

# DarkSide-20k and the Future Liquid Argon Dark Matter Program

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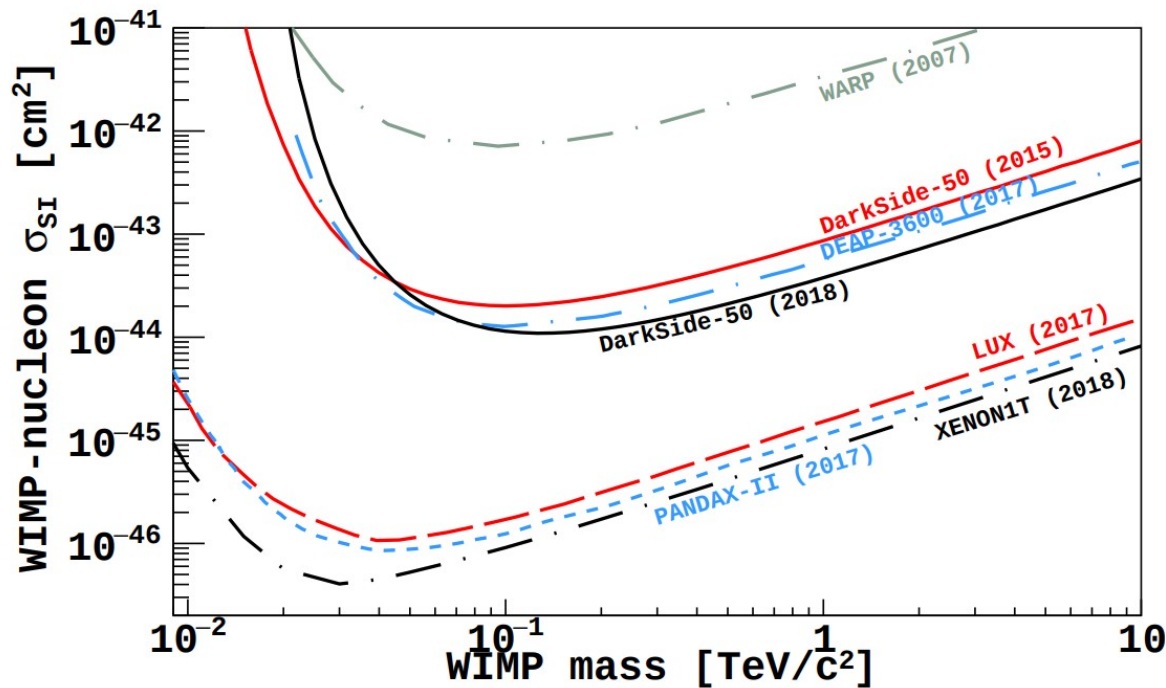
*Julie Rode (LPNHE/APC)  
on behalf of the DarkSide Collaboration*

*25/05/2022*

# Motivation

## Explore direct detection for WIMP Dark Matter down to the neutrino floor

- Build on existing DarkSide-50 technology
  - Low instrumental background
  - Dual phase argon time projection chamber



DOI: 10.48550/arXiv.1802.07198

### Noble liquid detectors assets:

- Efficient background discrimination
- Massive target
- High scintillation yield

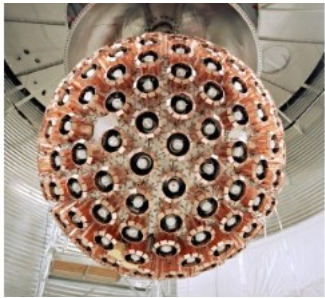
### Liquid argon detectors assets:

- Massive
- Radiopure
- High scintillation yield
- High ionization yield
- Low electron mobility
- Argon mass  $\ll$  Xenon mass
- Higher recoil energy (transferred momentum) wrt Xe at low energy

- Upgrade to a **20 t detector** to **gain exposure** and hence **improve sensitivity** to dark matter

# The Global Argon Dark Matter Collaboration

## Four Argon-based Experiments gathered in GADMC



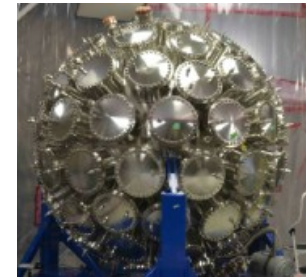
DEAP-3600



ArDM



DarkSide-50

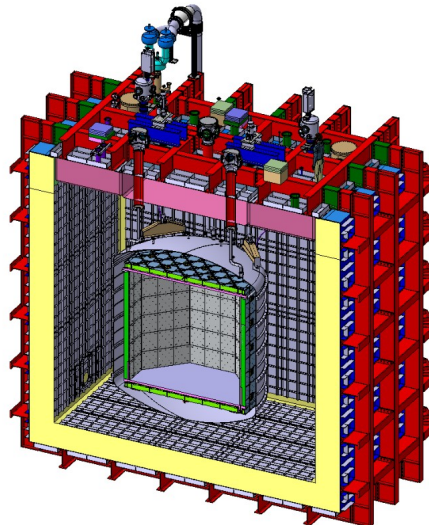


MiniClean

*~ 500 people  
100 institutions*



## DarkSide-20k

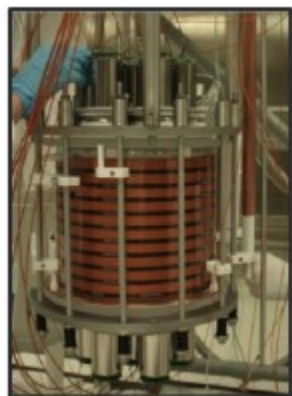


*Projects:*  
- DS-Lowmass  
- Argo (300 t)

# DarkSide Experiment

## Direct Detection of Dark Matter using Liquid Argon

- Argon dual-phase Time-Projection Chamber (TPC)
- Direct detection by nuclear or electronic scattering
- Background free at high WIMPs masses (Pulse Shape Discrimination)
- Modeled background for low DM masses studies ( $< 10 \text{ GeV}/c^2$ )



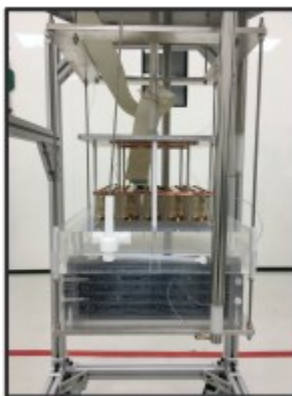
DS-10

2011 - 2013



DS-50

2013 - 2020



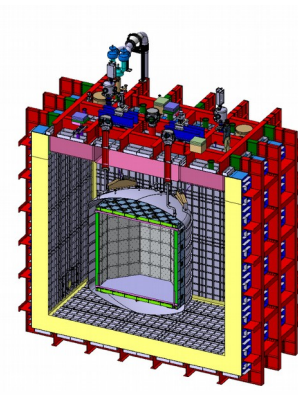
DS-Proto0

2019 - 2022



DS-Proto1T

2022 - ...



