

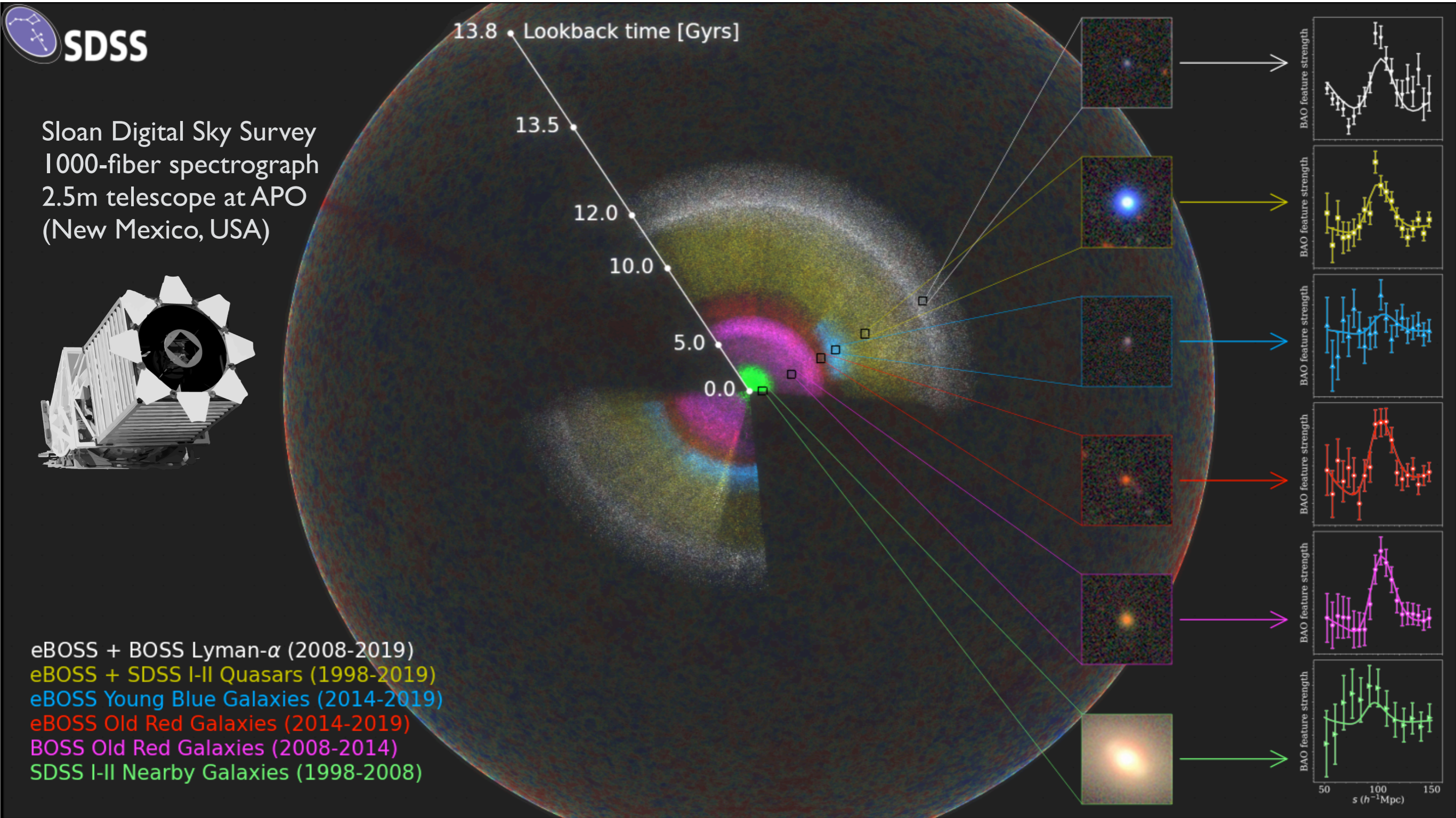


# Cosmological results from 20 years of the Sloan Digital Sky Survey

Andreu Font-Ribera

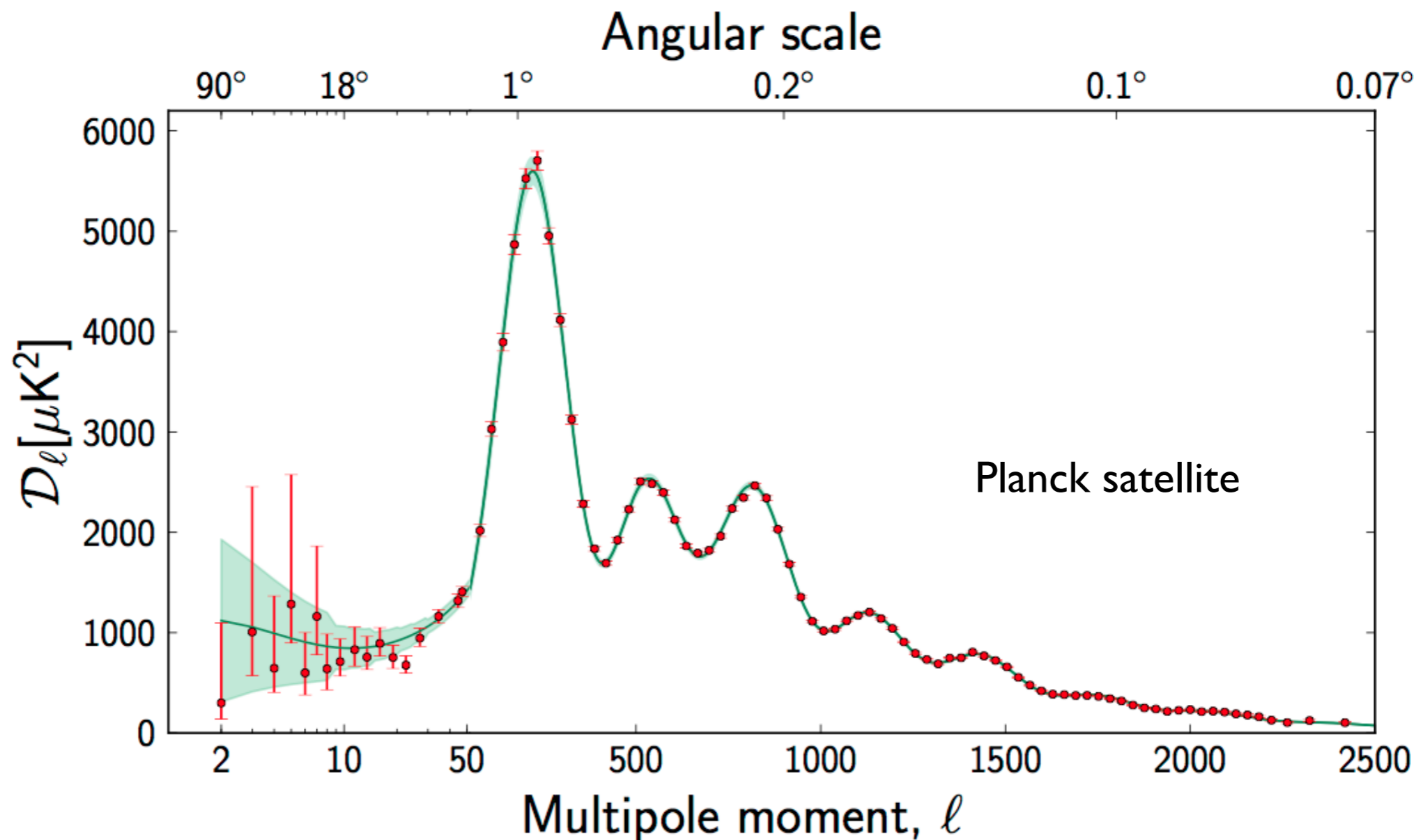
Institut de Física d'Altes Energies (IFAE, Barcelona)

On behalf of the eBOSS Collaboration



- Introduction to Baryon Acoustic Oscillations
- BOSS (2009-2014) and eBOSS (eBOSS, 2014-2019)
- Cosmology from 20 years of SDSS
  - Curvature, dark energy and neutrino masses
  - $H_0$  tension

