

ADRIAN CARMONA BERMUDEZ

THE ALPS FROM THE TOP

6th Red LHC workshop - Madrid IFT



OFPI
Oficina de Proyectos
Internacionales

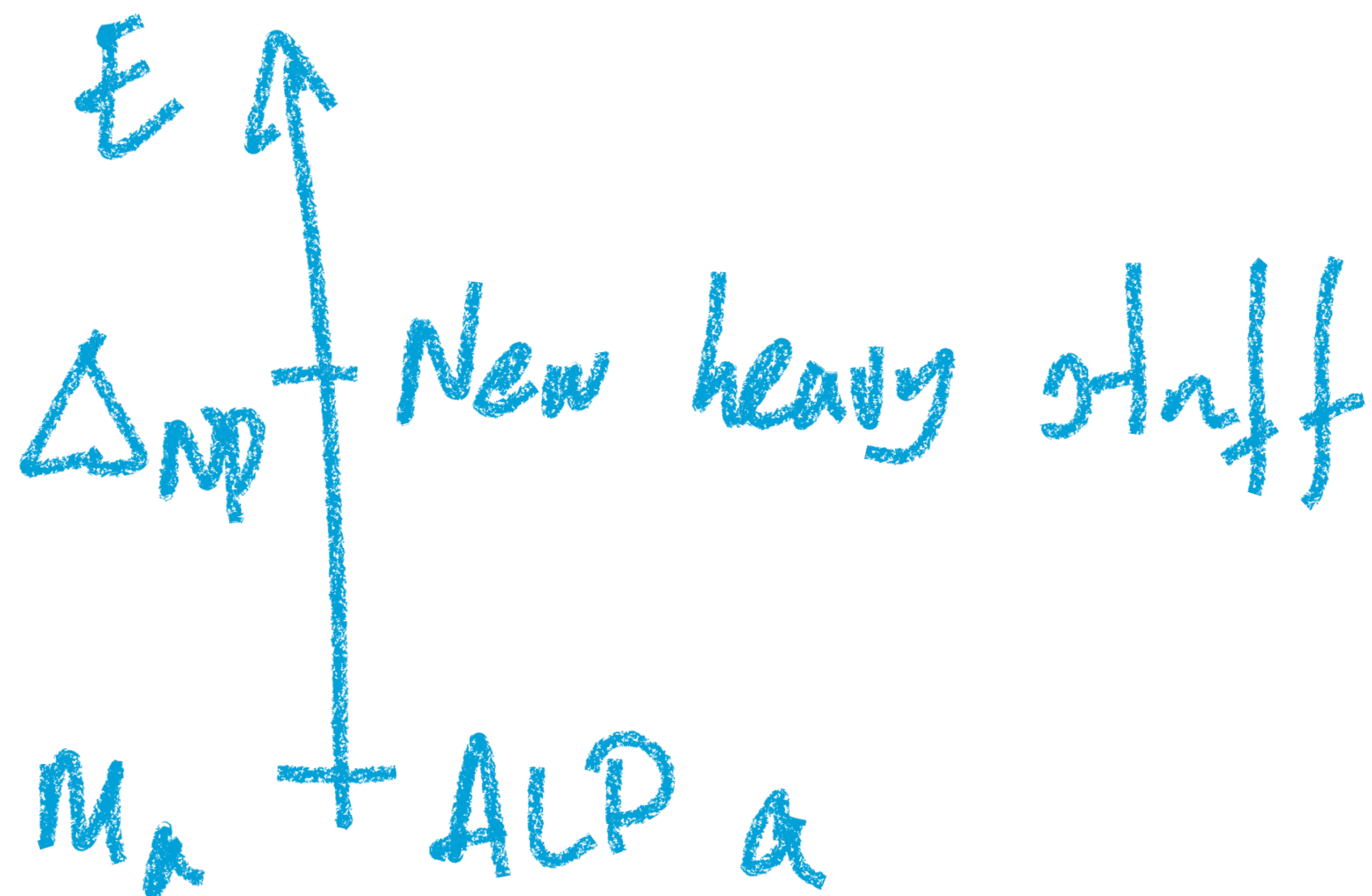


UNIVERSIDAD
DE GRANADA

ALPS

Axion-like particles (ALPs) are pseudo-Nambu-Goldstone bosons (pNGBs) of a spontaneously broken global symmetry

One typically assumes that CP is a good symmetry and that the ALP is CP-odd



$$a \rightarrow a + \theta$$

ALP shift symmetry

ALP + SMEFT

ALP EFT above the electroweak scale

$$\psi = q_L, l_L, u_R, d_R, e_R$$

$$\mathcal{L} = \frac{1}{2} \partial_\mu a \partial^\mu a - \frac{m_a^2}{2} a^2 + \frac{\partial_\mu a}{f_a} \sum_\psi \left(c_\psi \right)_{ij} \bar{\psi}_i \gamma^\mu \psi_j - \frac{a}{f_a} \left[c_{GG} \frac{g_3^2}{32\pi^2} G_{\mu\nu}^a \tilde{G}^{a\mu\nu} + c_{WW} \frac{g_2^2}{32\pi^2} W_{\mu\nu}^I \tilde{W}^{I\mu\nu} + c_{BB} \frac{g_1^2}{32\pi^2} B_{\mu\nu} \tilde{B}^{\mu\nu} \right]$$

We assume the following EFT at the UV (**CHARMING ALPS**)

$$\mathcal{L} = \frac{1}{2} (\partial_\mu a) (\partial^\mu a) - \frac{m_a^2}{2} a^2 + \frac{\partial_\mu a}{f_a} (c_{uR})_{ij} \left(\bar{u}_{Ri} \gamma^\mu u_{Rj} \right)$$

ALP + SMEFT

ALP EFT above the electroweak scale

$$\psi = q_L, l_L, u_R, d_R, e_R$$

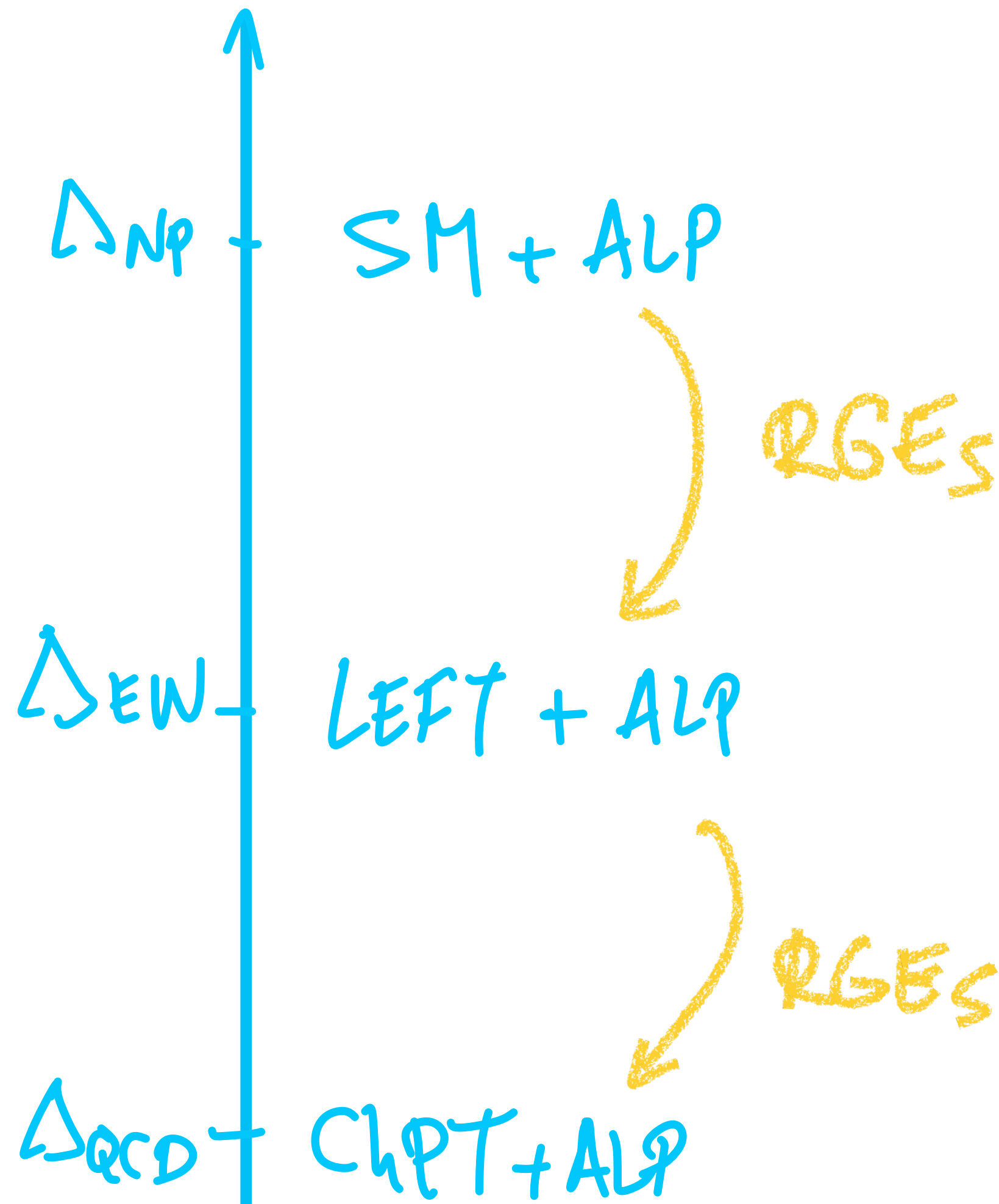
$$\mathcal{L} = \frac{1}{2} \partial_\mu a \partial^\mu a - \frac{m_a^2}{2} a^2 + \frac{\partial_\mu a}{f_a} \sum_\psi \left(c_\psi \right)_{ij} \bar{\psi}_i \gamma^\mu \psi_j - \frac{a}{f_a} \left[c_{GG} \frac{g_3^2}{32\pi^2} G_{\mu\nu}^a \tilde{G}^{a\mu\nu} + c_{WW} \frac{g_2^2}{32\pi^2} W_{\mu\nu}^I \tilde{W}^{I\mu\nu} + c_{BB} \frac{g_1^2}{32\pi^2} B_{\mu\nu} \tilde{B}^{\mu\nu} \right]$$

We assume the following EFT at the UV (**CHARMING ALPS**)

$$\mathcal{L} = \frac{1}{2} (\partial_\mu a) (\partial^\mu a) - \frac{m_a^2}{2} a^2 + \frac{\partial_\mu a}{f_a} (c_{uR})_{ij} \left(\bar{u}_{Ri} \gamma^\mu u_{Rj} \right) \quad \leftarrow \text{Dark QCD}$$

ONE NEEDS TO RUN

Choi et al, 1708.00021
 Chala et al, 2012.09017
 Bauer et al, 2012.12272
 Bonilla et al, 2107.11392



Even if some Wilson coefficients are zero at the UV they will be generated via the RGEs. For instance

$$c_{q_L} = \frac{Y_u c_{u_R} Y_u}{32\pi^2} \ln \left(\frac{\Lambda_{NP}}{\mu^2} \right), \quad c_H = \frac{3}{8\pi^2} \text{Tr} \left(Y_u c_{u_R} Y_u \right) \ln \left(\frac{\Lambda_{NP}}{\mu^2} \right)$$

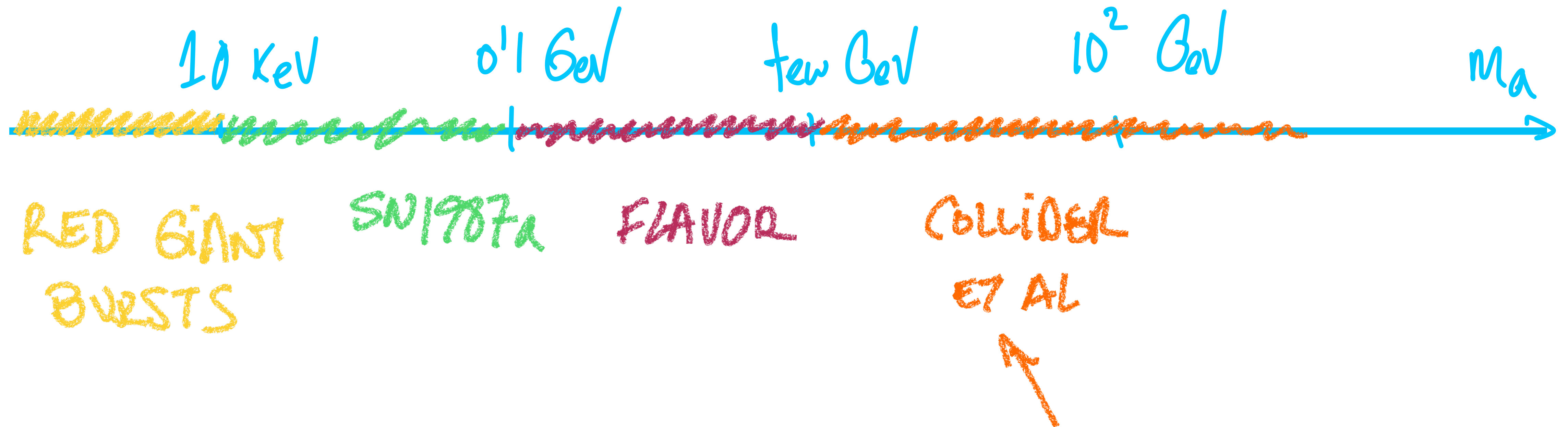
where

$$\left(\mathcal{O}_{q_L} \right)_{ij} = \frac{\partial_\mu a}{\Lambda_{NP}} \left(\bar{q}_{Li} \gamma^\mu q_{Lj} \right), \quad \mathcal{O}_H = \frac{\partial_\mu a}{\Lambda_{NP}} \left(H^\dagger i \overleftrightarrow{D}^\mu H \right)$$

Top couplings will make a difference!

ALPS PHENOMENOLOGY

Flavor probes will compete or be complemented by astrophysical or cosmological bounds as well as by collider or fixed target experiments



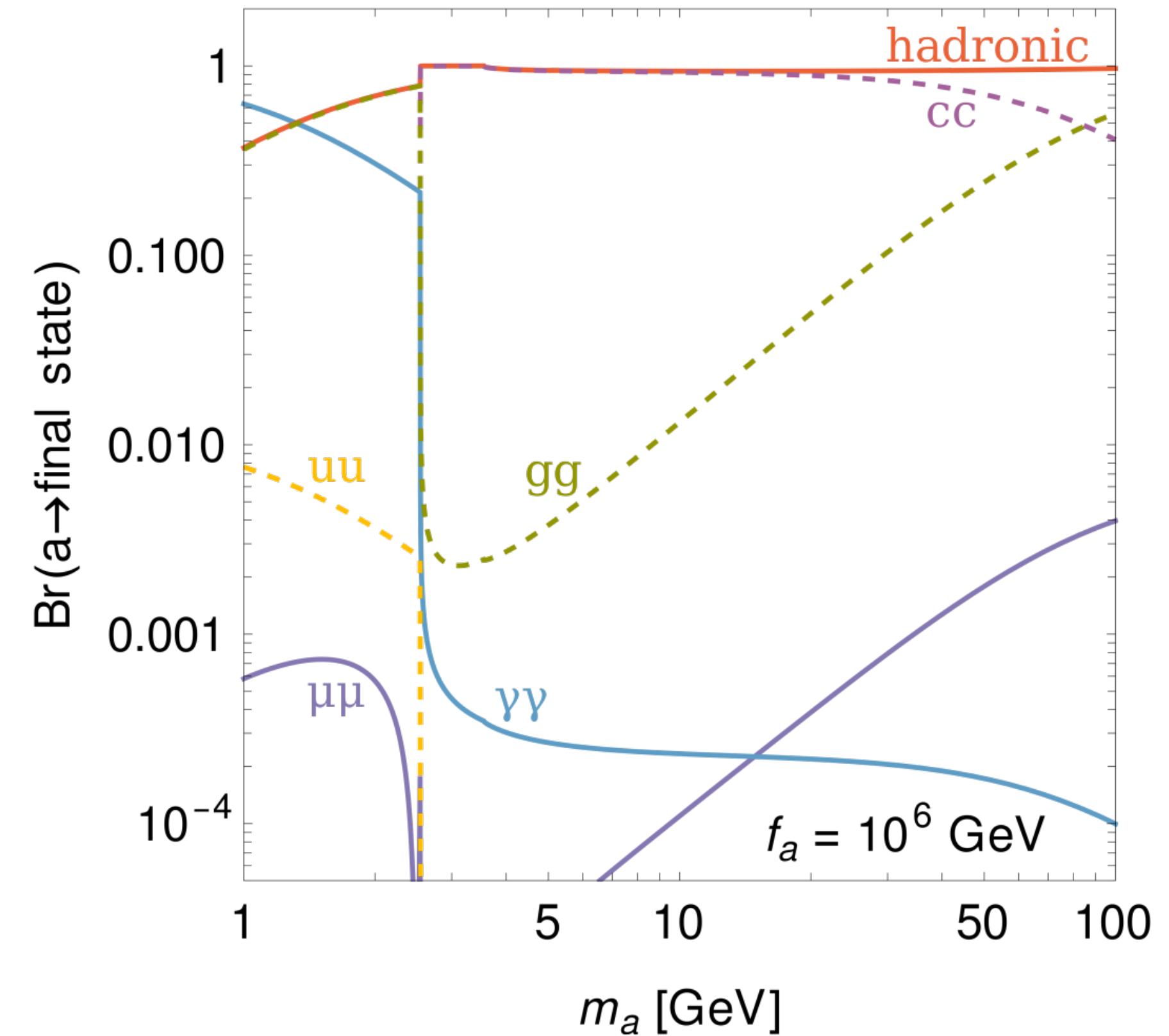
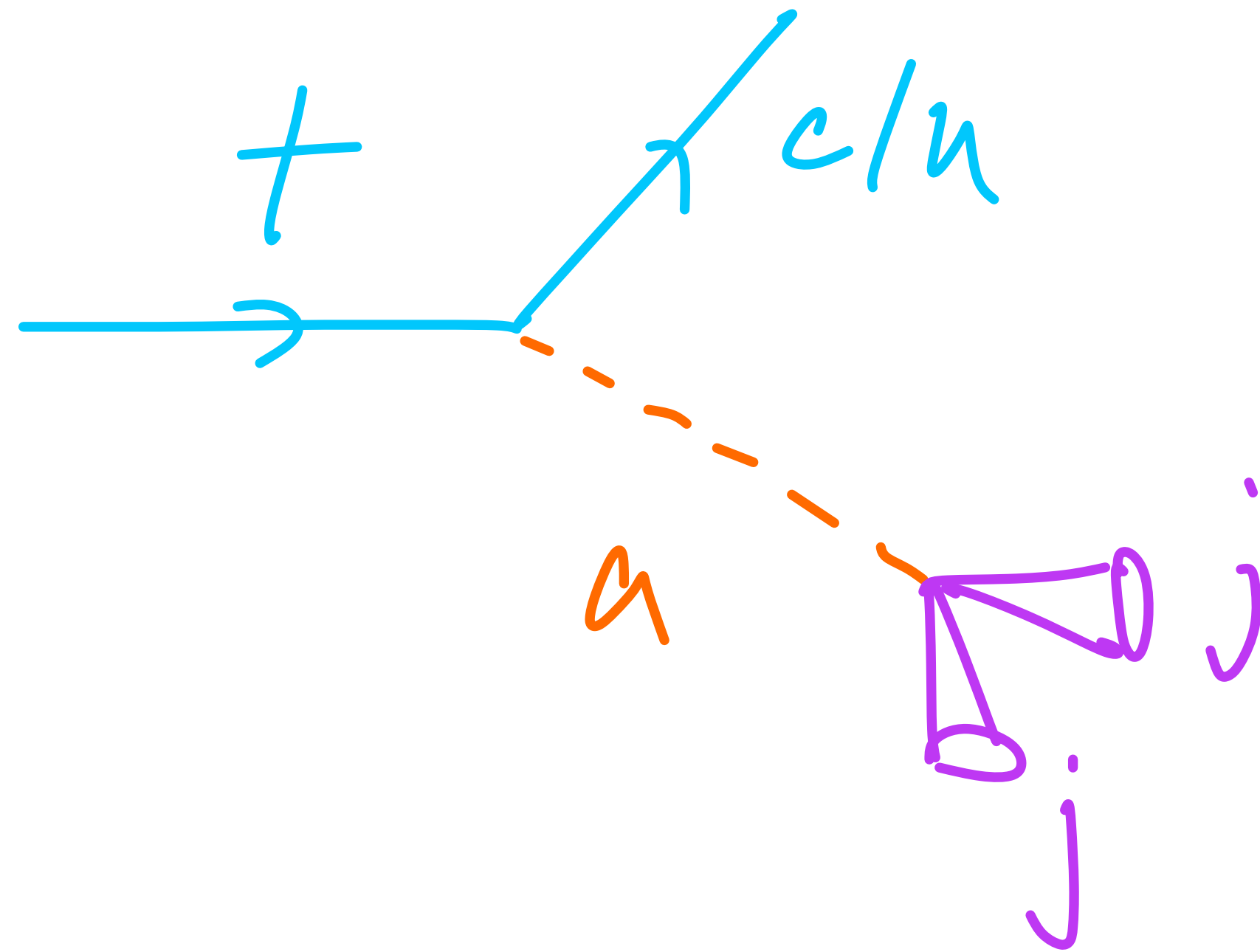
'Charming ALPs' AC, Scherb, Schwaller. JHEP 08 (2021) 121, arXiv: [2101.0783](https://arxiv.org/abs/2101.0783)

