

# Recent neutrino mass results from the KATRIN experiment

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Stephanie Hickford | Tuesday 24<sup>th</sup> May 2022



# Outline

## 1. Neutrino mass determination

- Massive neutrinos
- Tritium single  $\beta$ -decay

## 2. The KATRIN experiment

- Beamline
- Integrated spectrum measurement

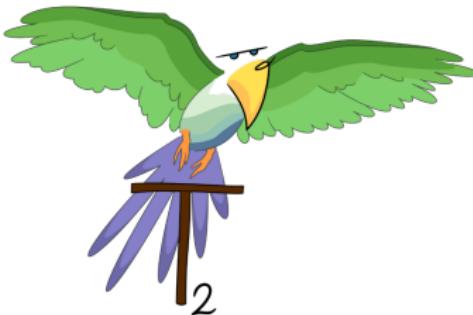
## 3. Analysis

- Spectra fitting
- Systematics

## 4. Recent results

- Measurement campaigns
- Individual campaign results
- Combined campaign results
- (Further) BSM analyses

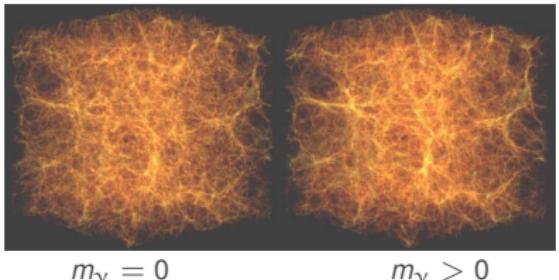
## 5. Summary



# Massive neutrinos

## Standard Model

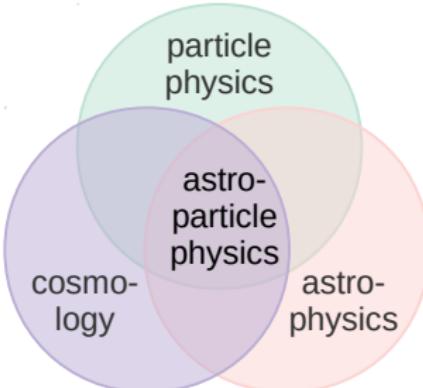
- Mass generation
- Weak interactions
- Oscillation
- Sterile
- LIV
- RH current



## Massive neutrinos as “cosmic architects”

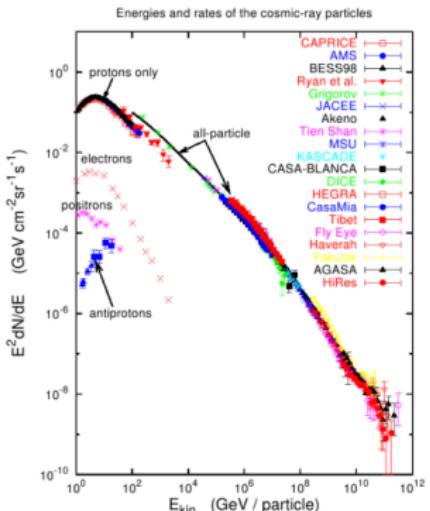
- $336 \text{ v}/\text{cm}^3$  in the universe
- Cosmic relic neutrinos

	I	II	III		
mass	2.2 MeV/c <sup>2</sup>	1.28 GeV/c <sup>2</sup>	173.1 GeV/c <sup>2</sup>	0	124.97 GeV/c <sup>2</sup>
spin	½	½	½	0	0
charge	½	½	½	0	0
quarks	u	C	t	g	H
	4.7 MeV/c <sup>2</sup>	96 MeV/c <sup>2</sup>	4.18 GeV/c <sup>2</sup>	0	scalar bosons
	-½	-½	-½	0	
	d	S	b	γ	
	0.511 MeV/c <sup>2</sup>	105.66 MeV/c <sup>2</sup>	1.78 GeV/c <sup>2</sup>	0	
	-1	-1	-1	1	
	e	μ	τ	Z	
	< 0.8 MeV/c <sup>2</sup>	< 0.17 MeV/c <sup>2</sup>	< 18.2 MeV/c <sup>2</sup>	80.39 GeV/c <sup>2</sup>	
	0	0	0	±1	
leptons	V <sub>e</sub>	V <sub>μ</sub>	V <sub>τ</sub>	W	gauge bosons
	½	½	½	1	