DS-20k

2025 - ...

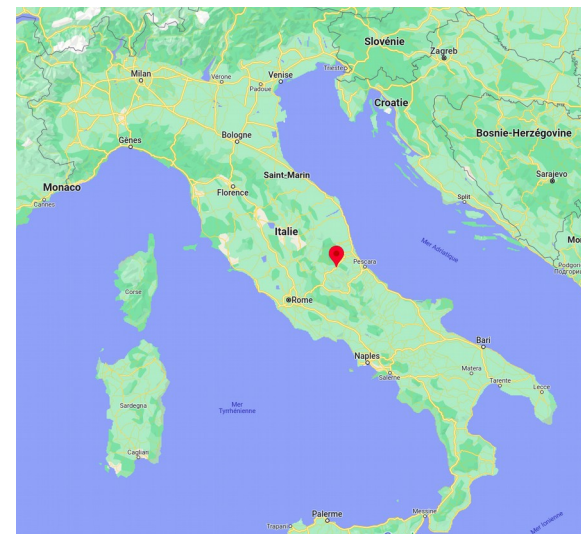
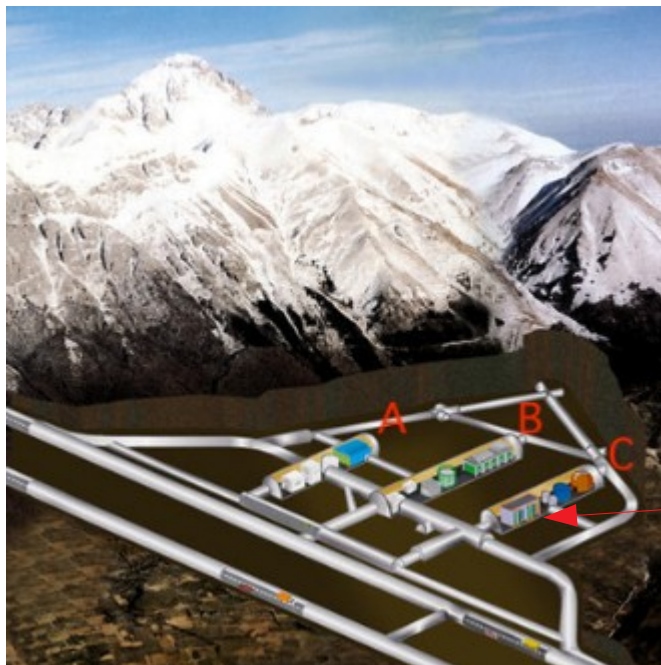
Argo

...

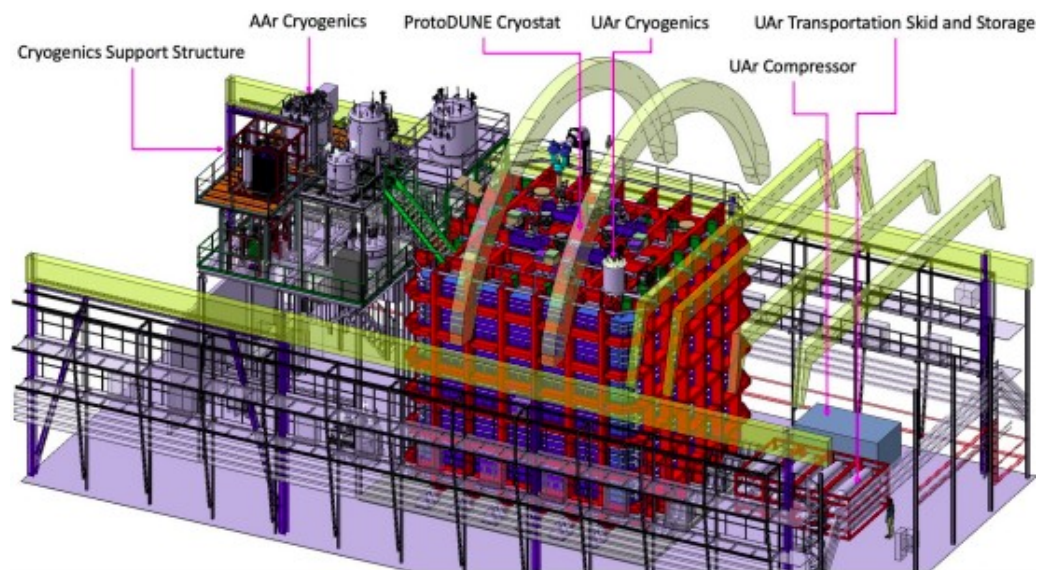
# Laboratori Nazionali del Gran Sasso (LNGS)

