

Exploring the Dark Universe with the Simons Observatory



Steve Choi (Cornell)
for the Simons Observatory Collaboration



Rencontres de Blois
May 25, 2022



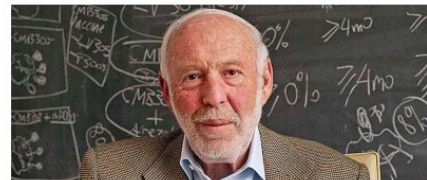
Photo: Debra Kellner

Simons Observatory (SO)

SIMONS
FOUNDATION

HEISING-SIMONS
FOUNDATION

>10 countries, >300 people



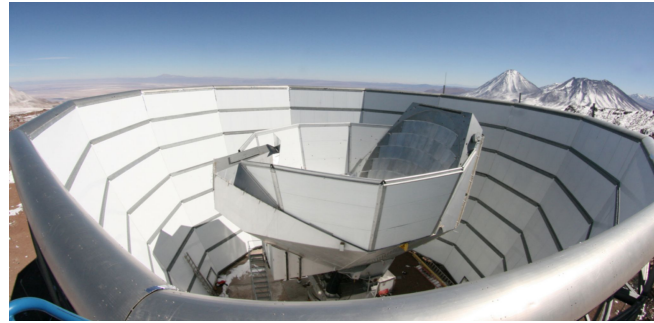
Simons Observatory Collaboration Summer 2018



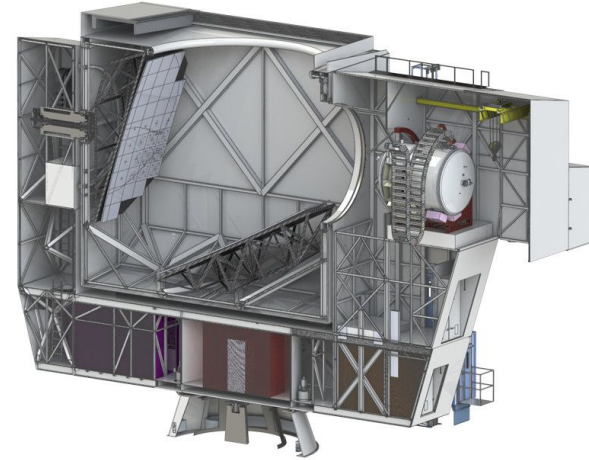
Planck



Atacama Cosmology Telescope



Simons Observatory



2009 – 2013

Full sky

0.35 – 10 mm (9 bands)

5 – 33' resolution

2008 – 2022+

~40% sky fraction

Noise ~1.5–3 times better than Planck

1.4 – 10 mm (5 bands)

1 – 7' resolution

2023 – 2030+

~40% sky fraction

Noise ~3 times better than ACT

1 – 10 mm (6 bands)

1 – 7' resolution

Summary of SO key science goals

Parameter	SO-Baseline ^b (no syst)	SO-Baseline ^c	SO-Goal ^d	Current ^e	Method	Sec.	
Primordial perturbations	r	0.0024	0.003	0.002	0.03	$BB + \text{ext delens}$	3.4
	$e^{-2\tau} \mathcal{P}(k = 0.2/\text{Mpc})$	0.4%	0.5%	0.4%	3%	$TT/TE/EE$	4.2
	$f_{\text{NL}}^{\text{local}}$	1.8	3	1	5	$\kappa\kappa \times \text{LSST-LSS} + 3\text{-pt}$	5.3
		1	2	1		$\text{kSZ} + \text{LSST-LSS}$	7.5
Relativistic species	N_{eff}	0.055	0.07	0.05	0.2	$TT/TE/EE + \kappa\kappa$	4.1
Neutrino mass	Σm_ν	0.033	0.04	0.03	0.1	$\kappa\kappa + \text{DESI-BAO}$	5.2
		0.035	0.04	0.03		$\text{tSZ-N} \times \text{LSST-WL}$	7.1
		0.036	0.05	0.04		$\text{tSZ-Y} + \text{DESI-BAO}$	7.2
Deviations from Λ	$\sigma_8(z = 1 - 2)$	1.2%	2%	1%	7%	$\kappa\kappa + \text{LSST-LSS}$	5.3
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	$H_0 (\Lambda\text{CDM})$	0.3	0.4	0.3		0.5	$TT/TE/EE + \kappa\kappa$
Galaxy evolution	η_{feedback}	2%	3%	2%	50-100%	$\text{kSZ} + \text{tSZ} + \text{DESI}$	7.3
	p_{nt}	6%	8%	5%		$\text{kSZ} + \text{tSZ} + \text{DESI}$	7.3
Reionization	Δz	0.4	0.6	0.3	1.4	$TT (\text{kSZ})$	7.6