## Understanding astro-physical processes

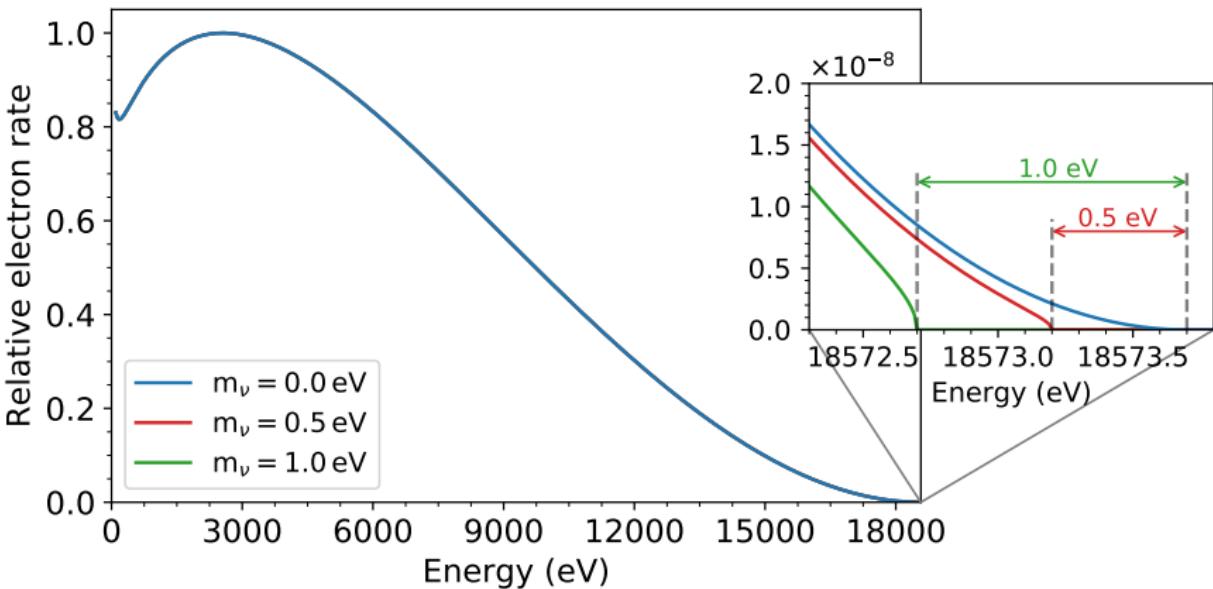
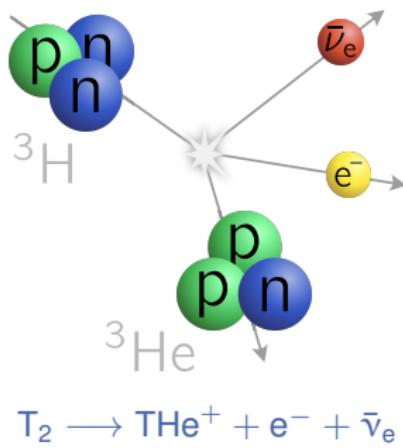
- Nuclear reactions in stars
- ν as probes for cosmic rays