Oscillations clearly seen in the CMB temperature power spectrum



Sound horizon at recombination (from Planck):  $r_d = 147.6 \pm 0.3$  Mpc

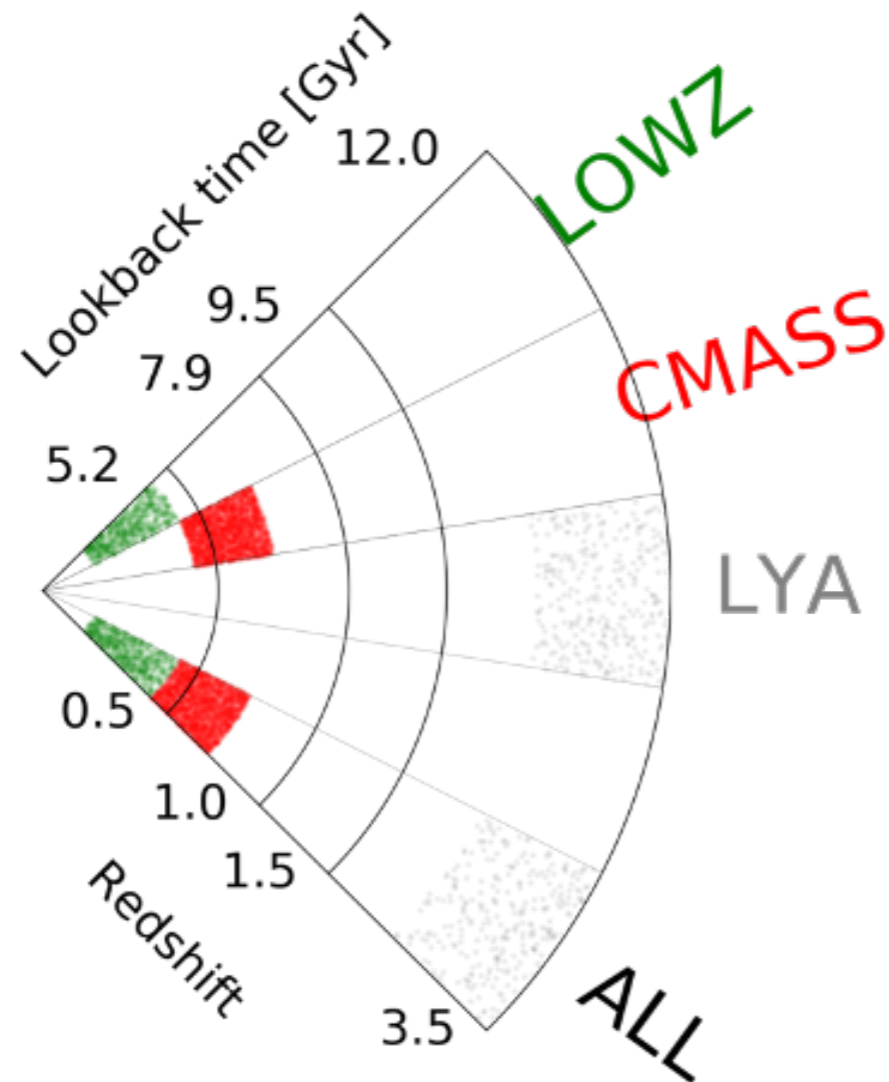
We measure BAO peak in the transverse direction in SDSS :  $\Delta\theta_{BAO}$

We measure BAO peak along the line of sight in SDSS :  $\Delta v_{BAO}$

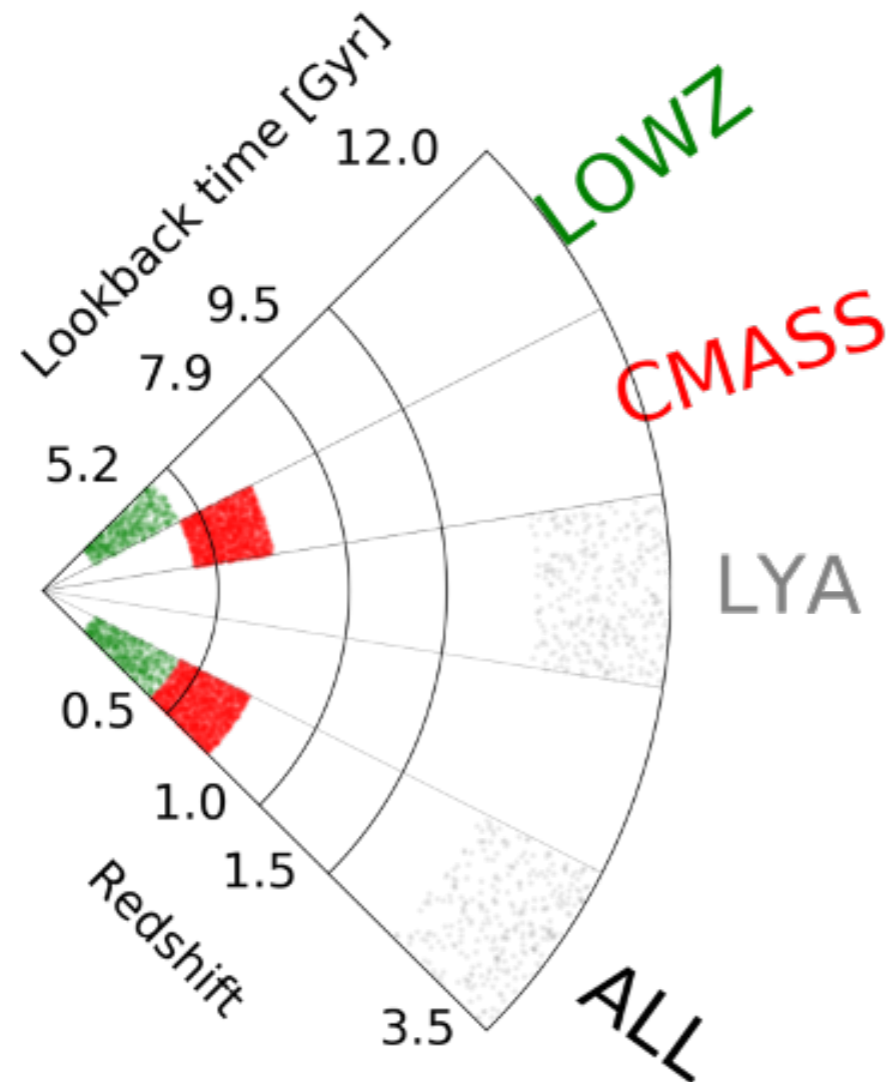
$$\Delta\theta_{BAO} = \frac{r_d}{1+z} \frac{1}{D_A(z)} \qquad \Delta v_{BAO} = \frac{r_d}{1+z} H(z)$$