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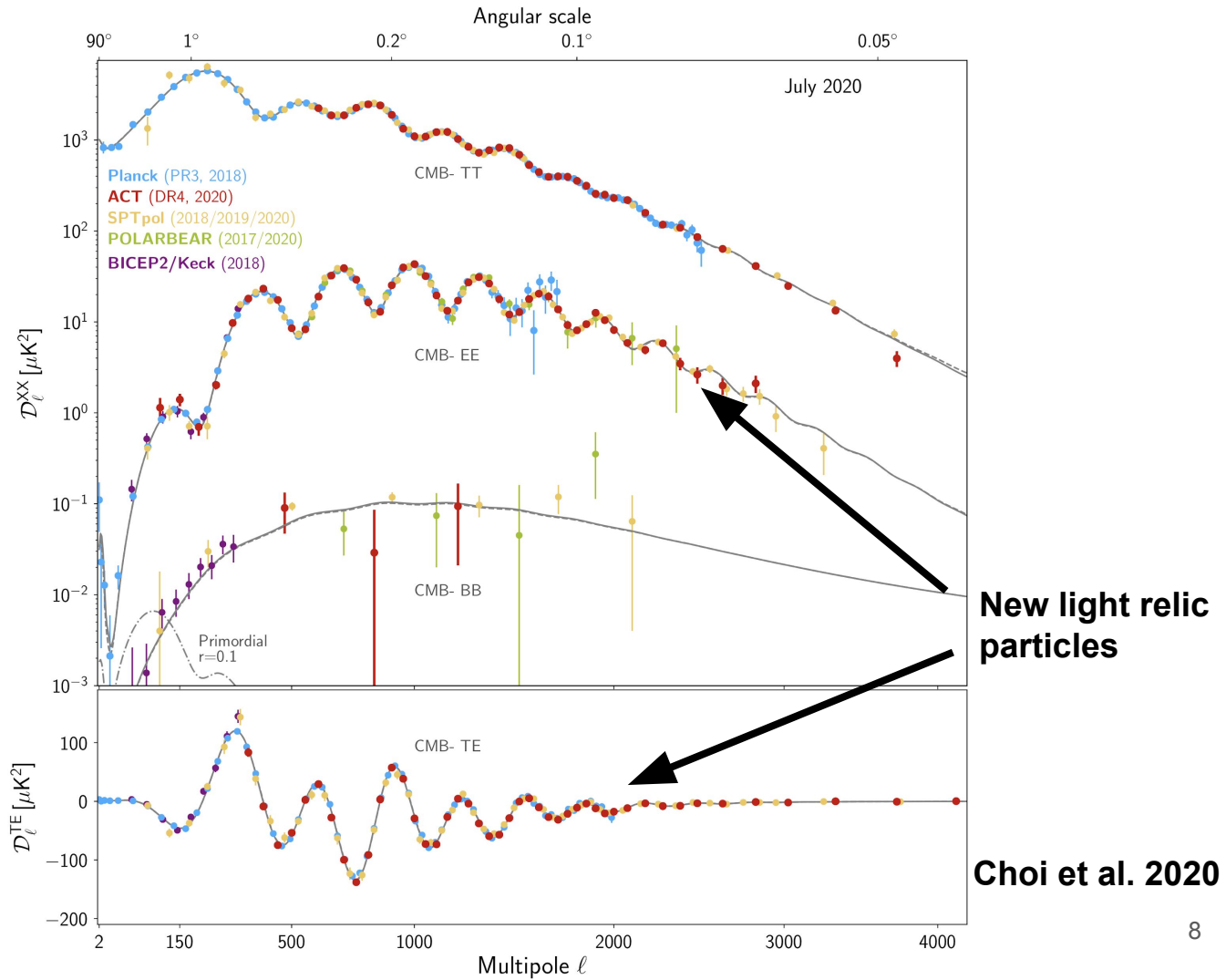
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How many relativistic species are there?

		1	2	1		kSZ + LSST-LSS	7.5
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Fundamental physics from the CMB with SO



Summary of SO key science goals

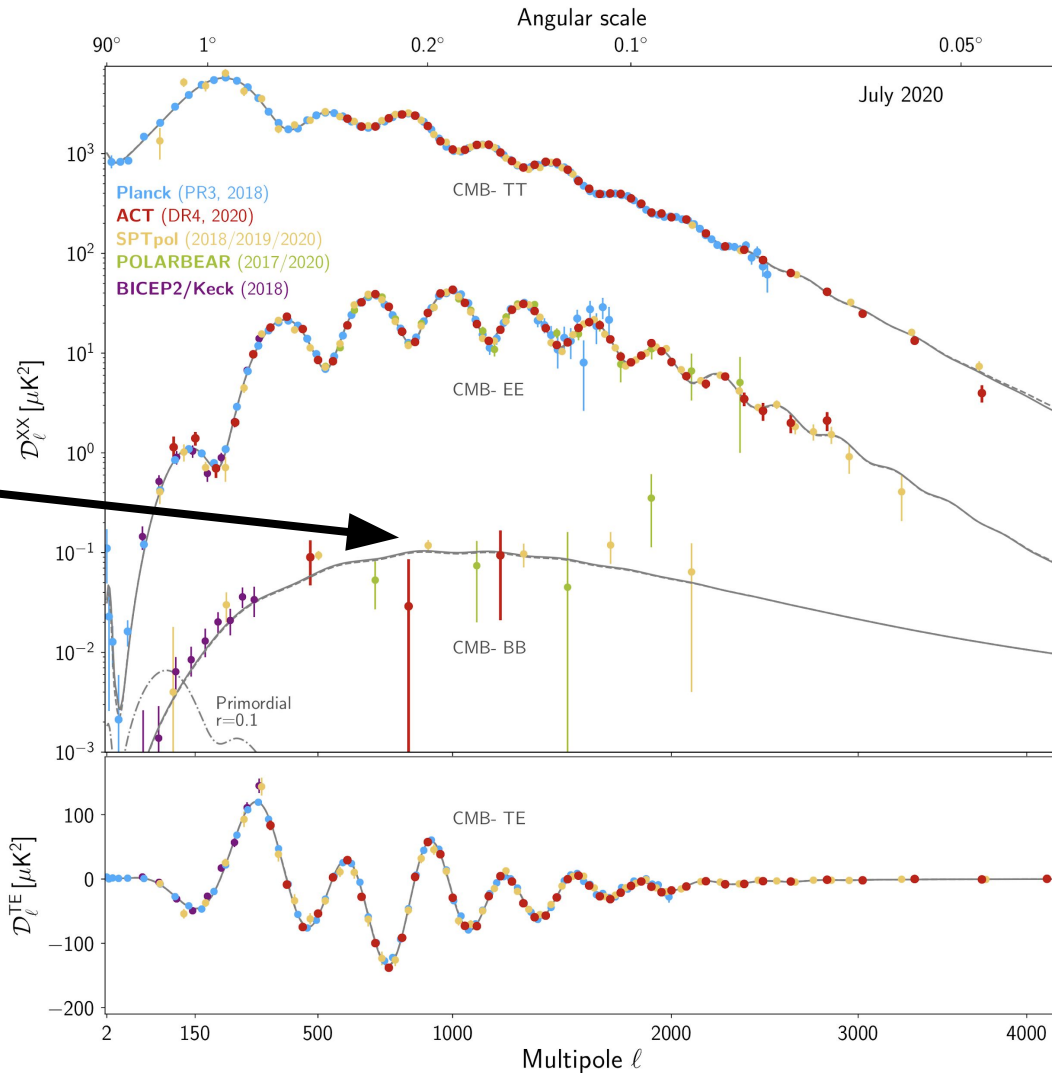
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What is the mass sum of the neutrinos?