## Underground Research Facility

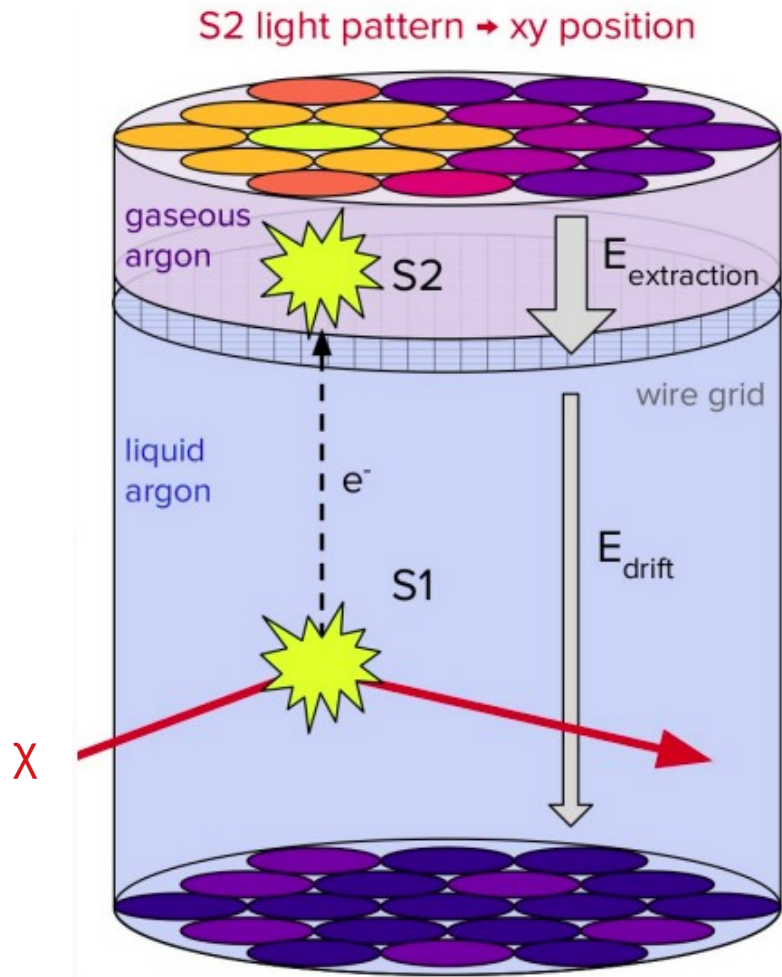
→ 1400 m of rock above the facility



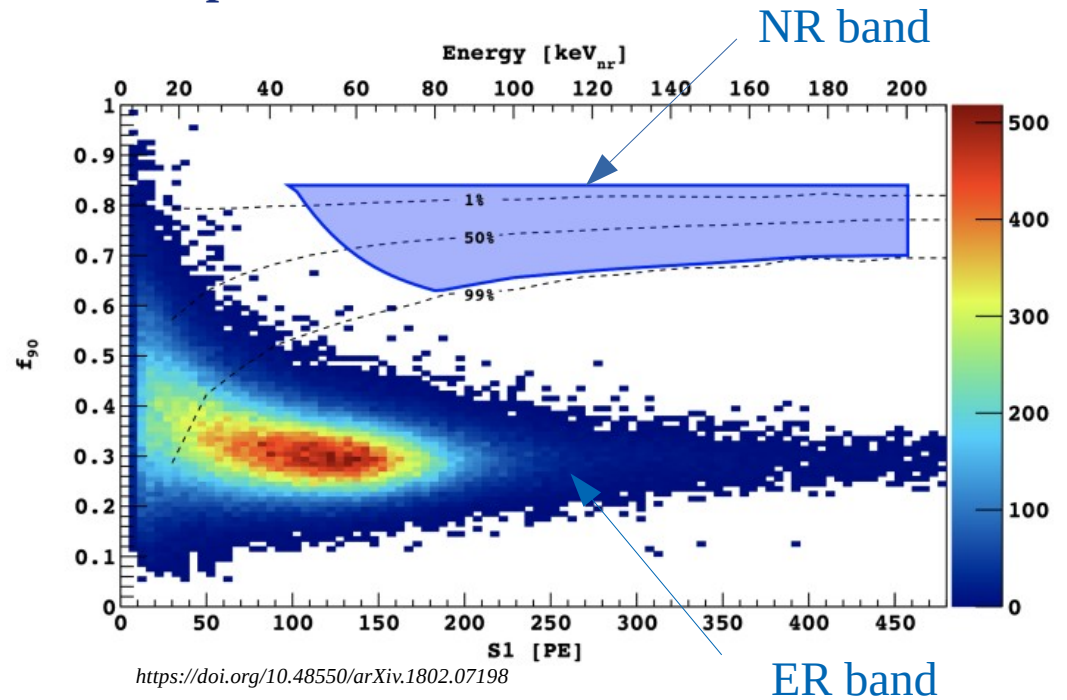
DarkSide-20k will be in Hall-C



# Time Projection Chamber



## Pulse Shape Discrimination

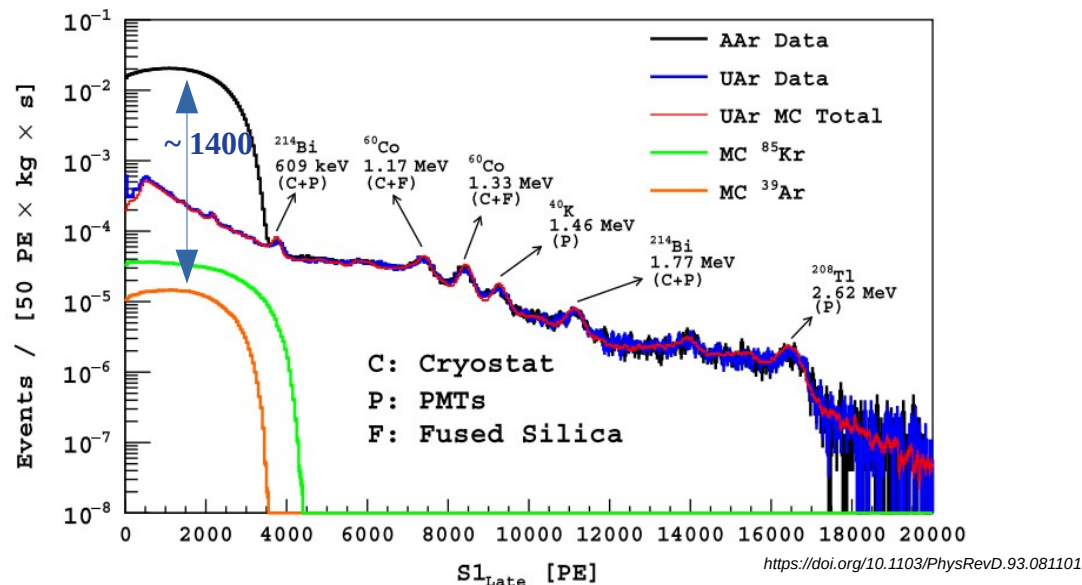


- Discrimination between electron recoils (ER) and nuclear recoils (NR)
- $\beta$ ,  $\gamma$  rejection background
- Based  $f_{90}$ : the ratio of scintillation light in the first 90 ns compared to the total
  - ER  $\sim$  0.3
  - NR  $\sim$  0.7

# Underground Argon (UAr)

- $^{39}\text{Ar}$  radioactivity in **atmospheric** argon:
  - $\beta$ -emitter with an endpoint of 565 keV
  - activity  $\sim 1\text{Bq/kg}$
- $^{39}\text{Ar}$  **cosmogenic** isotope
  - Lower  $^{39}\text{Ar}$  production rate in UAr

Using **UAr** allows to have a **reduction factor of background  $\sim 1400$** .