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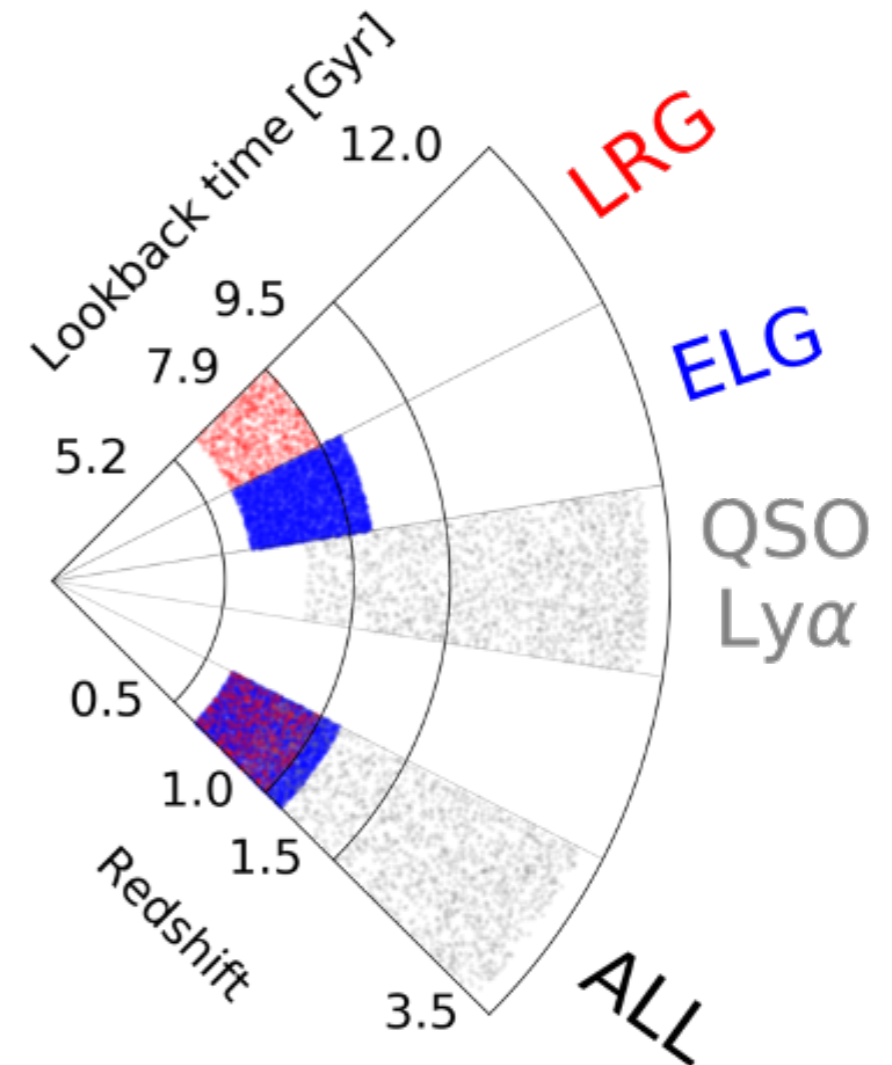
Baryon Oscillation Spectroscopic Survey  
(BOSS, SDSS III, 2009-2014)



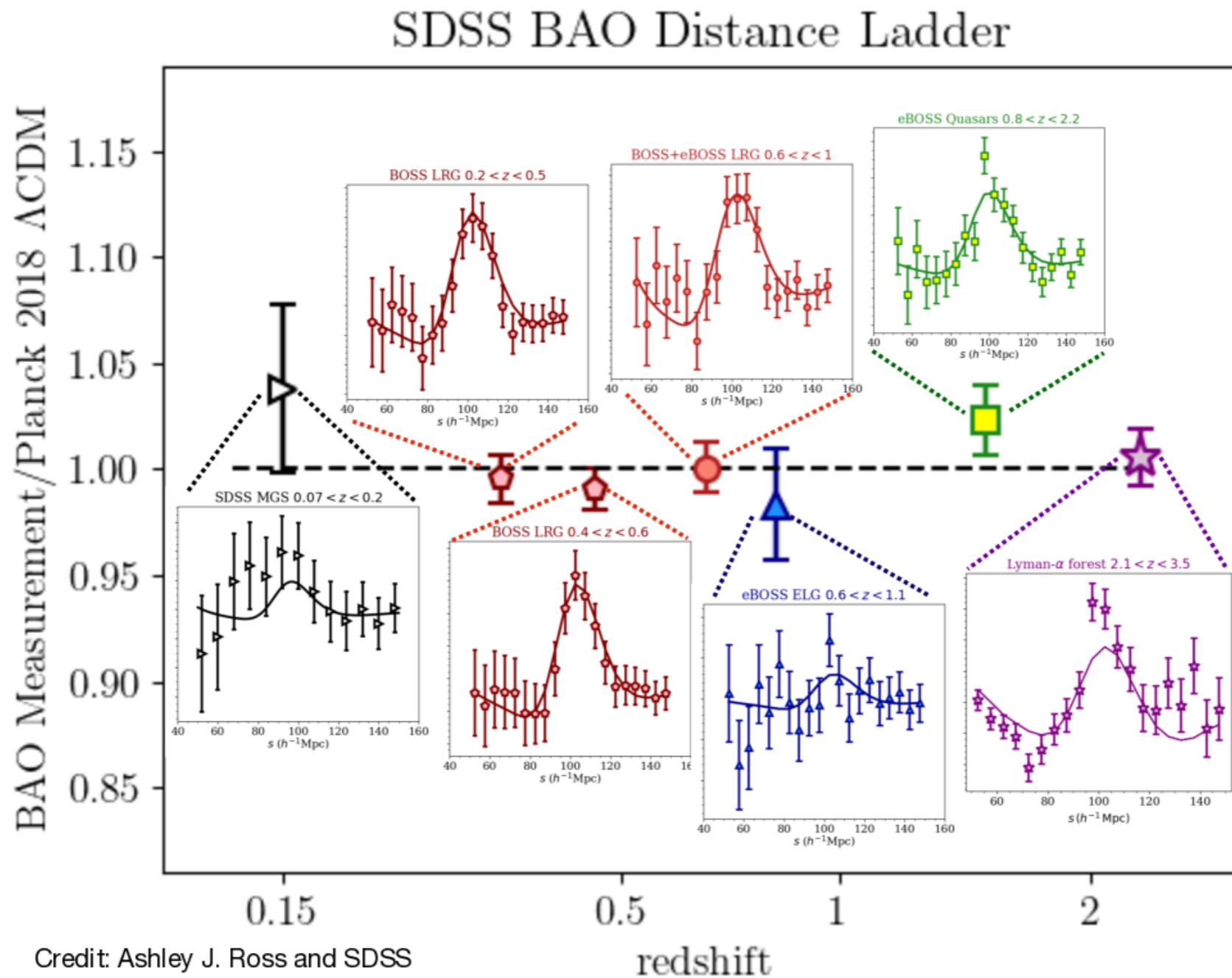
Baryon Oscillation Spectroscopic Survey  
(BOSS, SDSS III, 2009-2014)



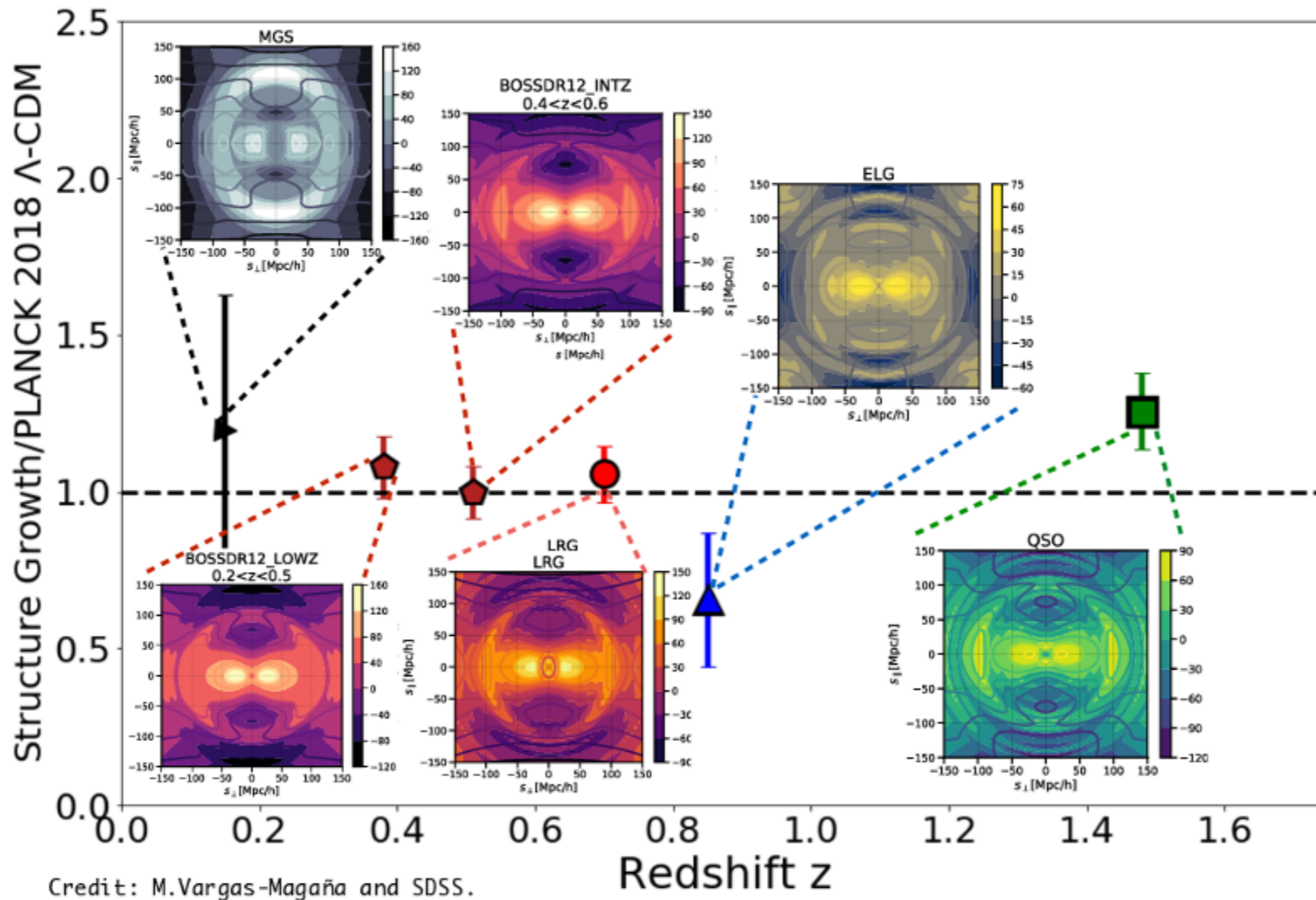
Extended BOSS  
(eBOSS, SDSS IV, 2014-2019)