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Fundamental physics from the CMB with SO

Neutrino mass sum from CMB lensing due to large scale structure



Choi et al. 2020

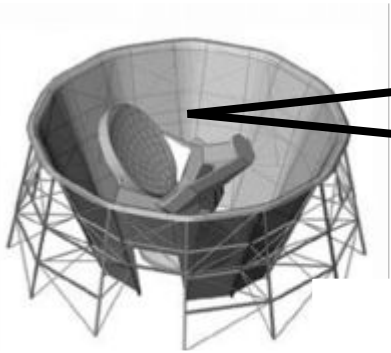
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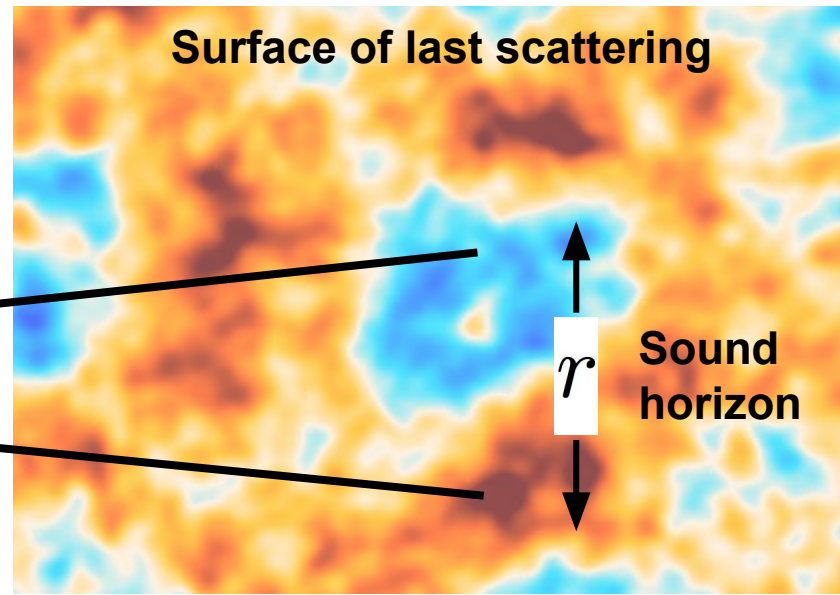
Is the accelerated expansion just Λ ?

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Hubble constant from the CMB

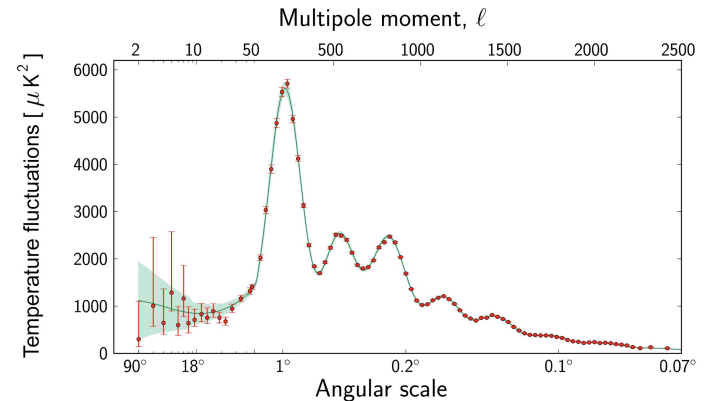


θ
 D



$$D_A^* = r_s^* / \theta^*$$

bigger distance = older universe
older universe = smaller expansion rate



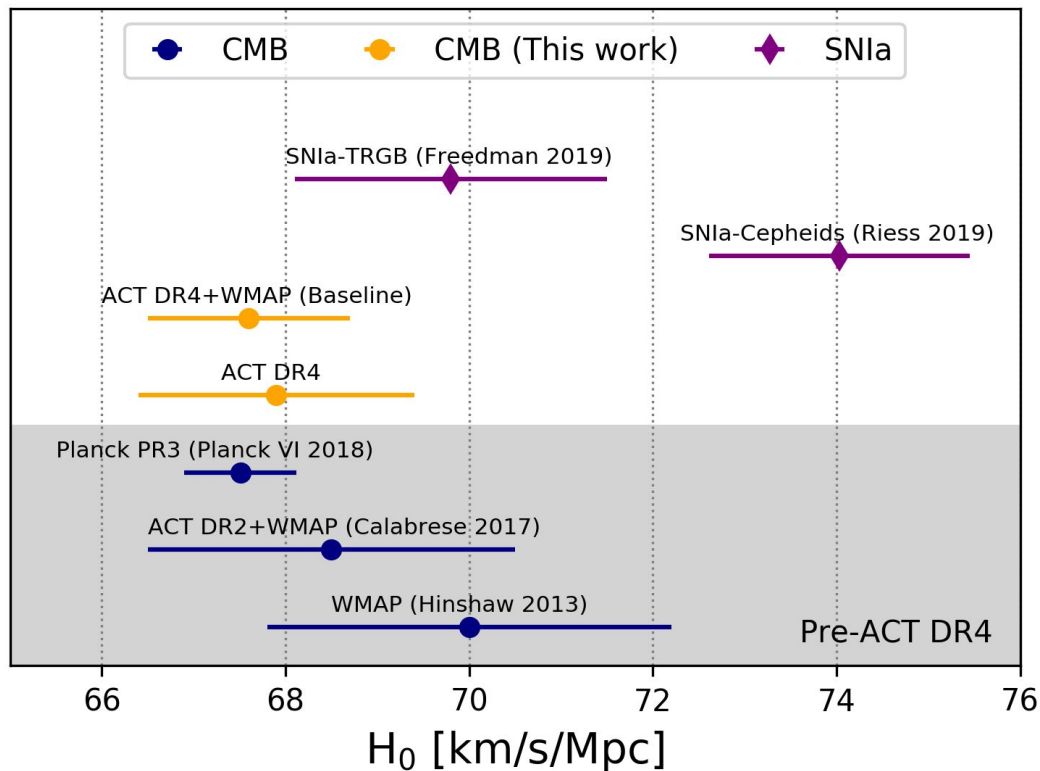
ACT DR4 H_0 Estimate (blind analysis)

