## Vacuum Design for PUMA

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#### Outline

- Introduction and motivation
- □ PUMA trap
- PUMA at ELENA
- □ PUMA at ISOLDE
- □ Summary



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#### What does PUMA mean?

## antiProton Unstable Matter Annihilation



#### Collaboration































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Vacuum Design for PUMA

## **Motivation and objectives**

Characterization of neutron halos and neutron skins

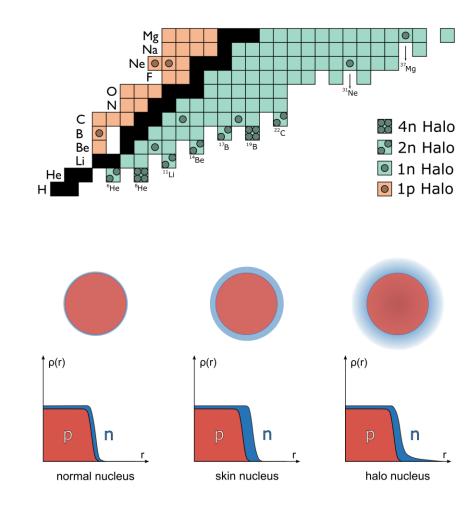
Objectives [1]:

□ to provide a new observable for <u>radioactive nuclei</u> that characterises the neutron to-proton asymmetry of their density tail, namely the neutron-to-proton annihilation ratio,

to characterize the density tail of known halos and neutron skins with this new method,

□ to evidence new proton and neutron halos,

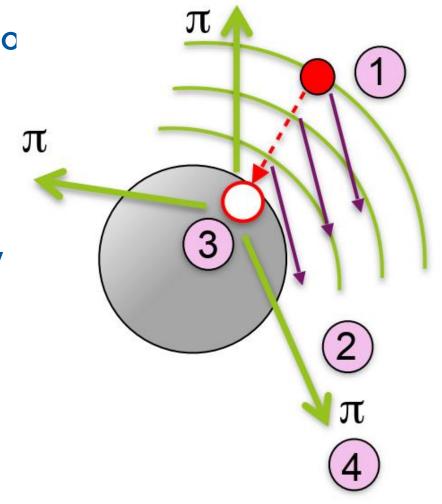
to understand the development of neutron skins in medium-mass nuclei along isotopic chains.





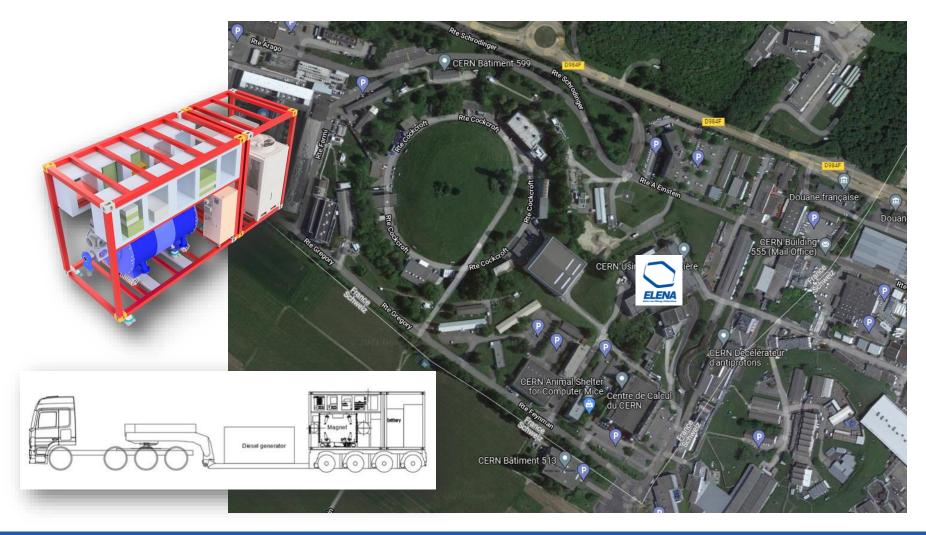
#### Antiproton annihilation with nuclei

- 1. Antiprotons captured into atomic orbital.
- 2. Decay into lower atomic orbital. X-rays & Auger electrons
- 3.  $\bar{p}$  reaches at the density tail of nuclei. Annihilate with p or n.
- 4. Annihilation product: Multiple <u>pions</u> and residual nucleus