BAO from BOSS/eBOSS provide accurate distances over wide redshift range



RSD from BOSS/eBOSS provide accurate measurements of growth of structure

- Introduction to Baryon Acoustic Oscillations
- BOSS (2009-2014) and eBOSS (eBOSS, 2014-2019)
- **Cosmology from 20 years of SDSS**
  - Curvature, dark energy and neutrino masses
  - $H_0$  tension

eBOSS Collaboration 2021 (arXiv:2007.08991)

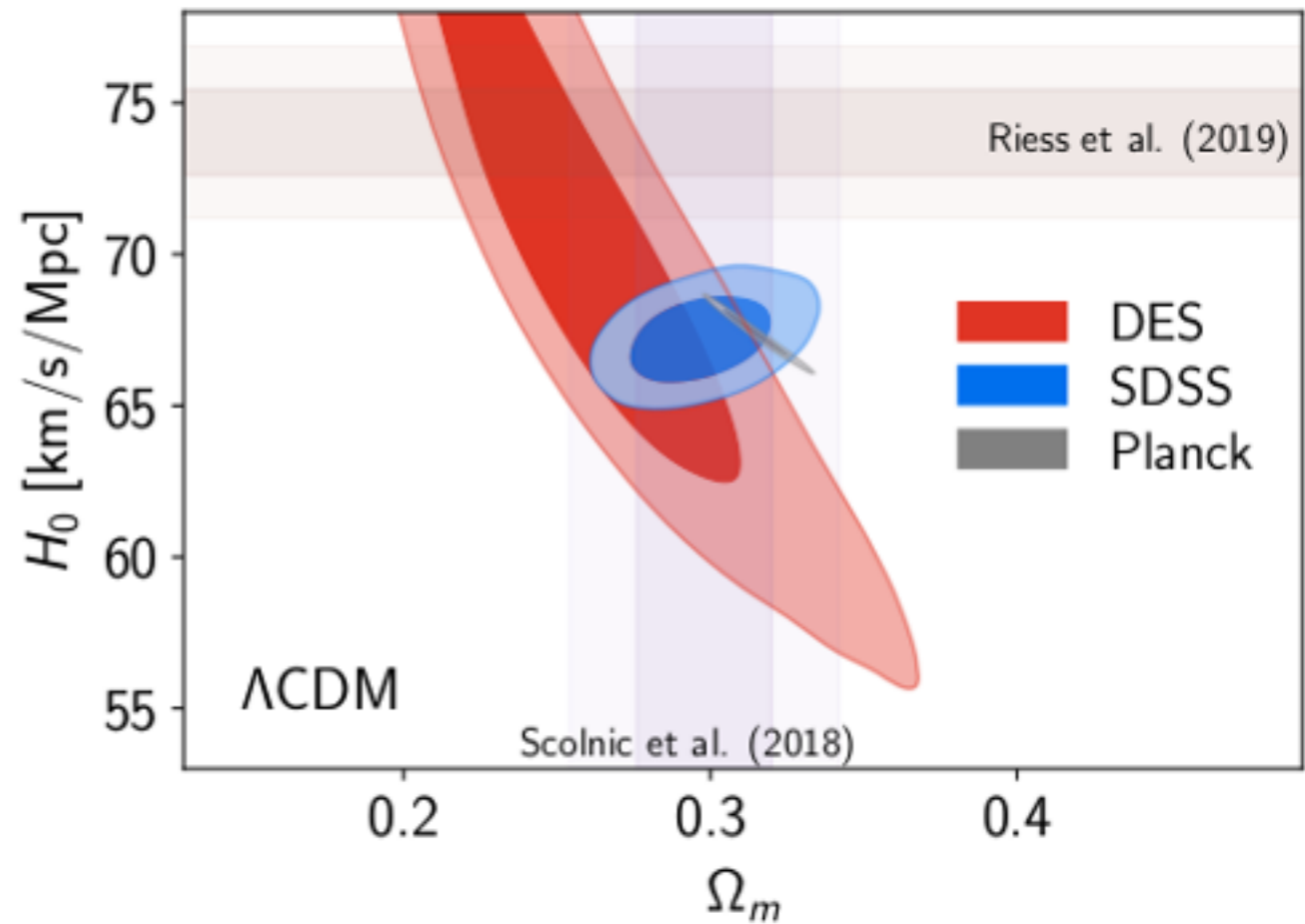
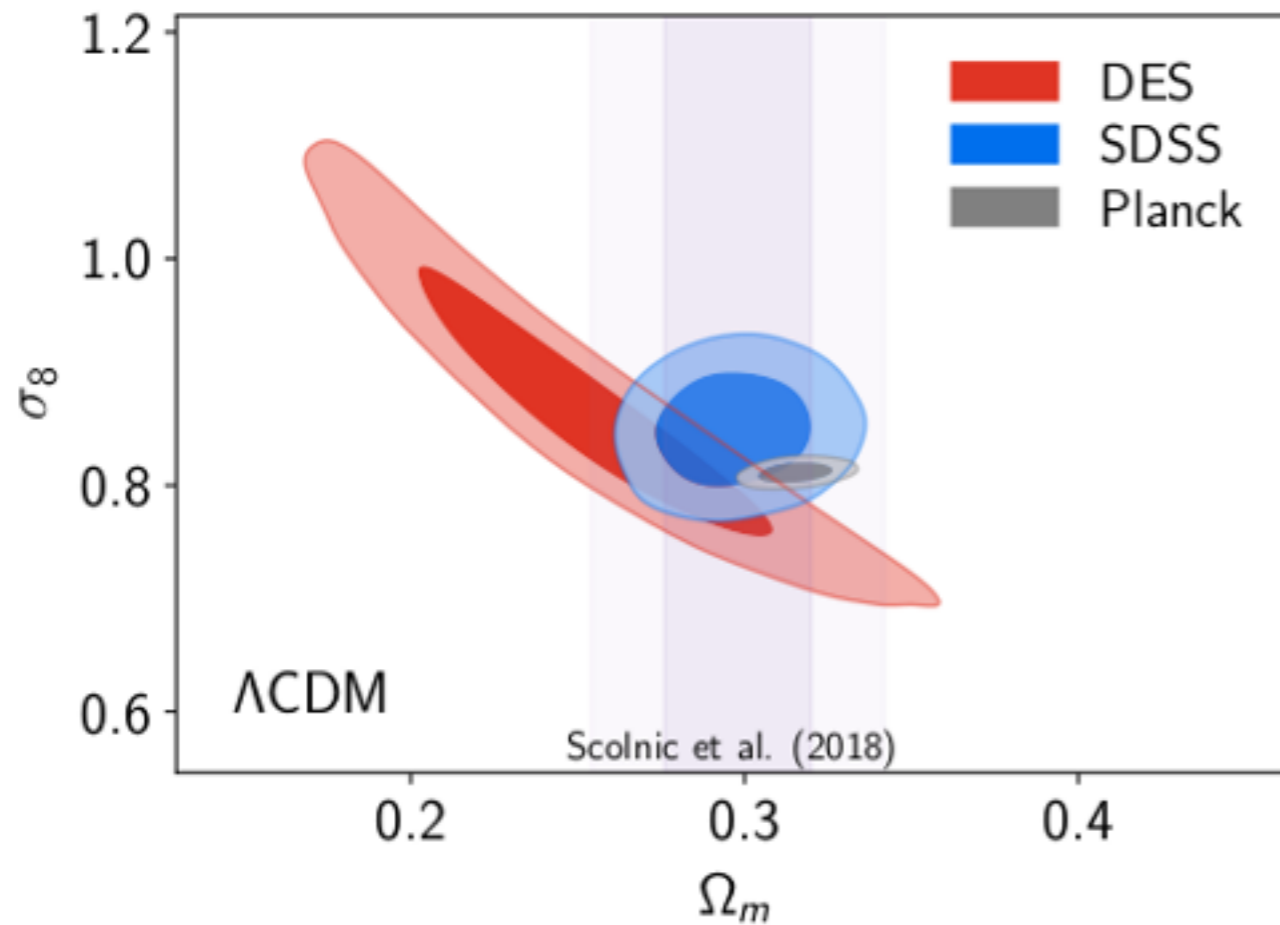
*“The Completed SDSS-IV extended Baryon Oscillation Spectroscopic Survey: Cosmological Implications from two Decades of Spectroscopic Surveys at the Apache Point observatory”*