'The ALPs from the Top: Searching for long-lived axion-like particles from exotic top decays' AC, Elahi, Scherb, Schwaller. arXiv: [2202.0973](https://arxiv.org/abs/2202.0973)

ALPS FROM THE TOP

'The ALPs from the Top: Searching for long-lived axion-like particles from exotic top decays' AC, Elahi, Scherb, Schwaller. arXiv: [2202.0973](https://arxiv.org/abs/2202.0973)

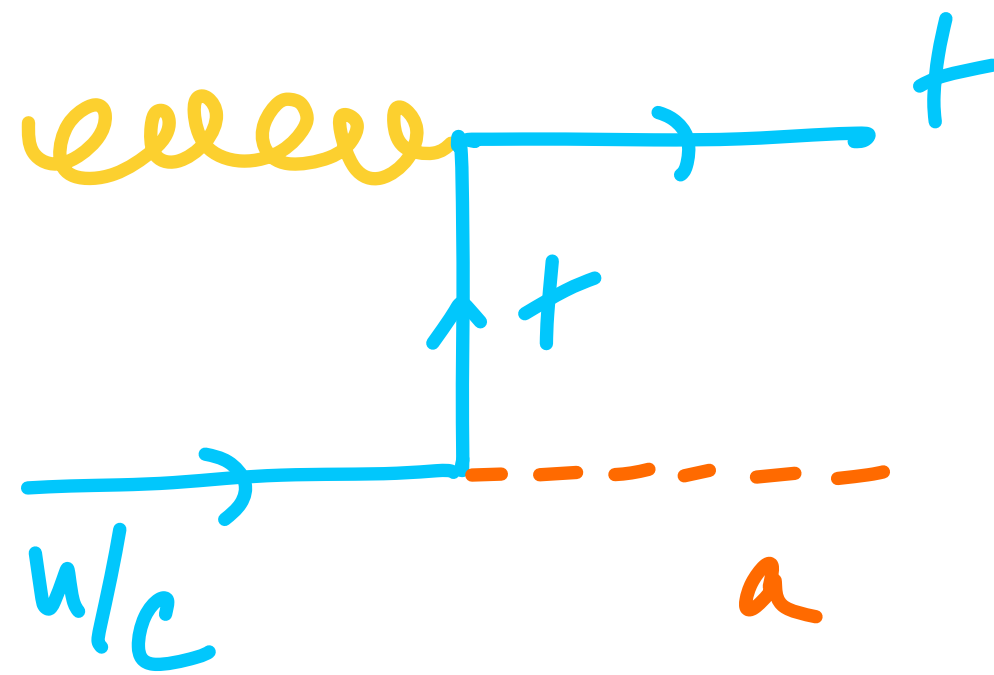
Probe charming ALPs above charm threshold



We trade diagonal and off-diagonal (equal) entries of c_{u_R} by $\text{Br}(t \rightarrow aq_i)$ and $c\tau$

ALPS FROM THE TOPS: CONSTRAINTS

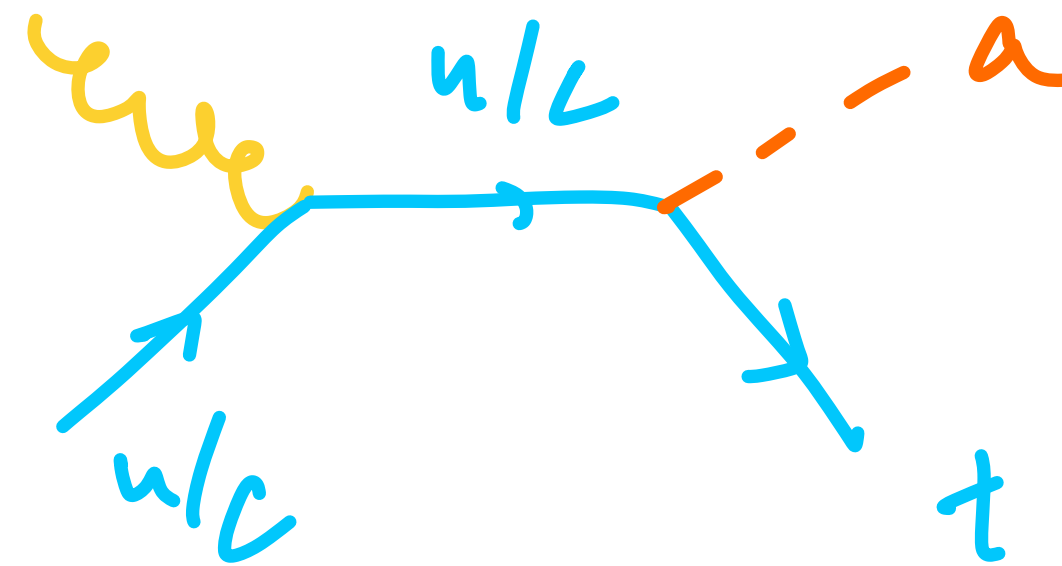
We recast searches from exotic top decays



PROMPT

$$r < 0.01 \text{ cm}$$

$$\int_0^{10^{-4} \text{ m}} (\gamma c \tau)^{-1} \exp\left(-\frac{ct}{\gamma c \tau}\right) d(ct)$$



LONG-LIVED

$$2.5 \text{ cm} < r < 2 \text{ m}$$

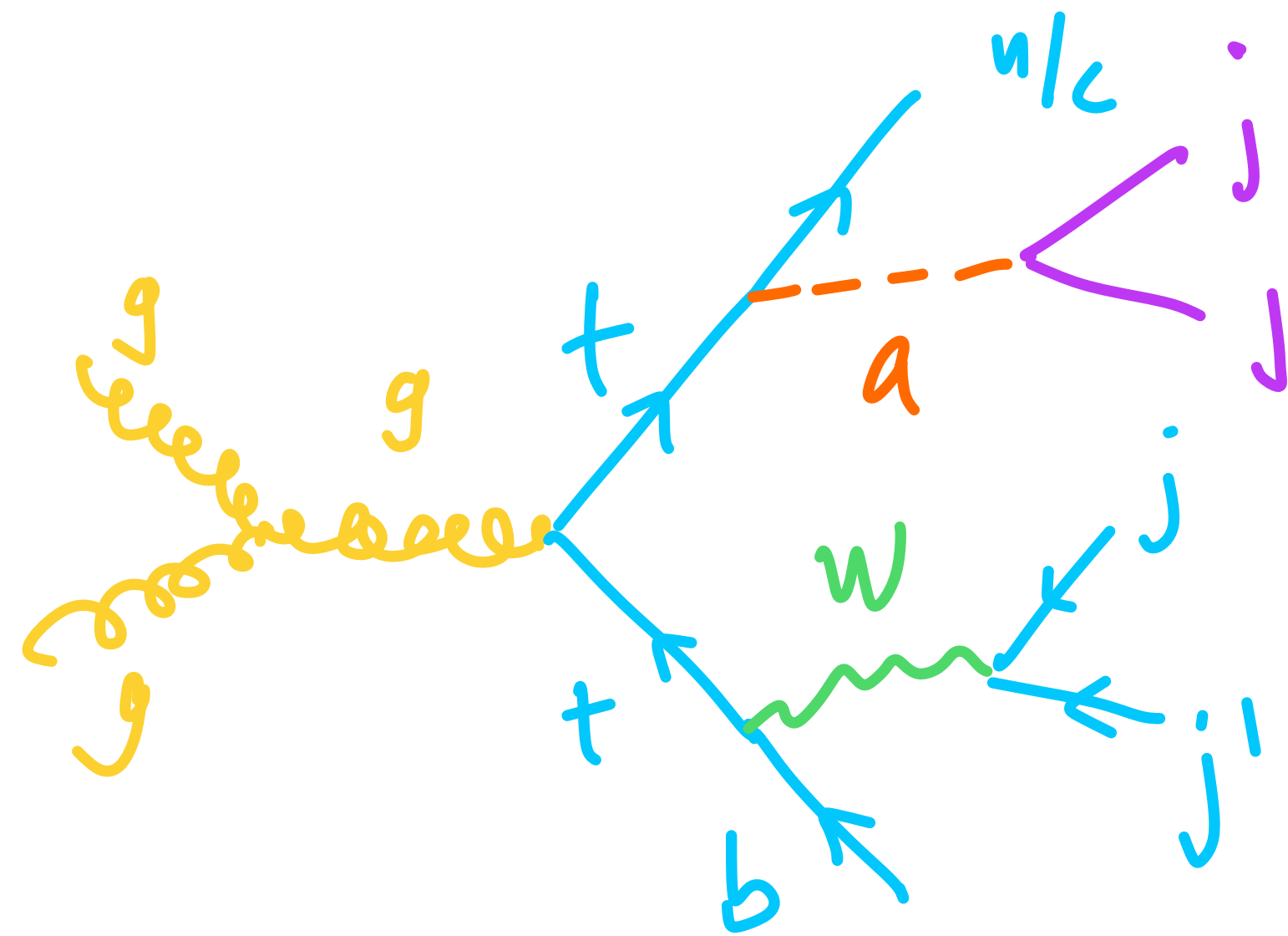
$$\int_{2.5 \cdot 10^{-2} \text{ m}}^{2 \text{ m}} (\gamma c \tau)^{-1} \exp\left(-\frac{ct}{\gamma c \tau}\right) d(ct)$$

'STABLE'

$$c\tau \geq 10 \text{ m}$$

$$\exp\left(-\frac{10 \text{ m}}{\gamma c \tau}\right)$$

SIGNAL



We consider two benchmarks with ALP masses

$$m_a = 2 \text{ GeV}, \quad m_a = 10 \text{ GeV}$$

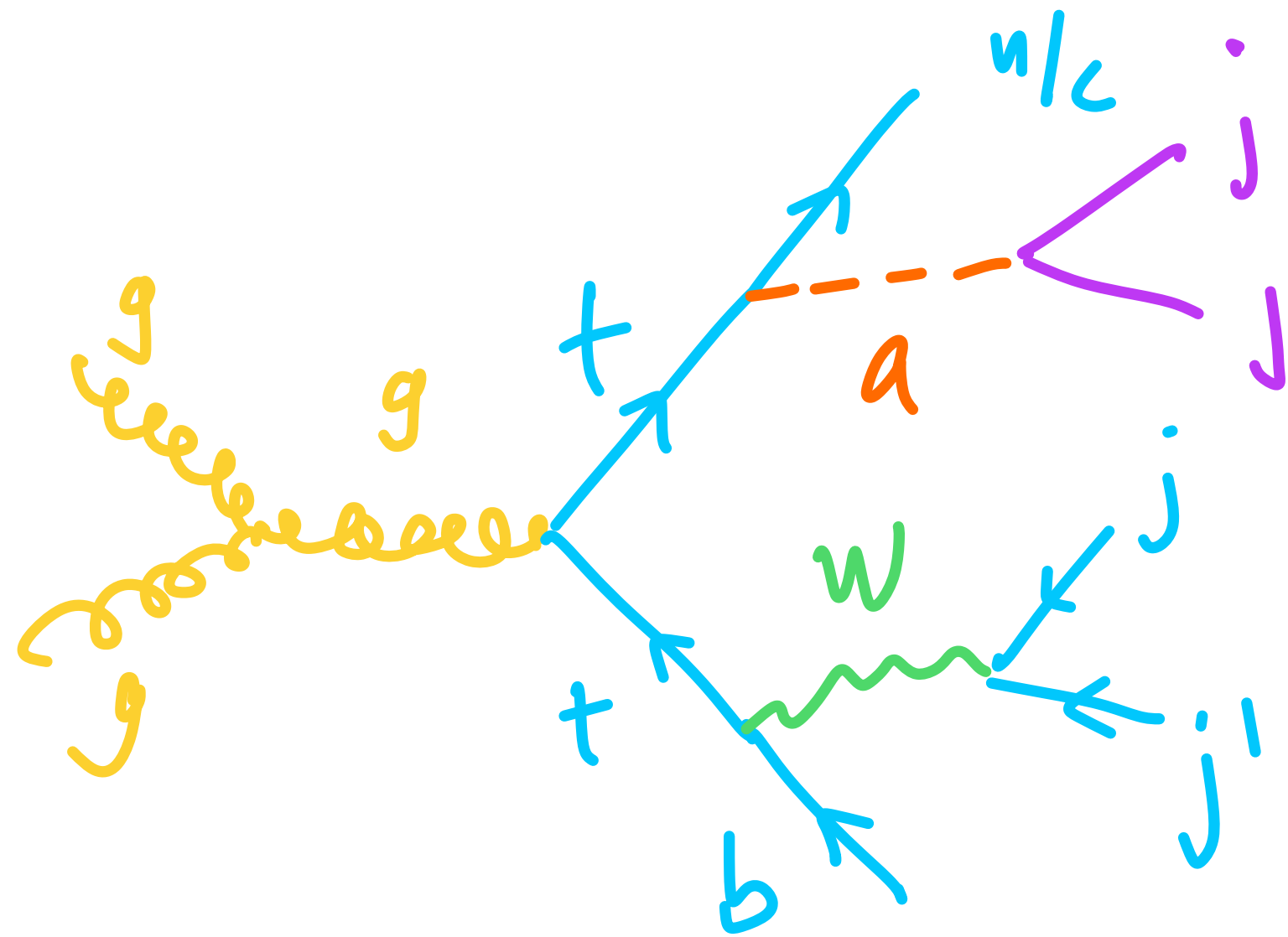
$$\sigma_{\text{signal}} = \sigma_{tt} \cdot \text{Br}(t \rightarrow Wb) \cdot \text{Br}(t \rightarrow aq)$$

$$(c_{u_R})_{ij} = \mathcal{O}(1), \quad f_a = \mathcal{O}(10^5 - 10^9) \text{ GeV} \Rightarrow c\tau \sim 1 \text{ mm} - 100 \text{ m}$$

while having $\text{Br}(t \rightarrow aq) \lesssim 10^{-3}$

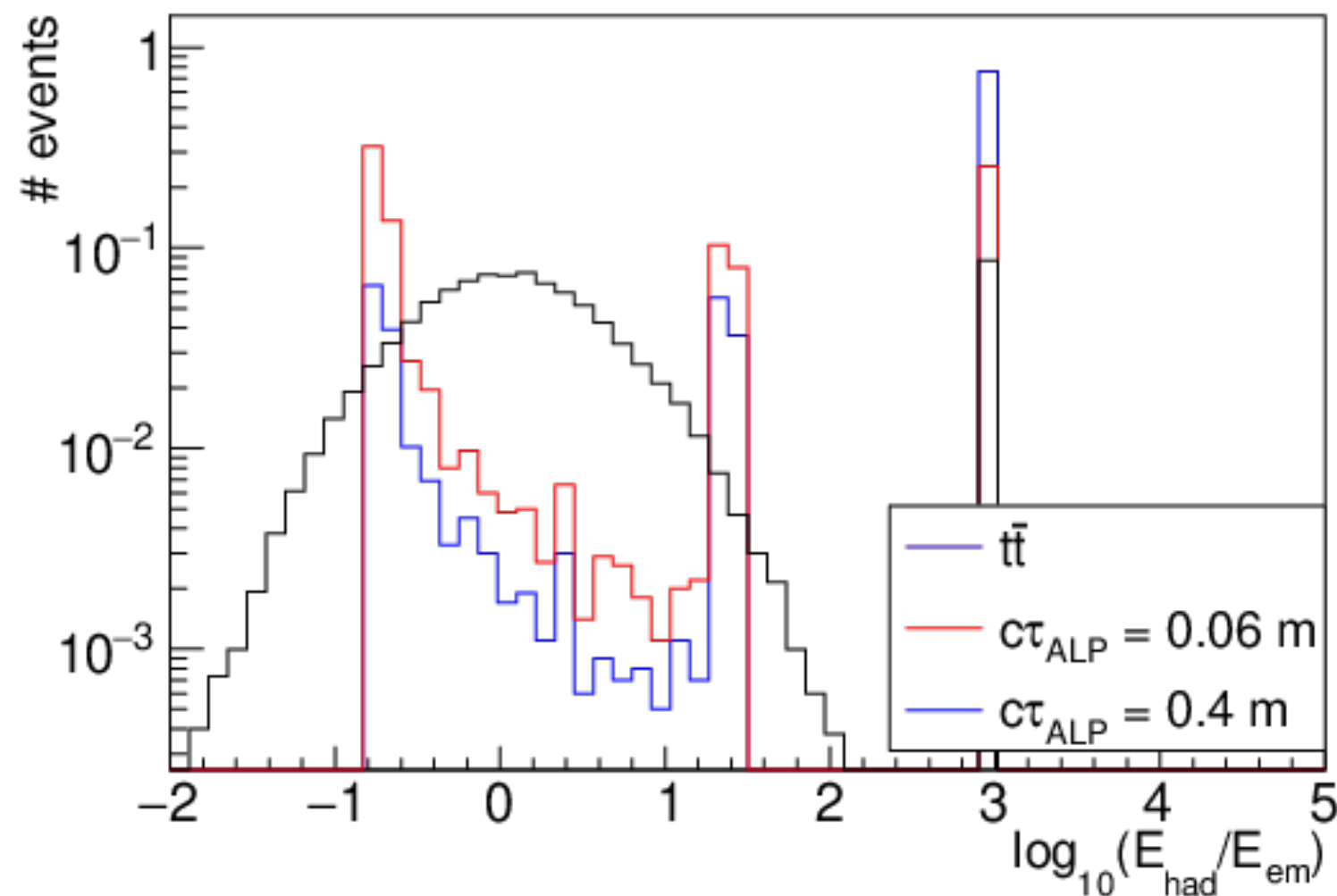
The ALP decay mostly in the $\begin{cases} \nearrow \text{hadronic calorimeter} \\ \searrow \text{muon spectrometer} \end{cases}$

