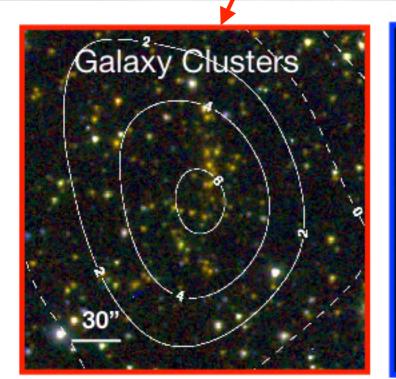


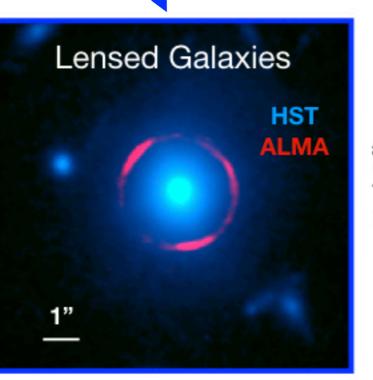


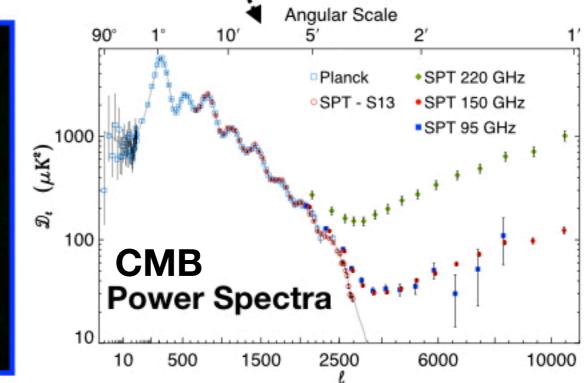
# SPTpol

6x deeper 6x finer angular resolution



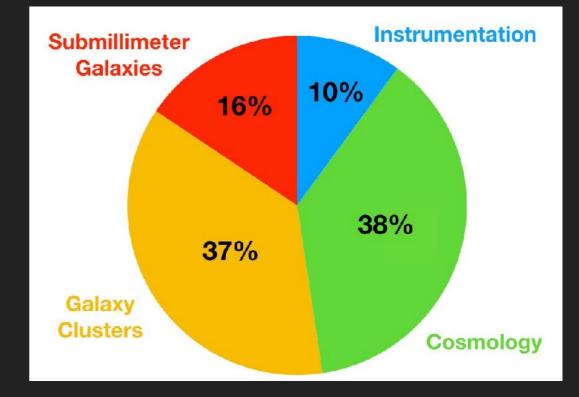






#### BREADTH OF SPT RESULTS (>130 SCIENCE PUBLICATIONS)

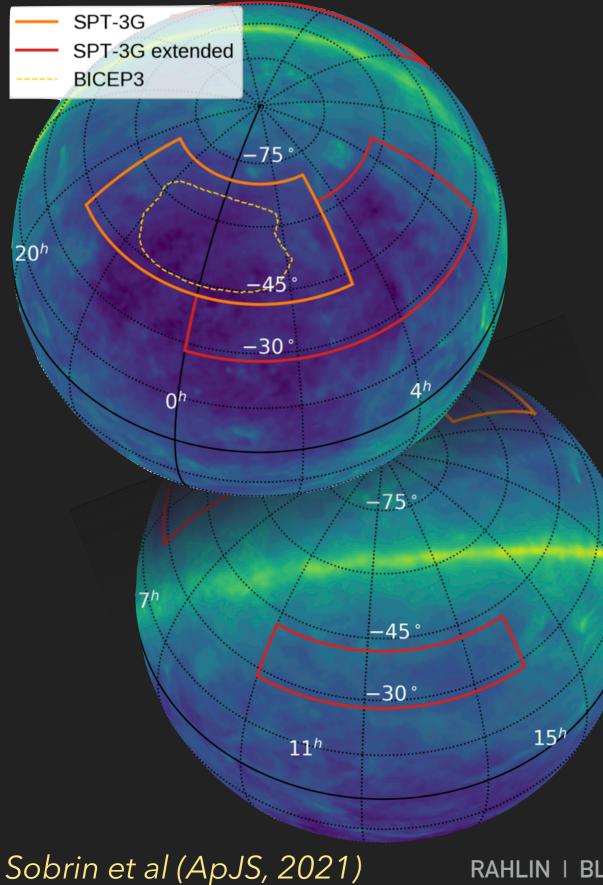
- CMB anisotropy: Power spectra and cosmological parameters
- CMB B-Modes: First detection of lensing B-mode polarization; demonstration of delensing for improved constraints on inflationary tensor-to-scalar ratio
- **CMB lensing:** power spectra; cross-correlations; cluster-lensing mass calibration
- Sunyaev-Zel'dovich (SZ): Diffuse kinematic and thermal SZ effect constraints: bispectrum, pairwise kSZ, patchy reionization
- Galaxy Clusters: First SZ discovery clusters, cluster catalog and cosmology
- High-Redshift Galaxies: Discovered population of lensed dusty star forming galaxies
- Transients: mm-wave phenomena (GRBs, FRBs), mJy-level monitoring of 1000s of blazars, AGNs
- Participating in the Event Horizon Telescope