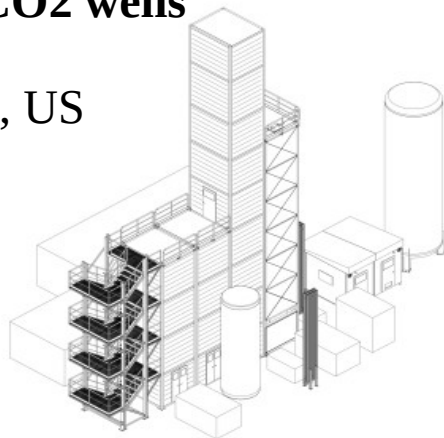
Thanks to UAr + PSD:  
We expect **< 0.1 events/(200t year)** of residual  
background due to electron recoils

# UAr: Urania / Aria / DArT

## Urania

### Extraction of UAr from CO2 wells

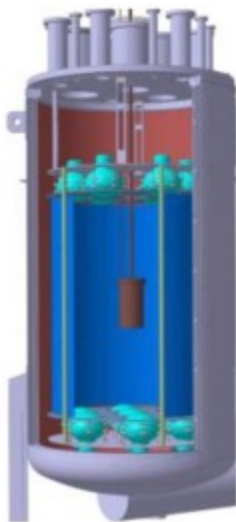
- Plant in Cortez, Colorado, US
- Can extract 330 kg/day
- Purity 99.99%



## DArT

### Measures $^{39}\text{Ar}$ depletion factor

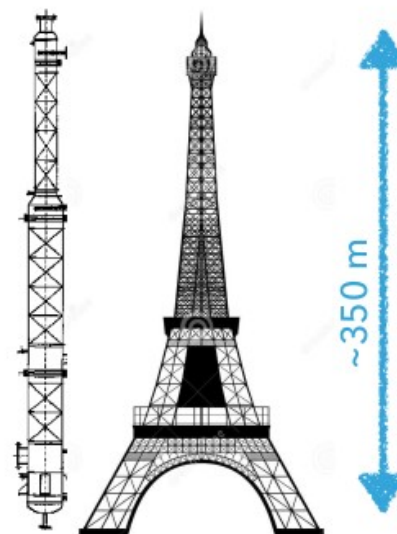
- At LSC, Canfranc, Spain
- Radiopure single phase LAr inner detector (1.42 kg LAr)
- Veto: Inside ArDM detector (1t LAr)



## Aria

### Perform chemical and isotopic purification of UAr

- At Seruci mine in Sardinia, Italy
- Distillation Column, rate 1t/day
- A  $^{39}\text{Ar}$  reduction factor 10 expected per pass
- Medical applications (oxygen isotope separation for instance)
- Assembly of the final column will be finish this year





# Photo Detection

Single-Photon  
Avalanche Diode  
(SPAD) – 25-30  $\mu\text{m}^2$



Silicon Photomultipliers  
(SiPM) – 7.9 x 11.7  $\text{mm}^2$   
*Composed by SPADs*

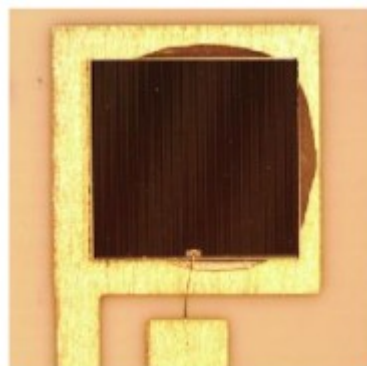


Photo Detector Module  
(PDM) – 5 x 5  $\text{cm}^2$   
*Composed by 24 SiPMs*

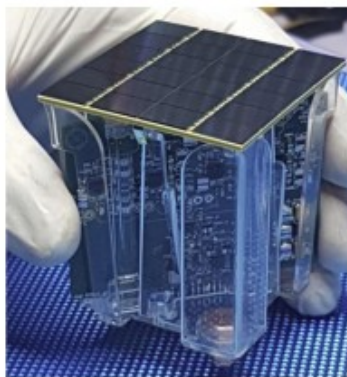
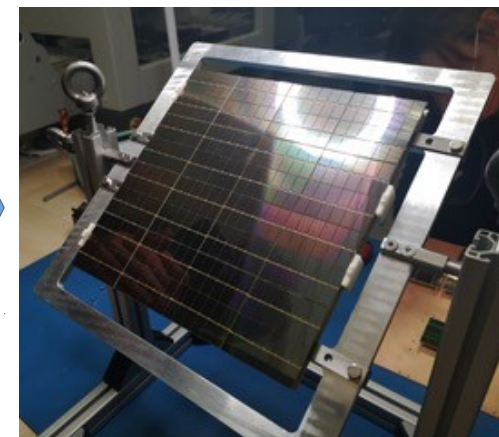