SIGNAL: HADRONIC CALORIMETER

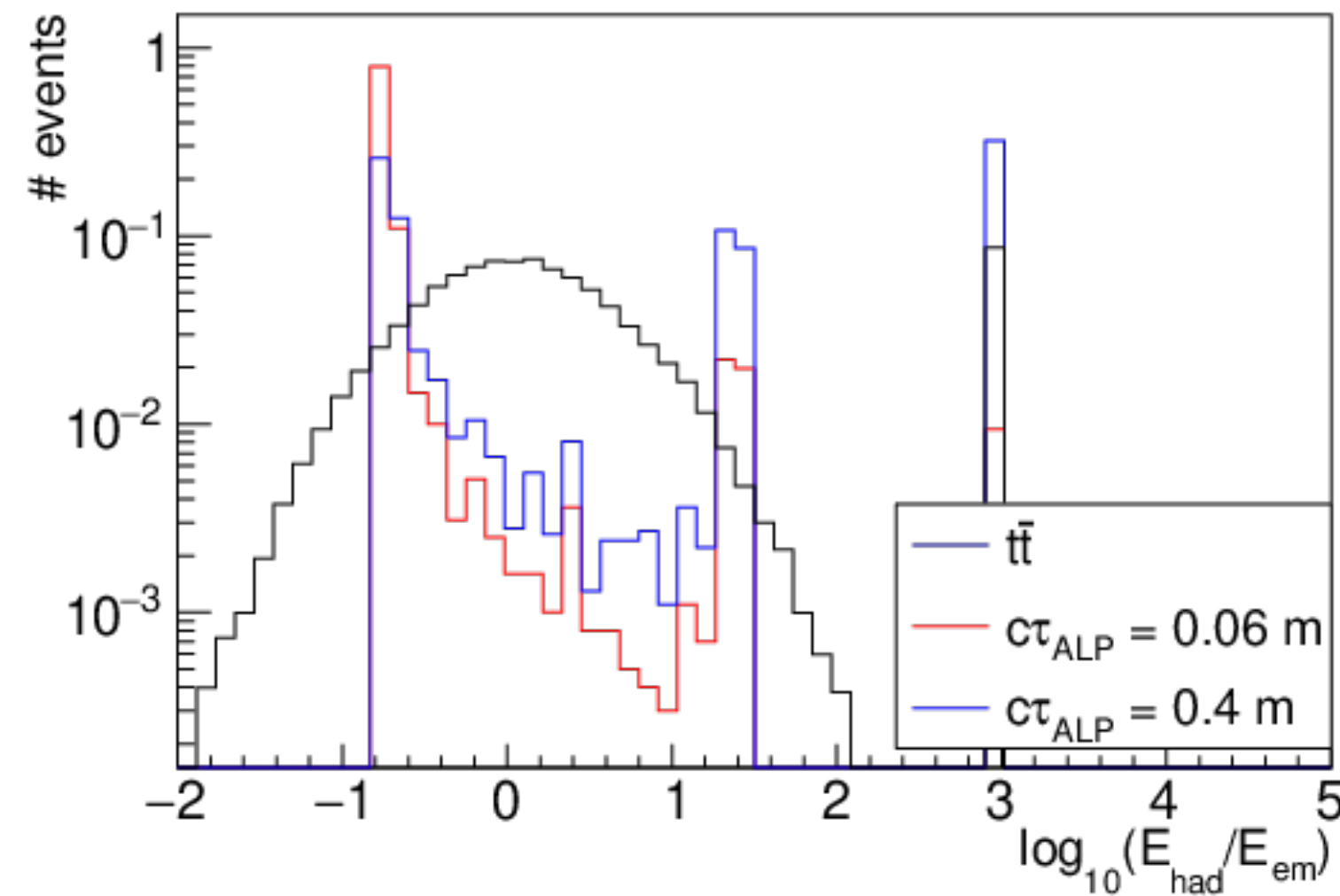


- ✿ Large E_{had}/E_{cal} ratio
- ✿ No tracks in the displaced jet
- ✿ 3-5(6) jets with 1(2) displaced and another b-tagged

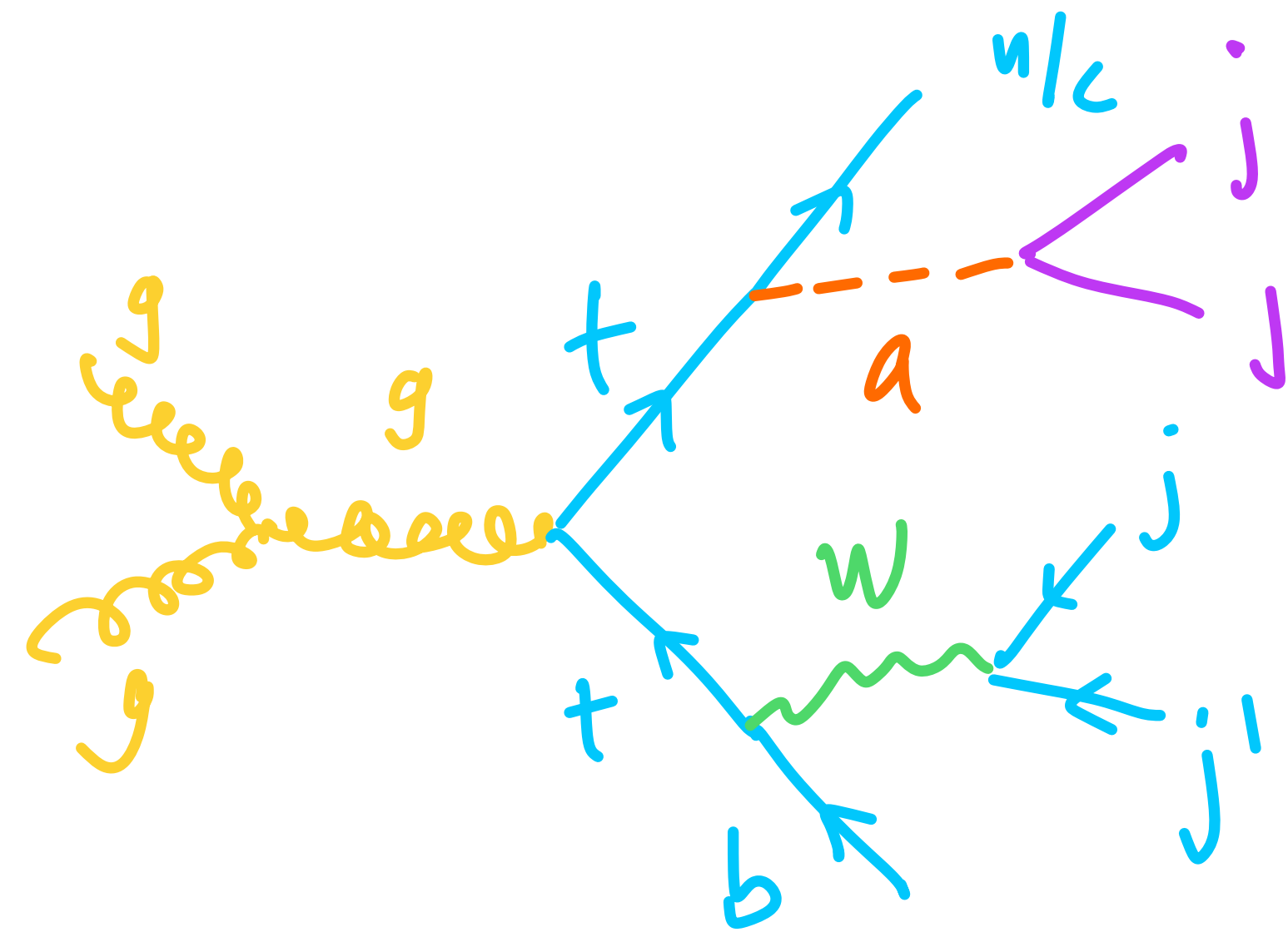
M_a = 2 GeV



M_a = 10 GeV

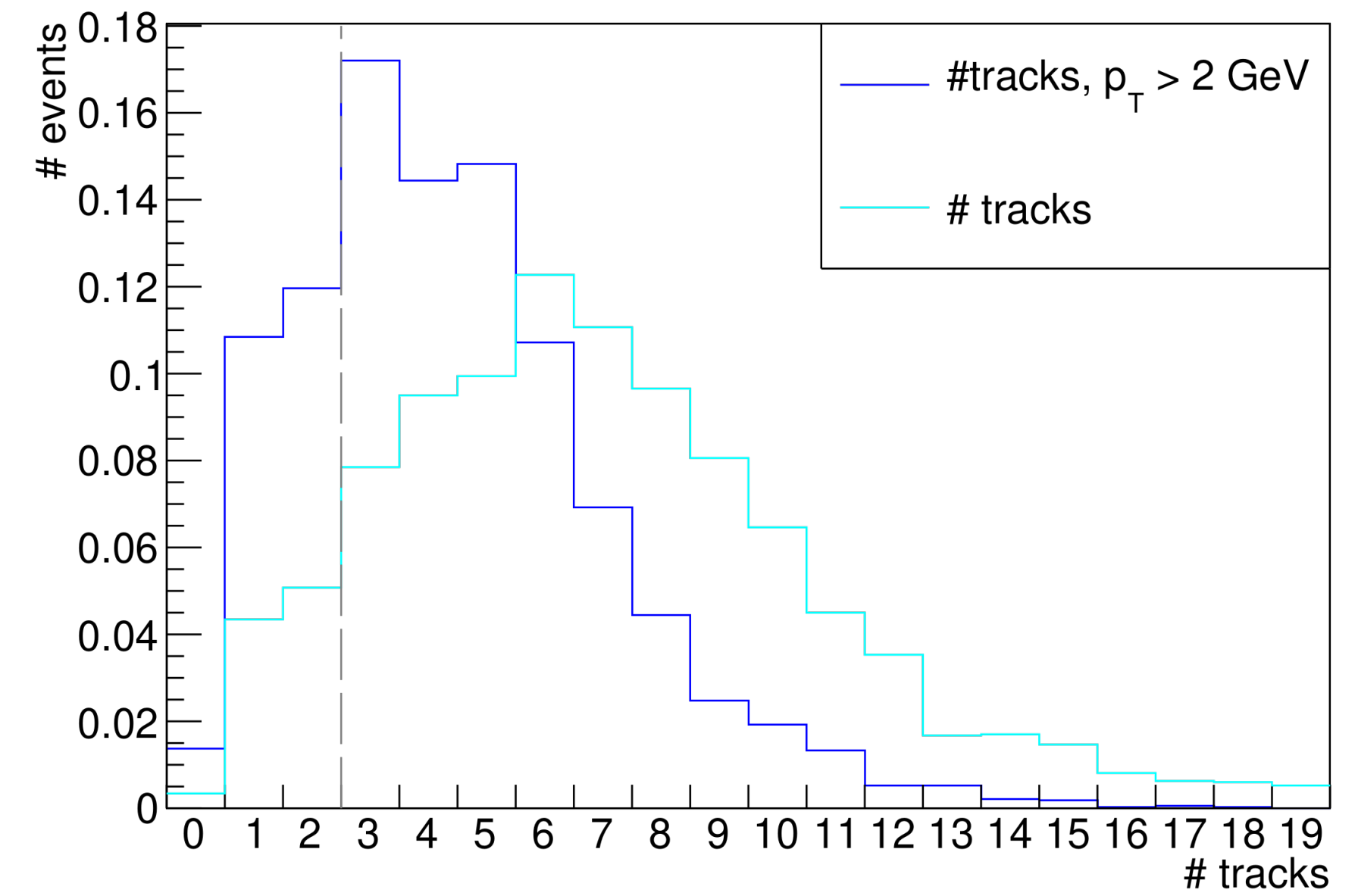


SIGNAL: HADRONIC CALORIMETER

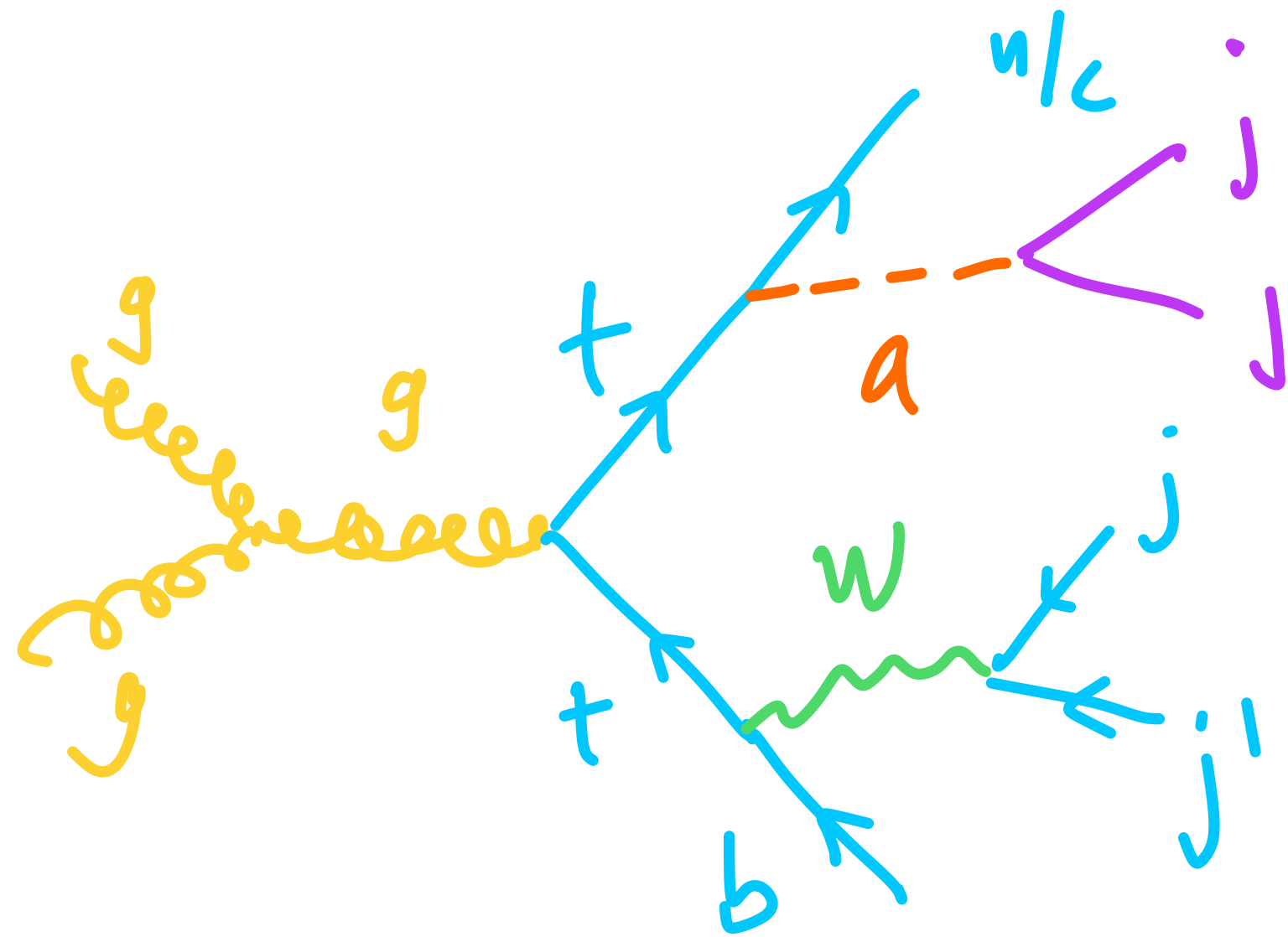


- ✦ Large E_{had}/E_{cal} ratio
- ✦ No tracks in the displaced jet
- ✦ 3-5(6) jets with 1(2) displaced and another b-tagged