#### Publications: pole.uchicago.edu

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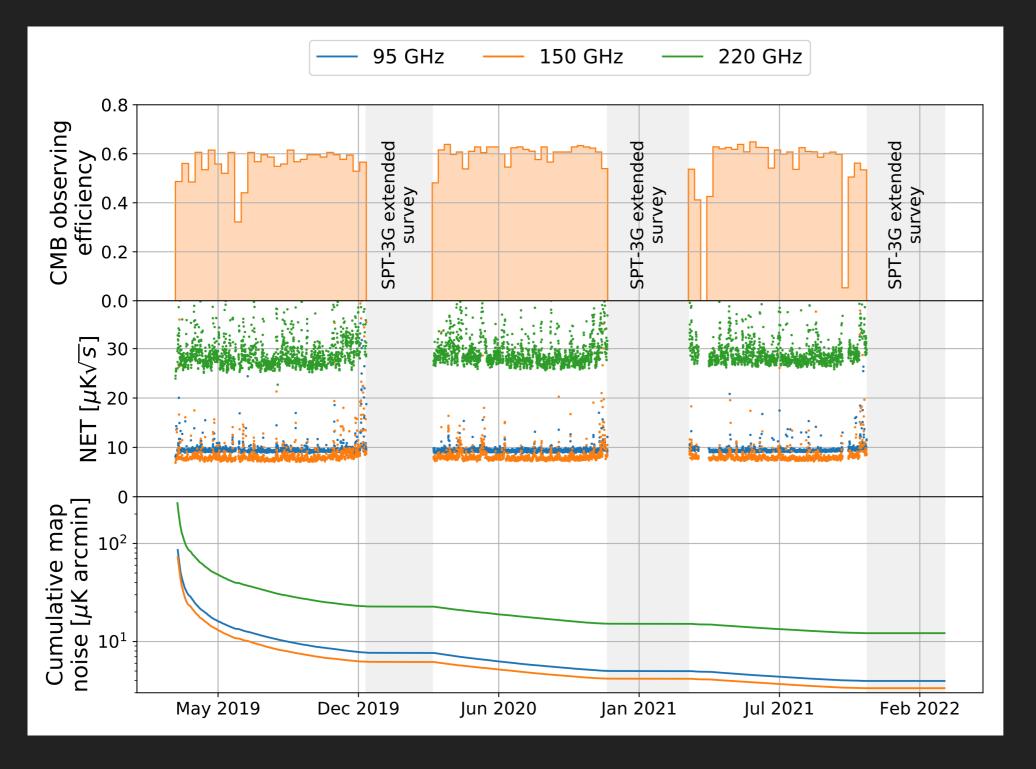
### THE SPT-3G 1500 DEG<sup>2</sup> SURVEY



- SPT-3G 1500 deg<sup>2</sup> survey will be ~10x deeper than SPT-SZ
- Overlaps BICEP Array, to optimize inflationary constraints from CMB delensing

	Obs. Years	Area (deg²)	95 GHz (uK- arcmin)	150 (uK- arcmin)	220 (uK- arcmin)
SPT-SZ	2007-11	2500	40	17	80
SPTpol- 500d	2012-16	500	13	6	-
<b>SPTpol-</b> 100d	2012-16	100	10	5	_
SPTpol- 2700d	2012-16	2700	47	28	-
SPT-3G (projected)	2018-23	1500	3.0	2.2	8.8

#### SPT-3G: DATA QUALITY



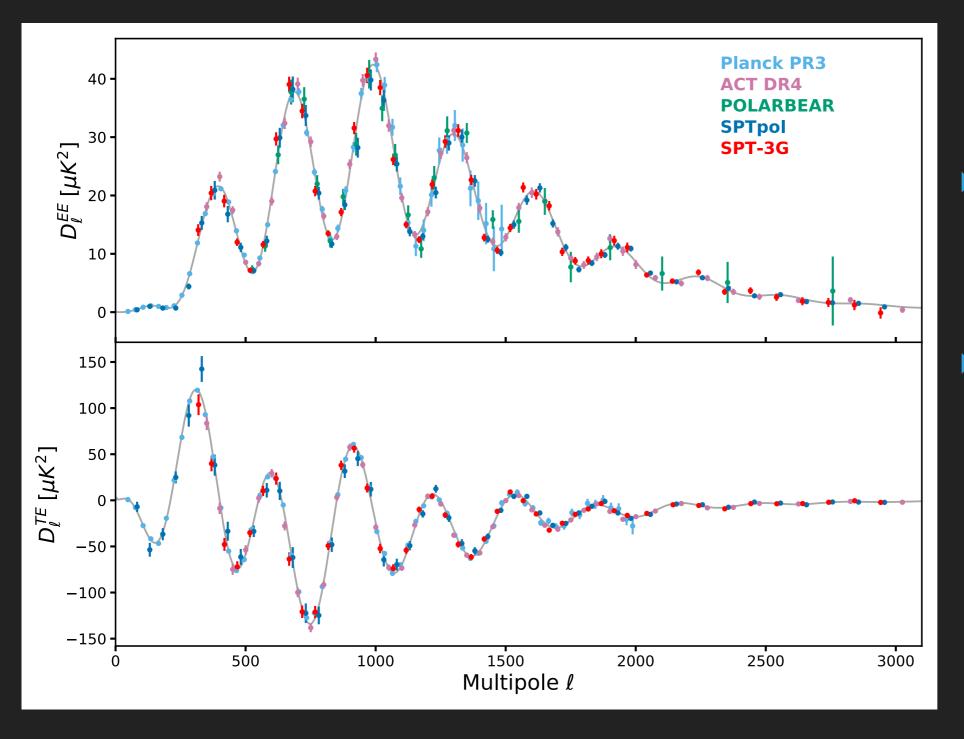
~60% observing efficiency in 9-month observing season

Daily camera sensitivity is stable over season

Expect to reach CMB-S4 wide survey depth by end of 2023

#### A. Anderson, W. Quan

### SPT3G: 2018 CMB POWER SPECTRA

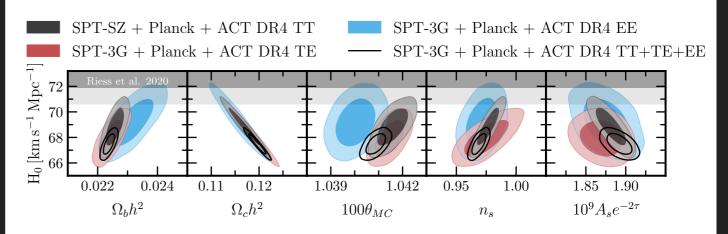


- 2018: half season with half of full detector count
- Constraints already comparable to or exceed leading measurements at intermediate scales