Photo Detection Unit  
(PDU) – 20x20  $\text{cm}^2$   
*composed by 4x4 PDMs*

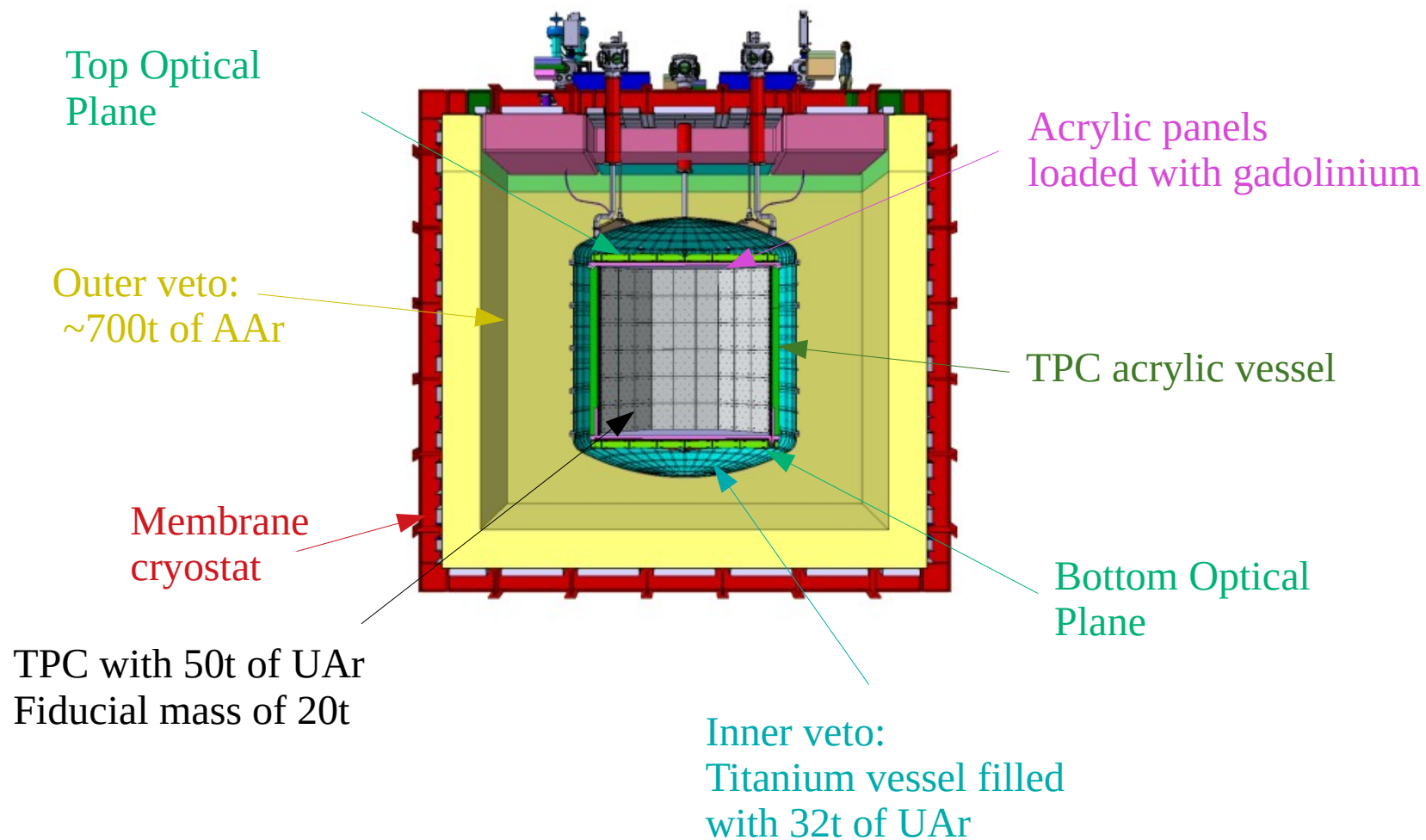


## SiPMs

- Developed with the Fondazione Bruno Kessler
  - 8448 tiles for the TPC
  - 1920 tiles for the Veto
- Key features:
  - Low dark count rate < 20 cps
  - Timing resolution ~ 10 ns
  - Photon detection efficiency ~ 45 %
  - Radiopure ~ 2mBq/PDM

# DarkSide-20k: Design

## Nested structure detector

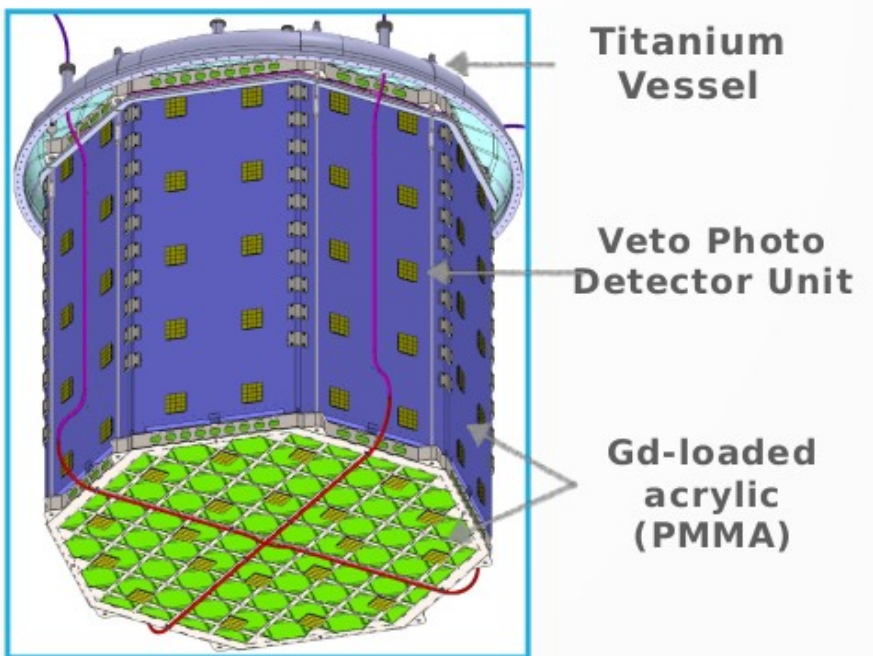
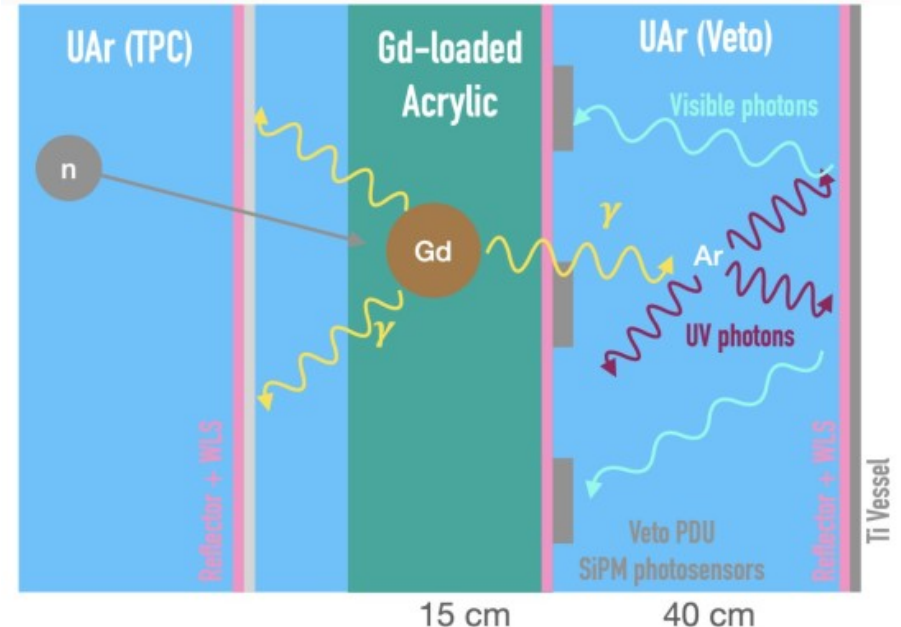