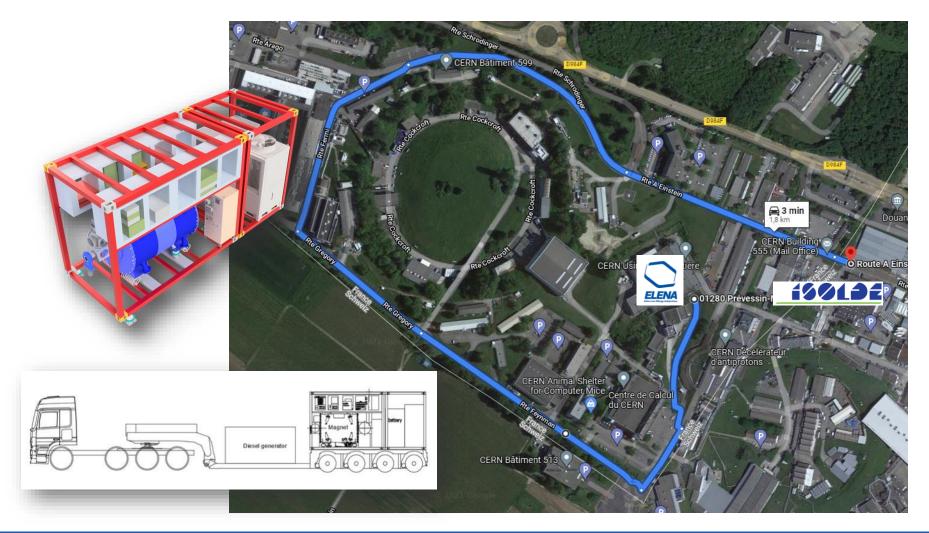
## $ar{p}$ Trip





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## $ar{p}$ Trip





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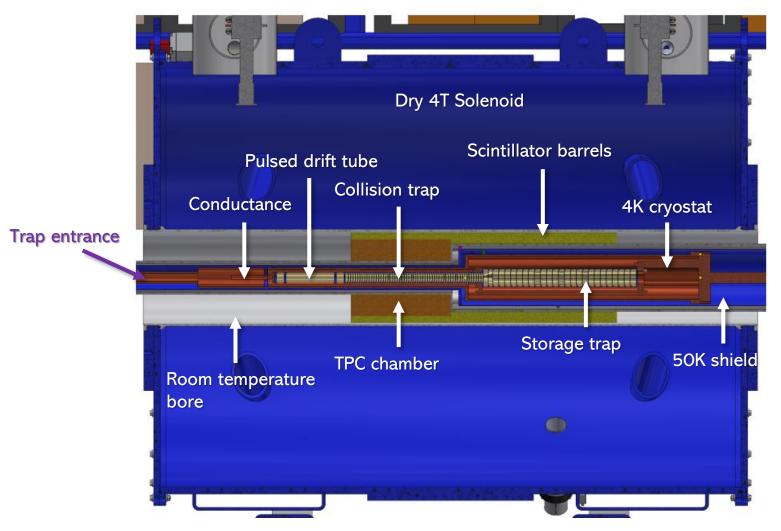


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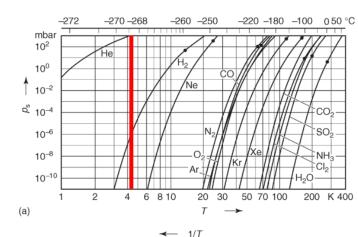
## **PUMA** Trap



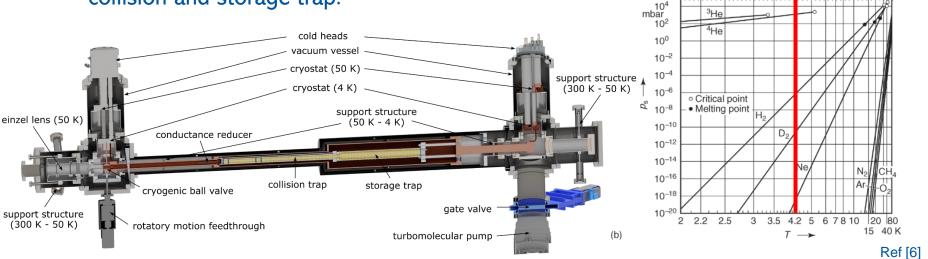


## Pumping concept

- Pumping on cold surfaces (4.2K) of gases entering the trap
- ❑ Maximum allowed pressure at the entrance of trap
   ⇒ Defined by expansion of gas entering the trap
- Small conductance between the entrance and collision and storage trap.



0.5 0.45 0.4 0.35 0.3 0.25 0.2 0.15 0.1 0.050.01 K<sup>-1</sup>





## PUMA vacuum requirements

- Storing 10<sup>9</sup> antiprotons (10<sup>7</sup> as first milestone)
- Vacuum level:
  - □ Antiproton lifetime (>30 days)
  - Low Signal background during measurement
- Specifications:
  - $\Box \quad T = 4K$
  - $\Box \quad \tau > 200 \ days$
  - $\Rightarrow \quad n_{H_2} \approx 20 \ cm^{-3} \rightarrow \underline{P} \approx 10^{-17} mbar$
  - $\Rightarrow$  0.5Hz background (10<sup>7</sup>  $\bar{p}$  in collision trap)

From [2]:  

$$\sigma_{H_2} = 3\pi a_0^2 \sqrt{\frac{27.2eV}{E_{CM}}} \qquad a_0 = 5.29 \cdot 10^{-11} m$$

$$\Gamma = \frac{1}{\tau} = n_{H_2} v_{rel} \sigma \qquad v_{rel} = \sqrt{\frac{4E_{CM}}{m_p}}$$

$$n_{H_2} = \frac{\left(6\pi a_0^2 \sqrt{\frac{27.2eV}{m_p}}\right)^{-1}}{\tau} = 3.91 \cdot 10^8 s \cdot cm^{-3} \frac{1}{\tau(s)}$$

Case: <sup>132</sup>Sn:  $\sigma$ =10<sup>-15</sup> cm<sup>2</sup>, 10<sup>5</sup> ions/bunch  $\Rightarrow$  100Hz Case: <sup>11</sup>Li:  $\sigma$ =10<sup>-16</sup> cm<sup>2</sup>, 200 ions/bunch  $\Rightarrow$  0.5Hz



#### Cross section for other gases (He)

Langevin cross section (ion-atom interaction) [2]:



## Adsorption Isotherm

 Low temperature gas adsorption modelled with Dubinin–Radushkevich (DR) isotherm

$$\theta = \exp\left(-\beta\left(RT\ln\left(\frac{P}{P_{sat}}\right)\right)^2\right) = \exp\left(-\left(\frac{T}{T_0}\ln\left(\frac{P}{P_{sat}}\right)\right)^2\right)$$

 $\theta$  relative surface coverage  $T_0$  adsorption energy (expressed in K)  $P_{sat}$  Saturation vapor pressure

- Available data
  - □  $H_2$  from [3]: T<sub>0</sub>=209 K ML=0.645×10<sup>14</sup> cm<sup>-2</sup>
  - He from [4]:

 $T_0$ =67.8 K ML= 1.27×10<sup>15</sup> cm<sup>-2</sup>

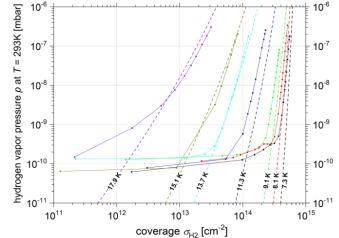
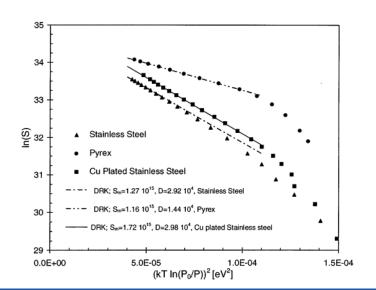


FIG. 3. Adsorption isotherms of H<sub>2</sub> on an electropolished stainless steel surface in the temperature range between 7.3 and 17.9 K. The dashed curves are the theoretical isotherms calculated according to the DRK [Eq. (2)] with the experimentally determined constants,  $B = 3075 \text{ eV}^{-2}$  and  $\sigma_{\text{mono}} = 6.45 \cdot 10^{14} \text{ cm}^{-2}$ .