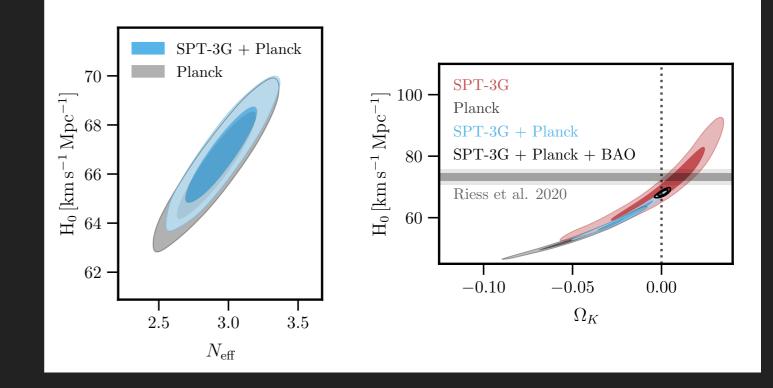
Dutcher et al (PRD, 2021)

### SPT3G: 2018 ACDM CONSTRAINTS

- Consistent H<sub>0</sub> values between
  T and P spectra when
  evaluated across experiments
- Tightest constraint on H<sub>0</sub> to date
  - 4.1 $\sigma$  tension with Riess et al
- Further model extensions:  $N_{eff}$ ,  $Y_p$ ,  $\sum m_v$ ,  $\Omega_K$ 
  - No clear evidence for any extensions
  - with Planck, reduce volume of parameter space by 1.3-2x

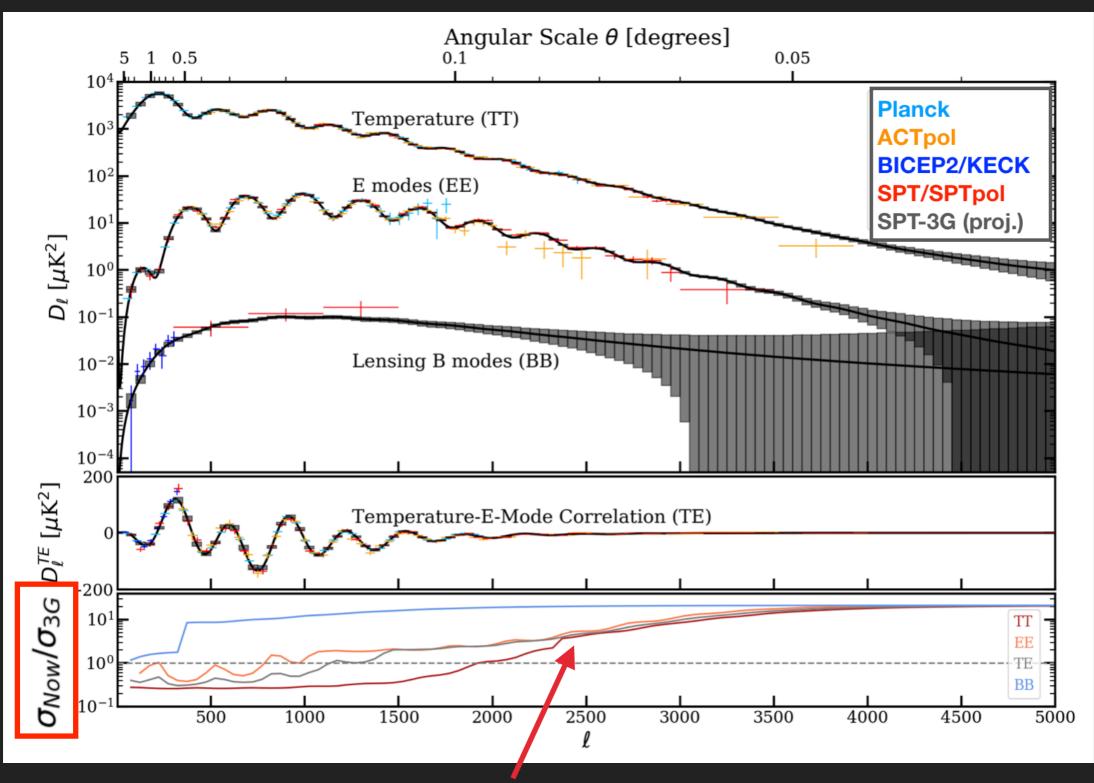


Spectra	Datasets	$H_0  [{\rm km  s^{-1}  Mpc^{-1}}]$
TT	$SPT-SZ + Planck + ACT DR4 \ (\ell > 1800)$	$68.85 \pm 0.97$
TE	SPT-3G 2018 + $Planck$ + ACT DR4	$67.95 \pm 0.94$
$E\!E$	SPT-3G 2018 + $Planck$ + ACT DR4	$69.2 \pm 1.2$
TT + TE + EE	SPT-3G 2018 + $Planck$ + ACT DR4	$67.49 \pm 0.53$



Balkenhol et al (PRD, 2021)