21 m<sup>2</sup> of SiPMs in the TPC  
5 m<sup>2</sup> cryogenic SiPMs in the Veto

# DarkSide-20k: Inner Veto

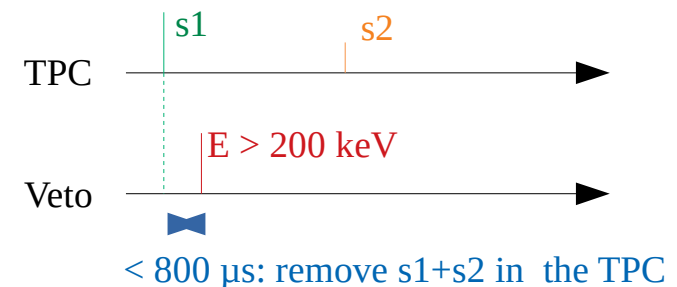
## Neutron Veto Principle

- Neutrons are captured by the Gd-loaded Acrylic
- Gd emits  $\gamma$  rays up to 8 MeV
- They interact with UAr
- The scintillation light is shifted using TPB
- And then detected by SiPMs



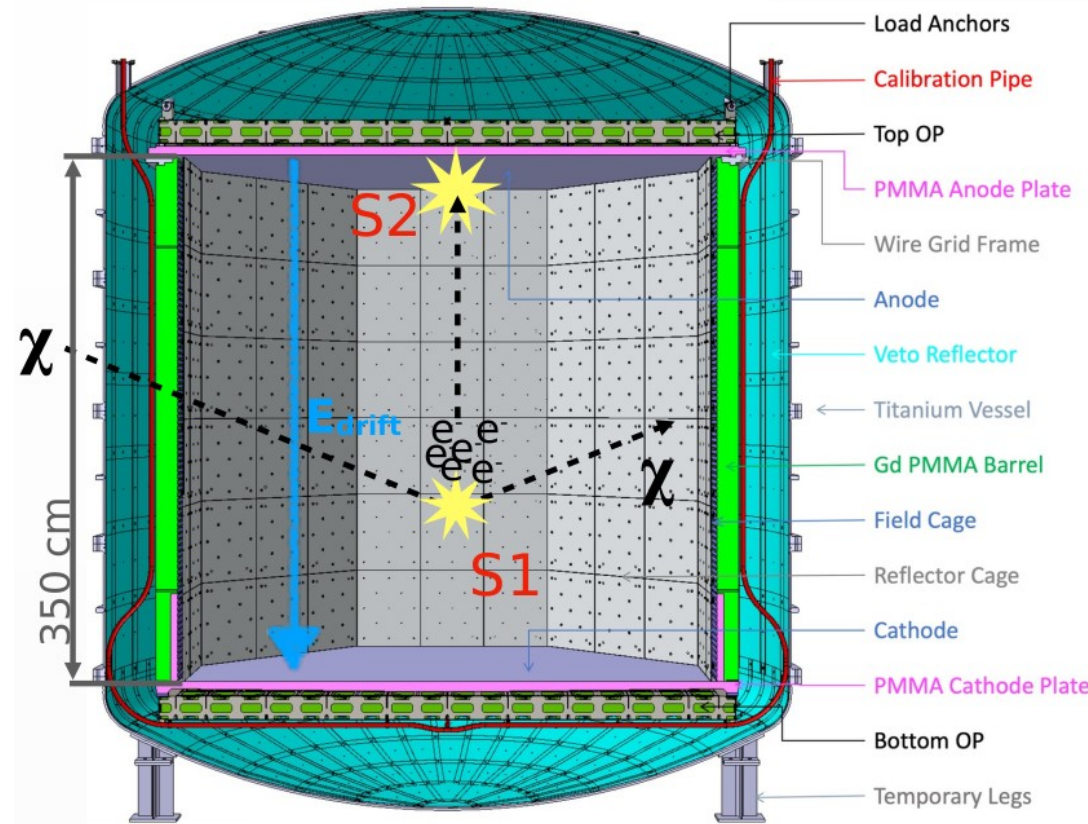
## Veto Condition

- Measured energy of the scintillation light in the veto  $> 200$  keV and within a  $800 \mu\text{s}$  time coincidence with an s1 in the TPC



# DarkSide-20k: TPC

## Design inside the Titanium Vessel:



## Features:

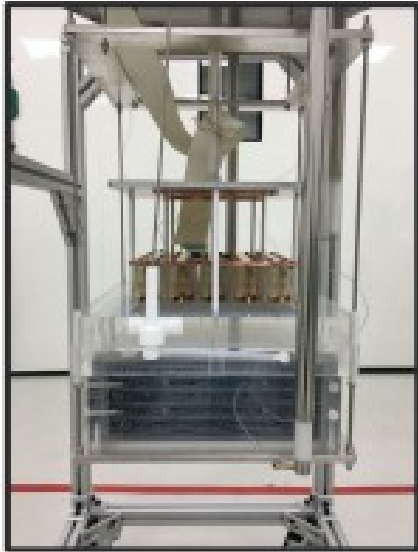
Maximum drift length: 348 cm  
Octagonal inscribed circle diameter: 350 cm  
Gas pocket width:  $7.0 \pm 0.5$  mm  
S1 Light Yield: 10 pe/keV  
S2 Yield:  $> 20$ pe/e<sup>-</sup>

Drift field: 200 V/cm  
Extraction field: 2.8kV/cm  
Luminescence field: 4.2kV/cm

XY resolution:  $< 5$ cm  
Z resolution: 1 mm

UAr mass in TPC: 51.1t  
Vertical fiducial cut: 70 cm  
Radial fiducial cut: 30 cm  
Fiducial UAr mass: 20.2t

# Prototypes



## DS-Proto0

→ At CERN (2019 - 2020) and then at Naples (2022)