Interpretation of 23-paper arXiv submission from July 20, 2020

Collaboration paper co-led by (left to right): Eva-Maria Mueller (Oxford), Kyle Dawson (Utah), Andreu Font-Ribera (IFAE), Zheng Zheng (Utah) and Anze Slosar (BNL)

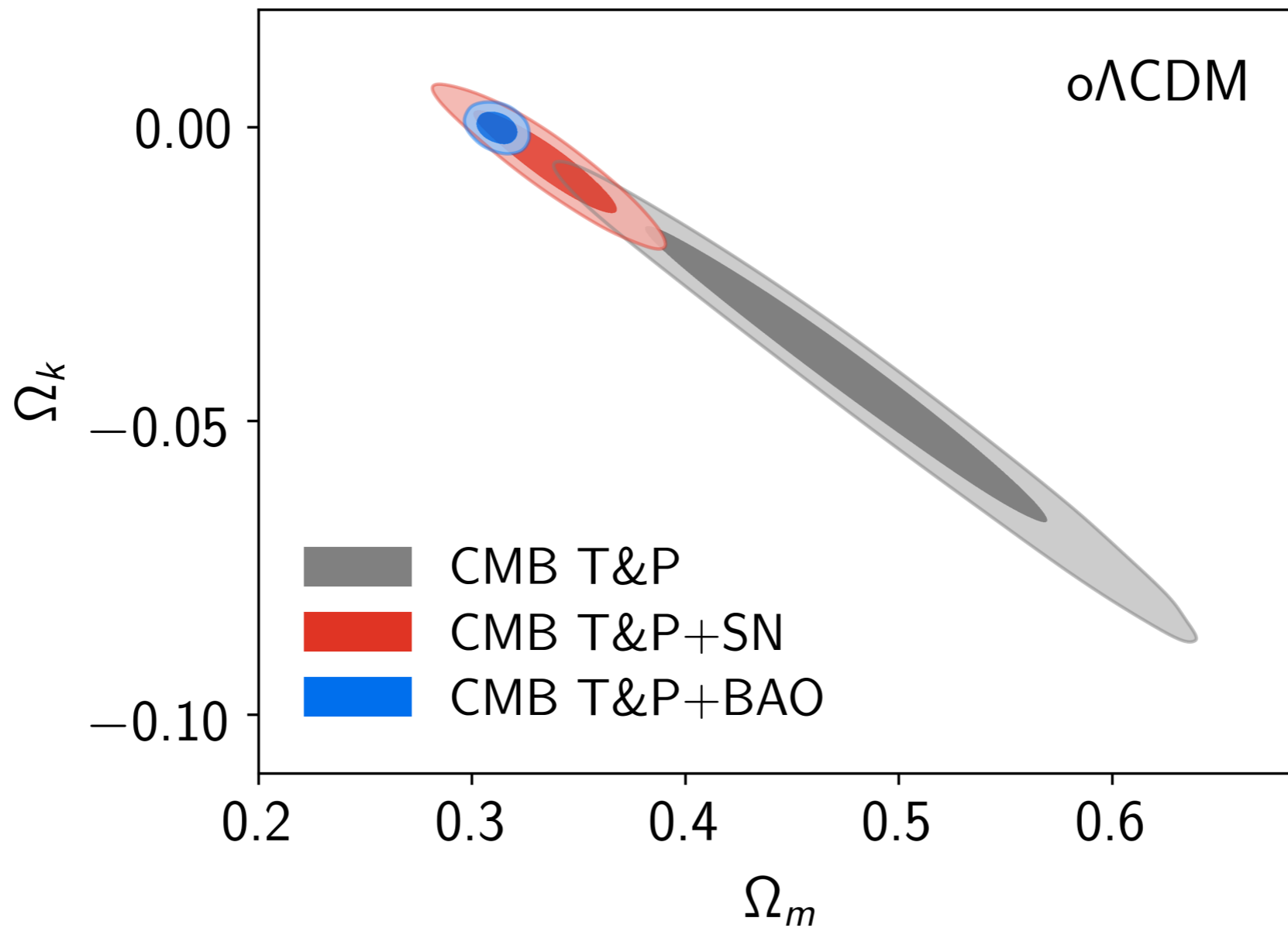


- SDSS fully consistent with Planck CMB, Pantheon SNe and DES 3x2
- Clear  $H_0$  tension with SHOES distance ladder (more latter)



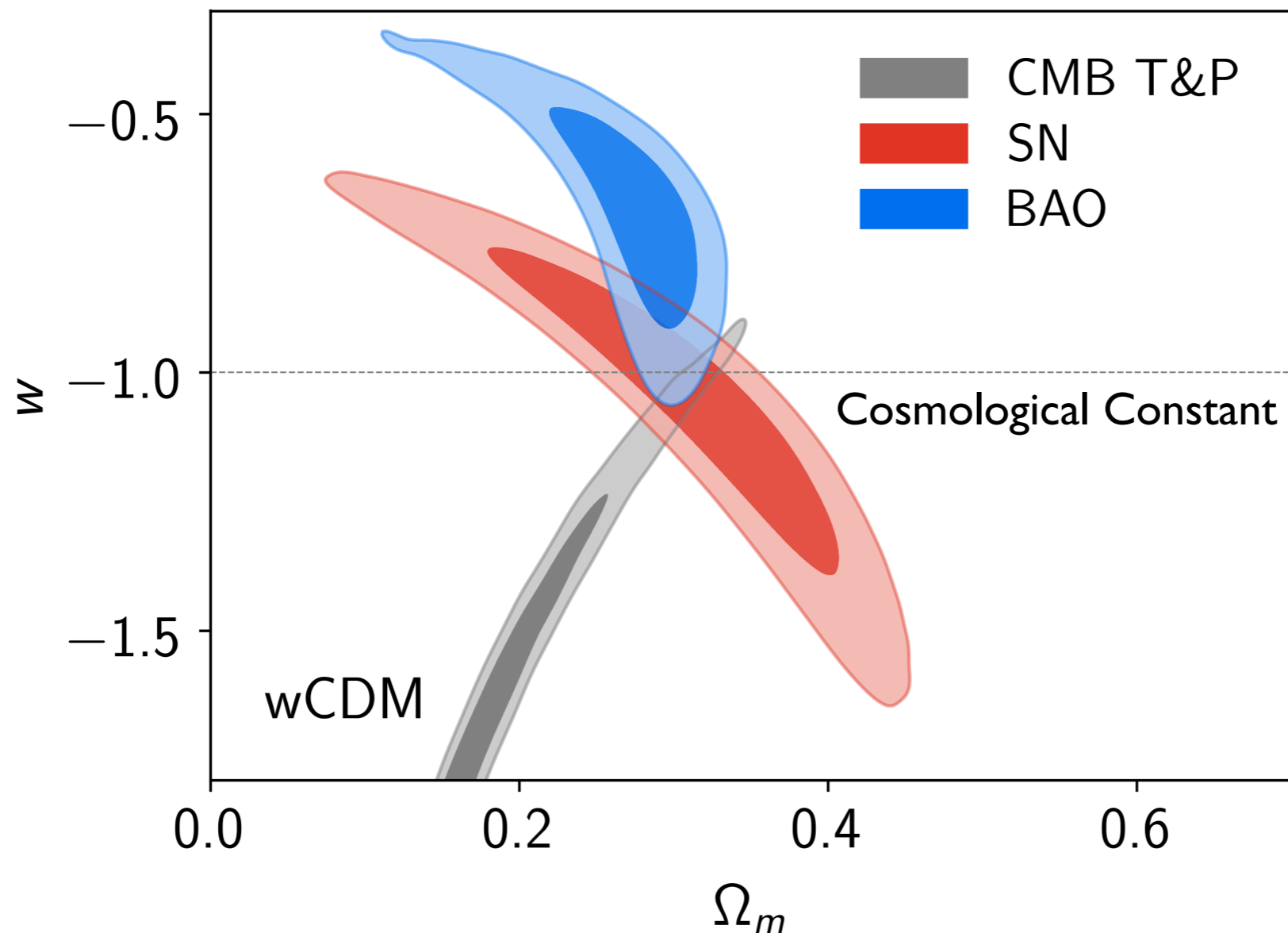
DES and SDSS results have BBN prior and weak prior on  $n_s$

