

Constraints on low-energy flavor from LFV and $(g-2)_\mu$

Oscar Vives

6th Red LHC workshop, 9 - 11 de Mayo



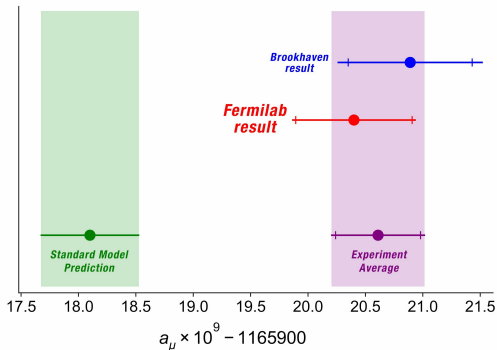
L Calibbi, M.L. López-Ibáñez, A. Melis and O.V., JHEP 06 (2020), 087

L Calibbi, M.L. López-Ibáñez, A. Melis and O.V., Eur. Phys. J. C 81 (2021) no.10, 929

M. L. López-Ibáñez, A. Melis, M. J. Pérez, M. H. Rahat and O. V., Phys. Rev. D 105 (2022) 035021

First results from the **Muon g-2** at Fermilab

Run1 in *Muon g-2* confirmed Brookhaven discrepancy:



$$a_\mu(\text{Exp}) - a_\mu(\text{SM}) = (251 \pm 59) \times 10^{-11} \Rightarrow \boxed{4.2 \sigma}$$

Very large discrepancy, compared with EW contribution,

$$a_{\mu}^{\text{EW}} = 153.6 (1.0) \times 10^{-11}$$

and at one loop ...

$$\begin{aligned} a_{\mu}^{\text{EW}(1)} &= \frac{G_F}{\sqrt{2}} \frac{m_{\mu}^2}{8\pi^2} \left[\frac{5}{3} + \frac{1}{3}(1 - 4 \sin^2 \theta_W) \right] \\ &= \frac{\alpha}{4\pi \sin^2 \theta_W} \frac{1.67 m_{\mu}^2}{4 M_W^2} = 194.79 (1) \times 10^{-11} \end{aligned}$$

any “natural” new physics with “minimal” flavor structure

$$a_{\mu}^{\text{NP}} \simeq \frac{\alpha}{4\pi \sin^2 \theta_W} \frac{m_{\mu}^2}{M_X^2} = 251 \times 10^{-11}$$

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$\Rightarrow M_X = 106 \text{ GeV}$. So ...

Where is New Physics??

Additional enhancing factors in NP contributions are possible ...

$$a_{\mu}^{\text{NP}} \simeq \frac{\alpha}{4\pi \sin^2 \theta_W} \frac{m_{\mu}^2}{M_X^2} k$$

- In SUSY: $k = \tan \beta$, Scalar Leptoquarks: $k = m_t/m_{\mu} \simeq 1600$

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But ...

Large enhancements in anomalous magnetic moments, produce enormous contributions to the fermion mass !!

$$\Delta a_{\mu}^{\text{NP}} \simeq \frac{\alpha}{4\pi \sin^2 \theta_W} \frac{m_{\mu}^2}{M_X^2} k \Rightarrow \Delta m_{\mu}^{\text{NP}} \simeq m_{\mu} \frac{\alpha k}{4\pi \sin^2 \theta_W}$$



Absence of fine-tuning requires $k \lesssim 4\pi \sin^2 \theta_W / \alpha \simeq 380$
largest possible (fully radiative m_{μ}) $M_X \lesssim 2 \text{ TeV} !!!$

Low-energy flavor symmetries

Flavour symmetry explains masses and mixings in Yukawas.

Small couplings generated in Froggatt-Nielsen, as function of small

vevs, $Y_{ij} = \left(\left(\frac{\langle \theta \rangle}{M} \right) \ll 1 \right)^n$.

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Magnetic moments and Yukawas have identical flavor charges.

\Rightarrow identical flavons enter the Dipole and Yukawa matrices !!

$$Y_\ell = y_{33} \begin{pmatrix} \lambda^{5+a} & \lambda^{2+b} & \lambda^c \\ \lambda^{2+d} & \lambda^{2+e} & \lambda^f \\ \lambda^g & \lambda^h & 1 \end{pmatrix}$$

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with $\lambda = 0.2$ and $a, b, c, d, e, f, g, h \geq 0$.