## **Adsorption Isotherm**

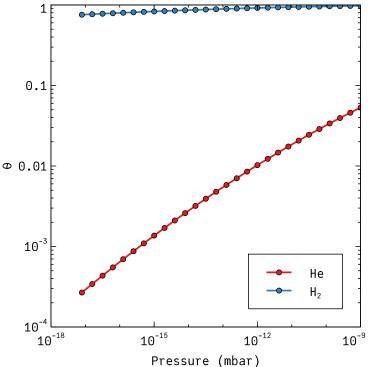
 Low temperature gas adsorption modelled with Dubinin–Radushkevich (DR) isotherm

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Is the extrapolation to very low equilibrium pressure valid?



## **COMSOL** model

- Molecular flow simulation using view factors between elements to simulate the reflexion of molecules
- COMSOL allows the study of an evolving wall following the surface coverage
- Quasi-static equilibrium

$$\frac{1}{P}\frac{\partial P}{\partial t} \ll s\frac{A}{V}\frac{\overline{v}}{4}$$

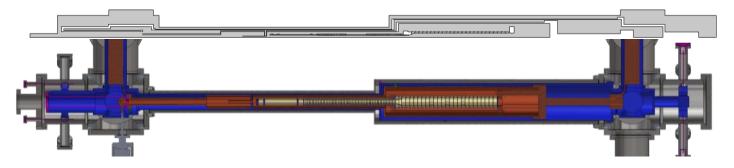
- 2D axisymmetric model
- Pressure at the entrance isotropic distribution (no beaming)
- DR isotherm model for 4.2K walls in COMSOL
  - Pumping:

$$\frac{1}{4}s_0 n_{gas} v_{th} \to \theta$$

✤ Gas Desorption:

$$\frac{1}{4}s_0 n_{eq}(\theta) v_{th} \to n_{eq}(\theta) = n_{sat} \cdot \exp\left(-\frac{T_0}{T}\sqrt{-\ln(\theta)}\right)$$

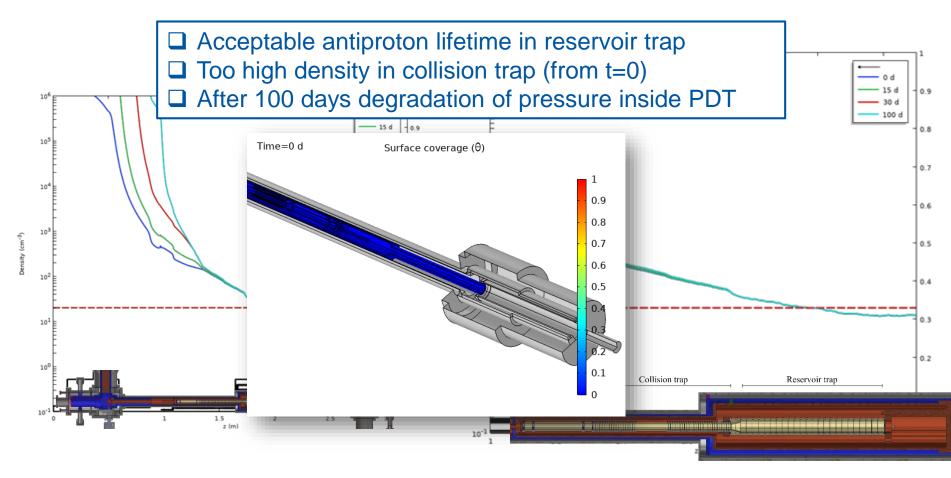
₩all	Time	
▼ Wall Type		
Wall type:		
Adsorption/Desorption 👻		
<ul> <li>Adsorption/Desorption</li> </ul>		
Sticking coefficient		
SG	stick	1
Desorption rate		
$D_{G}$	stick*flow(fmf.n_ads_G/n0,comp1.fmf.T)	mol/(m²·s)
Initial adsorbant concentration		
n <sub>ads,0,G</sub>	1e-9 [mol/m^2]	mol/m <sup>2</sup>
Additional molar flux		
$\Gamma_{\rm G}$	User defined	•
	0	mol/(m²·s)





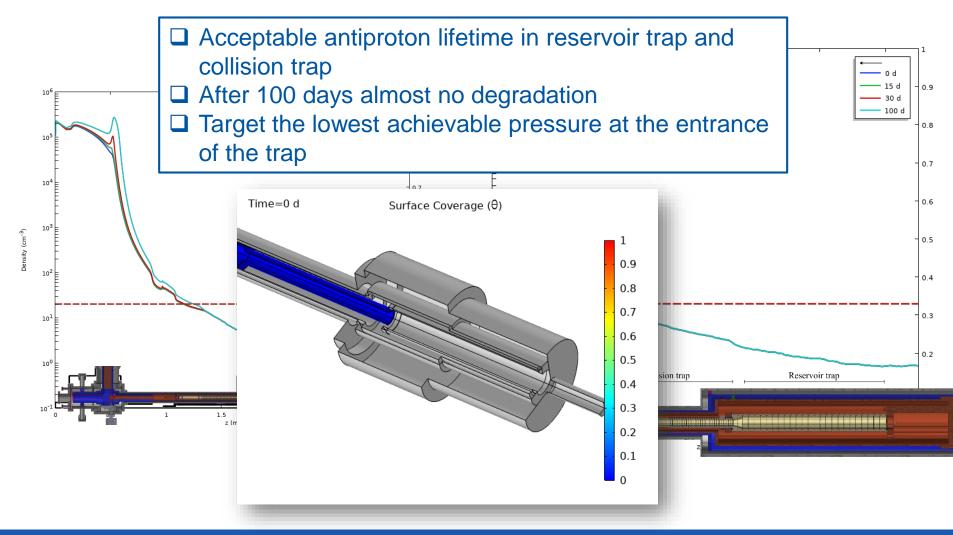
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## COMSOL model: P<sub>H2</sub>=10<sup>-10</sup>mbar