- Introduction to Baryon Acoustic Oscillations
- BOSS (2009-2014) and eBOSS (eBOSS, 2014-2019)
- Cosmology from 20 years of SDSS
  - Extensions: curvature, dark energy and neutrino masses
- $H_0$  tension



BAO + CMB has a clear preference for a flat Universe

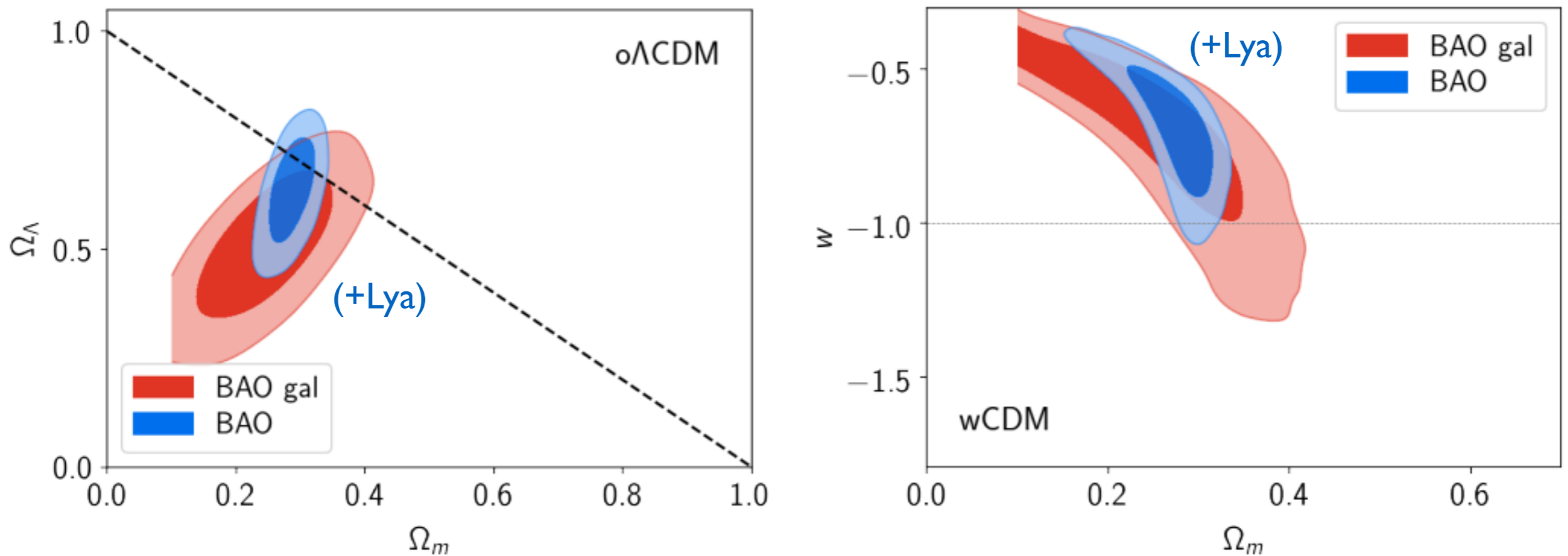
Equation of state of dark energy  $w = p_{\text{DE}}/\rho_{\text{DE}}$



Dark energy consistent with cosmological constant



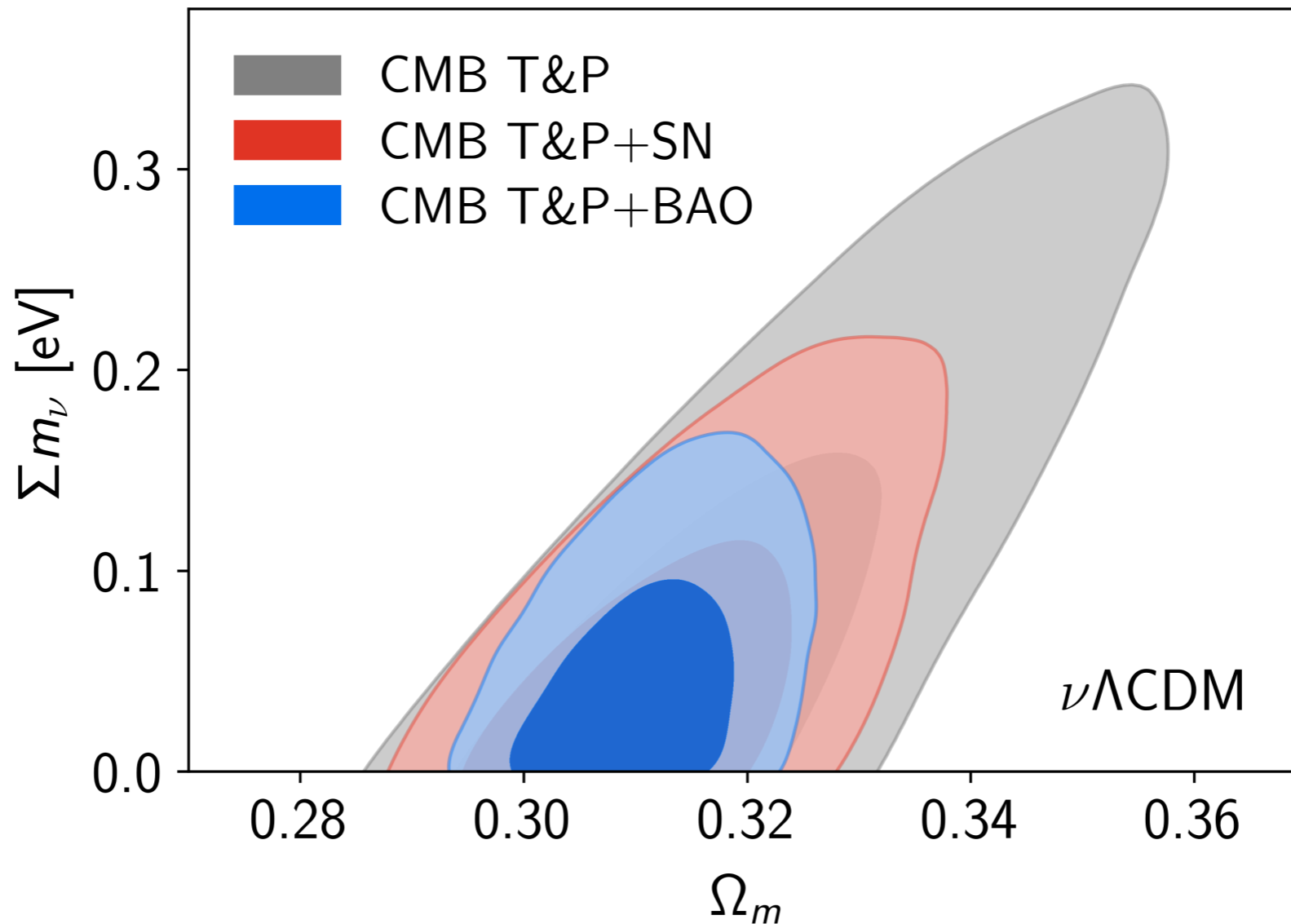
SDSS measures galaxy and quasar BAO at  $z \leq 1.5$  (BAO gal)  
 along with Lyman- $\alpha$  forest BAO at  $z = 2.33$