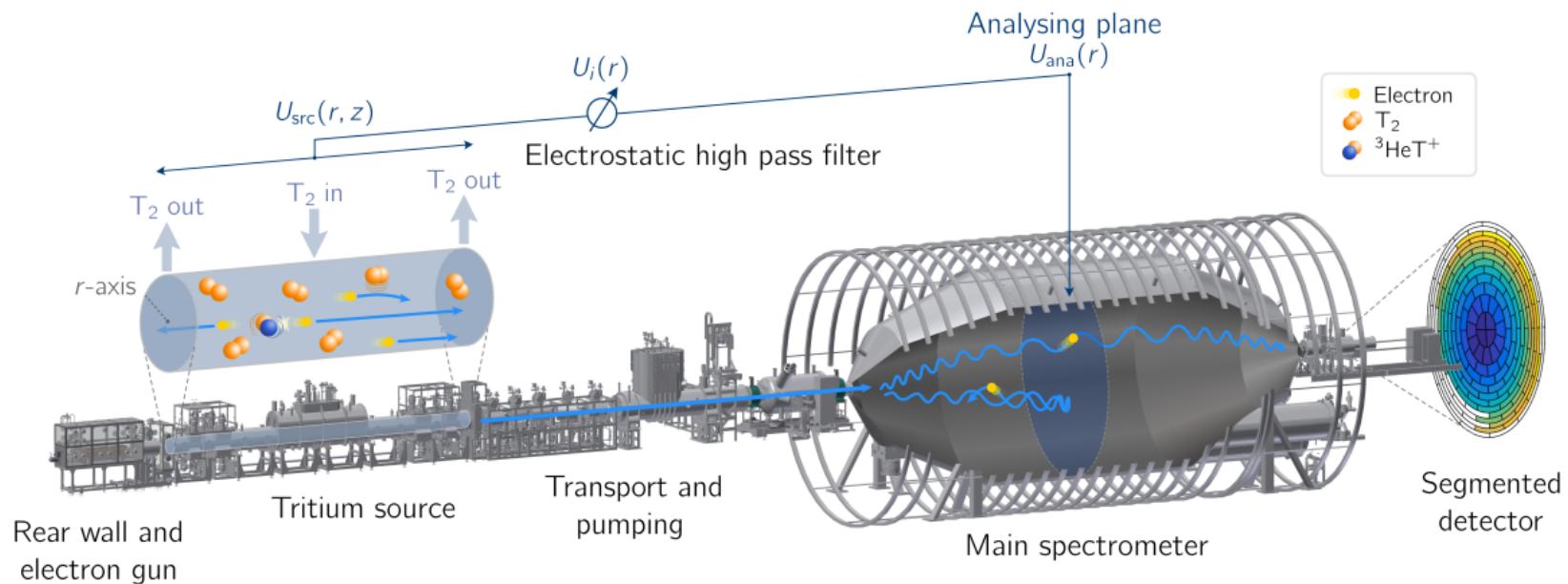
# Tritium single $\beta$ -decay

Decay of molecular tritium produces a  $\beta$ -electron spectrum

➡  $m_\nu^2$  can be determined with a precise measurement of the spectral shape near the endpoint