For $t\bar{t}$ with $\log_{10}(E_{had}/E_{cal}) > 1.2 \Rightarrow$



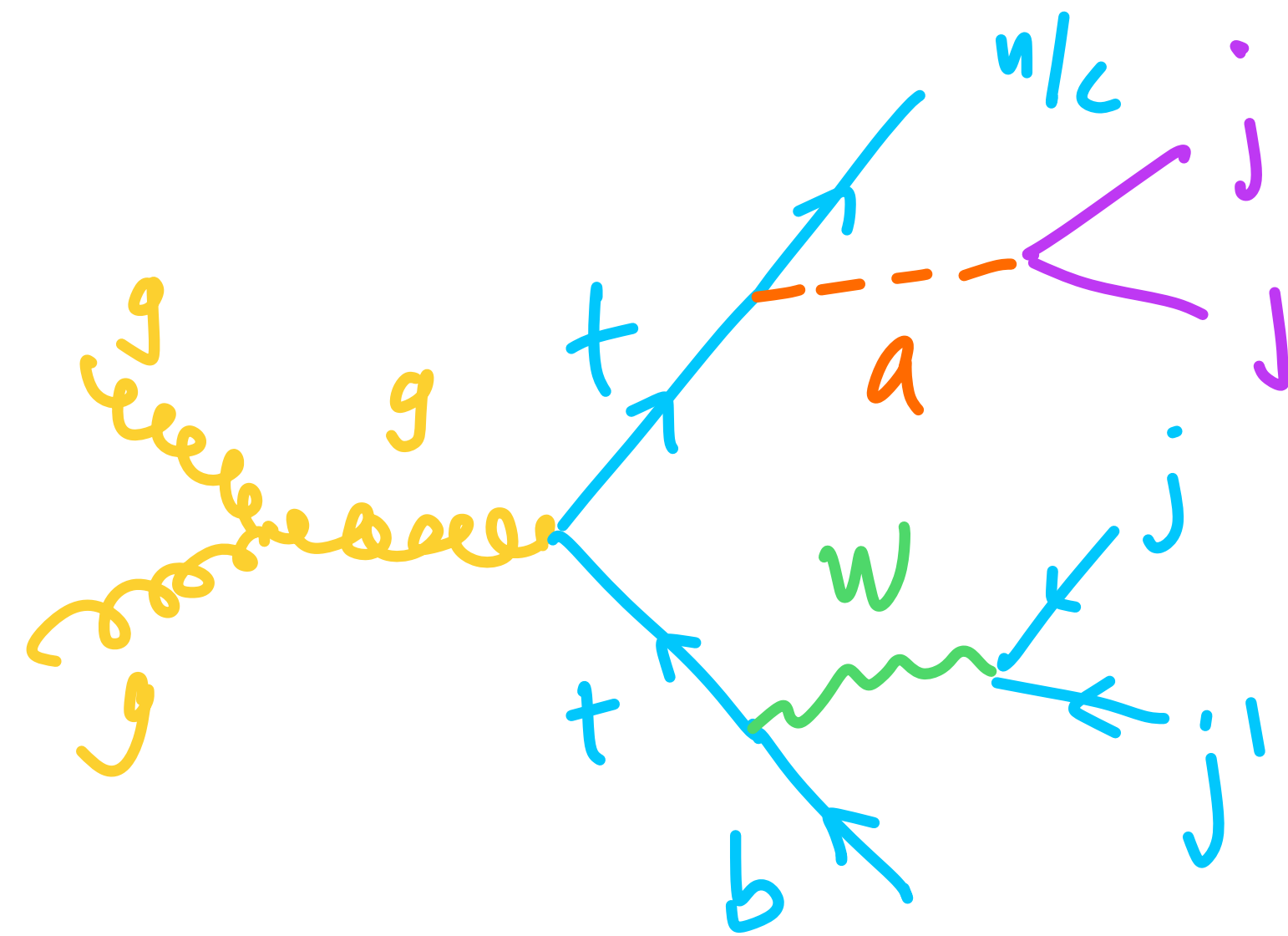
SIGNAL : HADRONIC CALORIMETER



- ✦ Large Ehad/Ecal ratio
- ✦ No tracks in the displaced jet
- ✦ 3-5(6) jets with 1(2) displaced and another b-tagged

	$m_a = 2 \text{ GeV}$	$m_a = 10 \text{ GeV}$	$t\bar{t}$
total	(1) 2.79×10^5	(1) 2.79×10^5	(1) 2.91×10^8
3 – 6 jets with $p_T > 40 \text{ GeV}$ & $ \eta < 2.5$	(0.8439) 2.35×10^5	(0.8414) 2.35×10^5	(0.71801) 2.09×10^8
1 jet with $\log_{10} \left(\frac{E_{\text{had}}}{E_{\text{em}}} \right) > 1.2$	(0.1436) 4.00×10^4	(0.0775) 2.16×10^4	(0.01244) 3.61×10^6
displaced jet has ≤ 2 tracks with $p_T > 2 \text{ GeV}$	(0.1436) 4.00×10^4	(0.0775) 2.16×10^4	(0.00022) 6.39×10^4

SIGNAL: MUON SPECTROMETER



- ✦ Event in the muon system
- ✦ No associated track pointing to the primary vertex
- ✦ 2-4(5) jets
- ✦ Signal is background free

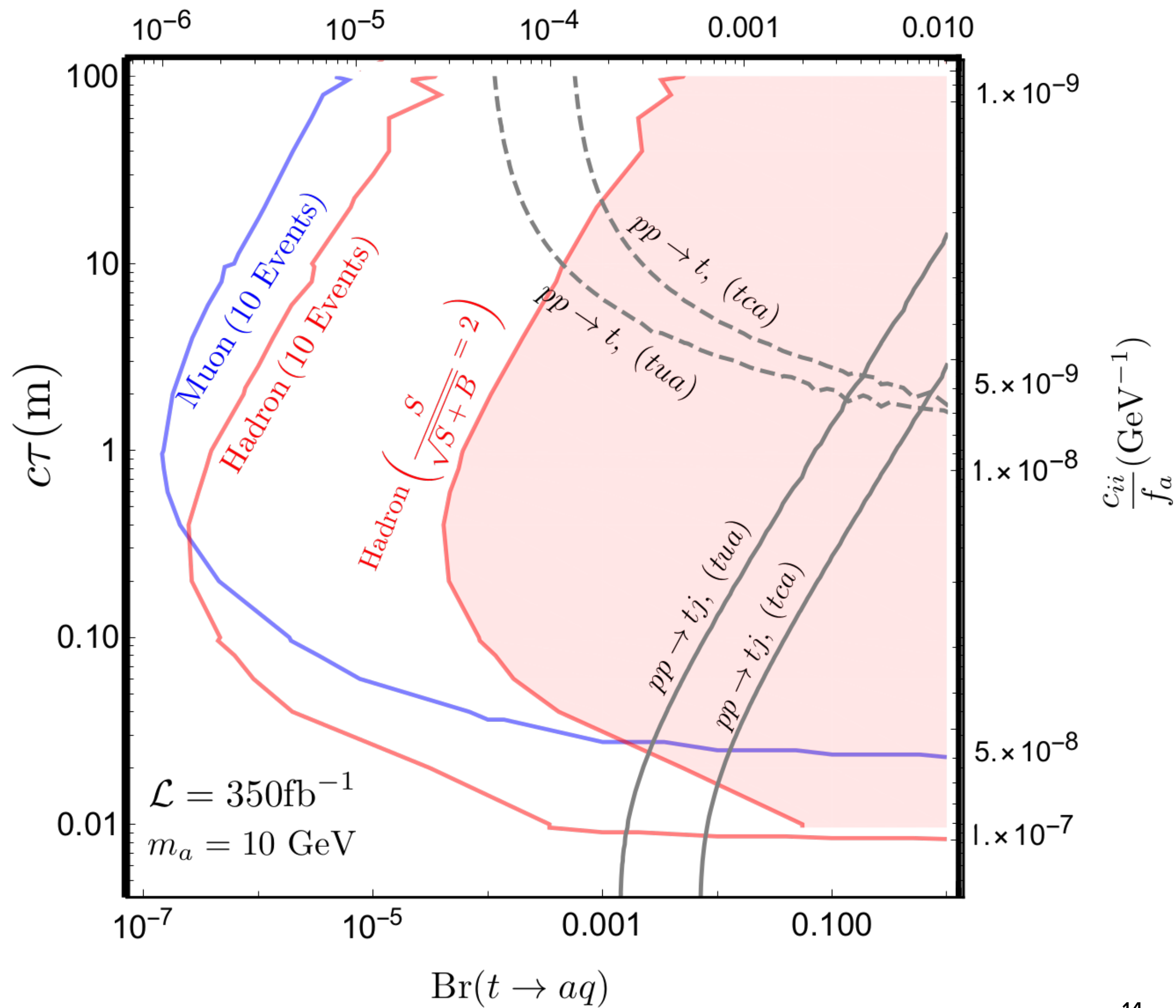
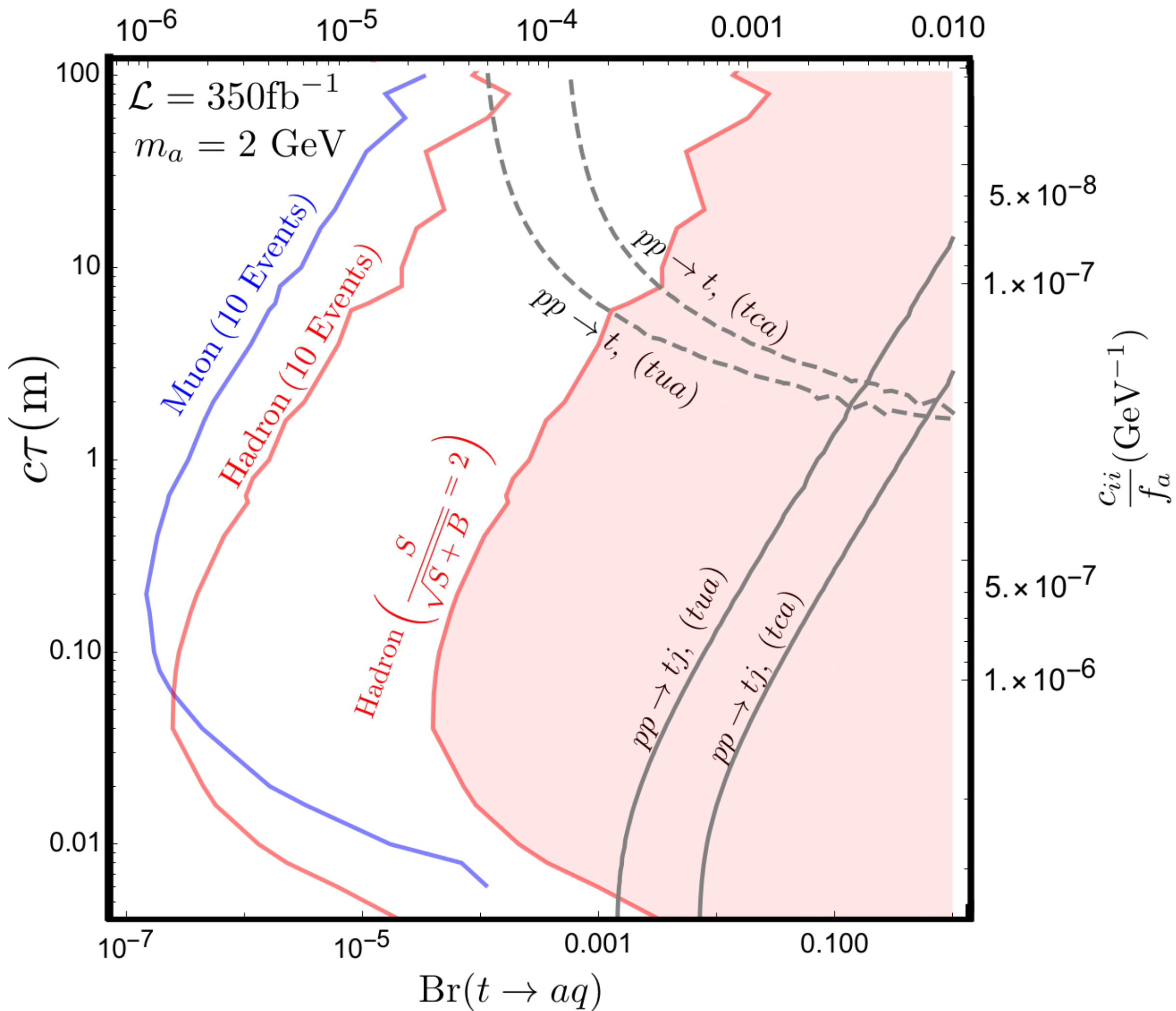
RESULTS

$$\frac{c_{tq}}{f_a} (\text{GeV}^{-1})$$

$m_a = 2 \text{ GeV}$

$m_a = 10 \text{ GeV}$

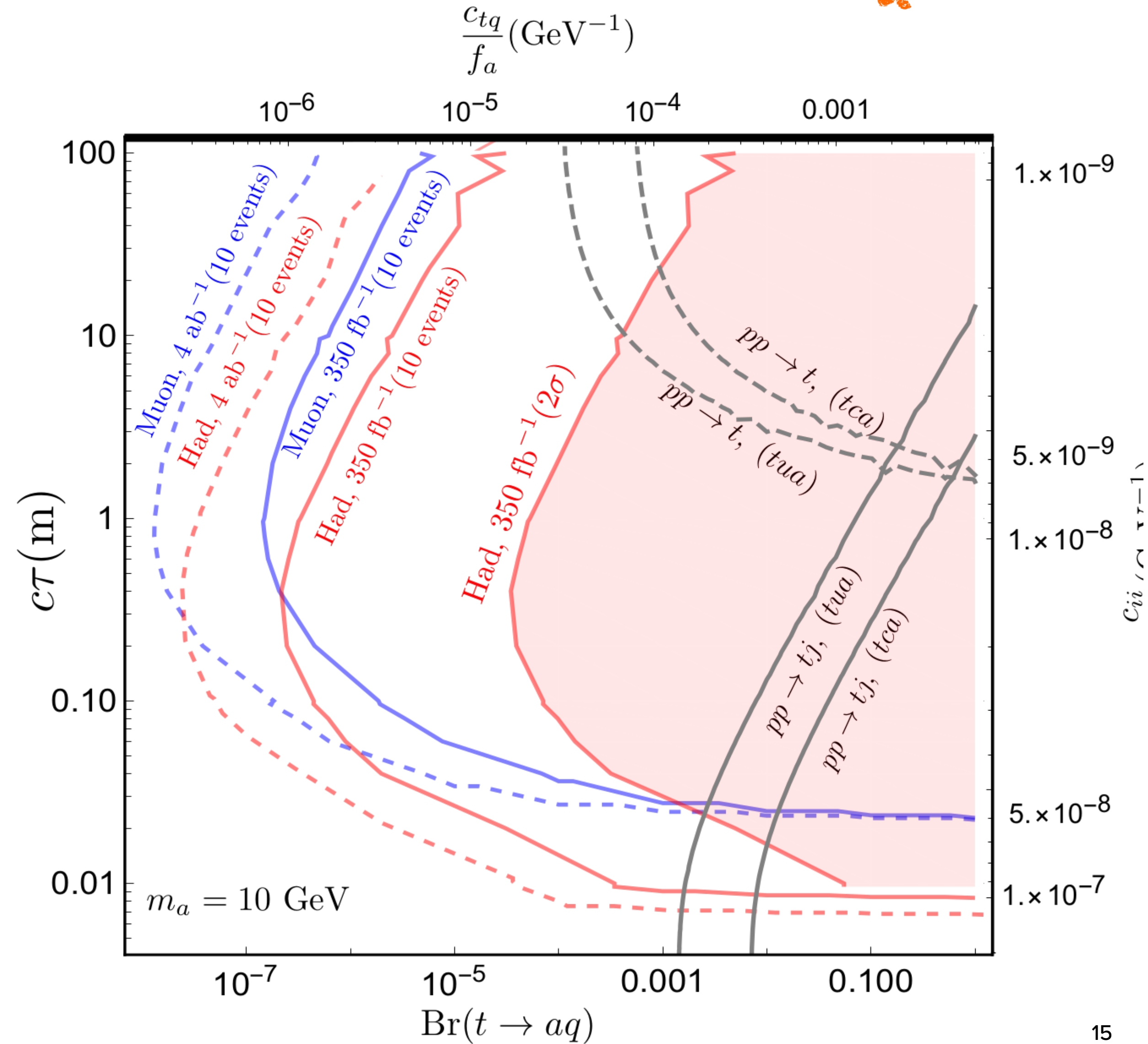
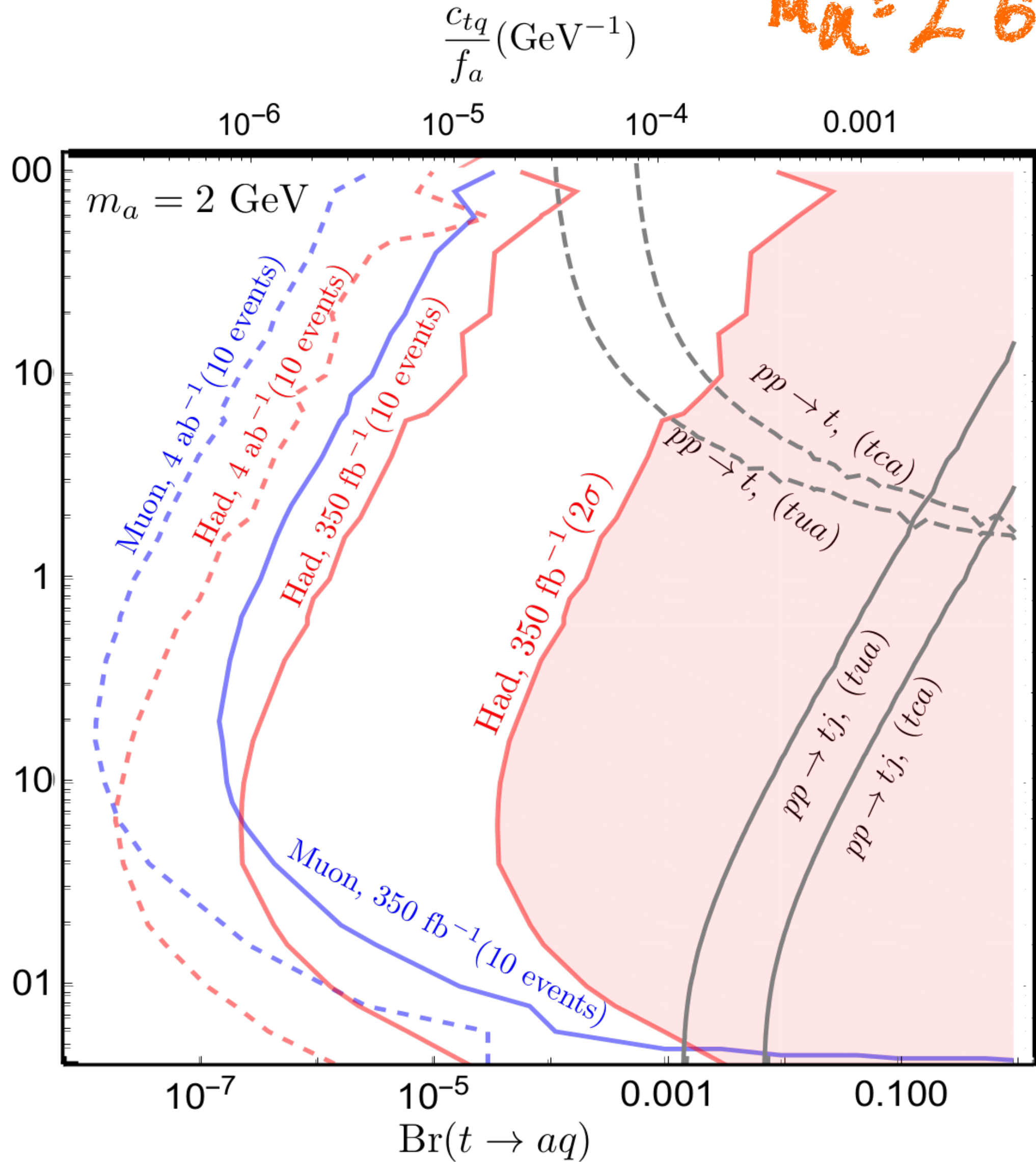
$$\frac{c_{tq}}{f_a} (\text{GeV}^{-1})$$



RESULTS

$m_a = 2 \text{ GeV}$

$m_a = 10 \text{ GeV}$



CONCLUSIONS

CONCLUSIONS

- ALPs are ubiquitous in beyond the SM physics
- They can be probed by very different and complementary experiments
- Exotic top decays provide a unique way of probing ALPs above the charm threshold
- We can probe $\text{Br}(t \rightarrow aq) \lesssim 10^{-4}$ and there is room for improvement!