ACT DR4 + WMAP: $H_0 = 67.6 \pm 1.1$ km/s/Mpc

- agrees with Planck within 1σ
- agrees with SNIa-TRGB within 1σ
- $\sim 4\sigma$ tension with SNIa-Cepheids

Since then, SPT also measured

$H_0 = 68.8 \pm 1.5$ km/s/Mpc (Dutcher et al. 2021)

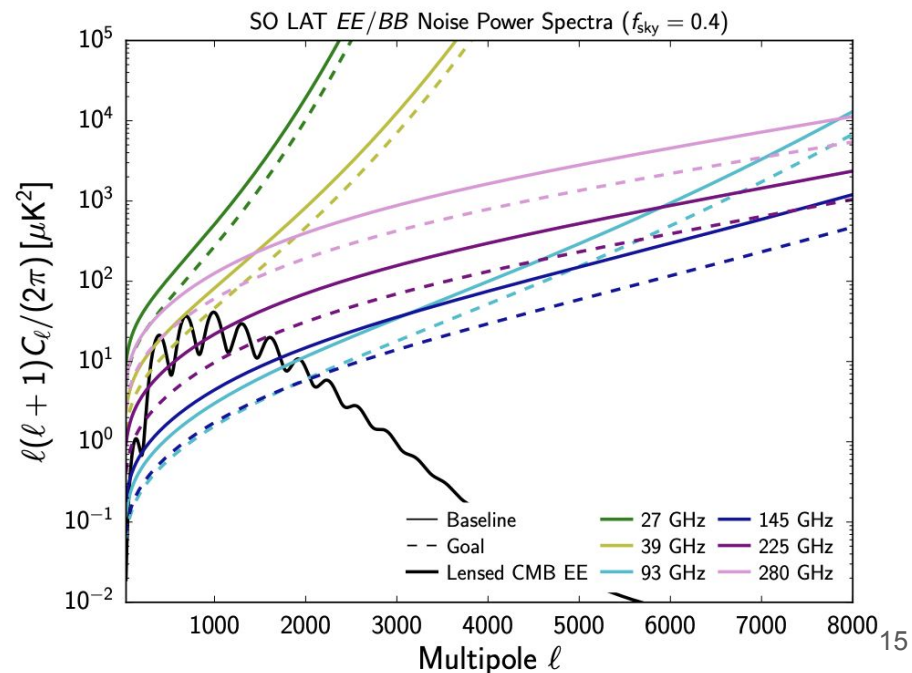
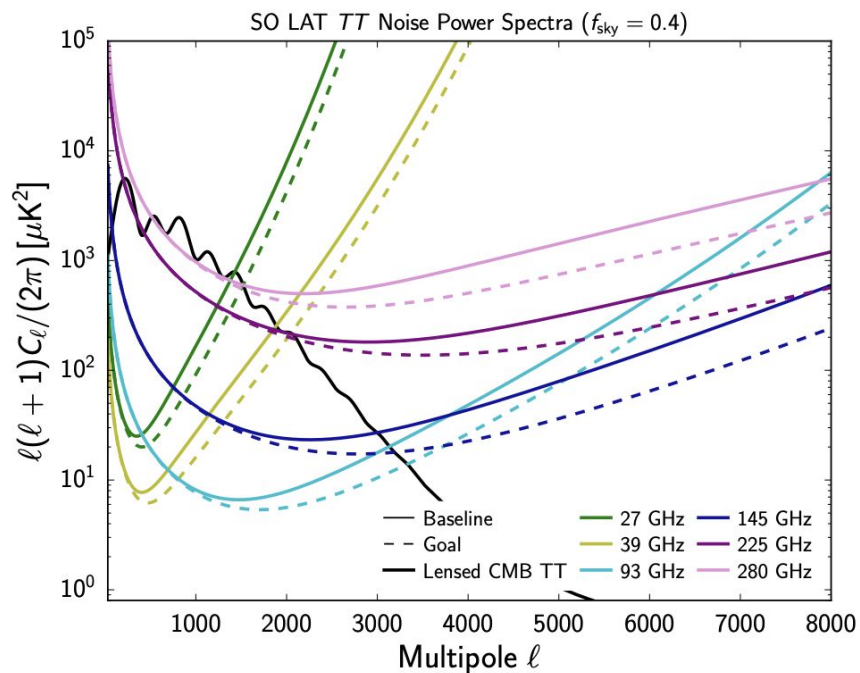


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SO noise curves

With realistic calculations/assumptions on instrument sensitivities and atmospheric noise.