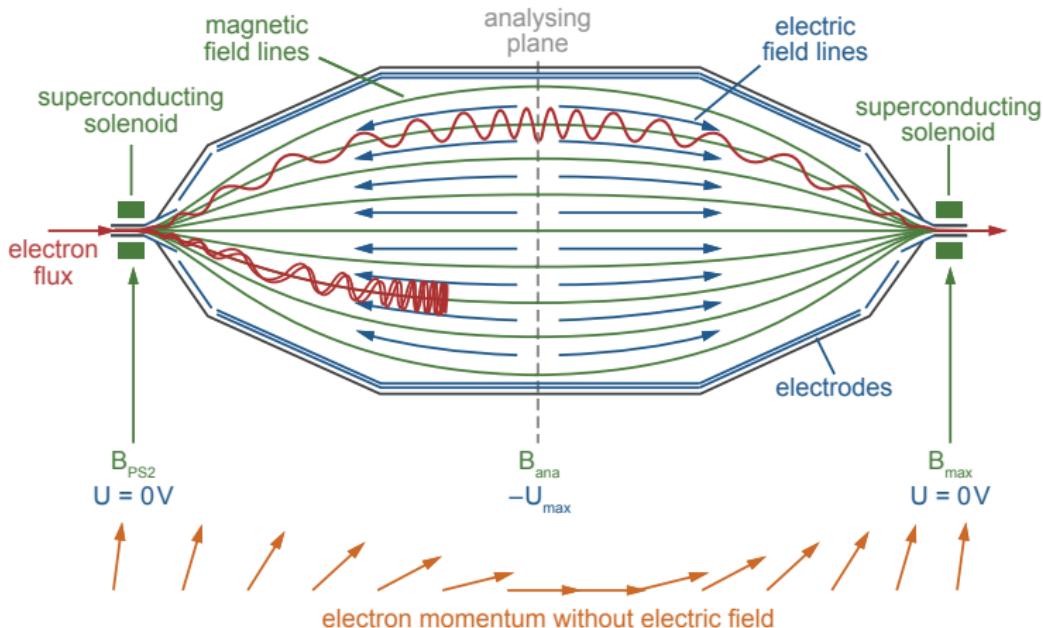
# Beamline



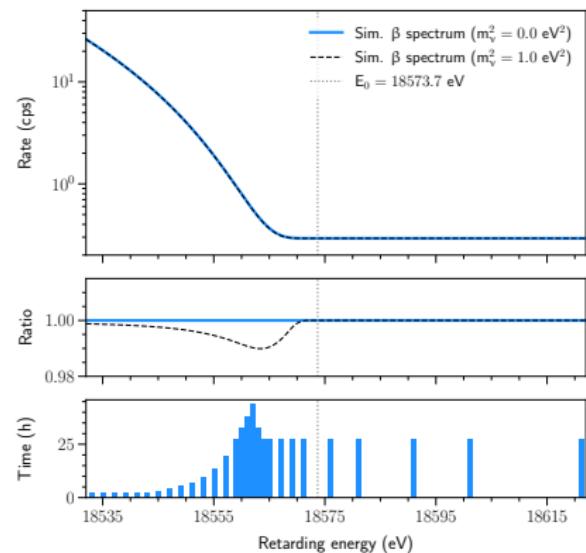
# Integrated spectrum measurement

Magnetic Adiabatic Collimation combined with an Electrostatic filter (MAC-E filter)

Beamson et al. (1980), Lobashev, Spivak (1985), Picard et al. (1992)



Voltage set points are scanned through to obtain an *integrated spectrum*



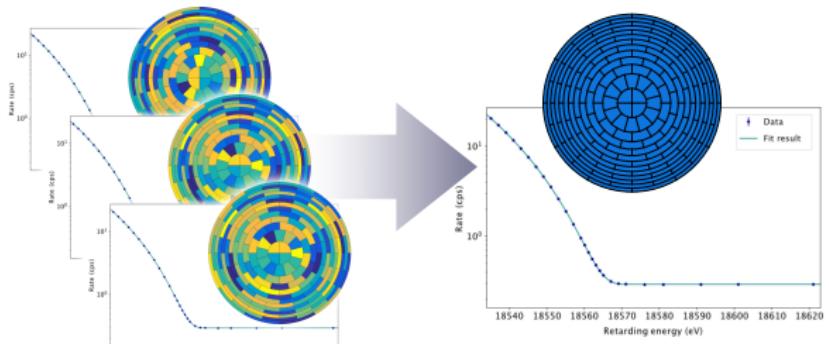
The Measurement Time Distribution (MTD) optimised for best sensitivity