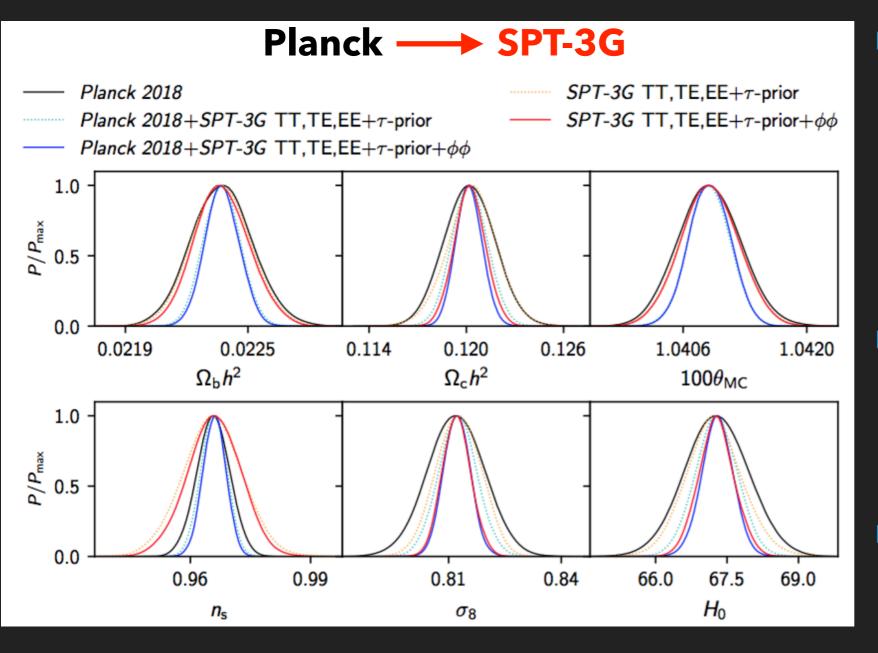
### **SPT3G FORECAST: CMB POWER SPECTRUM**



Improve S/N of CMB power spectra by factors of > ~10 at  $\ell$  > 2500 over current constraints