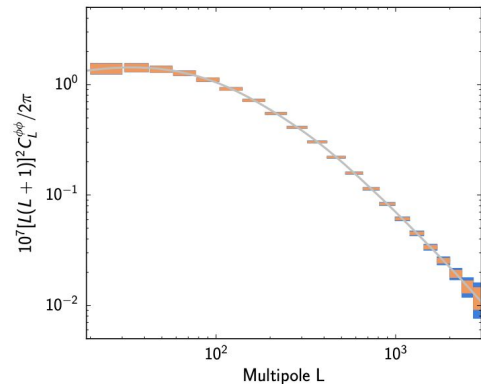
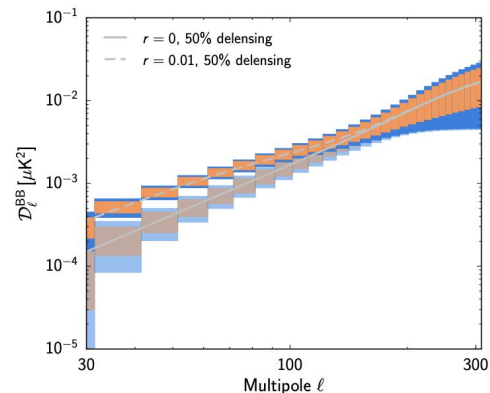
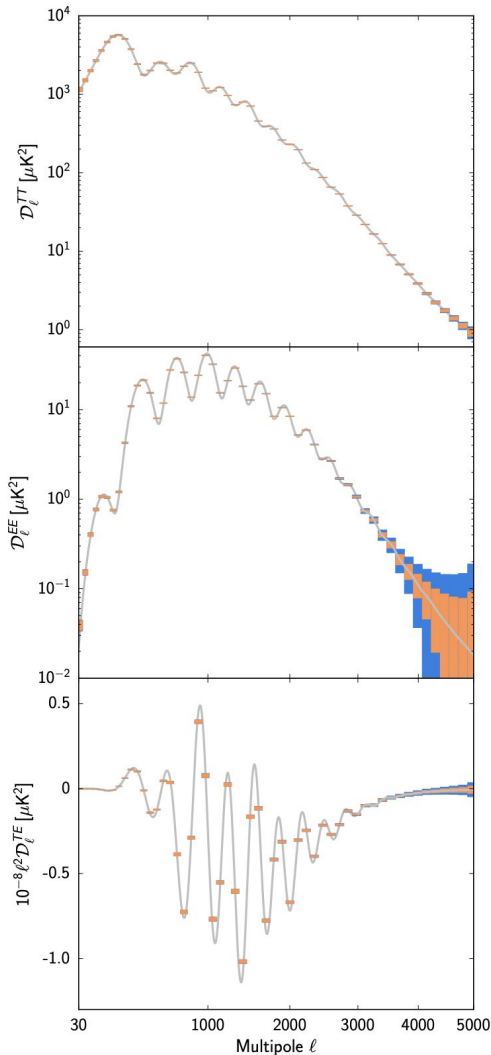


SO simulated power spectra

Need to measure many modes to high precision to study the dark universe with SO.

→ Need a large sky area survey with a large telescope.

SO science overview
(arXiv:1808.07445)

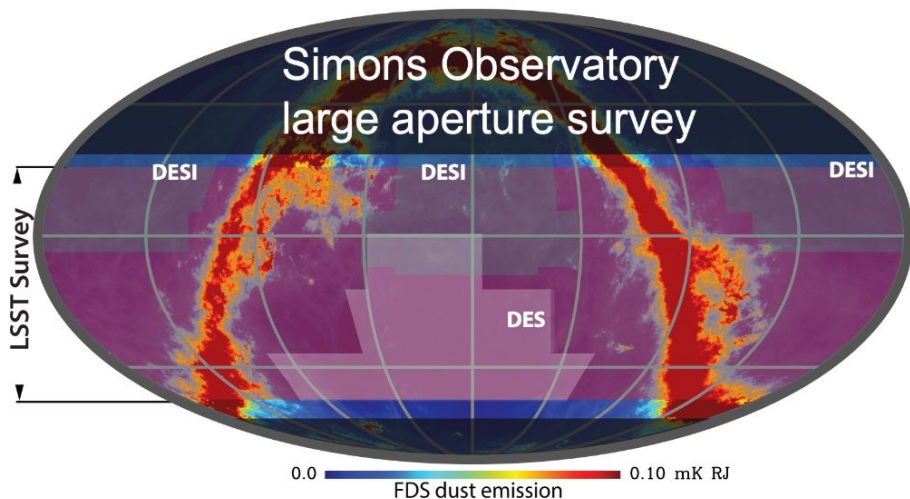


SO Baseline
SO Goal

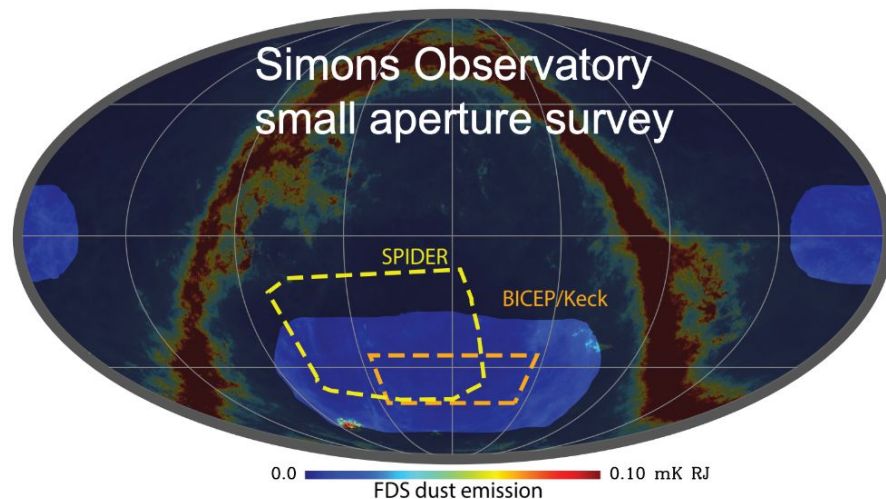
Simons Observatory Survey

SO science overview
(arXiv:1808.07445)