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For low Λ_f , lepton flavor violation, $\mu \rightarrow e\gamma$, require nearly diagonal charged-lepton Yukawa matrices, $a, e = 0, b, d \geq 6.6, g, h, c, f \geq 2.7$.

T₁₃ Model

T₁₃ = Z₁₃ × Z₃, O(39) finite subgroup of SU(3), with two ineq. complex triplets, 3₁, 3₂.

Fields	\bar{L}	l	H	φ_{33}	φ_{22}	φ_{13}	Δ	χ	χ'
$SU(2)_L$	2	1	2	1	1	1	1	1	1
T_{13}	3 ₁	3 ₁	1	$\bar{3}_1$	$\bar{3}_1$	3 ₁	$\bar{3}_1$	1	1
Z ₄	η^1	1	η^1	η^1	η^3	η^2	η^2	η^3	η^1

Diagonal and off-diagonal in different T₁₃ representations.

$$\begin{pmatrix} \bar{L}_1 \\ \bar{L}_2 \\ \bar{L}_3 \end{pmatrix}_{3_1} \otimes \begin{pmatrix} l_1 \\ l_2 \\ l_3 \end{pmatrix}_{3_1} = \begin{pmatrix} \bar{L}_1 l_1 \\ \bar{L}_2 l_2 \\ \bar{L}_3 l_3 \end{pmatrix}_{3_2} \oplus \begin{pmatrix} \bar{L}_2 l_3 \\ \bar{L}_3 l_1 \\ \bar{L}_1 l_2 \end{pmatrix}_{\bar{3}_1} \oplus \begin{pmatrix} \bar{L}_3 l_2 \\ \bar{L}_1 l_3 \\ \bar{L}_2 l_1 \end{pmatrix}_{\bar{3}_1} .$$

Three flavons: $\varphi_{33} \sim \bar{3}_1$, $\varphi_{22} \sim \bar{3}_1$, $\varphi_{13} \sim 3_1$,

$$\langle \varphi_{33} \rangle = \epsilon (0, 0, 1) \Lambda_f, \quad \langle \varphi_{22} \rangle = \epsilon^2 (0, 1, 0) \Lambda_f,$$

$$\langle \varphi_{13} \rangle = \epsilon^{5/2} (0, 1, 0) \Lambda_f,$$

Yukawas Lagrangian:

$$L_Y^e = \bar{L} H \left[\frac{1}{\Lambda_f} \varphi_{13} + \frac{1}{\Lambda_f^2} \varphi_{22} \varphi_{22} + \frac{1}{\Lambda_f^2} \varphi_{33} \varphi_{33} \right].$$

Tree-level + loop Yukawas:



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Tree-level + loop Yukawas:

$$Y_l/\epsilon^2 = \begin{pmatrix} 0 & 0 & \epsilon^{5/2} \\ 0 & \epsilon^2 & 0 \\ \epsilon^{5/2} & 0 & y_1 \end{pmatrix} + \frac{f_1(x_\varphi^2)}{32\pi^2} \begin{pmatrix} 0 & 0 & 2y_2 y_4 \epsilon^{5/2} \\ 0 & \beta_2 y_2 \epsilon^2 & 0 \\ 2y_2 y_4 \epsilon^{5/2} & 0 & y_1 \beta_3 y_3 \end{pmatrix}$$

But, Dipole matrix similar structure

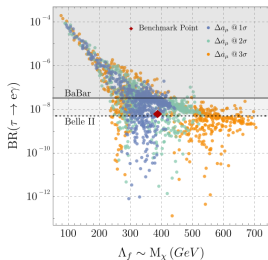
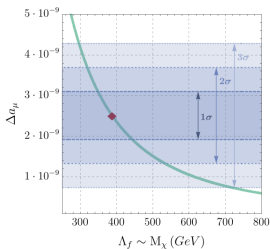
$$C_I = \frac{f_2(x_\varphi^2)}{16} \frac{\epsilon^2}{\Lambda_f^2} \begin{pmatrix} 0 & 0 & 2 y_2 y_4 \epsilon^{5/2} \\ 0 & \beta_2 y_2 \epsilon^2 & 0 \\ 2 y_2 y_4 \epsilon^{5/2} & 0 & y_1 \beta_3 y_3 \end{pmatrix},$$