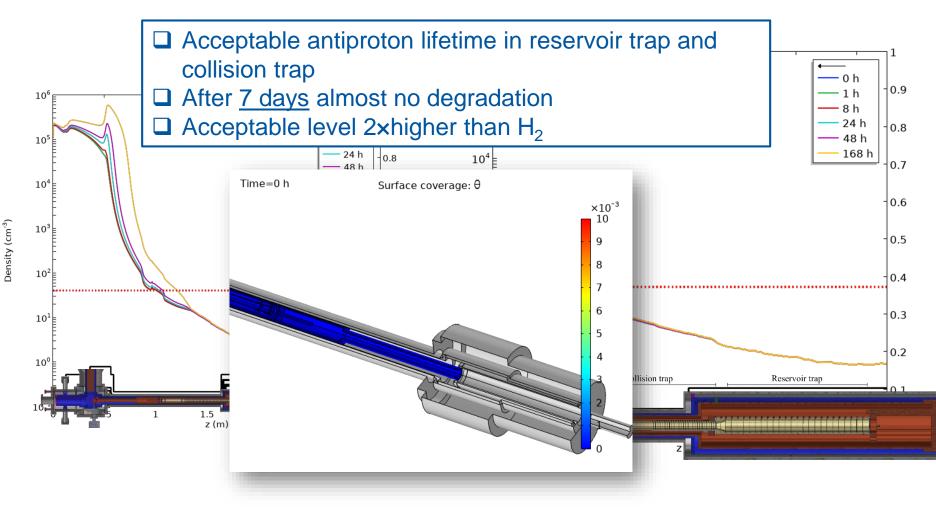
## COMSOL model: P<sub>H2</sub>=10<sup>-11</sup>mbar





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## COMSOL model: $P_{He} = 10^{-11}$ mbar

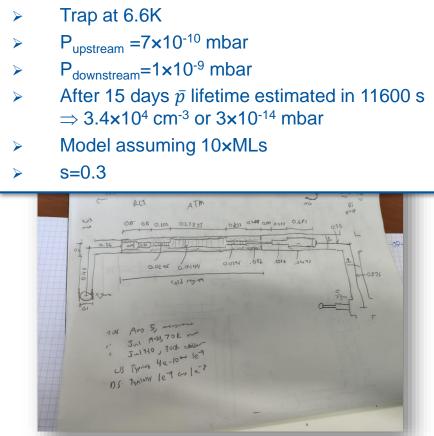




# Comparison with experimental results: Alpha experiment

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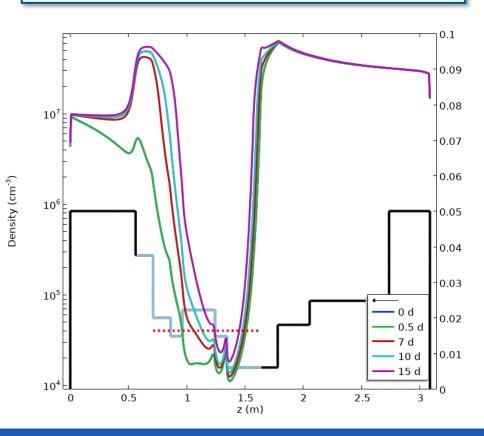




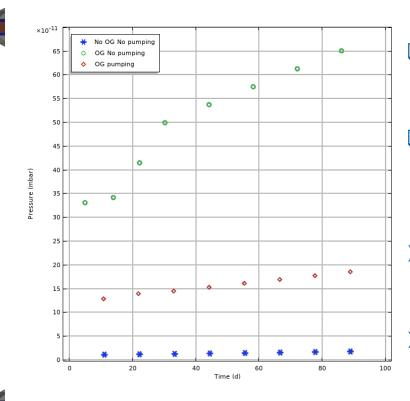
Data provided by Andrew Jordan Christensen



The model predicts a significant lower lifetime than observed  $\Rightarrow$  Conservative assumptions

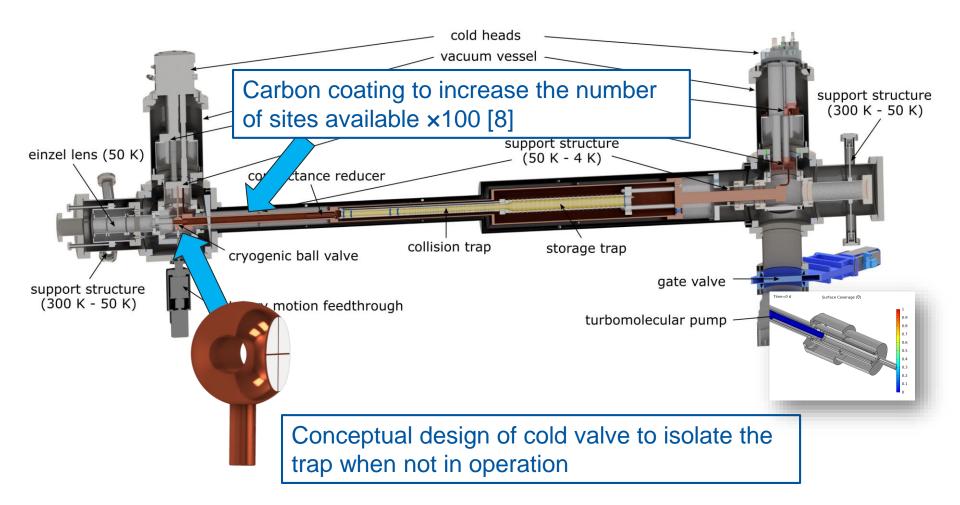


# Outgassing of room temperature material



- Insulation vacuum in communication with trap vacuum.
- H<sub>2</sub> from room temperature materials can be the dominant gas source unless is pumped
- Reduce conductance between trap and insulation vacuum
- Add pumping