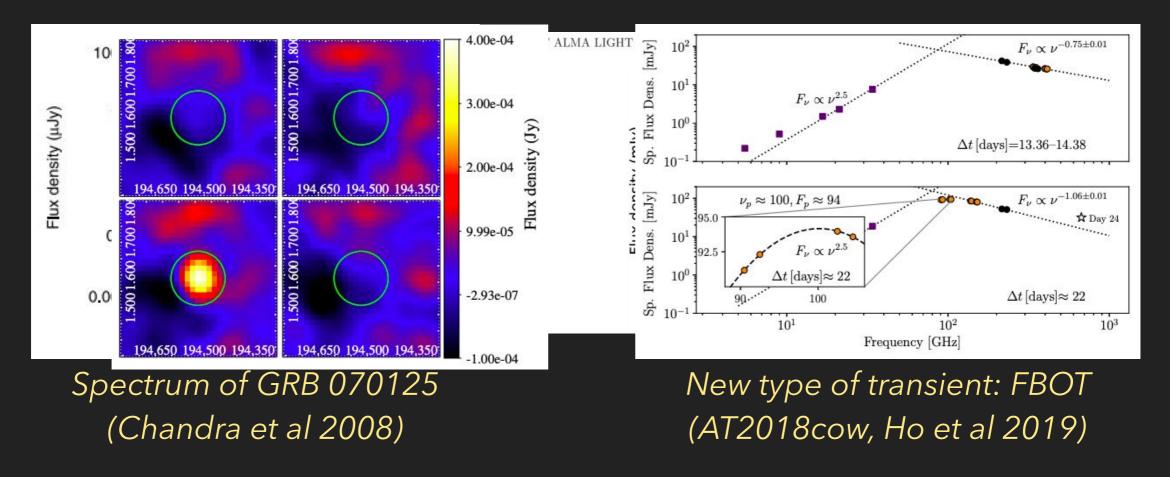
#### J. Henning

### SPT3G FORECAST: ACDM CONSTRAINTS

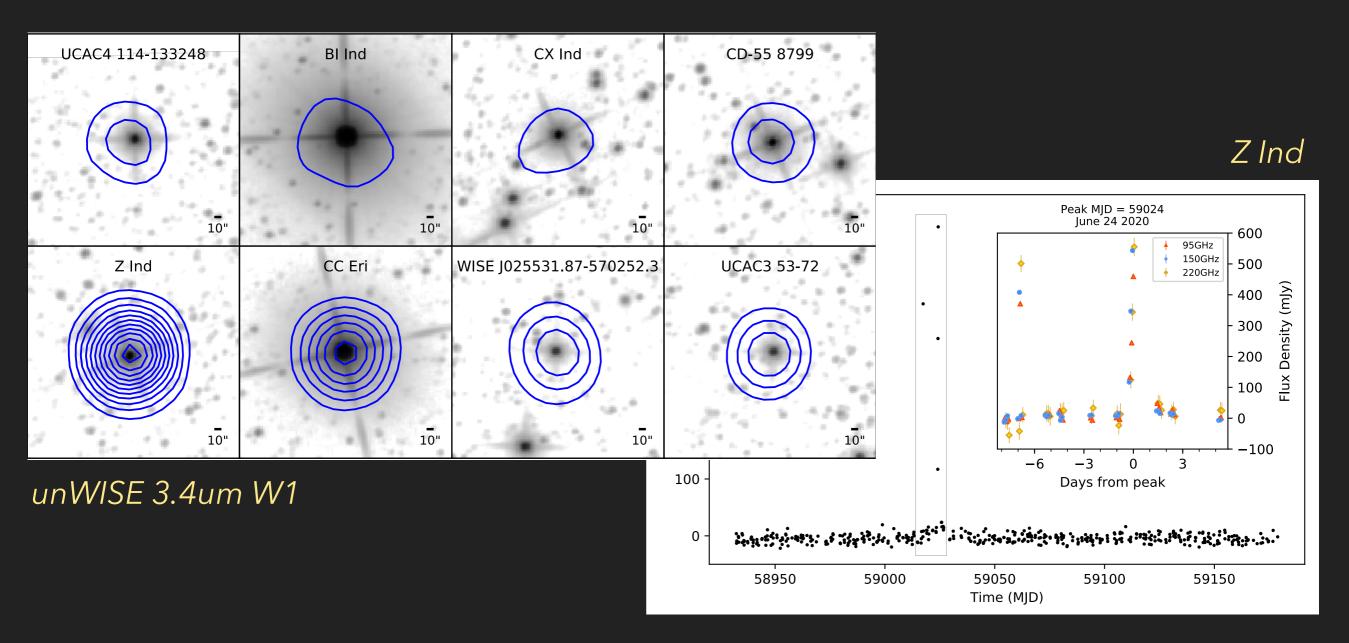


- SPT-3G alone will do as well as Planck on most parameters (except T, n<sub>s</sub>)
  - e.g., σ(H<sub>0</sub>)=0.6 km/s/Mpc from SPT-3G lensed TT,TE,EE spectra
- SPT-3G gives ~independent cosmological information from high-& CMB polarization spectrum
- SPT-3G + Planck will improve most parameters >2x over
   Planck alone

- Exploit CMB observating cadence: repeated daily imaging of large sky area at multiple frequencies
- Provide useful probe of high-energy astrophysics (AGN, GRB, FRB, stellar flares, ...)
- No dedicated transient surveys exist between 10GHz and IR frequencies
- Large discovery potential with online alert system

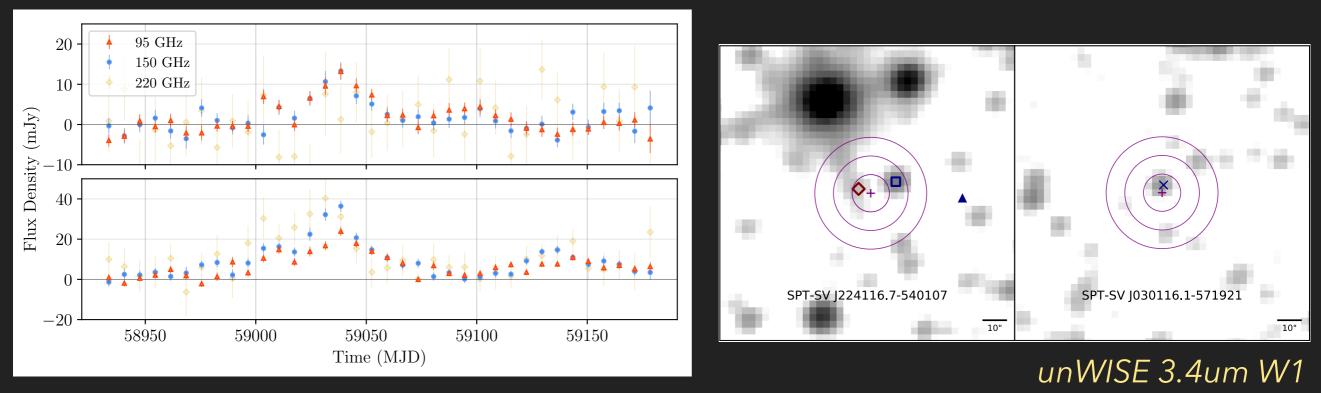


- First results: 15 transient events in 2019 dataset
- Most are stellar flares (likely magnetic reconnection events?)
  - up to 2 Jy in brightness, flares as short as 20 min, some flare more than once

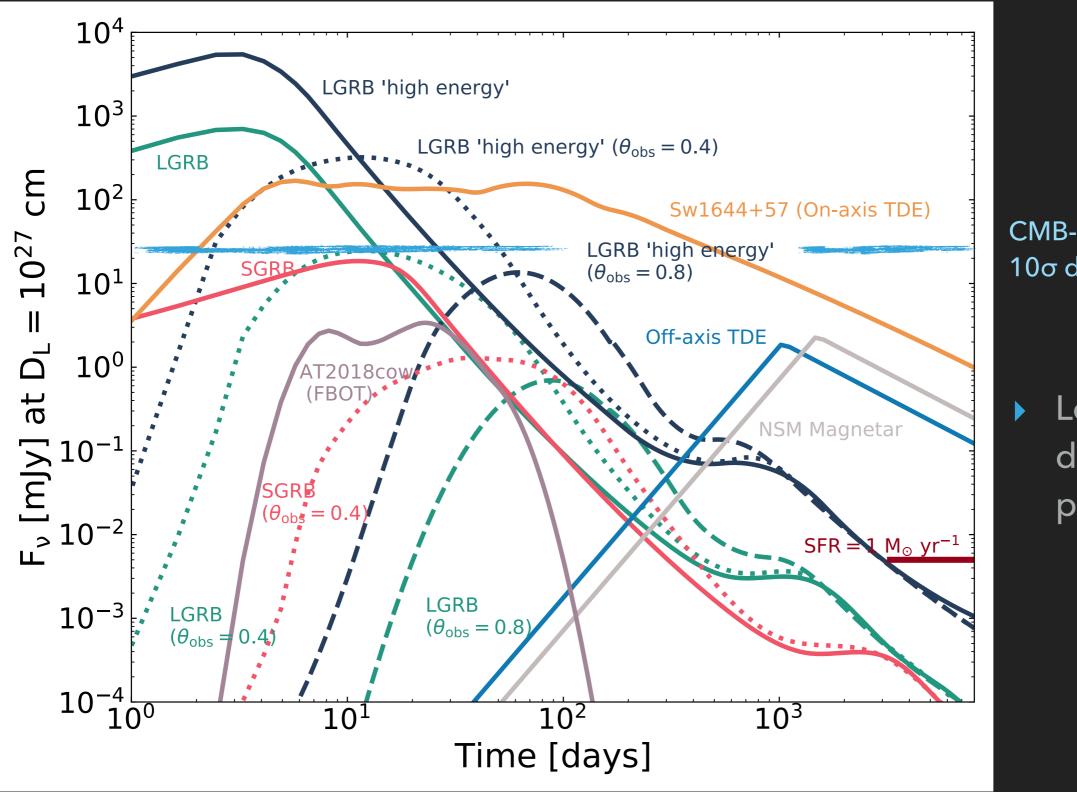


Guns et al (ApJ, 2021)