15 - 17 JUNE 2022

HEFT 2022

[HTTPS://FTA.E.UGR.ES/HEFT2022](https://ftae.ugr.es/heft2022)

Register!!

Wed 20 Apr

FTA.E
High Energy Theory

Speakers

Committee

Program

Participants

Registration

Code of Conduct

Previous editions



HIGGS AND EFFECTIVE FIELD THEORY - HEFT 2022

📅 15-17 June 2022

📍 Salon de Actos del Carmen de la Victoria, Cuesta del Chapiz, 9 (Granada)

HEFT is an annual workshop focusing on the use of effective field theories to search for physics beyond the Standard Model. A broad range of topics are encouraged, ranging from collider phenomenology and formal aspects to the latest experimental updates on dedicated searches. The meeting aims to foster discussions between theorists and phenomenologists from varied backgrounds as well as with experimental colleagues.

We would like to achieve a balance of senior and junior speakers, enhancing the visibility of younger scientists while keeping some overview talks.

This year's edition is organized by (and will be held in) the Dpto. Física Teórica y del Cosmos of the Universidad de Granada. The format of the event will be in-person only.

The workshop will begin on June 15th morning and end after lunch on June 17th.

This project has received funding from the European Union's Horizon 2020 research and innovation programme under the Marie Skłodowska-Curie grant agreement and UGR Research and Knowledge Transfer Found - [Athenea3i](#)

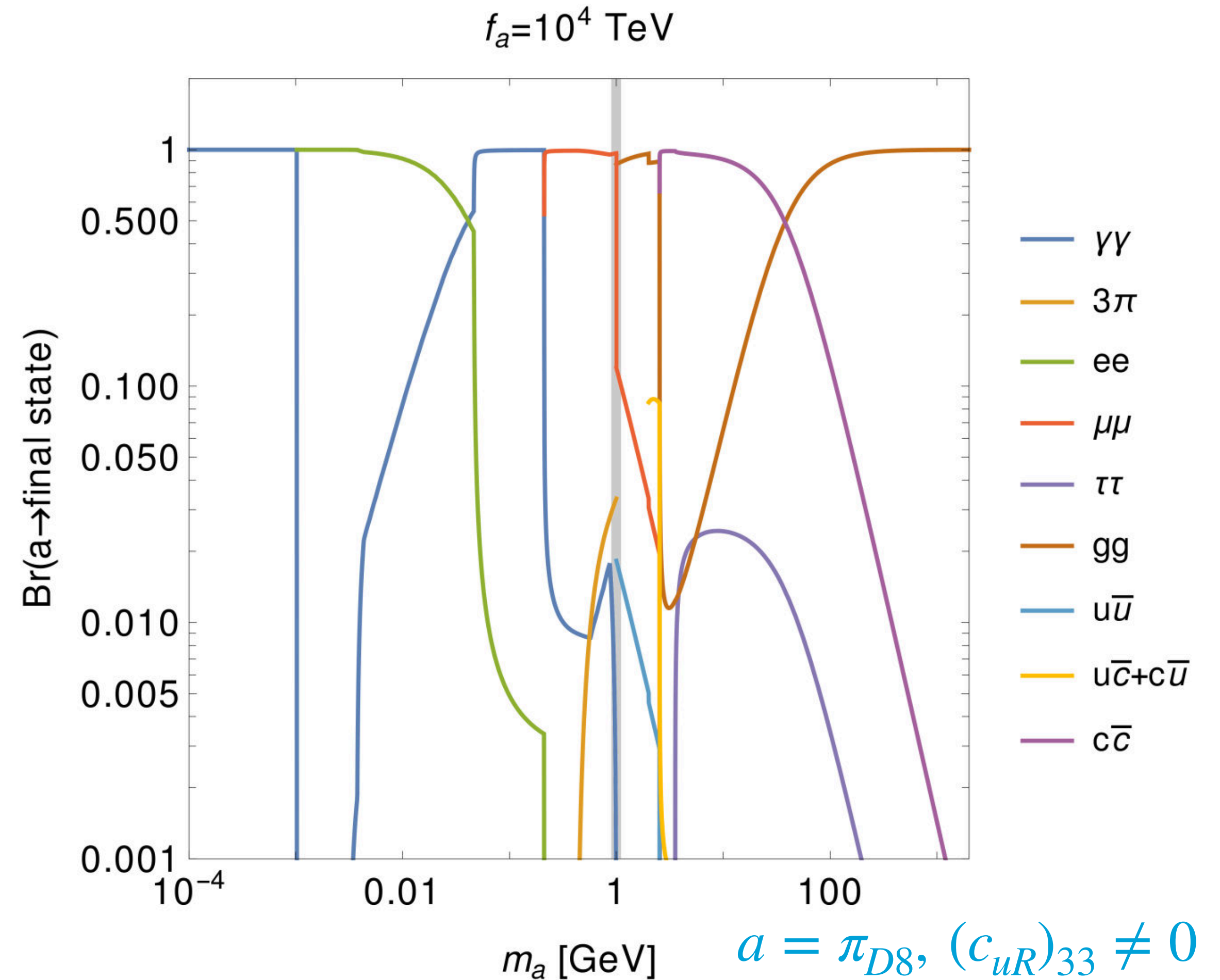
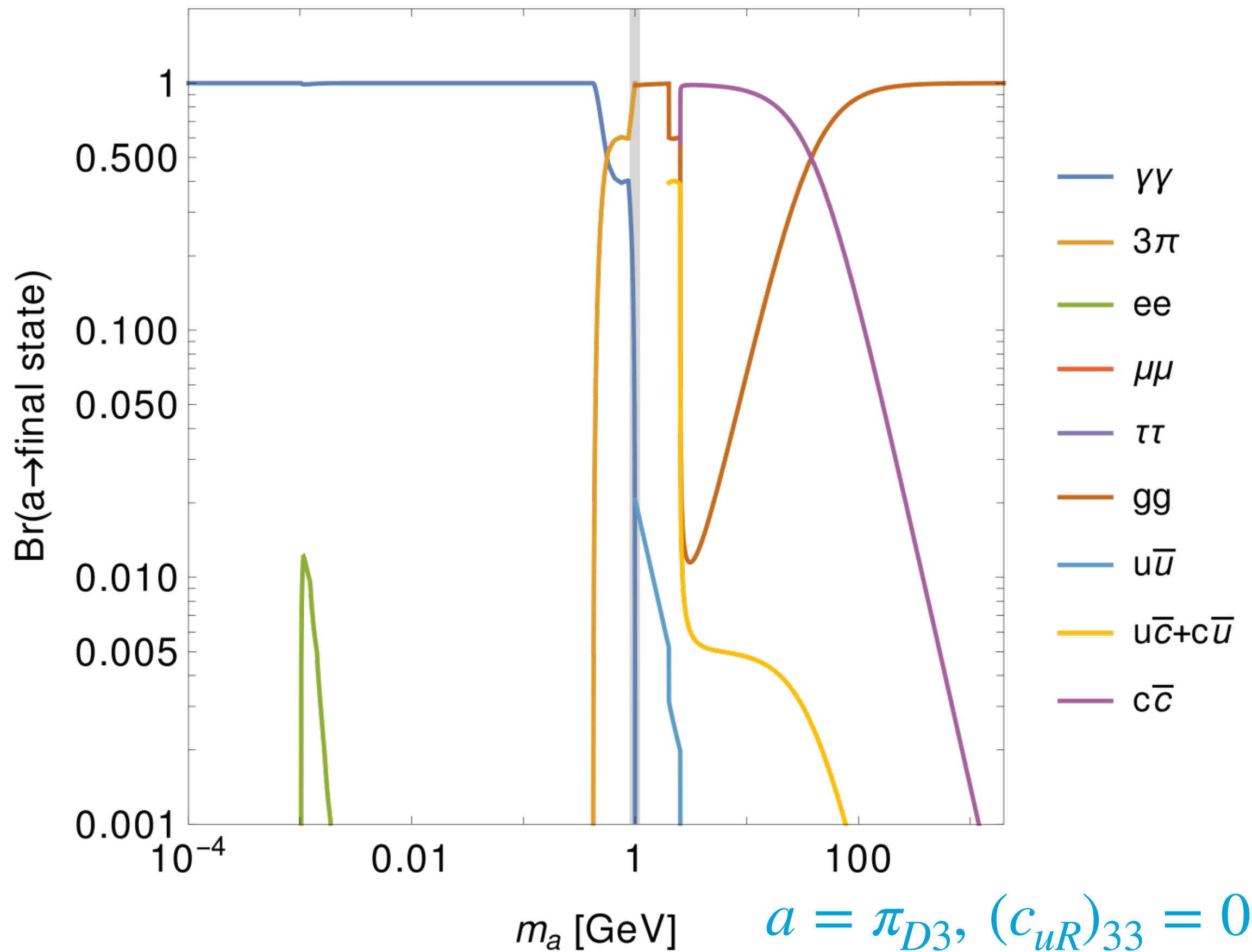
BACK UP

CHARMING ALPS

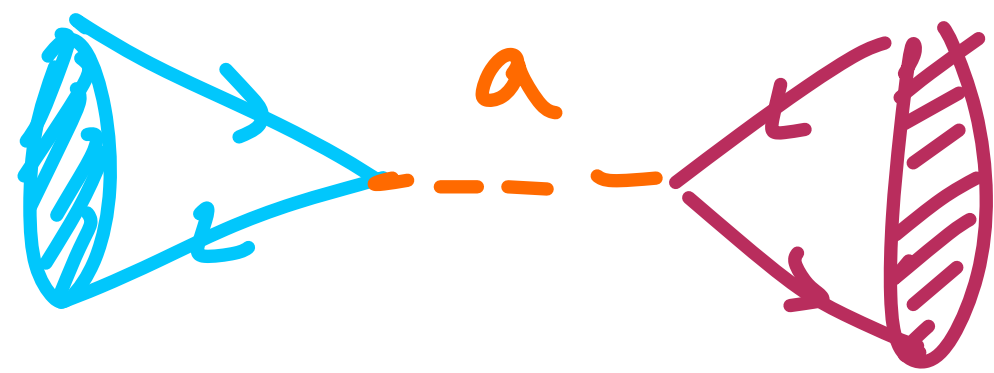
RGEs and heavy-quark loop induced processes will lead to

$$a \rightarrow \ell^+ \ell^-, a \rightarrow \bar{d}_i d_j, a \rightarrow \gamma\gamma, a \rightarrow gg$$

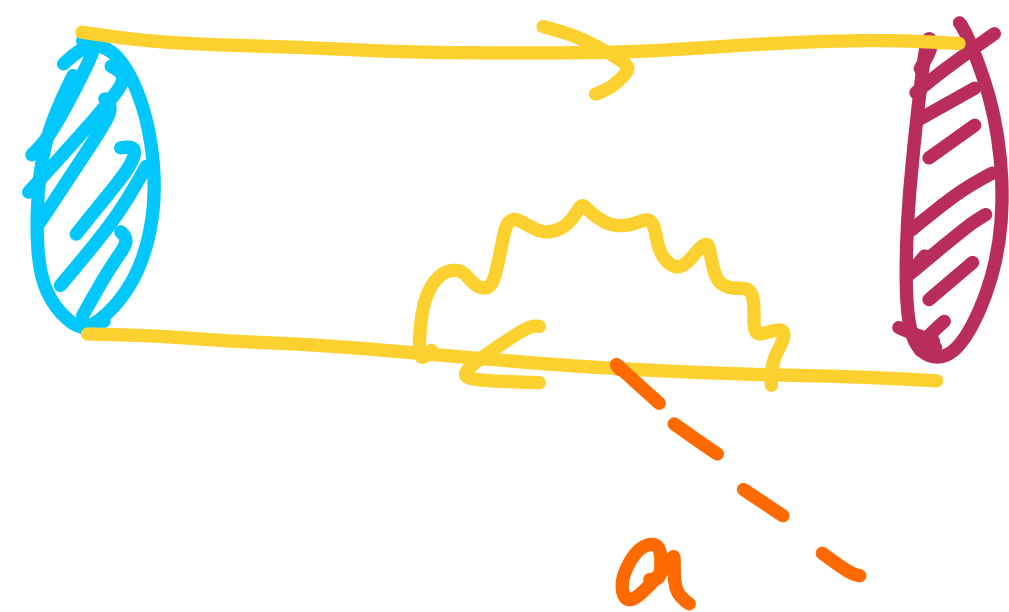
$f_a = 10^4$ TeV



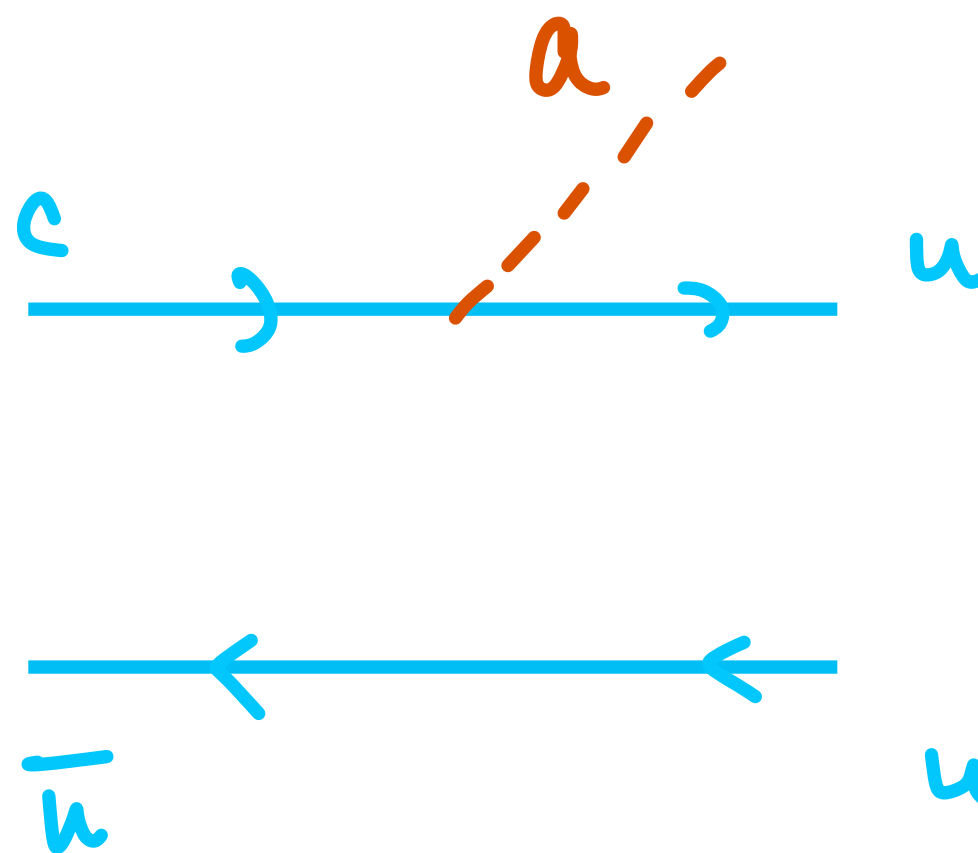
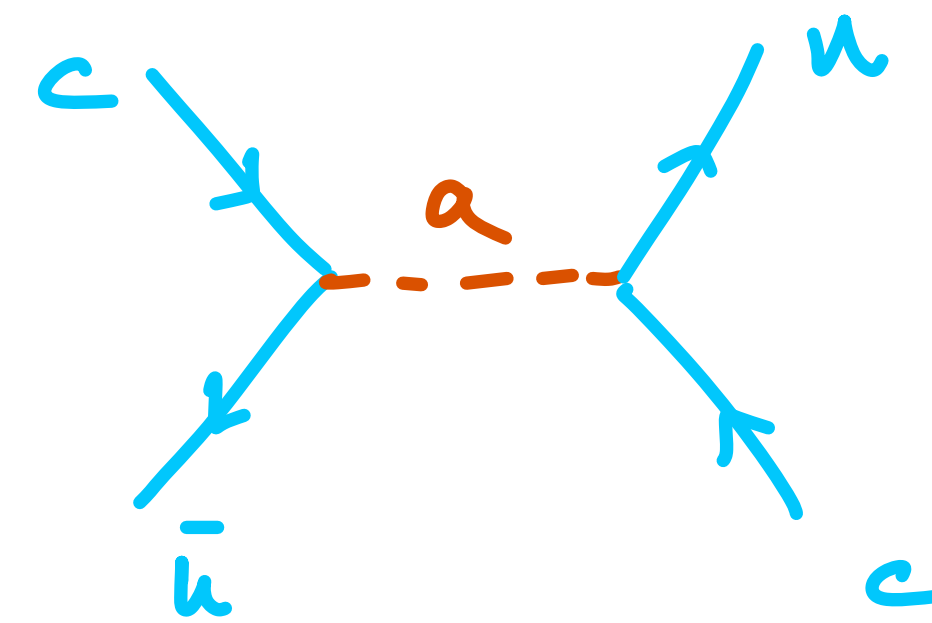
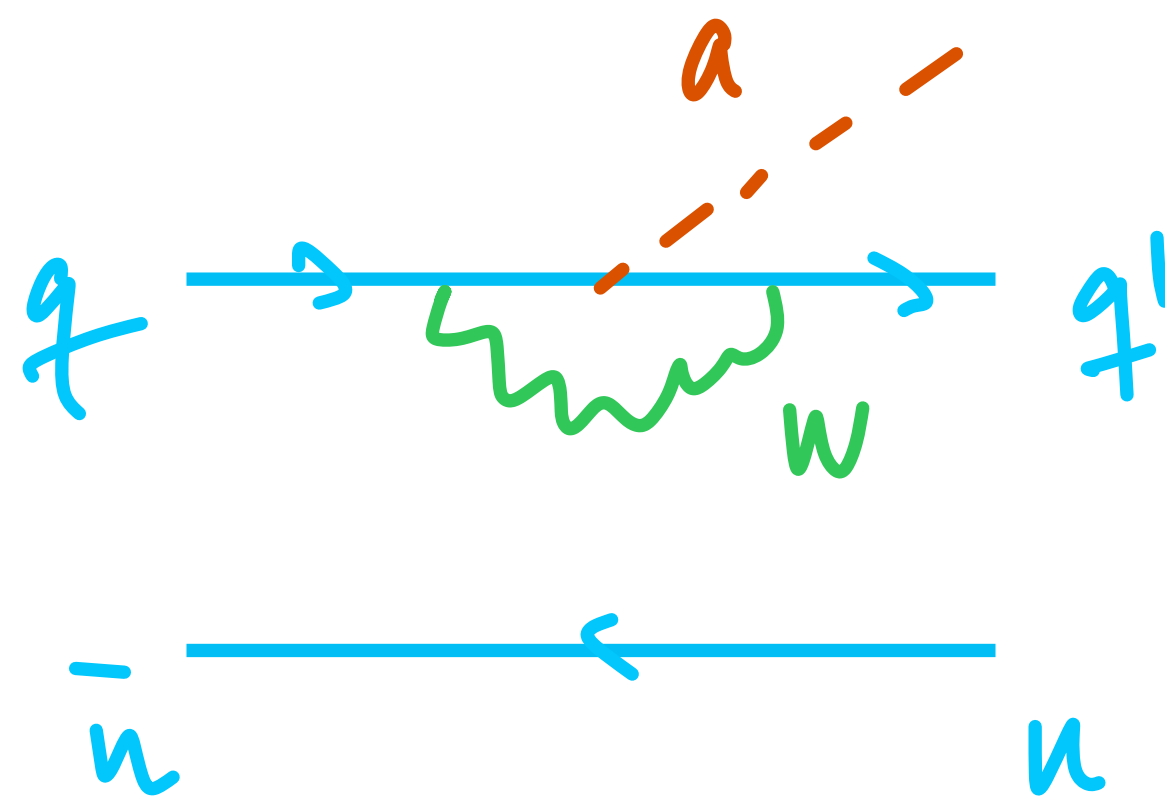
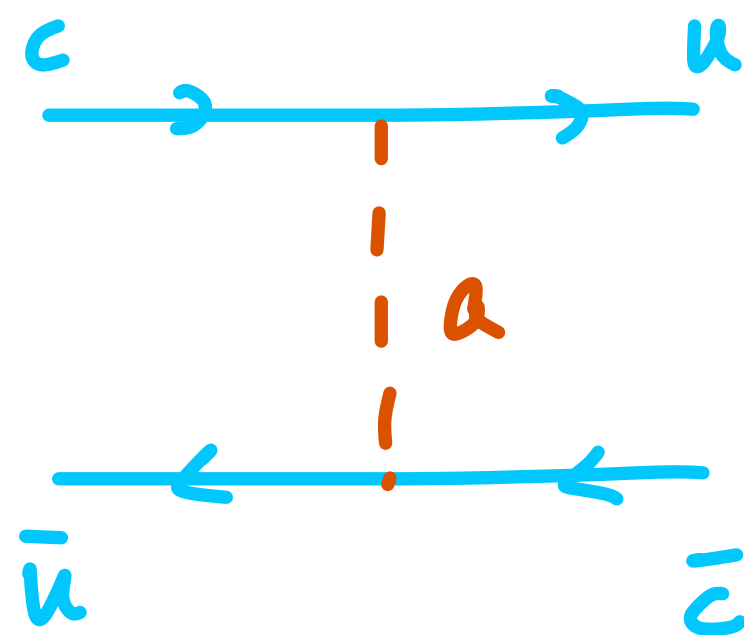
FLAVOR BOUNDS