### **Goal:**

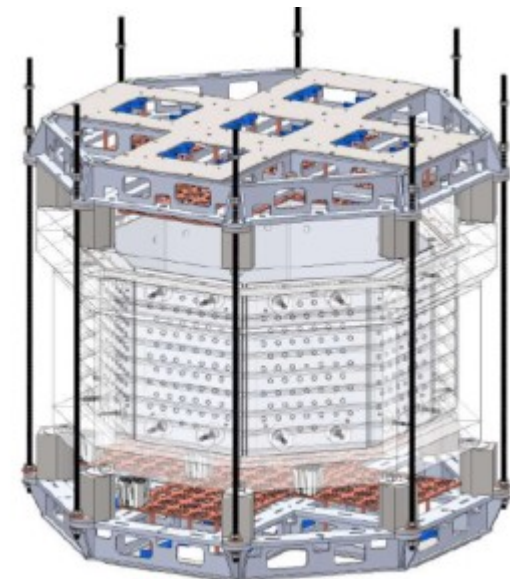
- First test of a full motherboard with SiPMs
- Test of the formation of the gas pocket
- Development of the offline calibration and reconstruction software

## DS-Proto1T

→ At CERN, started in 2022

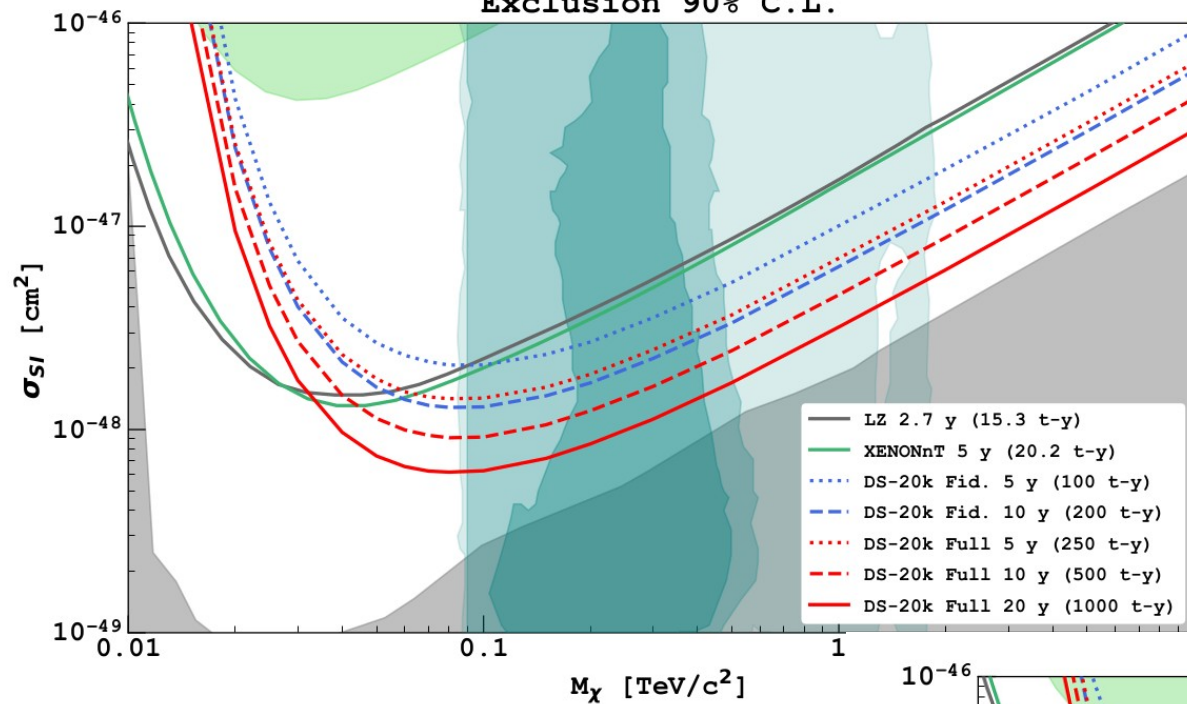
### **Goal:**

- Test of a down-scaled version of DarkSide-20k
- Test of PDUs and of the octagonal acrylic vessel



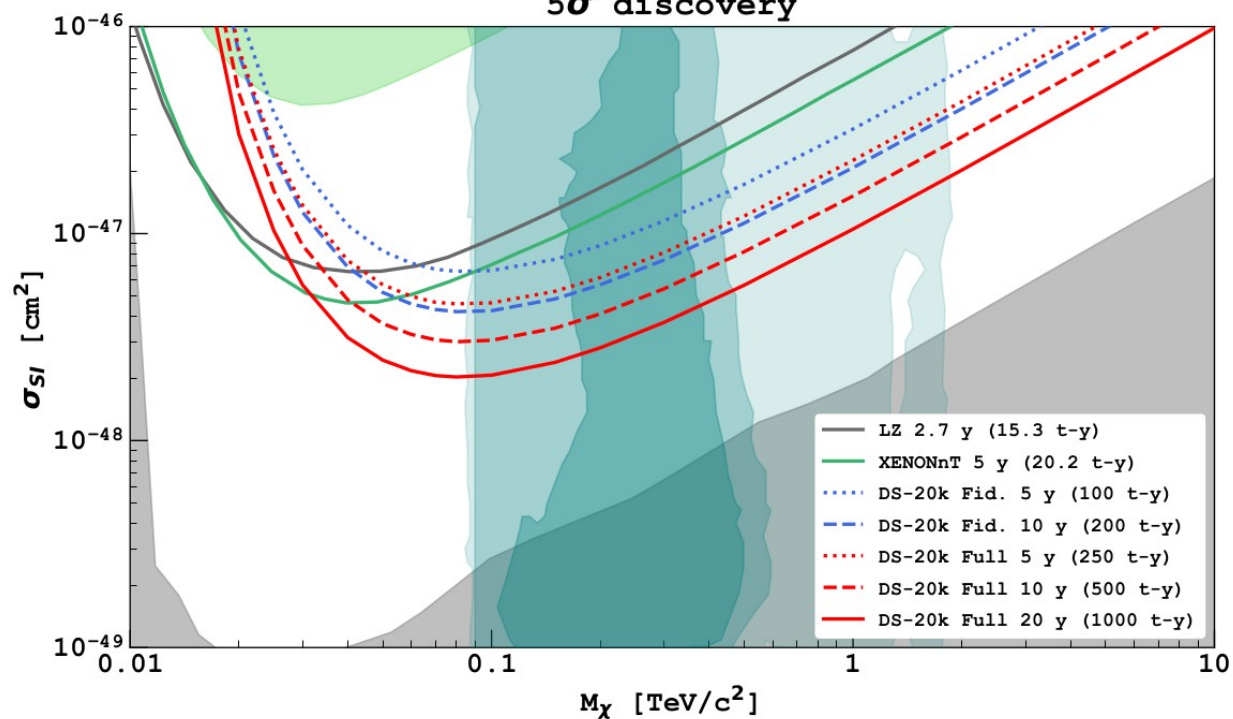
# DarkSide-20k: Sensitivity

Exclusion 90% C.L.