- First results: 15 transient events in 2019 dataset
- Two extragalactic long-duration flares
  - Super variable AGN? Something else? Follow-up in progress



4x / 15x increased flux relative to 2019 average



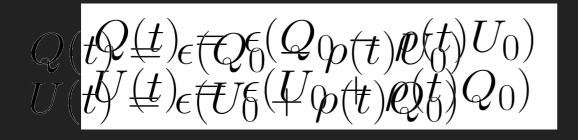
CMB-S4 Deep / SPT-3G 10σ daily limit

Lots of discovery potential!

Eftekhari et al (2021)

# SPT3G: SEARCH FOR ULTRA-LIGHT AXION-LIKE PARTICLES

- Exploit axion-photon coupling to search for evidence of axion-like dark matter candidates
- Oscillation of axion field results in rotation of linear polarization of photons along line of sight

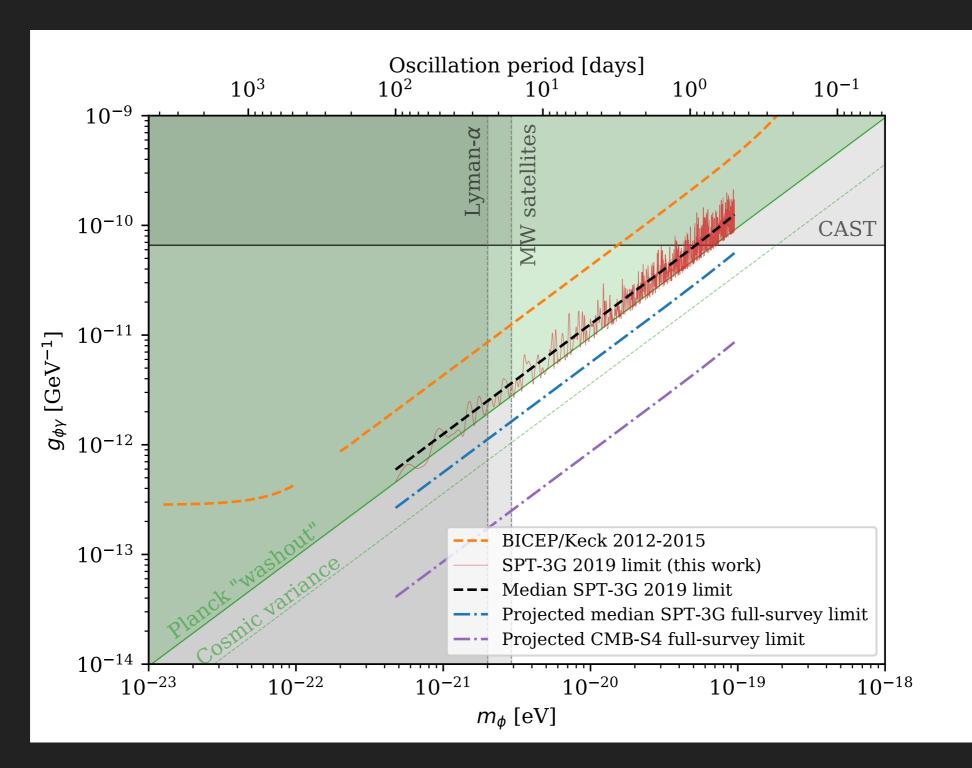


$$\rho(t) = g_{\phi\gamma}\phi_0 \cos\left(m_\phi t + \alpha\right) + \alpha$$

- Axion signatures in CMB polarization:
  - Polarization washout: average axion field oscillations to reduce polarization intensity of CMB photons -- expected to be small
  - AC birefringence: Local axion field oscillations create time-dependent polarization rotation in CMB photons
    - SPT3G observing cadence probes time scales from hours to days!

Ferguson et al (2022)