Combined they allow precise constraints on curvature and dark energy

Lower limits from neutrino oscillation experiments

$\sum m_\nu > 0.0588 \text{ eV}$     normal hierarchy,  
 $\sum m_\nu > 0.0995 \text{ eV}$     inverted hierarchy.

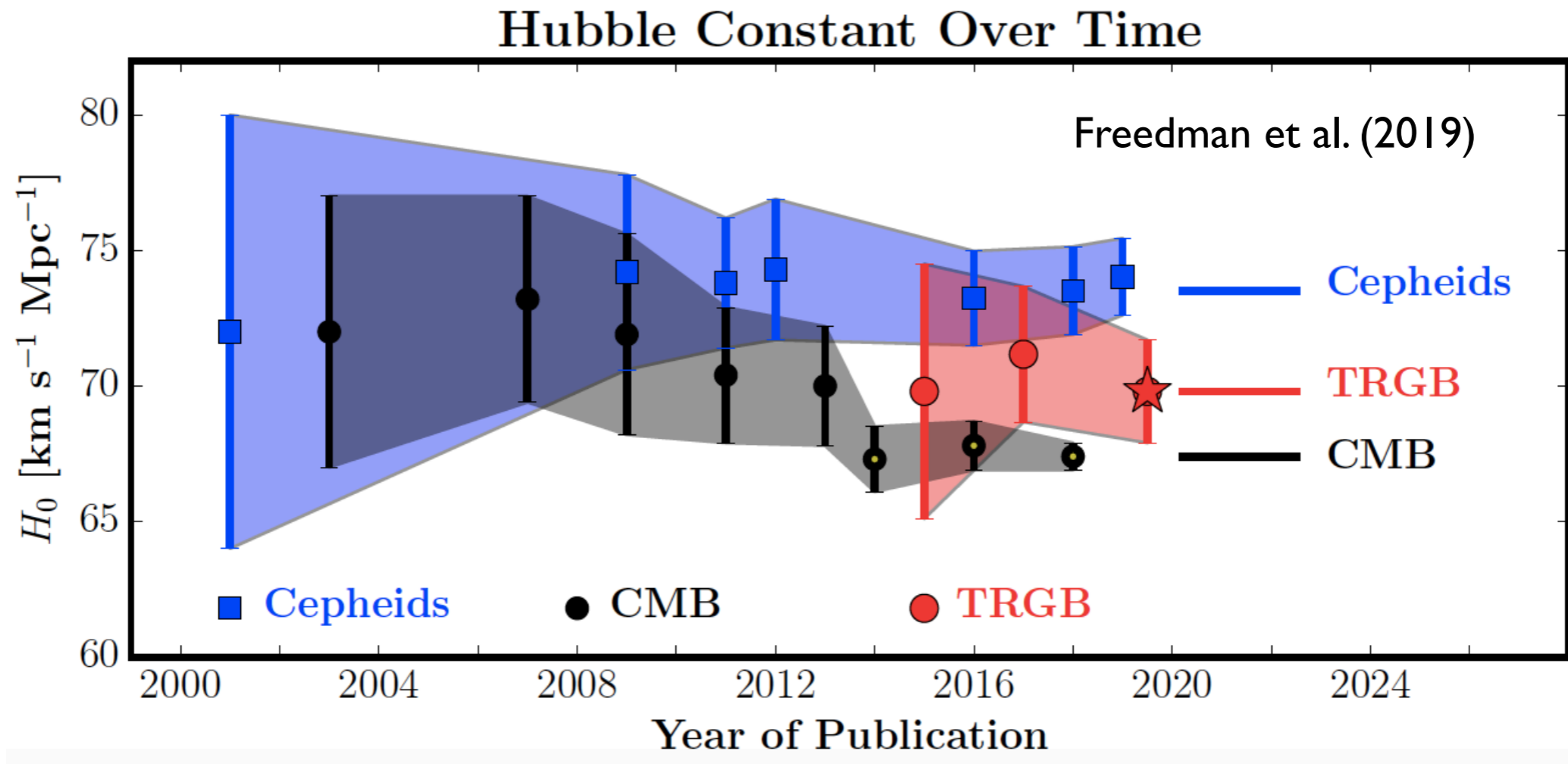


BAO breaks degeneracy with matter density in CMB

- Introduction to Baryon Acoustic Oscillations
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- Cosmology from 20 years of SDSS
  - Curvature, dark energy and neutrino masses
  - $H_0$  tension

How fast is the Universe currently expanding?

One of the key cosmological parameters has been historically controversial



Early Universe physics  $\rightarrow$  low  $H_0$

Local distances  $\rightarrow$  high  $H_0$

Prefer low  $H_0$

Planck  
Planck + SPT  
Planck + SDSS  
WMAP + SDSS  
SPT + SDSS  
ACT + SDSS  
 $r_d$  + SN + SDSS  
BBN + SDSS

Consistent with both

SPT  
TRGB  
LIGO GW

Prefer high  $H_0$

SHOES  
HoLICOW

Systematics on either side? Problems with flat  $\Lambda$ CDM?

Prefer low  $H_0$

Consistent with both

Prefer high  $H_0$

- Planck
- Planck + SPT
- Planck + SDSS
- WMAP + SDSS
- SPT + SDSS
- ACT + SDSS
- $r_d + \text{SN} + \text{SDSS}$
- BBN + SDSS