Large Aperture Telescope (LAT)

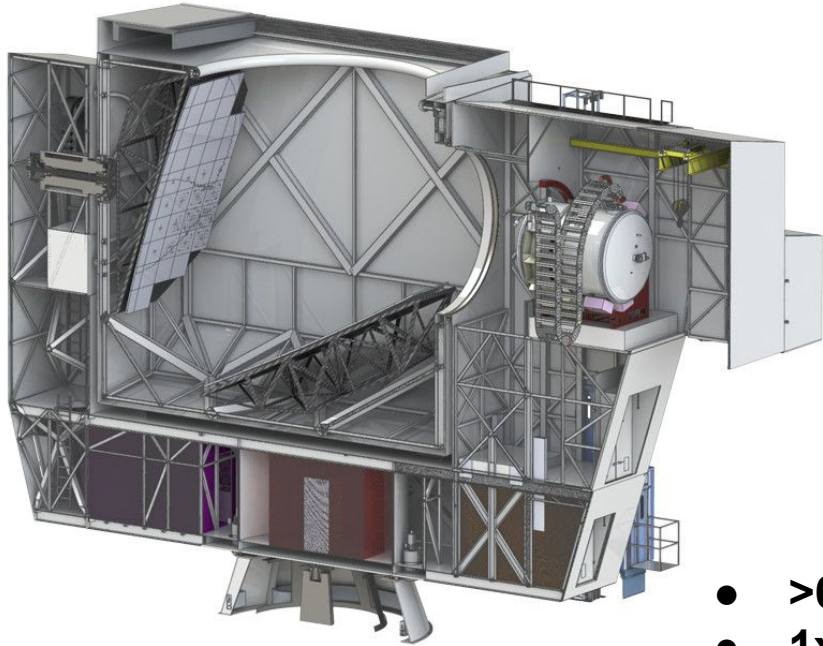


Small Aperture Telescope (SAT)

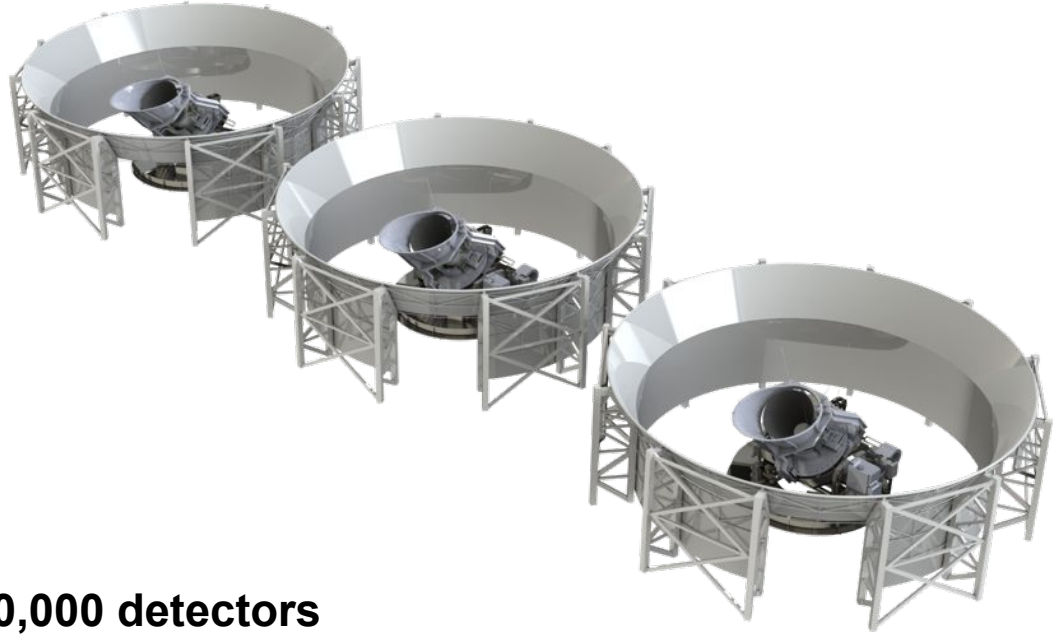


SO telescopes

Large Aperture Telescope



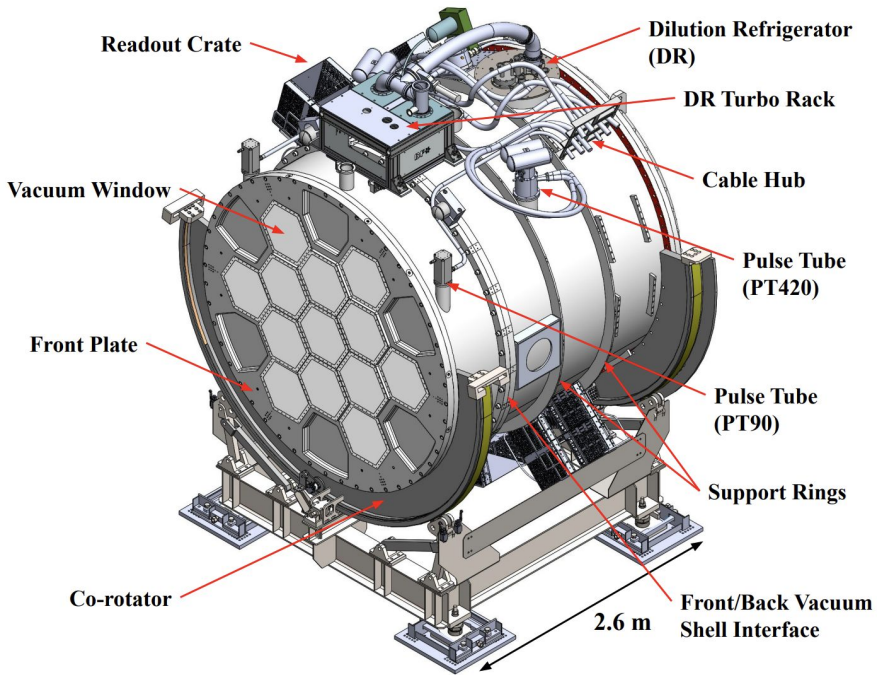
Small Aperture Telescopes



- >60,000 detectors
- 1x 6-m large aperture telescope
- 3x 42-cm small aperture telescopes