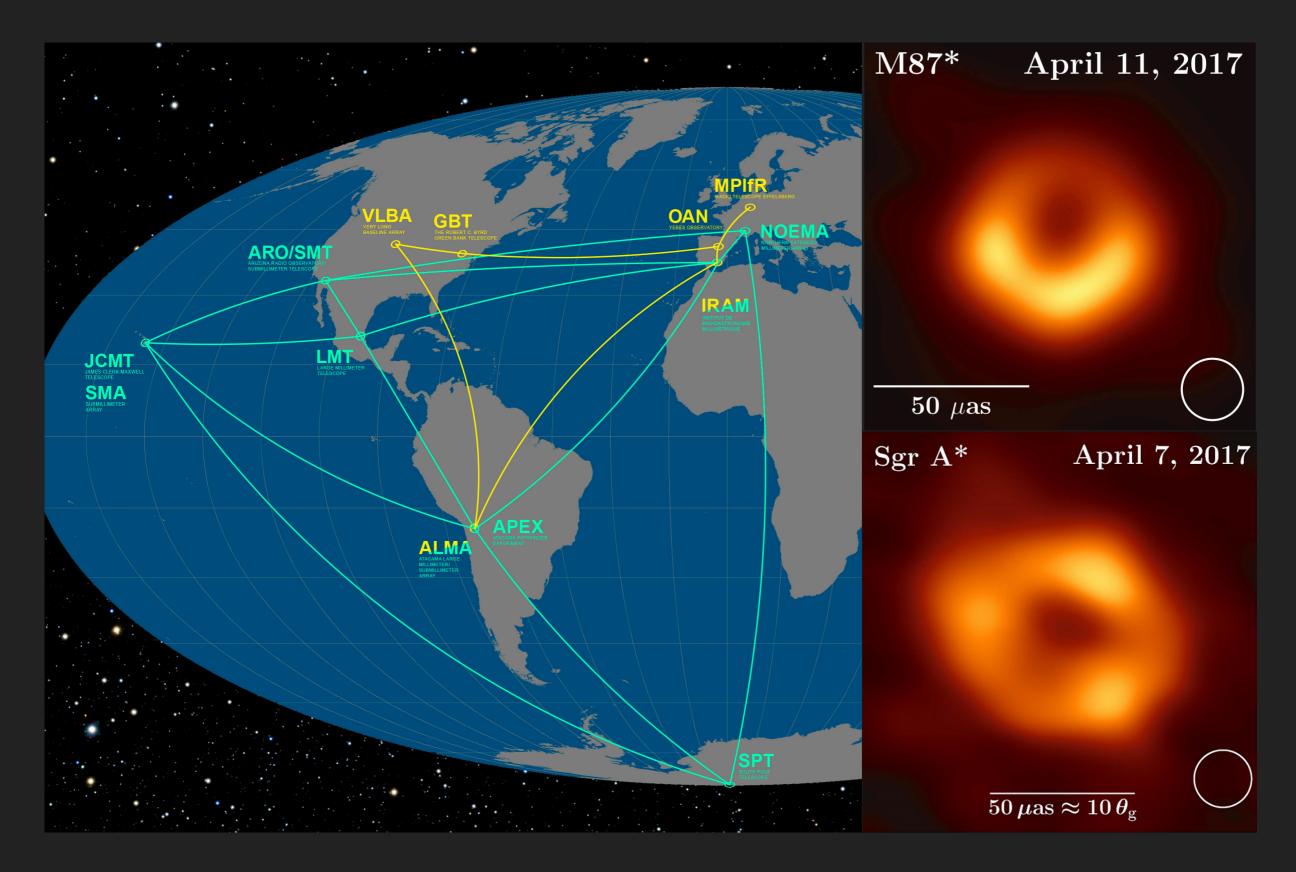
### SPT3G: SEARCH FOR ULTRA-LIGHT AXION-LIKE PARTICLES



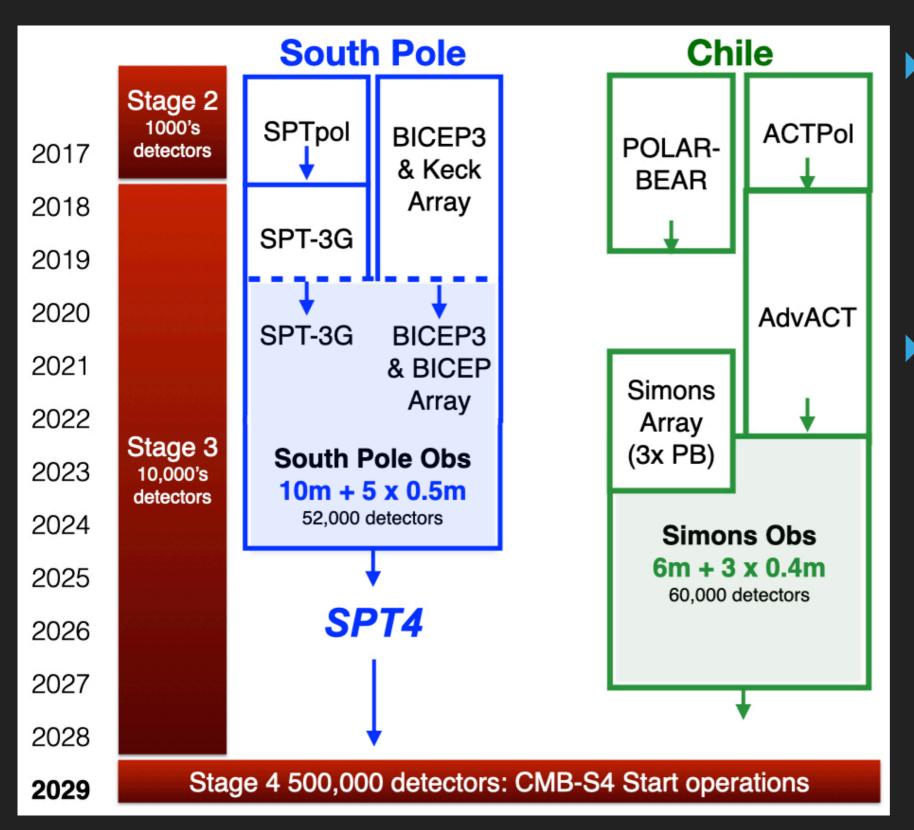
g<sub>φ<sub>y</sub></sub> < 1.18x10<sup>-12</sup> GeV<sup>-1</sup> x (m<sub>φ</sub> / 10<sup>-21</sup> eV) [95% CL] -- 3.8x improved limit

Ferguson et al (2022)