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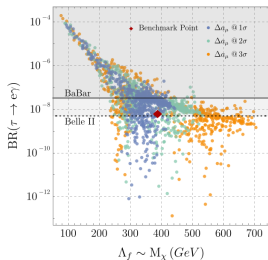
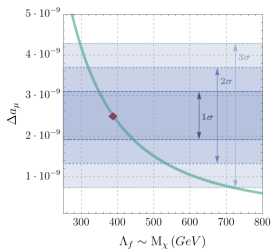
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y_1	y_2	y_3	y_4	β_2	β_3	β_4	ϵ	x_φ
0.50	0.50	3.50	0.70	0.25	1.56	-6.24	0.15	0.08

Conclusions

- Low scale New Physics required to explain the muon $(g-2)$ anomaly.
- Absence of fine-tuning requires electroweak-charged particles with $M_X \lesssim 1$ TeV
- Low-energy flavor symmetries can explain the muon (and electron anomalies).
- Realistic low-energy T_{13} (others...) possible to explain $(g-2)_\mu$.
- Vector-like fermions with muon quantum numbers and new scalars at reach in LHC.

Backup

LHC searches

- **Vector-like** fermions with muon quantum numbers required with $M_x \lesssim 1$ TeV (in absence of fine-tuning) .
- Weak-production at LHC: $p p \rightarrow \chi^+ \chi^-$
- Large couplings, $\mathcal{O}(1)$, to muon and scalar flavons
- Flavons, ϕ , decay to heavier fermions, $\mu^+ \mu^-$ (τ with three generations)

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Typical signature

$$p p \rightarrow \chi^+ \chi^- \rightarrow \phi_1(\rightarrow \mu^+ \mu^-) \ell^+ \phi_1(\rightarrow \mu^+ \mu^-) \ell^-, \quad \text{with } \ell = e, \mu$$

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**New Physics associated with $(g - 2)_\mu$
should be visible at LHC.**

LFV Constraints

$$\mathcal{L} \supset \frac{e\nu}{8\pi^2} C_{\ell\ell'} (\bar{\ell}\sigma_{\mu\nu}P_R\ell') F^{\mu\nu} + \text{h.c.} \quad \ell, \ell' = e, \mu, \tau$$

- $\Delta a_\ell = \frac{m_\ell\nu}{2\pi^2} \text{Re}(C_{\ell\ell}), \quad \text{B}_{\ell\rightarrow\ell'\gamma} = \frac{3\alpha}{\sqrt{2}\pi G_F^3 m_\ell^2} (|C_{\ell\ell'}|^2 + |C_{\ell'e\ell}|^2) \text{B}_{\ell\rightarrow\ell'\nu\bar{\nu}'}$

	$M_X = 520\sqrt{\kappa} \text{ GeV}$	
$\text{Re}(C_{\mu\mu})$	$[1.5, 2.4] \times 10^{-9} \text{ GeV}^{-2}$	λ^2
$\text{Re}(C_{ee})$	$[-1.9, 1.7] \times 10^{-10} \text{ GeV}^{-2}$	λ^5
$\text{Re}(C_{\tau\tau})$	$[-3.7, 1.7] \times 10^{-4} \text{ GeV}^{-2}$	1
$ C_{e\mu} , C_{\mu e} $	$\lesssim 3.9 \times 10^{-14} \text{ GeV}^{-2}$	$\lambda \geq 8.6$
$ C_{\tau\mu} , C_{\mu\tau} $	$\lesssim 5.0 \times 10^{-10} \text{ GeV}^{-2}$	$\lambda \geq 2.6$
$ C_{\tau e} , C_{e\tau} $	$\lesssim 4.3 \times 10^{-10} \text{ GeV}^{-2}$	$\lambda \geq 2.7$

with $\kappa \simeq 20$ if mass is fully radiative