# Spectra fitting

Data combination: Counts are summed, experimental parameters are averaged

## Pixel combination

- ➡ Uniform
- ➡ Multi-patch
- ➡ Multi-pixel



## Run combination

- ➡ Stacked
- ➡ Multi-period
- ➡ Multi-run

Fit is performed with many contributing spectra

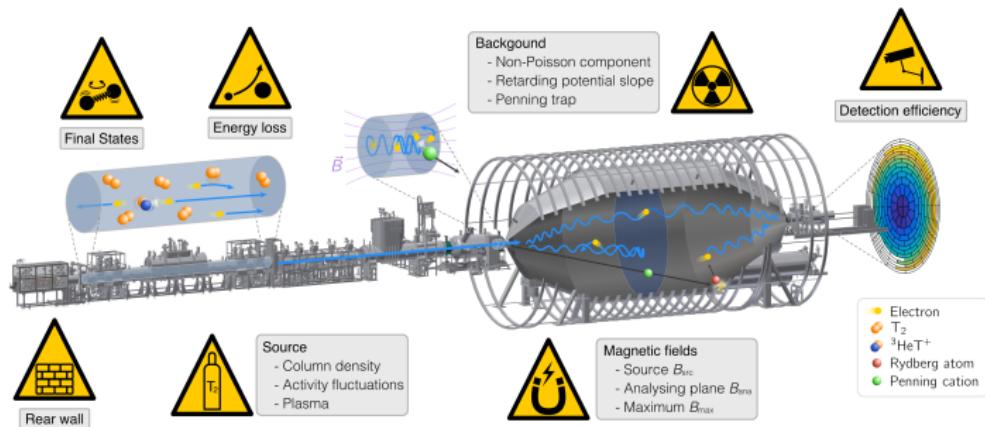
- ➡ One minimisation
- ➡ One combined likelihood,  $\mathcal{L}$

$$-\log \mathcal{L} = \sum_i -\log \mathcal{L}_i(m_\nu^2, E_{0i}, \text{Sig}_i, \text{Bg}_i)$$

Many parameters

- ➡ One common neutrino mass,  $m_\nu^2$
- ➡ Multiple  $E_0$ ,  $\text{Sig}$ , and  $\text{Bg}$
- ➡ Systematic parameters either common or multiple

# Systematics

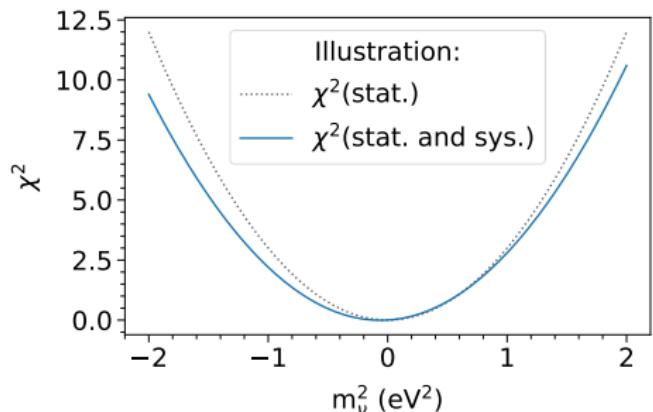


Adding pull terms widens the  $\chi^2$  distribution

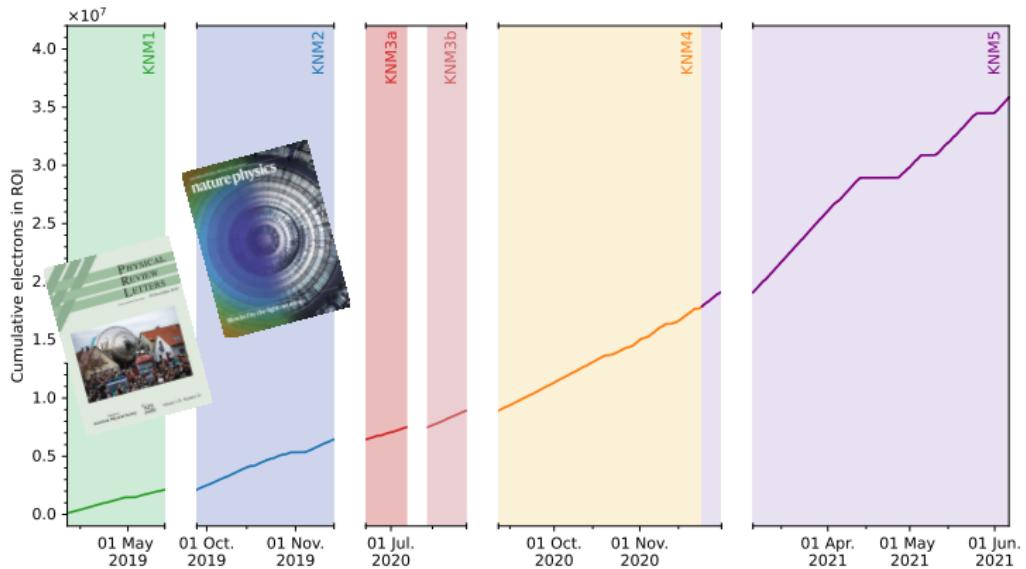
$$\chi^2(m_\nu^2, E_0, \text{Sig}, \text{Bg}, \theta_1, \dots) + \frac{(\theta_1 - \hat{\theta}_1)^2}{\sigma_{\theta_1}^2} + \dots$$