### **EVENT HORIZON TELESCOPE**



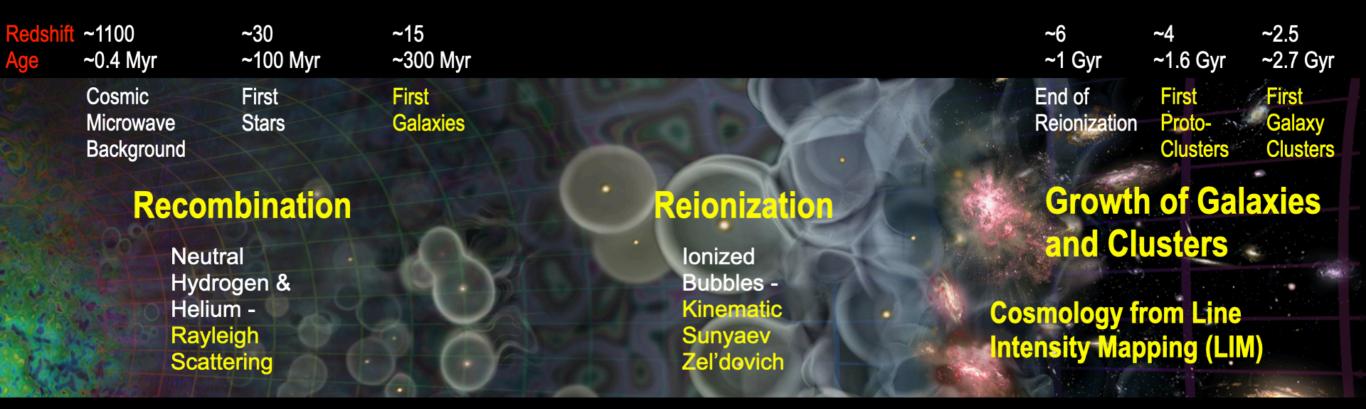
### **FUTURE DIRECTIONS**



5-year gap
 between end of
 SPT3G and start
 of CMB-S4

 SPT4 proposed to utilize the SPT submm-quality telescope during the gap

### **SPT4 SCIENCE**



#### Recombination

First detection of CMB Rayleigh Scattering, new probe of cosmic expansion history, and cosmology. (Dibert et al [2205.04494], Zhu et al [2205.04496])

#### Reionization

- Constrain duration and redshift of reionization via kSZ
- Growth of galaxies and clusters
  - Detect growth of massive galaxies and clusters from 1 < z < 9</p>
- Dust and star formation in the galaxy

### SUMMARY

#### SPT-3G survey is ongoing

- First results from 2018 season published, with impressive CMB sensitivity and performance at large scales
- SPT-3G science publications from 2019-2020 seasons coming soon
- Rich astrophysical transient program

#### SPT-3G science forecasts through 2023 season

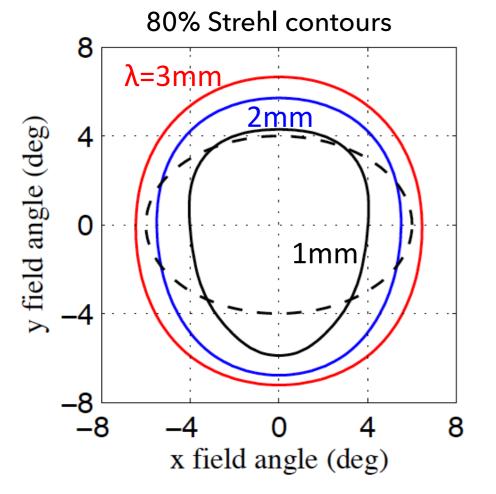
ACDM constraints as good as Planck on most cosmological parameters
 With Planck, factor of ~2x improved constraints on N<sub>eff</sub> and neutrino mass
 With BICEP Array, factor of ~10x improvement on current BK15 tensor-to-scalar ratio constraint

#### Future plans for CMB science at the South Pole

- SPT4 to enable first detection of CMB Rayleigh Scattering, and to constrain reionization via the kSZ effect
  - CMB-S4 coming online at the end of the decade

#### **SP-TMA: SOUTH POLE THREE MIRROR ANASTIGMAT**

- Degree B-modes and arc minute resolution with extremely high throughput.
- 5-meter Three Mirror Anastigmat (TMA) telescope design:
  - 75 deg<sup>2</sup> field of view
  - 424k/136k/63k Fλ pixels at λ=1/2/3mm
  - Monolithic mirrors (low scattering)
  - Boresight rotation for polarization systematics

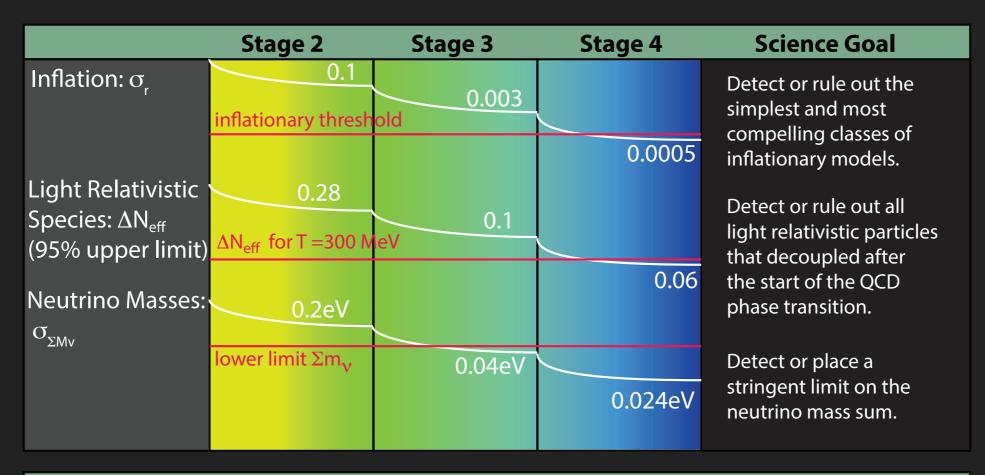


3.5 m diameter focal plane

Primary mirror prototyping in progress

Padin, Applied Optics, 57, 9, 2314 2018

#### **CMBS4 FORECASTS**



	Stage 2	Stage 3	Stage 4	Requirements
Survey Weight [µK <sup>-2</sup> ]	10 <sup>5</sup>	10 <sup>6</sup>	10 <sup>8</sup> ~500,000	500,000 detectors on multiple platforms with sensitivity from 10° to 1' scales
Detector Count	~1,000	~10,000		