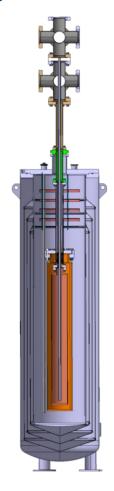
## Improvements of the trap

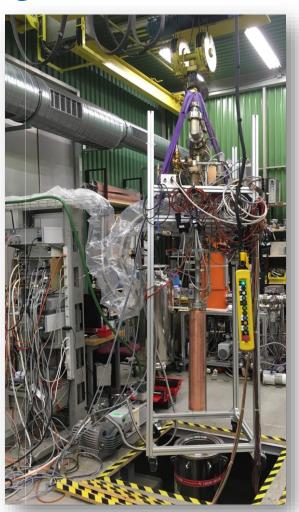


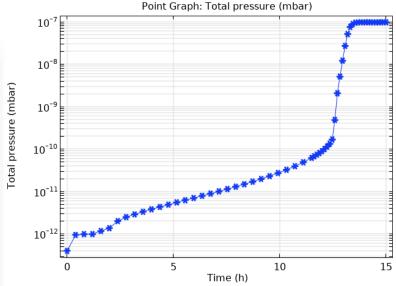


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#### H<sub>2</sub> and He isotherm data at low pressure







- New setup to measure isotherms at low pressure
- COMSOL to extract information from experiment
- Data to validate trap models