## Treatment of systematics

- Pull term
- Covariance matrix
- Monte Carlo propagation
- Markov Chain Monte Carlo (MCMC)



# Measurement campaigns



KATRIN Neutrino mass Measurements

	Time (hrs)	$\rho d\sigma$ (m <sup>-2</sup> )	Bg (mcps)
KNM1	522	$1.11 \times 10^{21}$	370
KNM2	294	$4.23 \times 10^{21}$	278
KNM3a	220	$2.08 \times 10^{21}$	137
KNM3b	224	$3.75 \times 10^{21}$	258
KNM4	1267	$3.77 \times 10^{21}$	150
KNM5	1232	$3.78 \times 10^{21}$	160

- Published results: KNM1 and KNM2
- Current analysis: KNM1 – KNM5
- Data-taking: KNM6, KNM7, ...

# Individual campaign results

## First measurement campaign

Fit strategy: Stacked uniform fit

⇒ One spectrum with 27 data points

⇒ Four free fit parameters ( $m_\nu^2$ ,  $E_0$ , Sig, Bg)

Statistics dominated fit result

$$m_\nu^2 = -1.0 \pm 1.0 \text{ eV}^2$$

## Second measurement campaign

Fit strategy: Stacked multi-ring fit

⇒ 12 spectra with  $12 \times 28 = 336$  data points

⇒ 37 free fit parameters ( $m_\nu^2$ ,  $12 \cdot E_0$ ,  $12 \cdot \text{Sig}$ ,  $12 \cdot \text{Bg}$ )

Statistics dominated fit result

$$m_\nu^2 = 0.26 \pm 0.34 \text{ eV}^2$$

Factor of  $\sim 2$  improvement on previous  $m_\nu$  limit

$$m_\nu < 1.1 \text{ eV (90 \% CL)}$$

Phys. Rev. Lett. **123** (2019) 221802

Statistics dominated fit result

$$m_\nu < 0.9 \text{ eV (90 \% CL)}$$

New sub-eV neutrino mass limit

$$m_\nu < 0.9 \text{ eV (90 \% CL)}$$

Nat. Phys. **18**, 160–166 (2022)

# Combined campaign results

Fit strategy: Multi-period uniform fit

- Data is stacked *within* the measurement phases
- Two spectra with  $27+28 = 55$  data points
- 7 free fit parameters ( $m_\nu^2$ ,  $2 \cdot E_0$ ,  $2 \cdot \text{Sig}$ ,  $2 \cdot \text{Bg}$ )