$D - \bar{D}$ mixing



$B \rightarrow Ka, B \rightarrow \pi a, K \rightarrow \pi a$



$D \rightarrow \pi a$

FLAVOR BOUNDS

- $D^+ \rightarrow (\tau^+ \rightarrow \pi^+ \nu) \bar{\nu}$ recasted with M_{miss}^2 for $D^+ \rightarrow \pi^+ a$ **CLEO 0806.2112**
- $B^+ \rightarrow K^+ \bar{\nu} \nu, B^0 \rightarrow K^0 \bar{\nu} \nu$ recasted with $s_B = k^2/m_B^2$ for $B \rightarrow Ka$ **BaBar 1303.7465**
- $B^+ \rightarrow \pi^+ \bar{\nu} \nu$ recasted with $\sqrt{\vec{p}_\pi^2}$ for $B \rightarrow \pi a$ **BaBar hep-ex/0411061**
- $K^+ \rightarrow \pi^+ a$ for $m_a > 0$ **NA62 2011.11329**
- $B^\pm \rightarrow K^\pm \bar{\nu} \nu$ expected at Belle II with 50 ab^{-1} and $K^\pm \rightarrow \pi^\pm \bar{\nu} \nu$ at **NA62**
- Recasts done with the CLs method

ASTRO & COSMO BOUNDS

Red Giant bursts

$$\mathcal{L} \supset i a g_{alt}(\bar{\ell} \gamma_5 \ell), \quad g_{alt} = \frac{3m_e}{8\pi v^2 f_a} \ln \left(\frac{f_a^2}{m_t^2} \right) \sum_{i=1}^3 (\mathcal{M}_u)_{ii} (c_{u_R})_{ii}$$

$g_{aee} < 1.6 \cdot 10^{-13}$

SN1987a

Bremsstrahlung

$$L_a \leq L_\nu = 3 \cdot 10^{52} \text{ erg/s}$$

$$N + N \rightarrow N + N + a$$

$$c_{app} = (c_{u_R})_{11} (0.75 \pm 0.03)$$

$$c_{ann} = (c_{u_R})_{11} (-0.51 \pm 0.03)$$

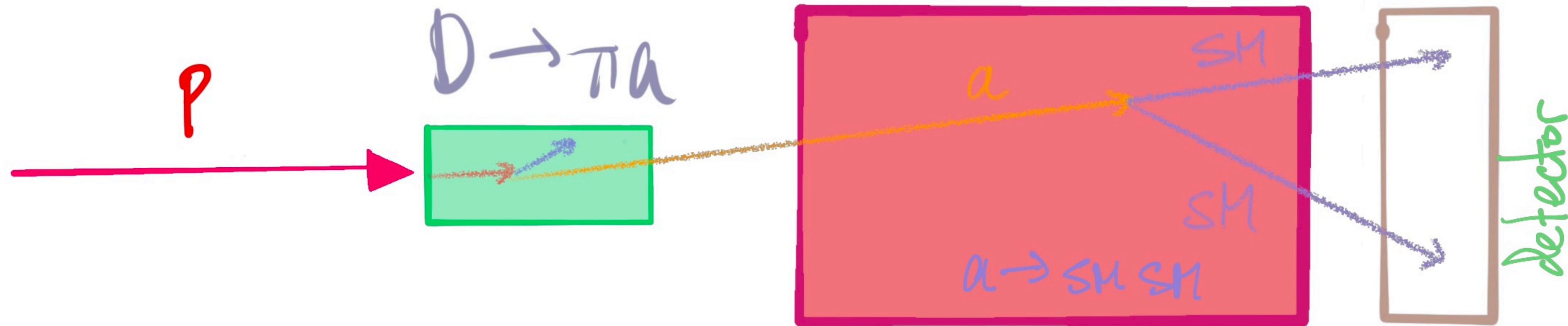
N_{eff} , distortion of CMB, BBN, ...

Cadamuro, Redondo '12
Millea, Knox '15
Depta et al '20

Most of the bounds derived assumed only couplings to photons but they can still be recasted

COLLIDER AND FIXED TARGET EXPERIMENTS

Fixed target experiments: NA62, SHiP, CHARM



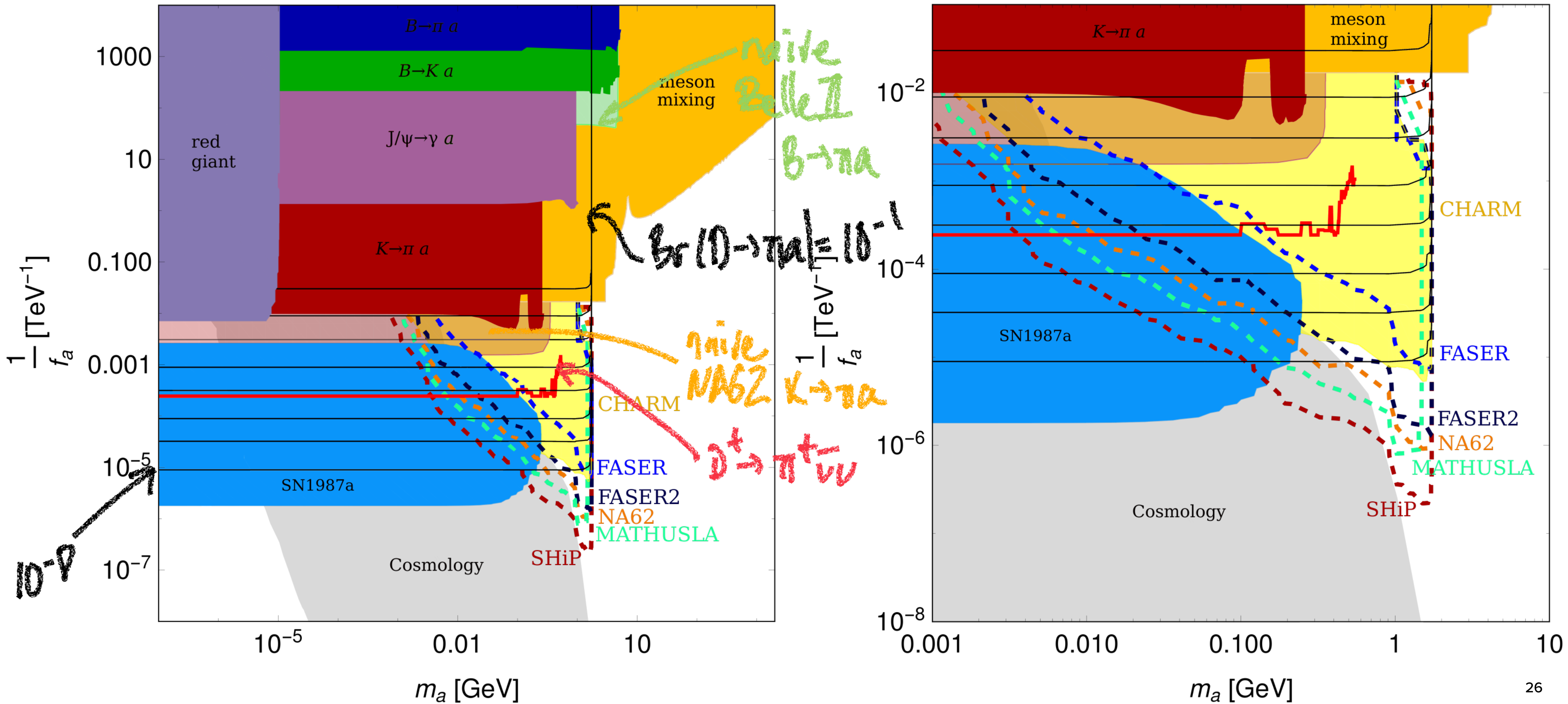
$$N_a = N_D \cdot \text{Br}(D \rightarrow \pi a) \cdot \epsilon_{\text{geom}} \cdot F_{\text{decay}}$$

decay volume

LHC forward detectors: FASER, FASER II, MATUSHLA

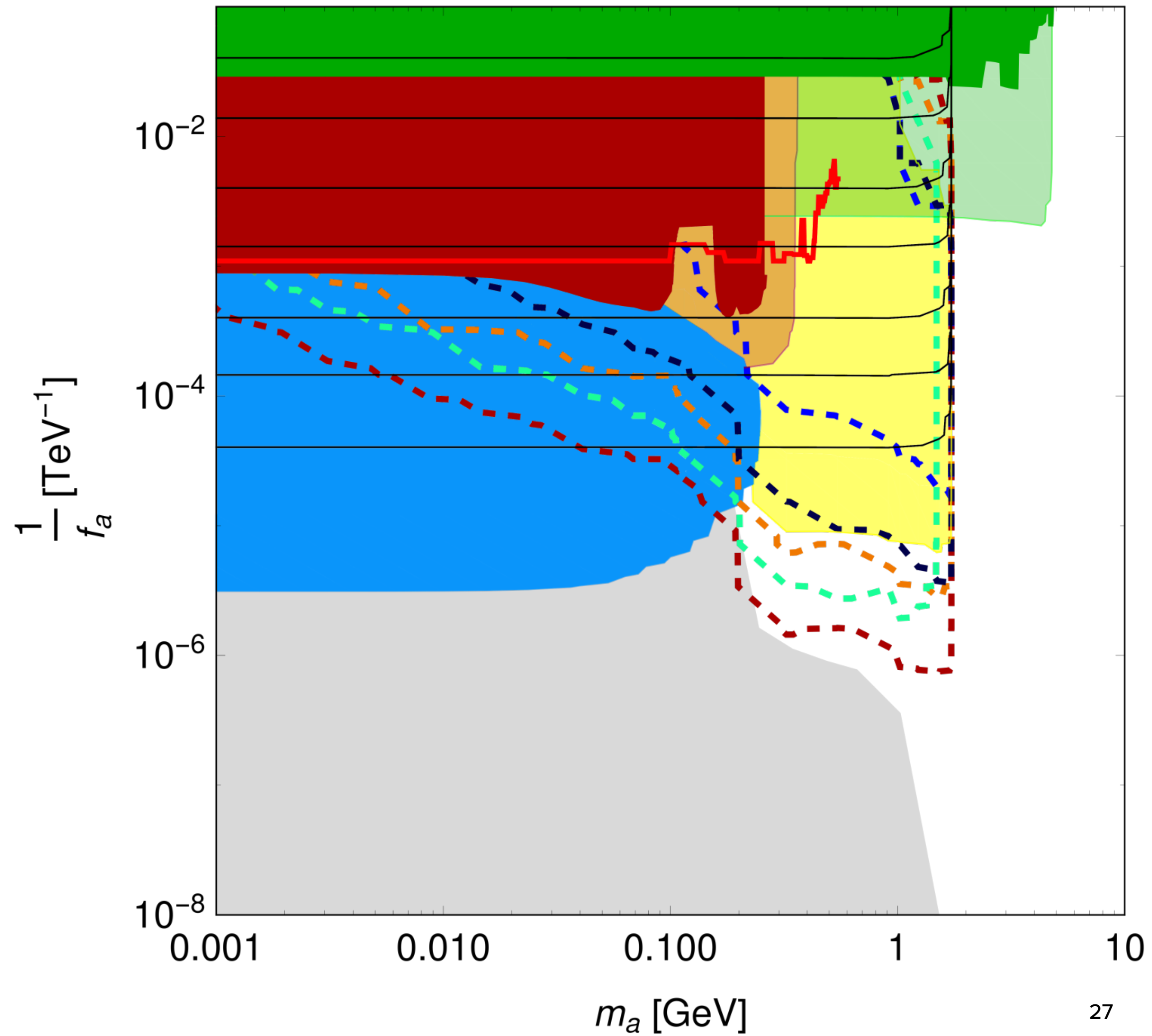
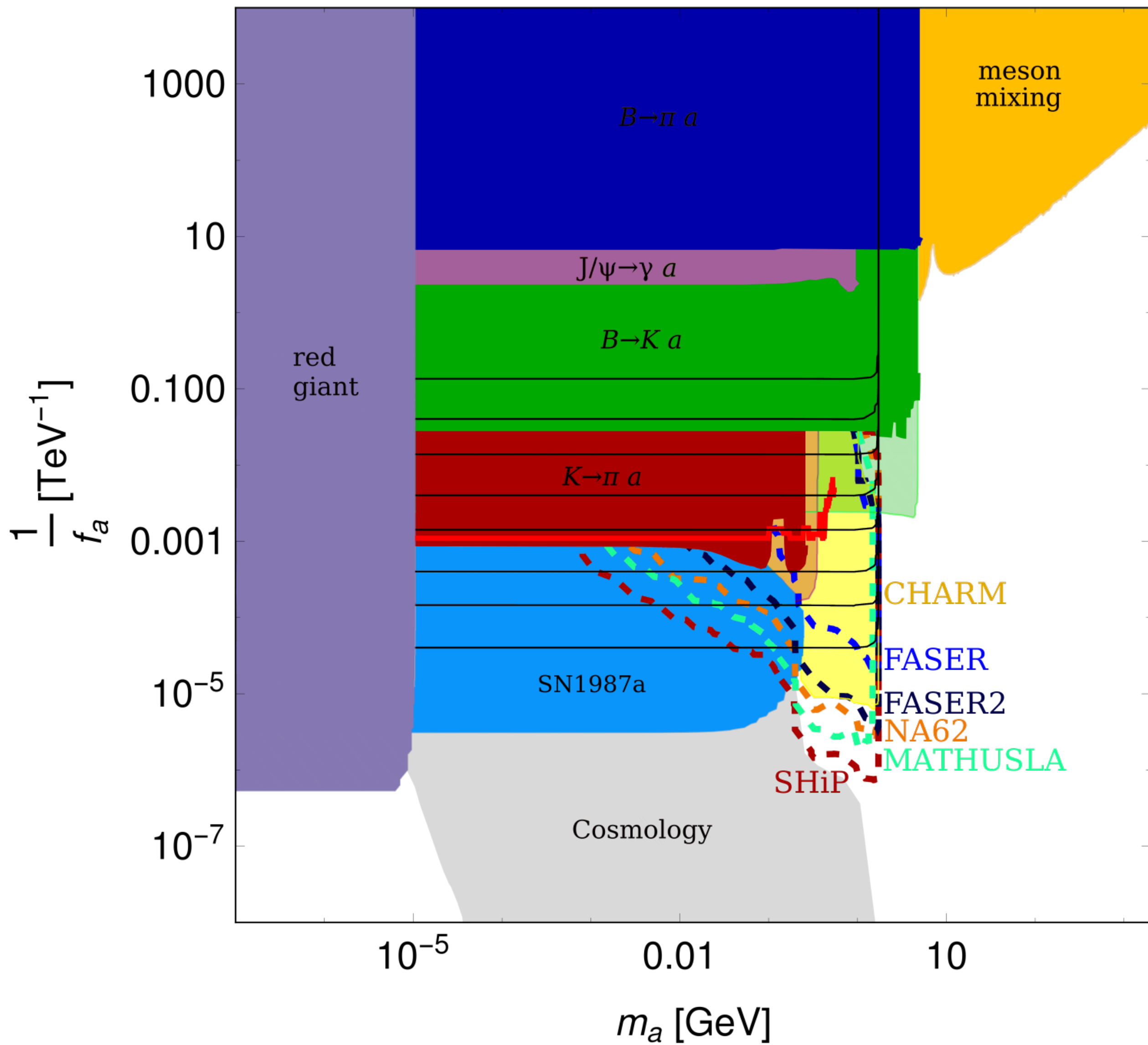
CHARMING ALPS PARAMETER SPACE