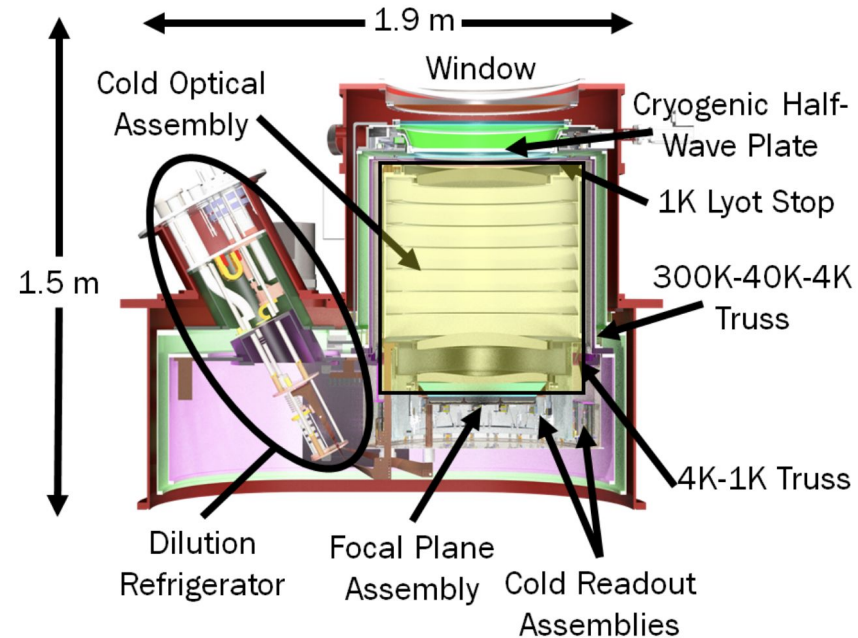
SO telescope receivers

Large Aperture Telescope Receiver



Zhu et al. 2020

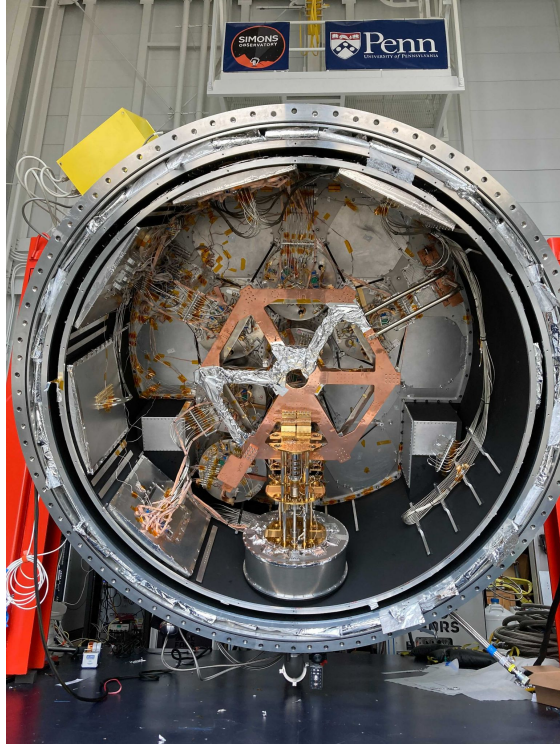
Small Aperture Telescope



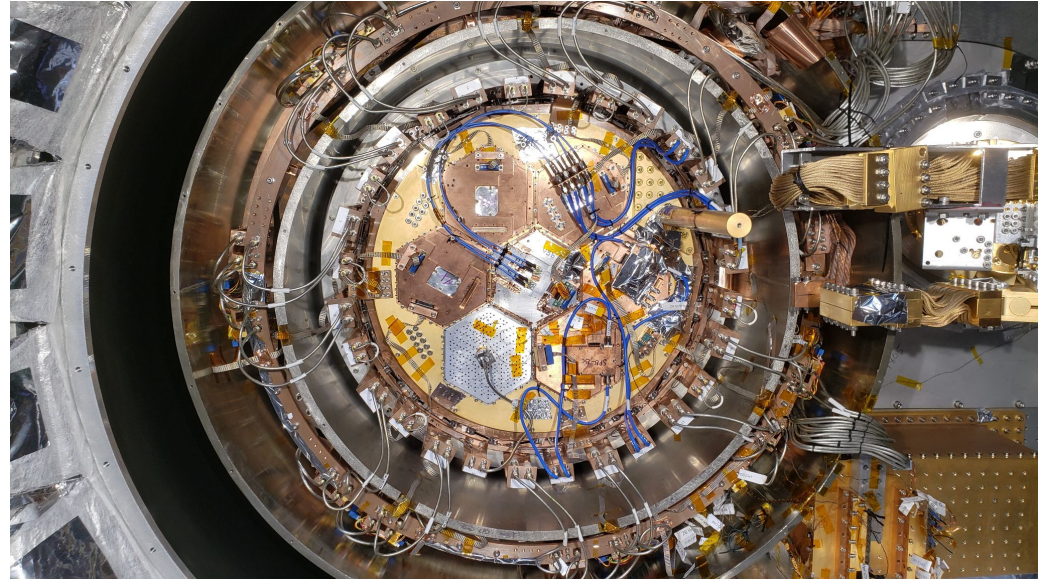
Ali et al. 2020

S0 integration and testing

LATR at UPenn

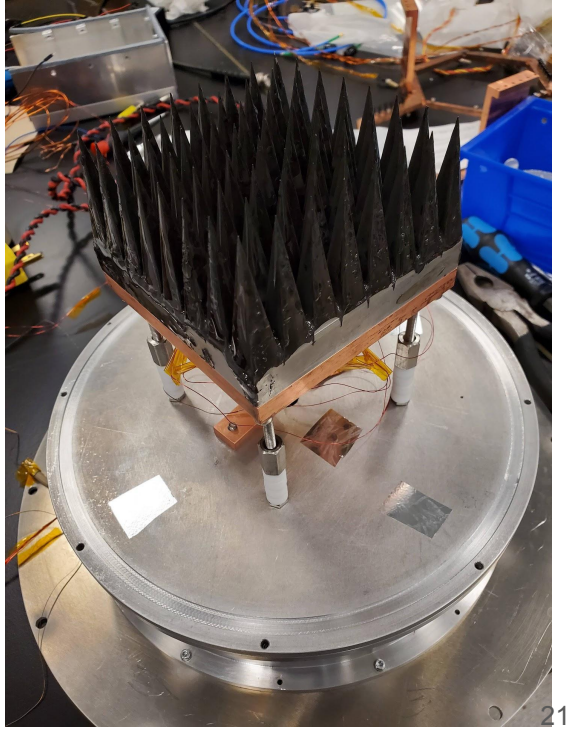
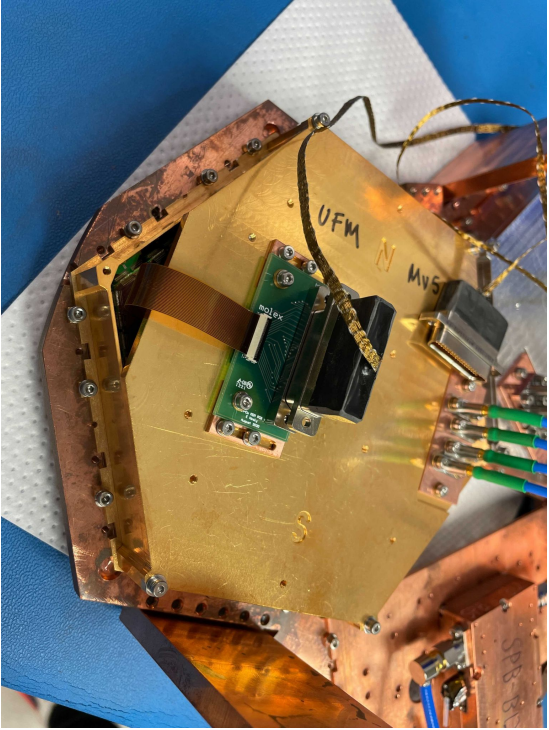


SAT-MF1 at UCSD



S0 integration and testing

Detector testing at Princeton and Cornell



S0 site and telescope/platform status



S0 site and telescope/platform status



S0 site and telescope/platform status





Photo: Debra Kellner

CCAT-prime

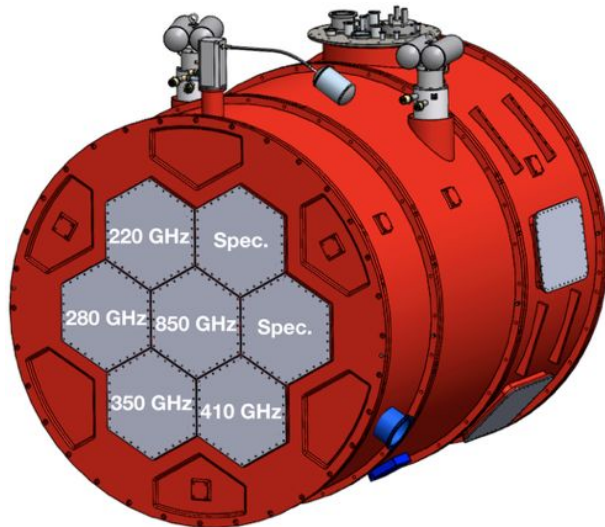
Fred Young Sub-mm Telescope



Better telescope: same as SO LAT but **better mirror surface accuracy** for high frequency observations. Observing at a superior 5,600-m site.