Statistics dominated fit result

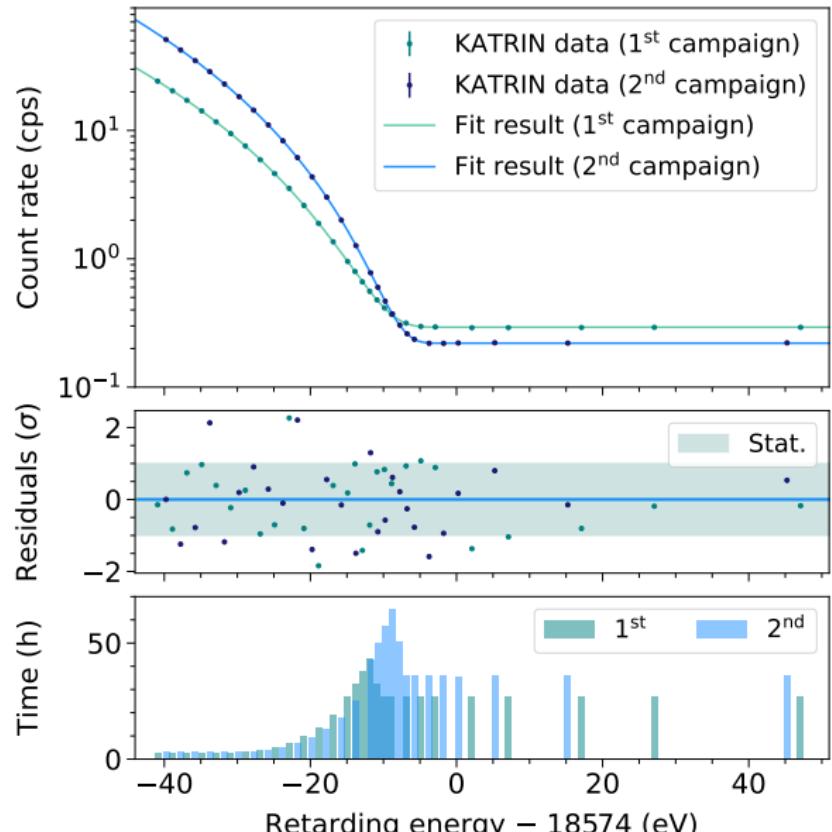
$$m_\nu^2 = 0.08 \pm 0.32 \text{ eV}^2$$

Sub-eV upper limit on the neutrino mass

$$m_\nu < 0.75 \text{ eV (90 \% CL)}$$

Nat. Phys. **18**, 160–166 (2022)

Supplementary material



# (Further) BSM analyses

## eV-sterile neutrinos

Constrain parameter space of eV-steriles from spectrum shape

- First measurement campaign: Phys. Rev. Lett. **126** (2021) 2011.05087
- First and second measurement campaigns: Phys. Rev. D **105** 072004

## keV-sterile neutrinos

Constrain parameter space of keV-steriles with dedicated measurements over larger energy range

## Cosmic relic neutrinos

Constrain local overdensity of relic neutrinos from peak search

- First and second measurement campaigns: arXiv: 2202.04587

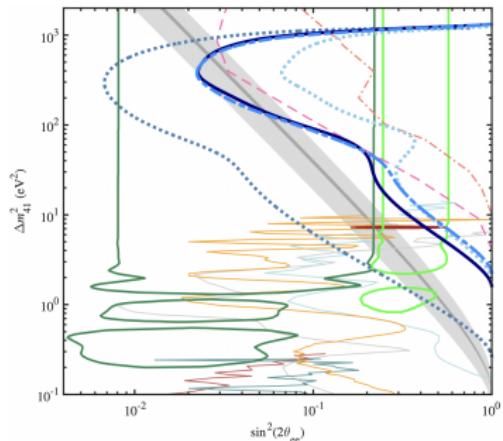
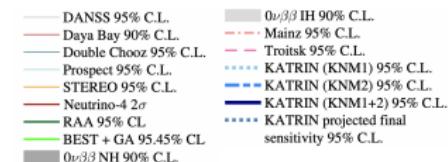
## Lorentz invariance violation

Constrain LIV from sidereal modulation of tritium endpoint

- First measurement campaign: In preparation

## Right-handed currents

Constrain exotic weak interactions from spectrum shape



Exclusion curve (mass vs. mixing angle parameter space) for eV-sterile neutrinos

# Summary

Leading upper limit on the neutrino mass from direct single  $\beta$ -decay measurements

→ KATRIN combined analysis of KNM1 and KNM2 measurement campaigns

$$m_\nu < 0.75 \text{ eV} \text{ (90 \% CL)}$$

Nat. Phys. 18, 160–166 (2022)

Towards improved sensitivity

→ KATRIN combined analysis of KNM1 to KNM5 measurement campaigns is ongoing

→ Expected sensitivity  $< 0.5 \text{ eV}$