$$a = \pi_{D_3}, (c_{u_R})_{33} = 0$$



CHARMING ALPS PARAMETER SPACE

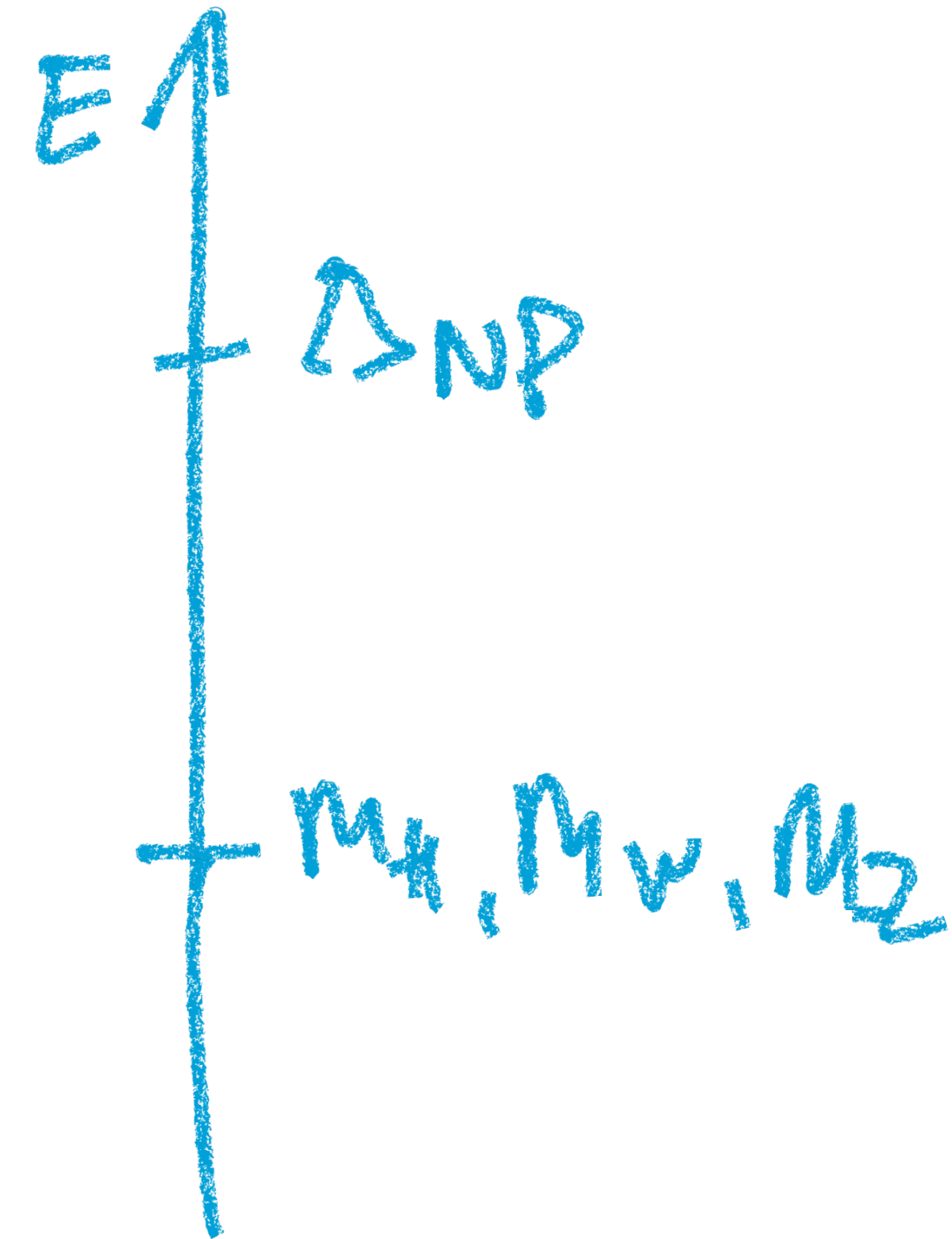
$$a = \pi_{D_8}, (c_{u_R})_{33} \neq 0$$



NATURALLY LIGHT SCALARS

The scale of new physics seems to be rather heavy

A natural way of obtaining light scalar degrees of freedom is through the Goldstone theorem

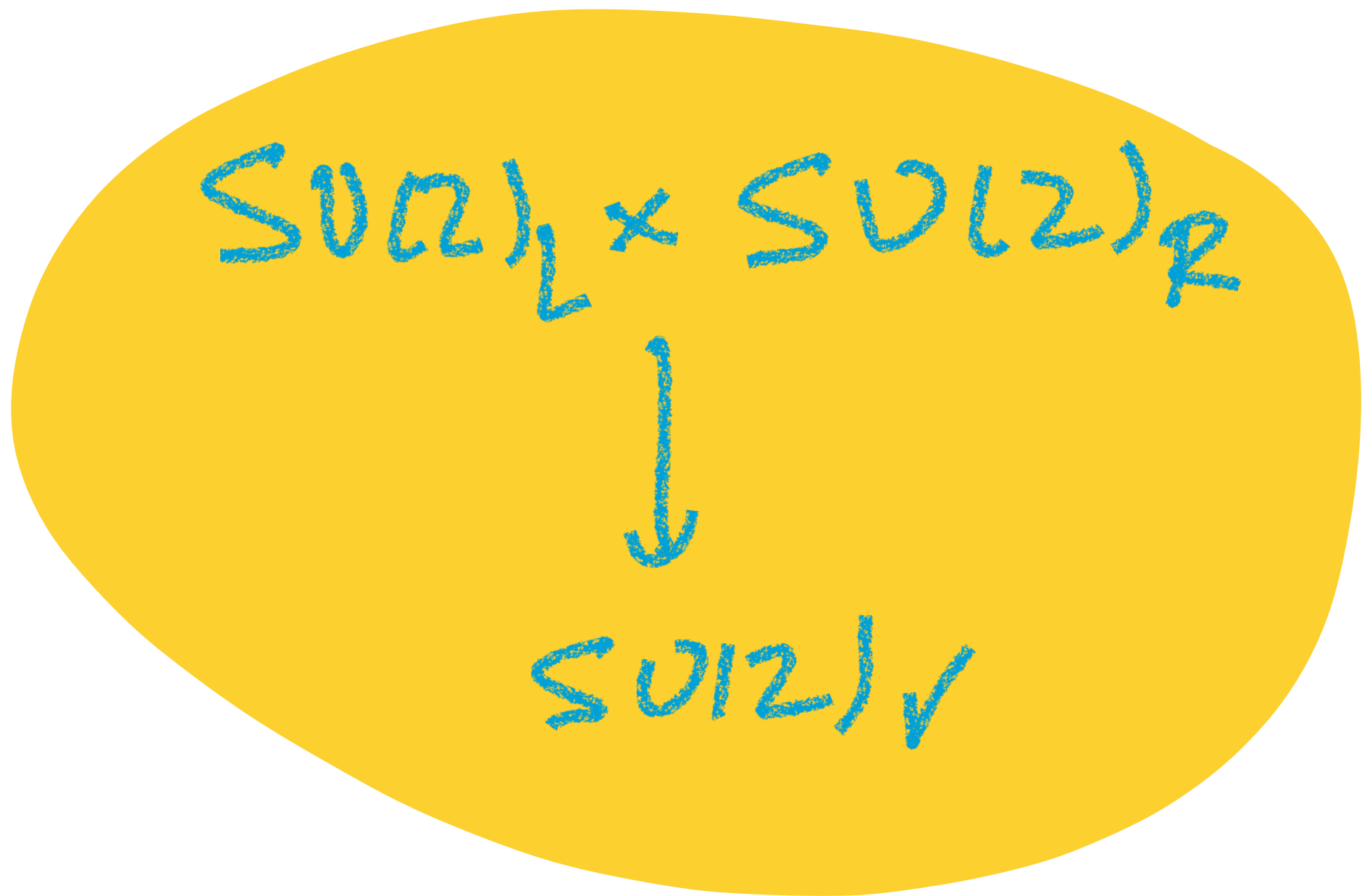


$$G \xrightarrow{\Delta_{NP}} H \quad \pi^i \in \text{Alg} \{G/H\}$$

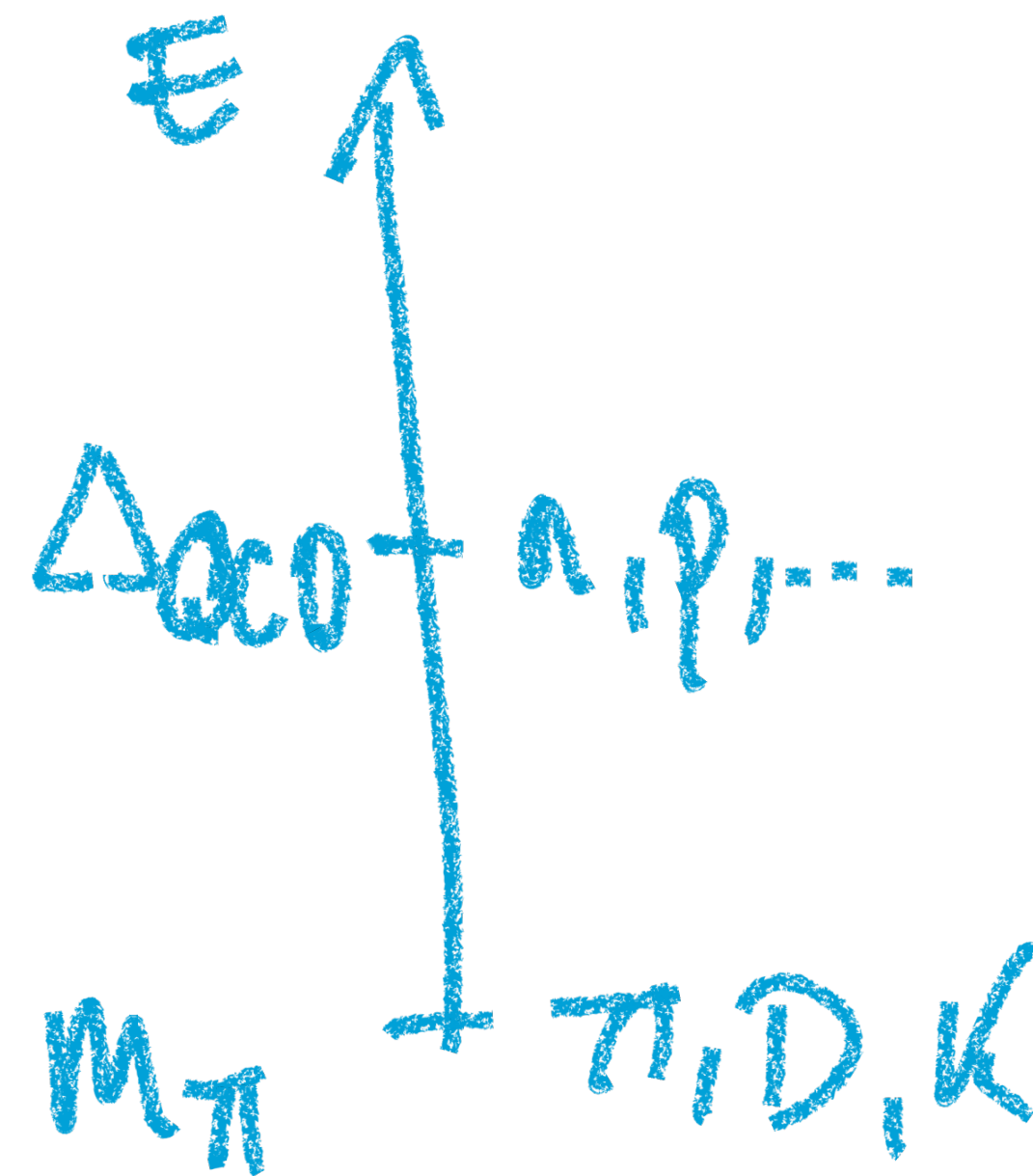
$$m_{\pi^i} \ll \Delta_{NP}$$

NATURALLY LIGHT SCALARS

QCD gives us a beautiful example



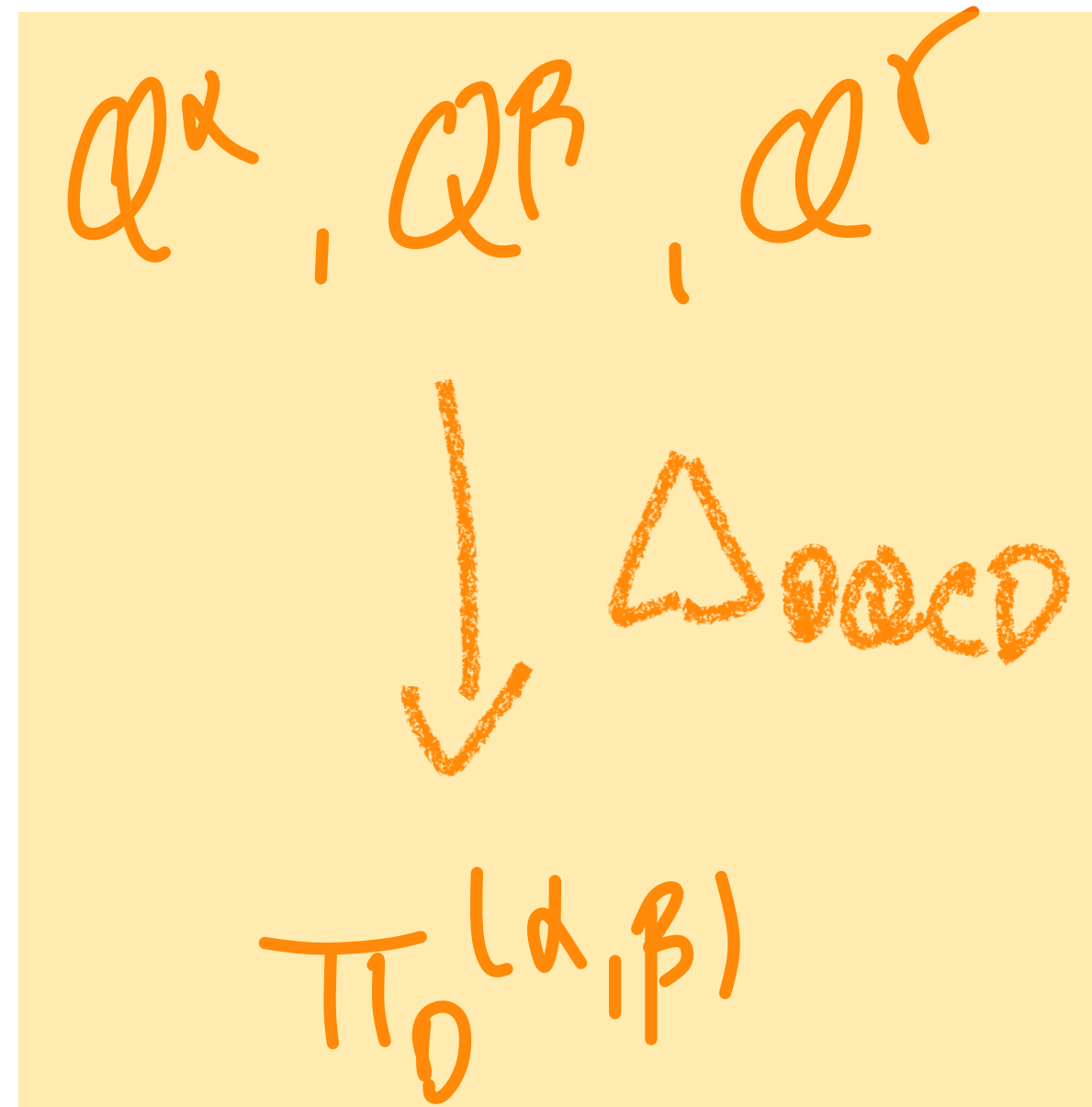
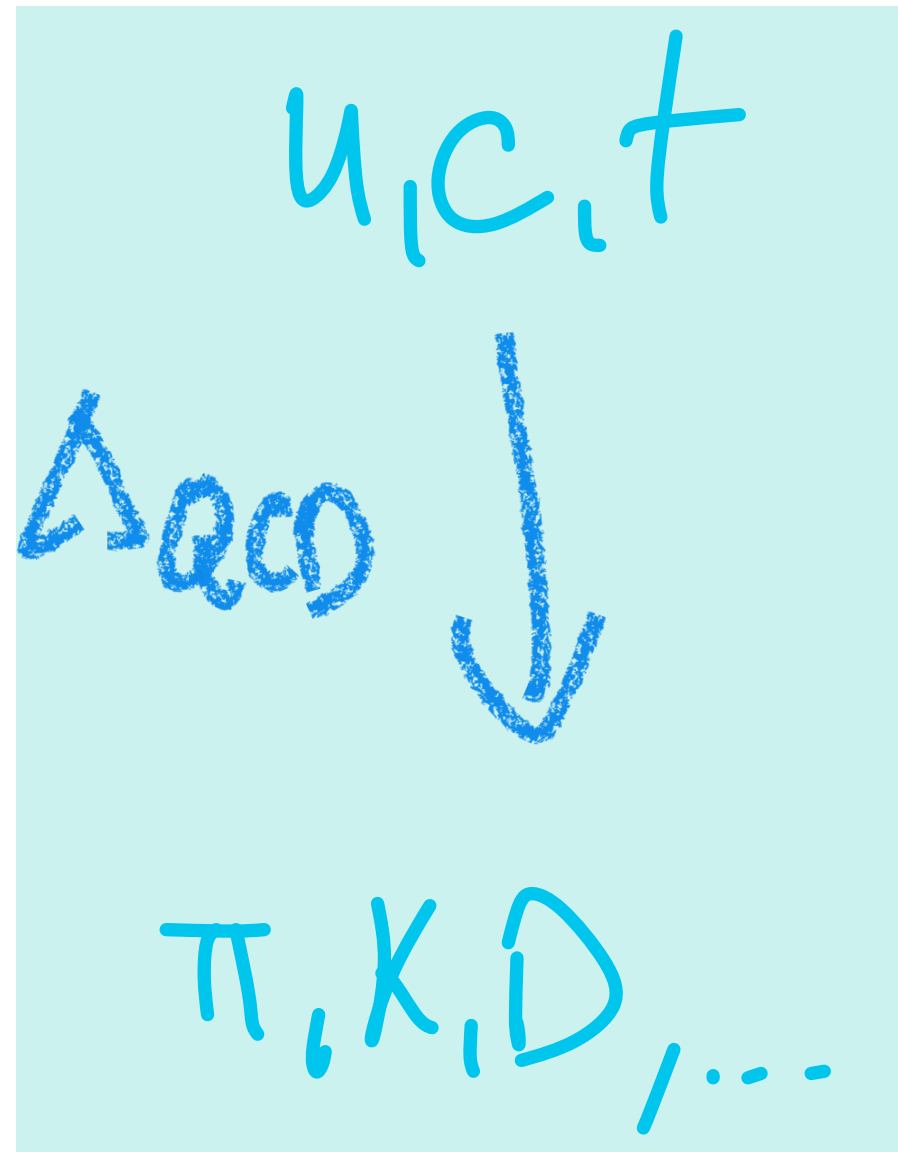
$$m_\pi \ll \Delta_{QCD}$$