Introduction and motivation

PUMA trap

PUMA at ELENA

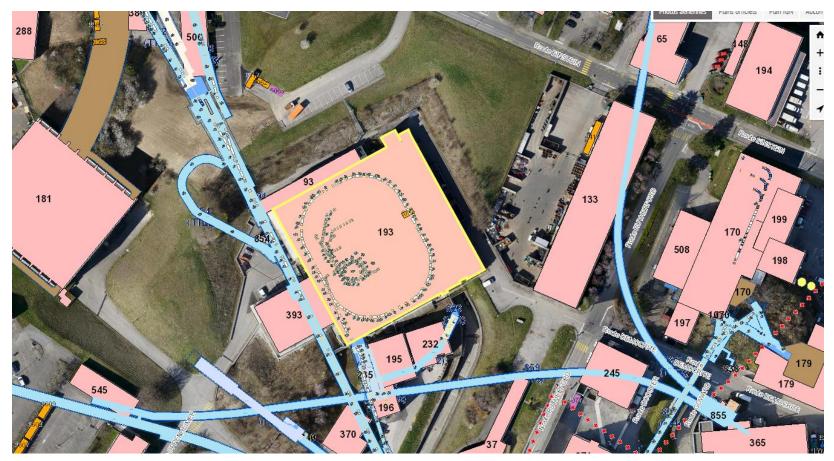
PUMA at ISOLDE

□ Summary



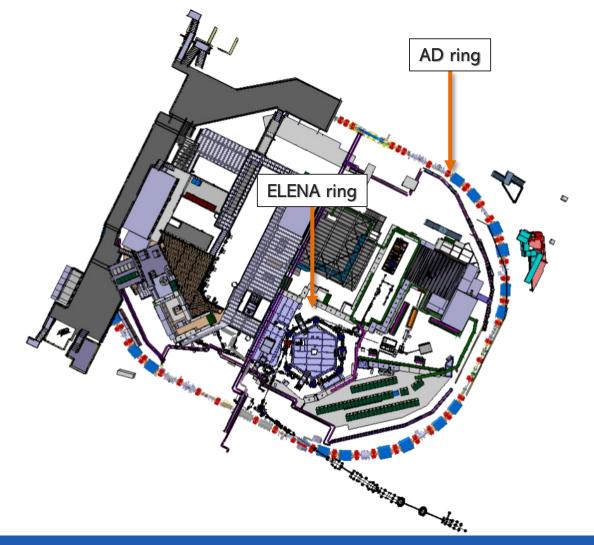








## **ELENA** layout

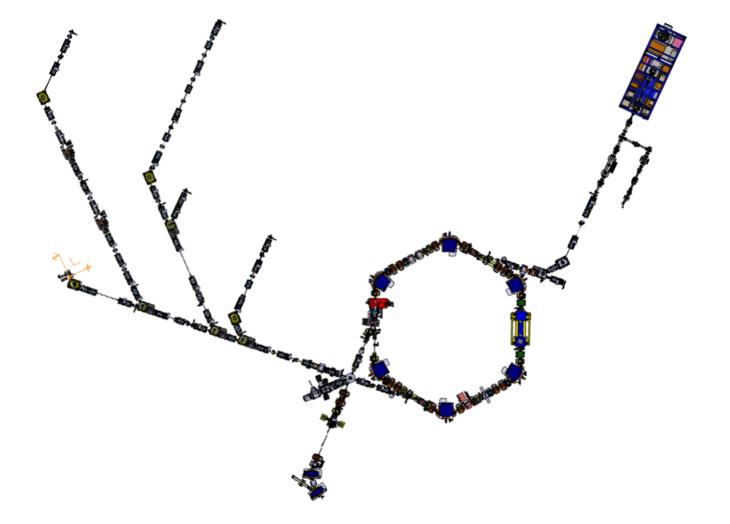




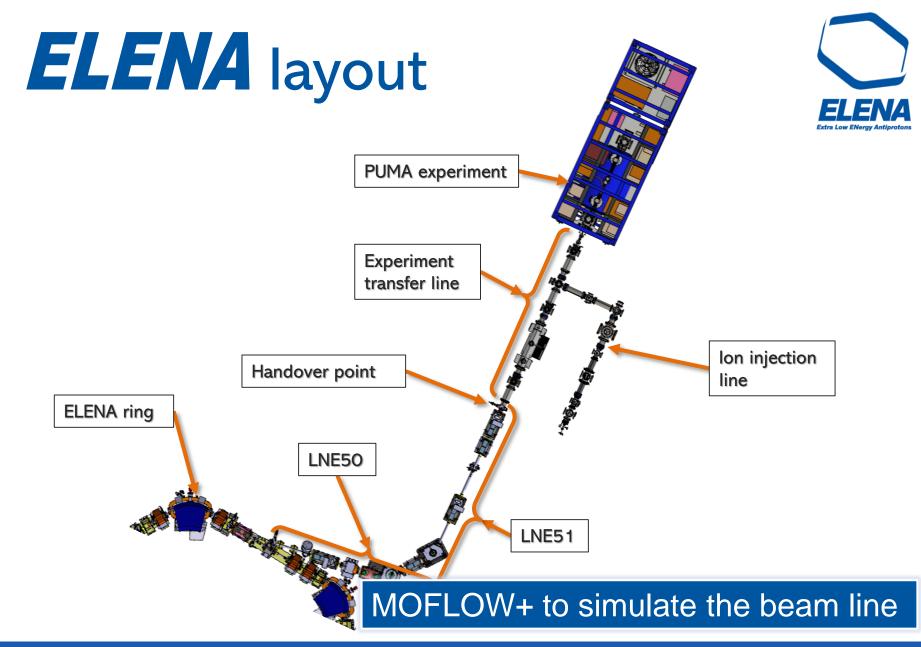




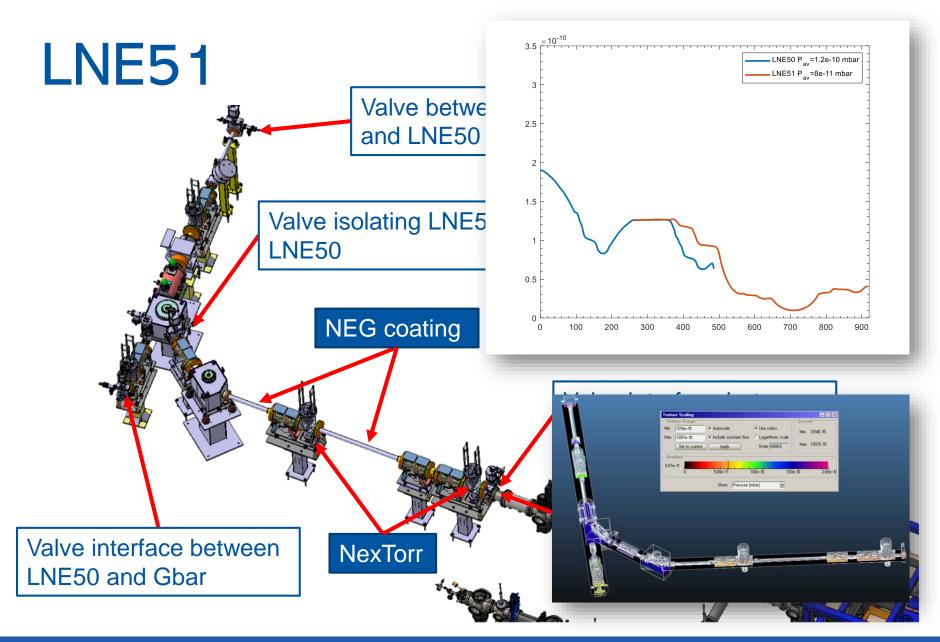










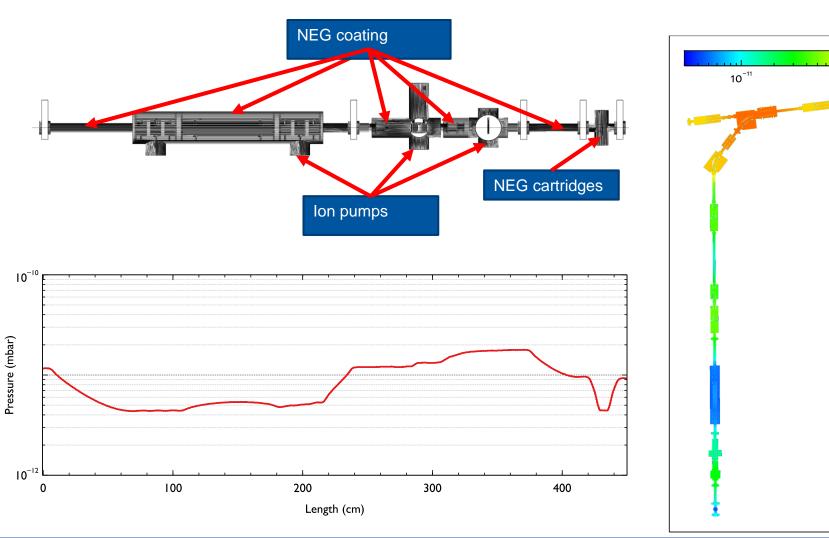




## **PUMA Transfer line**



10<sup>-10</sup>





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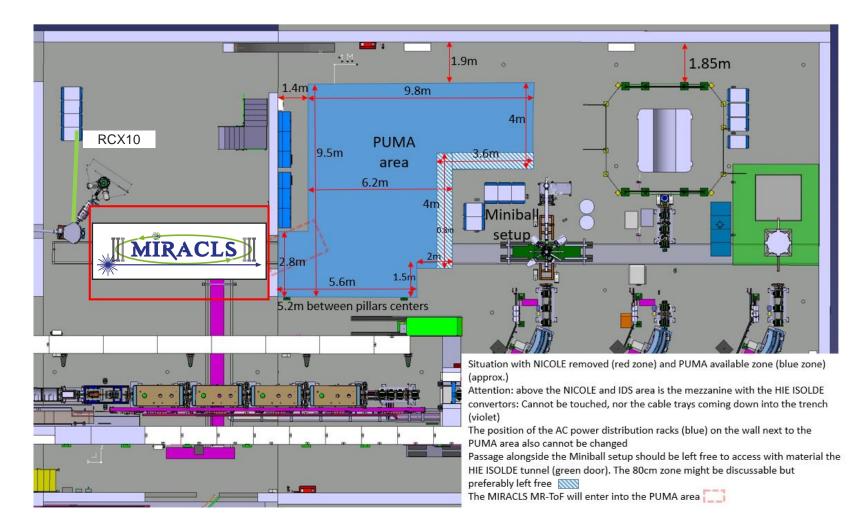