Parshley et al. 2018
Stacey et al. 2018

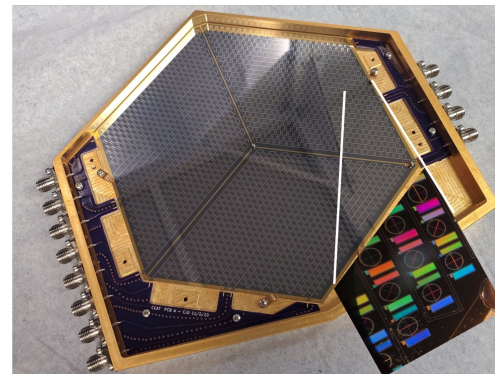
Prime-Cam receiver



Unique instrument: 1.8-m diameter, 7-module receiver under construction. **High frequency** broadband channels & **spectrometers**.

Vavagiakis et al. 2018
Choi et al. 2020

280 GHz MKID array



New detectors: Microwave kinetic inductance detectors (MKIDs)
→ **Allows for >100,000 detectors** in the Prime-Cam receiver.

Duell et al. 2020
Choi et al. 2021

Conclusion

- State-of-the-art cosmology from Planck.
- Existing data in hand (e.g., ACT) will soon improve on constraining cosmological parameters beyond Planck.
- With rapid progress in mm and sub-mm instrumentation, Simons Observatory (and CCAT-prime) will soon enable next generation cosmological studies of the dark universe.