- SPT
- TRGB
- LIGO GW

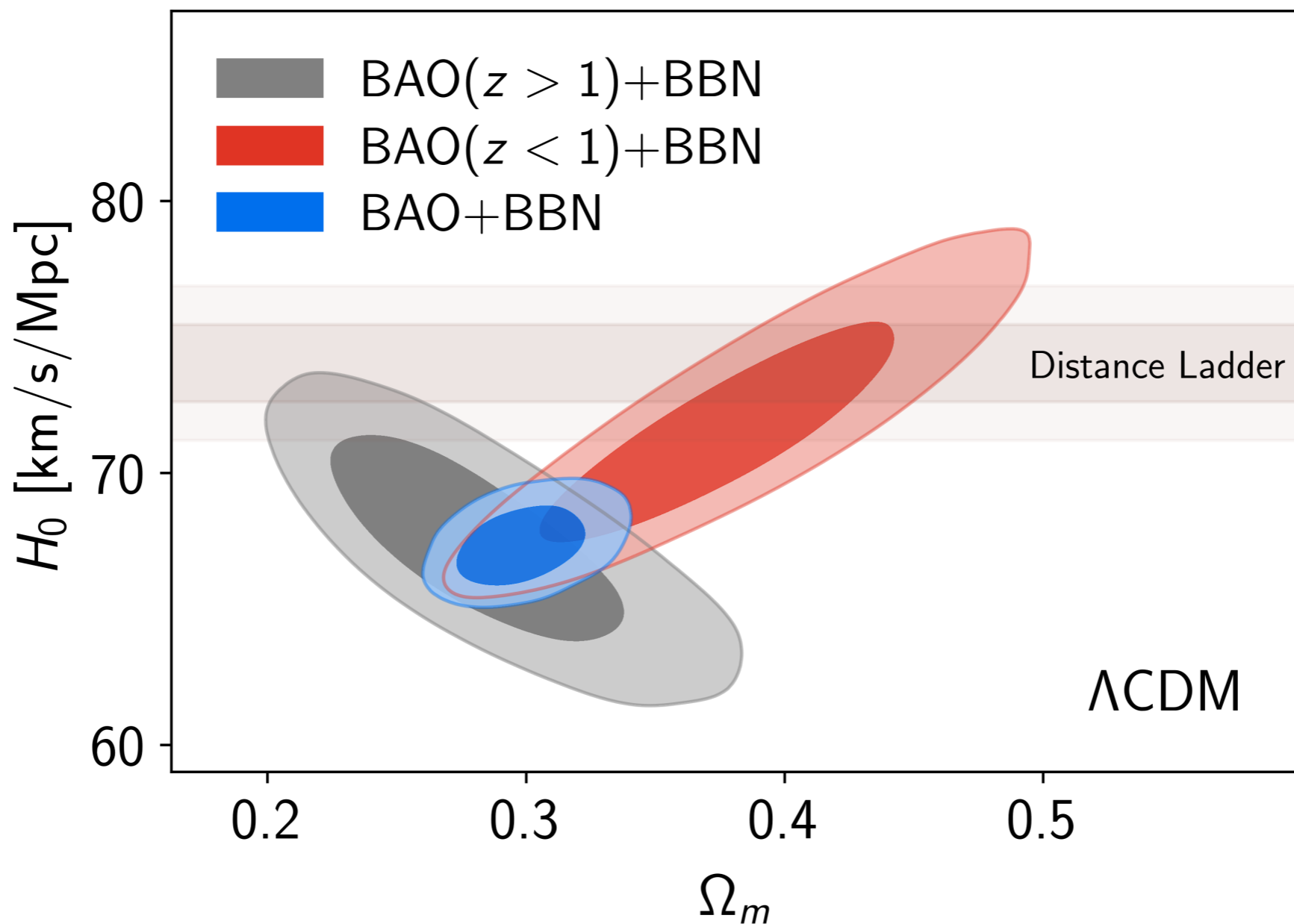
- SHOES
- HoLICOW

$r_d + \text{SN} + \text{SDSS}$  → This does not assume  $\Lambda$   
 BBN + SDSS → This does not use CMB

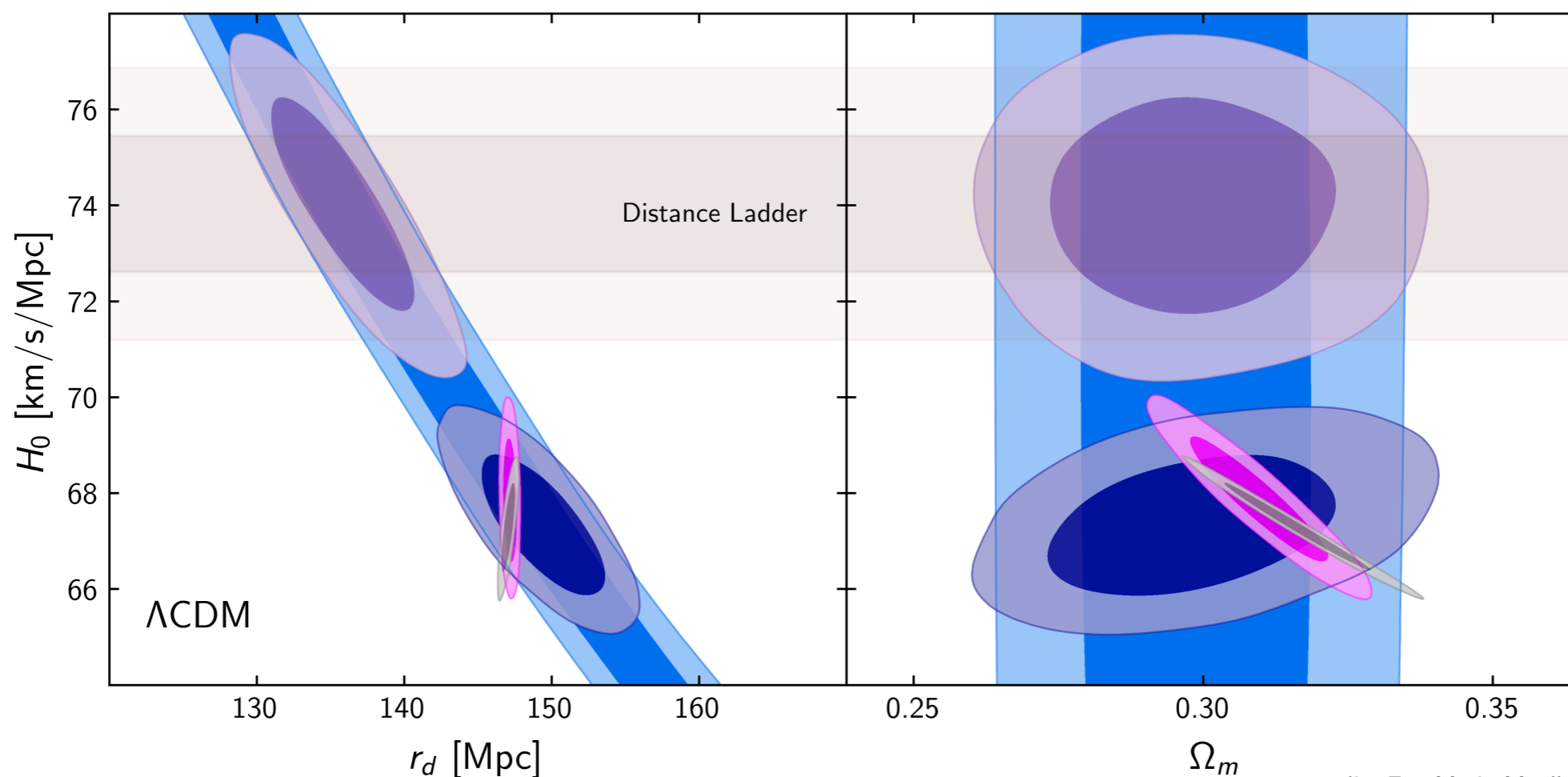
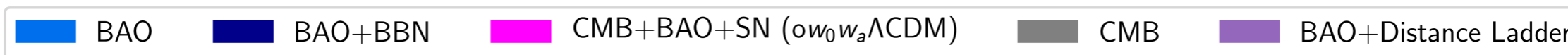
Systematics on either side? Problems with flat  $\Lambda$ CDM?

BAO + LCDM constraint  $\Omega_m$  and  $H_0 r_d$  (sound horizon, size of ruler)

BBN prior on  $\Omega_b$  can break degeneracy and measure  $H_0$  from BAO



## Inverse Distance Ladder



credit: Eva-Maria Mueller and SDSS

BAO constraints  $\Omega_m$  and product  $H_0 r_d$   
(sound horizon, size of ruler)

$$r_d = \int_{z_d}^{\infty} \frac{c_s(z)}{H(z)} dz$$



Prefer low  $H_0$

- Planck
- Planck + SPT
- Planck + SDSS
- WMAP + SDSS
- SPT + SDSS
- ACT + SDSS
- $r_d$  + SN + SDSS
- BBN + SDSS



They all assume we understand early universe physics (to compute  $r_d$ )

Consistent with both

- SPT
- TRGB
- LIGO GW

Prefer high  $H_0$

- SHOES
- HoLICOW

$$r_d = \int_{z_d}^{\infty} \frac{c_s(z)}{H(z)} dz$$

$$c_s(z) = 3^{-1/2} c \left[ 1 + \frac{3}{4} \rho_b(z) / \rho_\gamma(z) \right]^{-1/2}$$

- BOSS/eBOSS measured BAO at  $<1\%$  accuracy using galaxies
- 1.4% measurement at  $z=2.3$  using quasars and Ly- $\alpha$  forest
- Independent ( $8\text{-}\sigma$ ) detection of Dark Energy
- Order-of-magnitude better constraint on spatial curvature
- Tightest constraints on neutrino mass (CMB+BAO)
- Alternative  $H_0$  measurements consistent with CMB ones

# Extra slides