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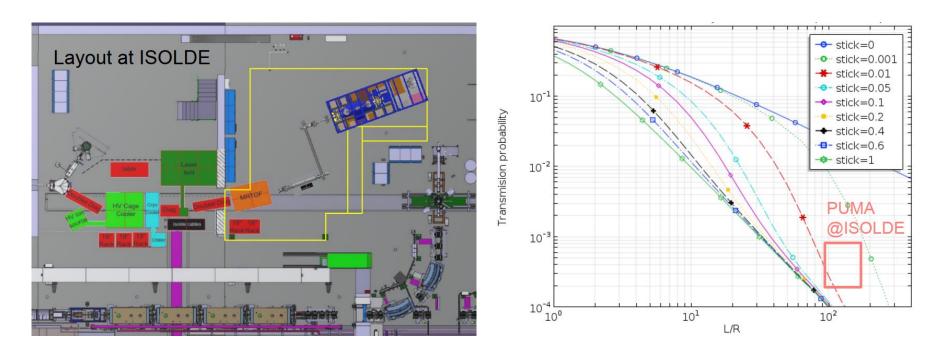




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Vacuum Design for PUMA

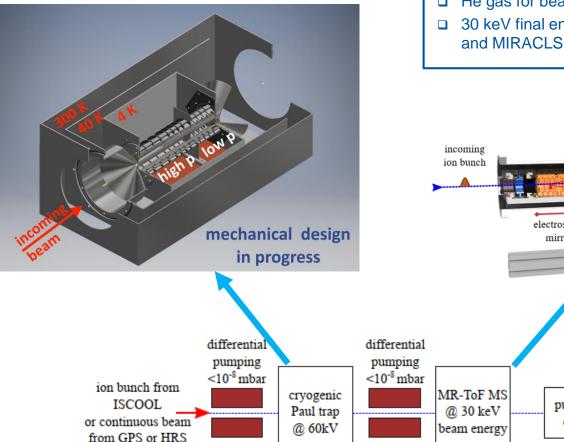




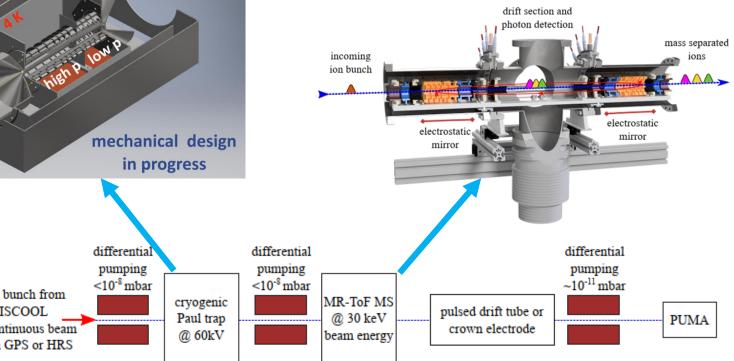
- □ PUMA transfer line connected to MIRACLS  $\Rightarrow$  Handover Point <10<sup>-8</sup> mbar
- Conductance reduction over ~2 meters, smallest diameter 20 mm. Chicane to reduce beaming.
- Line still under definition







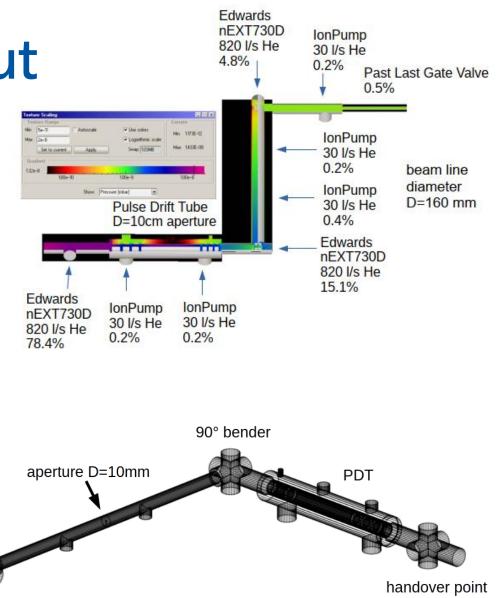
- MIRACLS: Multi Ion Reflection Apparatus for Collinear Laser Spectroscopy
- □ Bunches trapped in MR-ToF MS (Multi-Reflection Time-of-Flight Mass Spectrometer)
- Previous cryogenic Paul Trap for beam cooling with buffer gas for optimal longitudinal emittance.
- □ He gas for beam cooling!
- □ 30 keV final energy  $\rightarrow$  Pulsed drift tube between PUMA and MIRACLS (<100 eV at PUMA)







- Light ions (A<15) can only be efficiently cooled down with He gas
- O.5% transmission probability for He
- PUMA beam line at ISOLDE still to be optimised.
- Minimize the time He is injected while not in operation (open valve only when required) to PUMA



90° bender



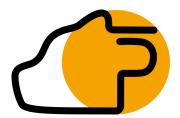
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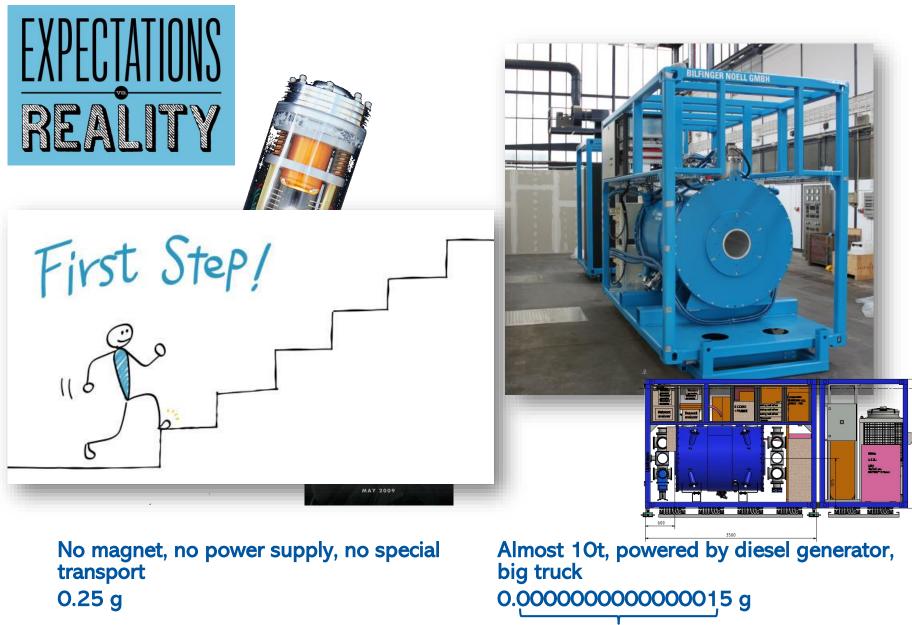
## Summary