Sensitivity to spin independent  
WIMPs

$5\sigma$  discovery

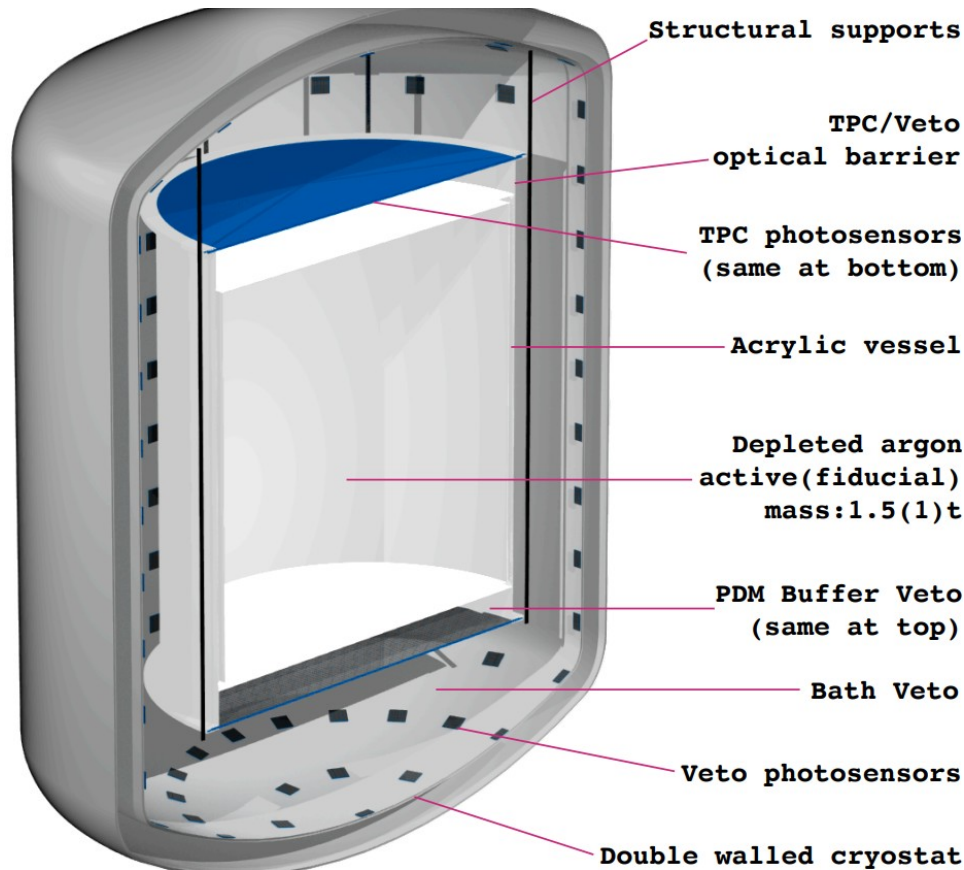


# Side Project: DarkSide-Lowmass

## Design

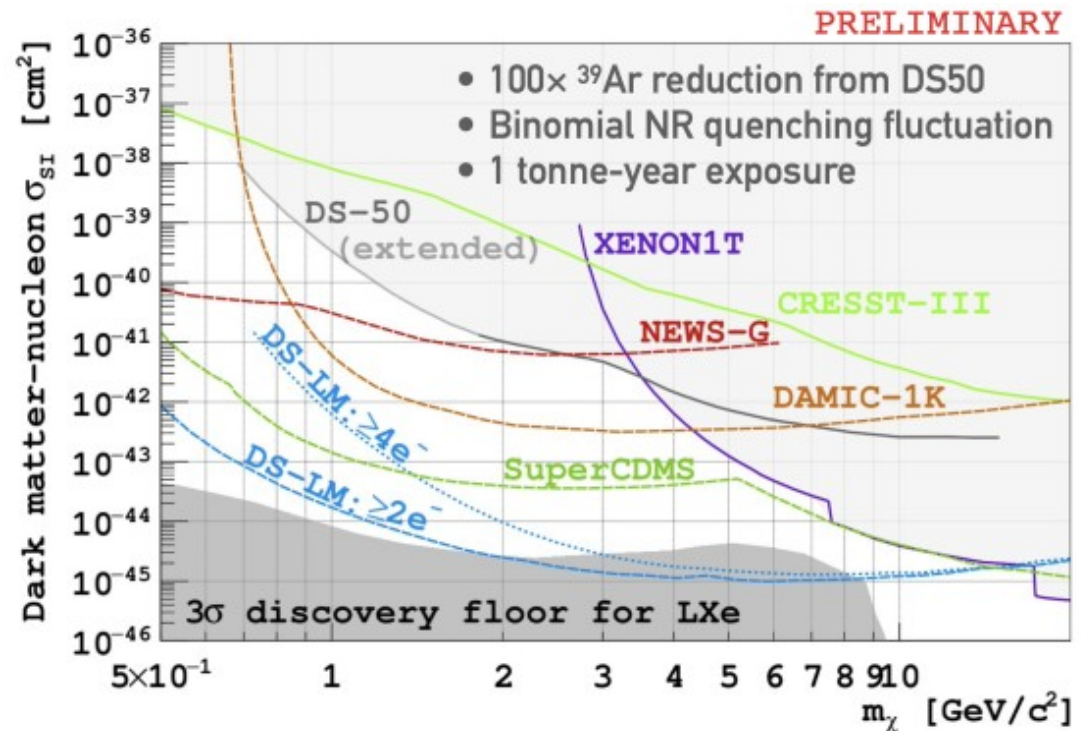
- Dual-Phase TPC
- Cylinder or octagonal shape
- 864 PDMs

**Aim:** Gaining sensitivity to **light dark matter** candidates



Based on the **s2 only signal** as for DarkSide-50 low mass analysis

## Discovery sensitivity



→ Sensitive **up to the neutrino floor**

# Summary

- **UAr** allows to perform efficient WIMP searches both a low and high masses
- A dedicated process **Urania/Aria/DArT** is under development to scale up UAr production
- A new detector aiming at reaching a better sensitivity to dark matter :

## **DarkSide-20k**

- A large collaboration (GADMC) with **expertise** on argon detectors
- **Technological developments**, for instance for photo-sensors
- Construction of DarkSide-20k started in **2022** and the data taking will start in **2025**