- Dependence of p and  $\bar{p}$  as a high precision probe. The manipulation and transport of  $\bar{p}$  requires extreme low vacuum only achievable with cryo-pumping
- COMSOL and MOFLOW models to design the cold trap and the transfer lines
- □ The pressure at the entrance of the trap is crucial to reach the objective → Extensive use of NEG coatings and strict outgassing budget
- □ Isotherm data at very low pressure required to refine the models
- □ Aspects that require further development:
  - □ The trap needs to be loaded with e<sup>-</sup> using a cold field emission source. At the end of the manipulation, they should be extracted out of the trap
  - **Effect** of vibrations during transport
- $\Box$  PUMA and BASE-STEP are targeting the transport of  $\bar{p} \Rightarrow$  It will open new physics opportunities





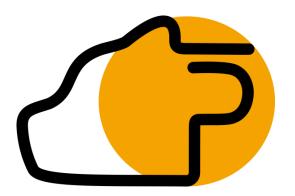




×10<sup>-15</sup>



## Thank you for your attention!!





#### References

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[2] X. Fei, Trapping Low-Energy Antiprotons in an Ion Trap., Harvard University, 1990. https://ui.adsabs.harvard.edu/abs/1990PhDT......101F

[3] F. Chill, S. Wilfert, and L. Bozyk, J. of Vac. Sci. Technol. A 37, 031601 (2019)

[4] E. Wallén, J. Vac. Sci. Technol. A 15, 265 (1997)

[5] N. Fray and B. Schmitt, Planet. Space Sci. 57, 2053 (2009)

[6] Handbook of Vacuum Technology, 1st ed. (John Wiley & Sons, Ltd, 2016)

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[8] Salemme, Roberto, et al. "Vacuum Performance of Amorphous Carbon Coating at Cryogenic Temperature with Presence of Proton Beams." 7th Int. Particle Accelerator Conf.(IPAC'16), Busan, Korea, May 8-13, 2016. JACOW, Geneva, Switzerland, 2016.
[9] C. Benvenuti, R.S. Calder, and G. Passardi, J. Vac. Sci. Technol. 13, 1172 (1976).

[10] J.W. Leachman, R.T. Jacobsen, S.G. Penoncello, and E.W. Lemmon, Journal of Physical and Chemical Reference Data **38**, 721 (2009).





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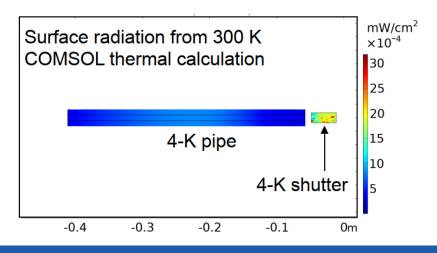
#### Milestones

- Install and commissioning 100 kV Pulse Drift Tube at ELENA
- $\hfill\square$  Build ELENA  $\bar{p}$  transfer line and commissioning
- $\Box$  Install trap and first  $\bar{p}$  trapped
- □ Install ion line at ELENA
- □ First ion trapped and  $\bar{p}$  annihilation with stable ions (first physics run)
- Build ISOLDE transfer line
- $\hfill\square$  Transport of  $\bar{p}$  to ISOLDE
- Physics with unstable ions



## Influence of thermal radiation

- H<sub>2</sub> desorption can be stimulated by thermal radiation [9]
- First simulations show <<10<sup>-3</sup> mW·cm<sup>-2</sup>
- Shielding of warm areas is mandatory



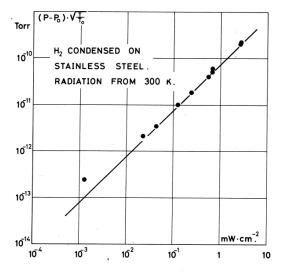
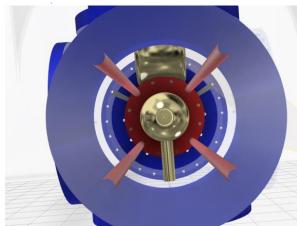


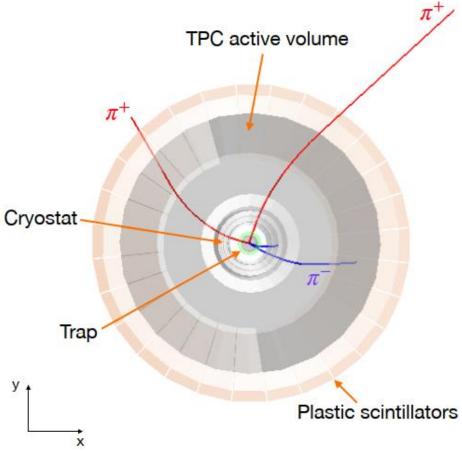
Fig. 11. Dependence of the H<sub>2</sub> saturated vapour pressure at 2.3 K on the thermal radiation power absorbed by the cryosurface.





#### Detection

- Measure the number of annihilated neutrons to the number of annihilated proton
- □ Electric charge conservation during the annihilation process → total charge of pions −1 <sup>y</sup> for a neutron and 0 for a proton.





## **Density vs Pressure**

- □ The flow through orifices in molecular flow disturbs the speed distribution → No longer isotropic.
- PV = nRT doesn't apply in those conditions
- It is preferred to use density for cryogenic systems and/or anisotropic distributions