A QCD-LIKE DARK SECTOR

A QCD-LIKE DARK SECTOR

Schwaller, Bai, '14



* $SU(N_D)$ gauge group

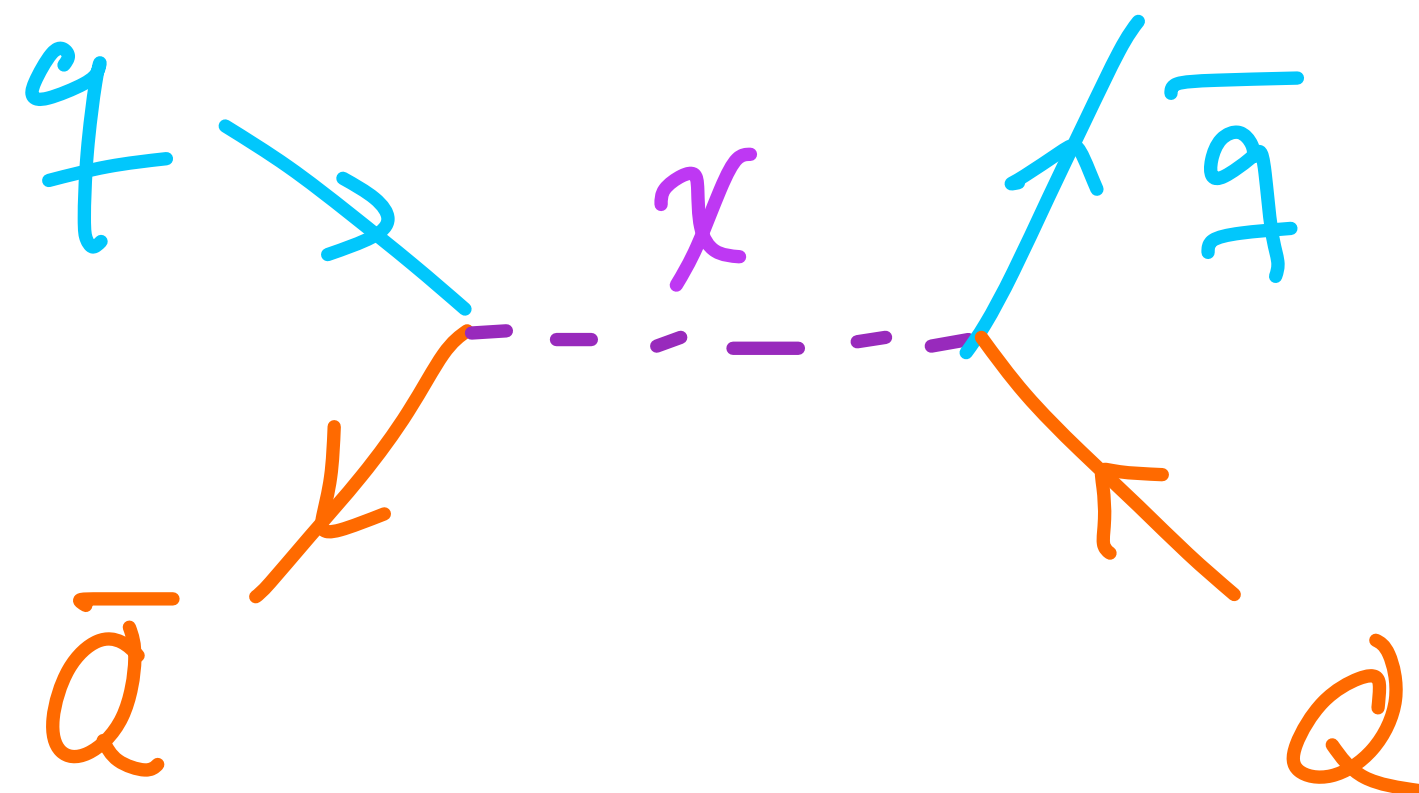
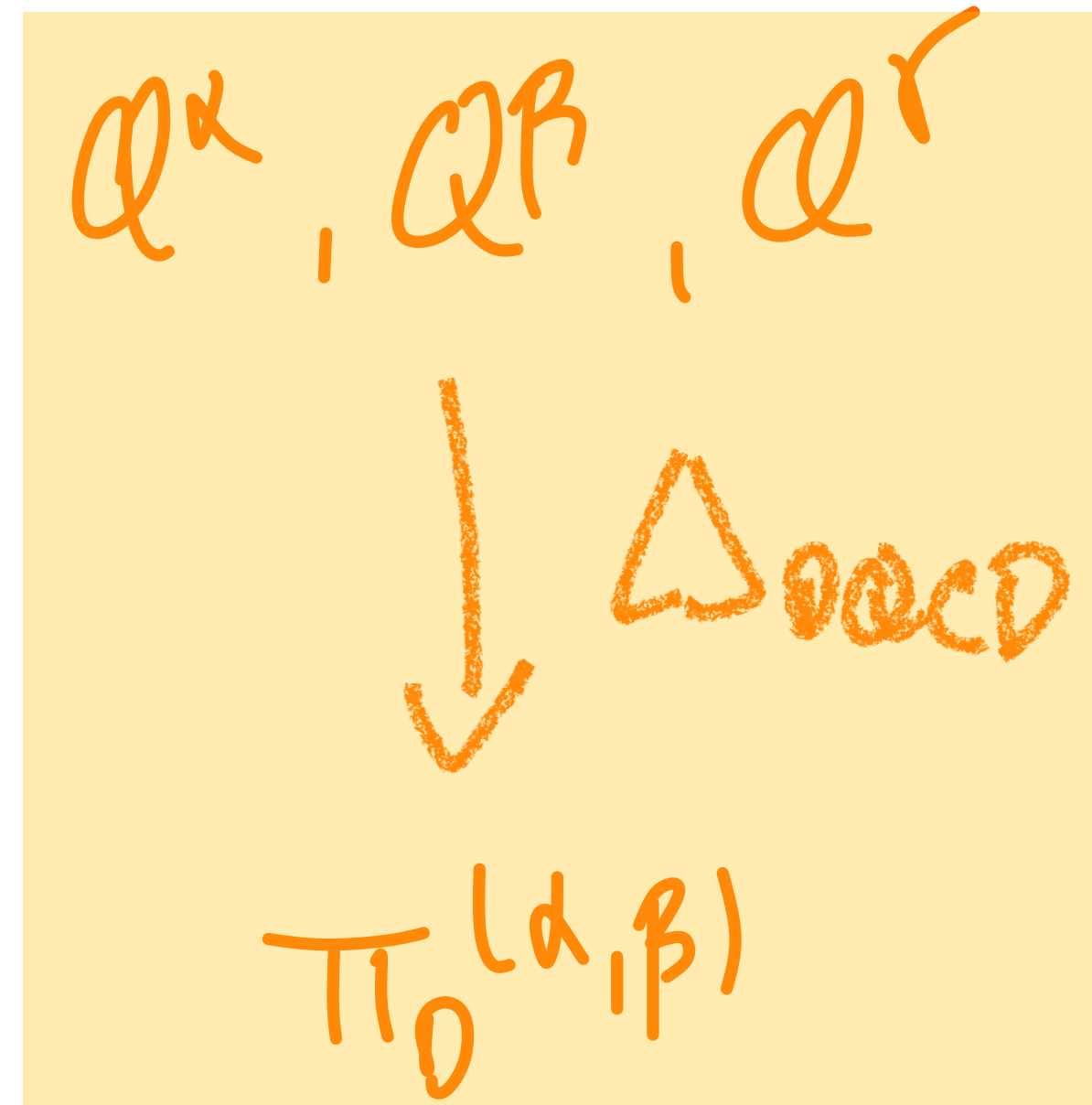
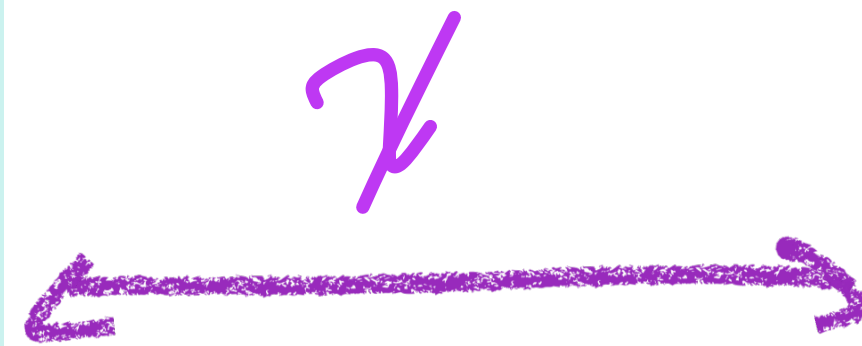
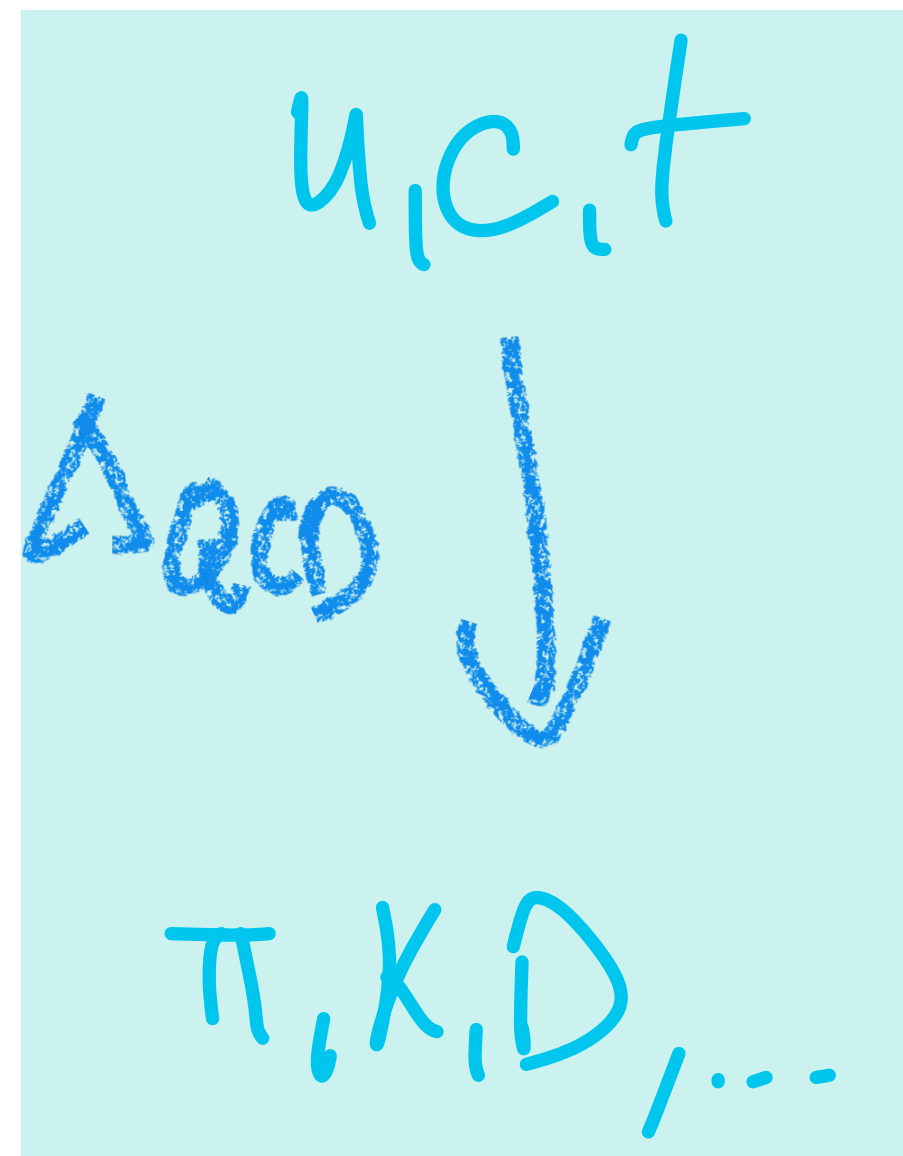
* N_f Dirac fermions

* $M_Q \ll \Delta_{DQCD}$

* $SU(N_f) \otimes SU(N_f)$
 \downarrow
 $SU(N_f)$

A QCD-LIKE DARK SECTOR

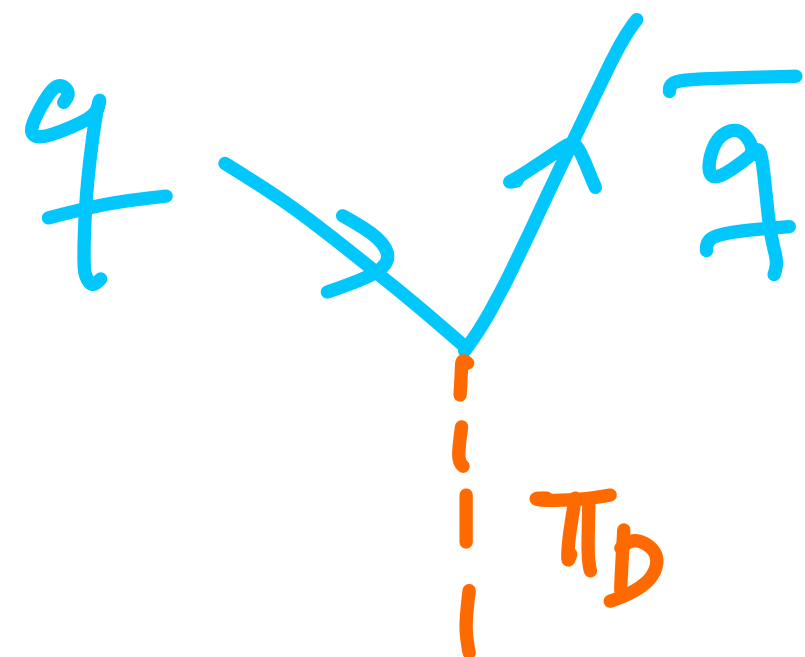
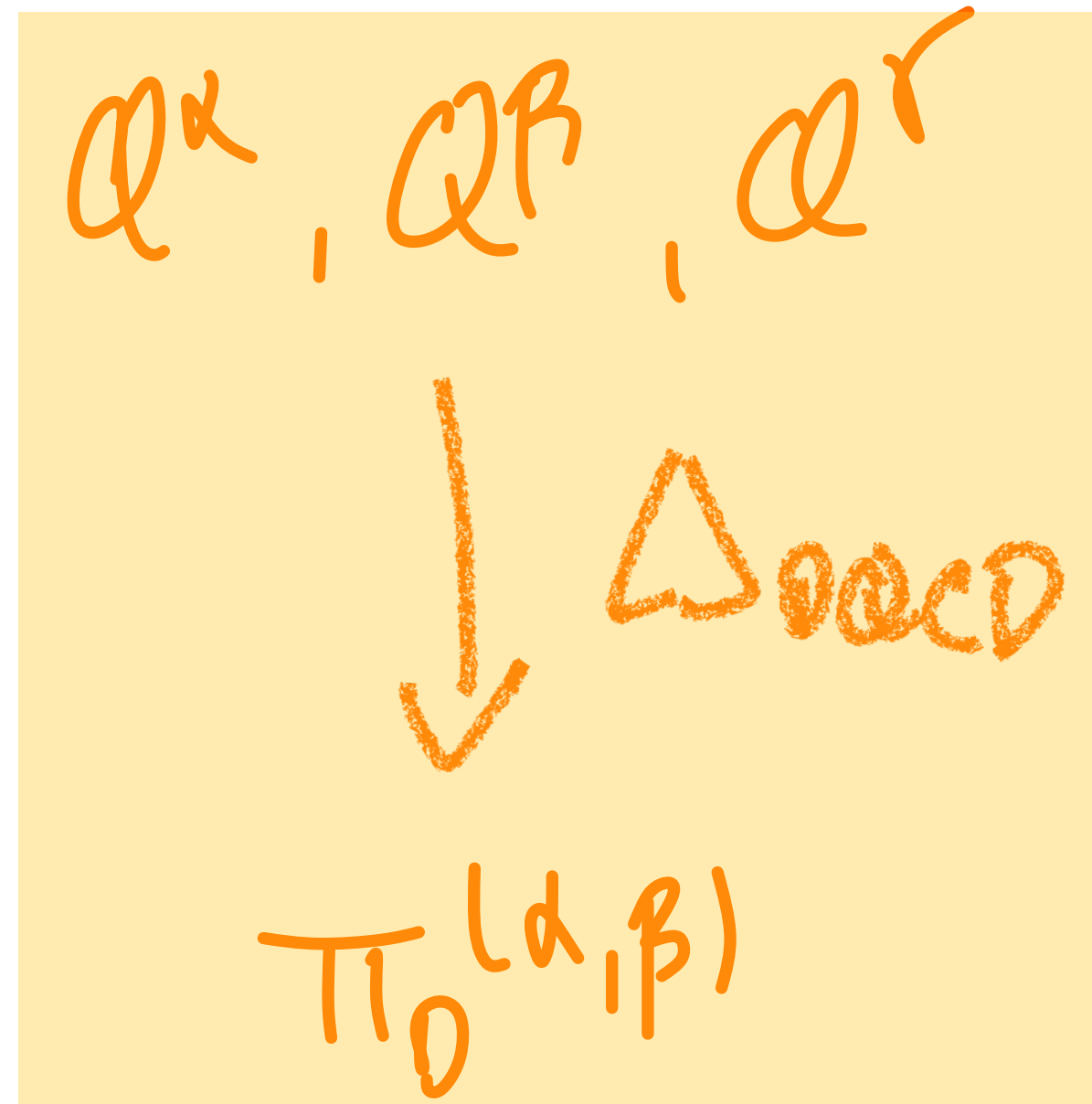
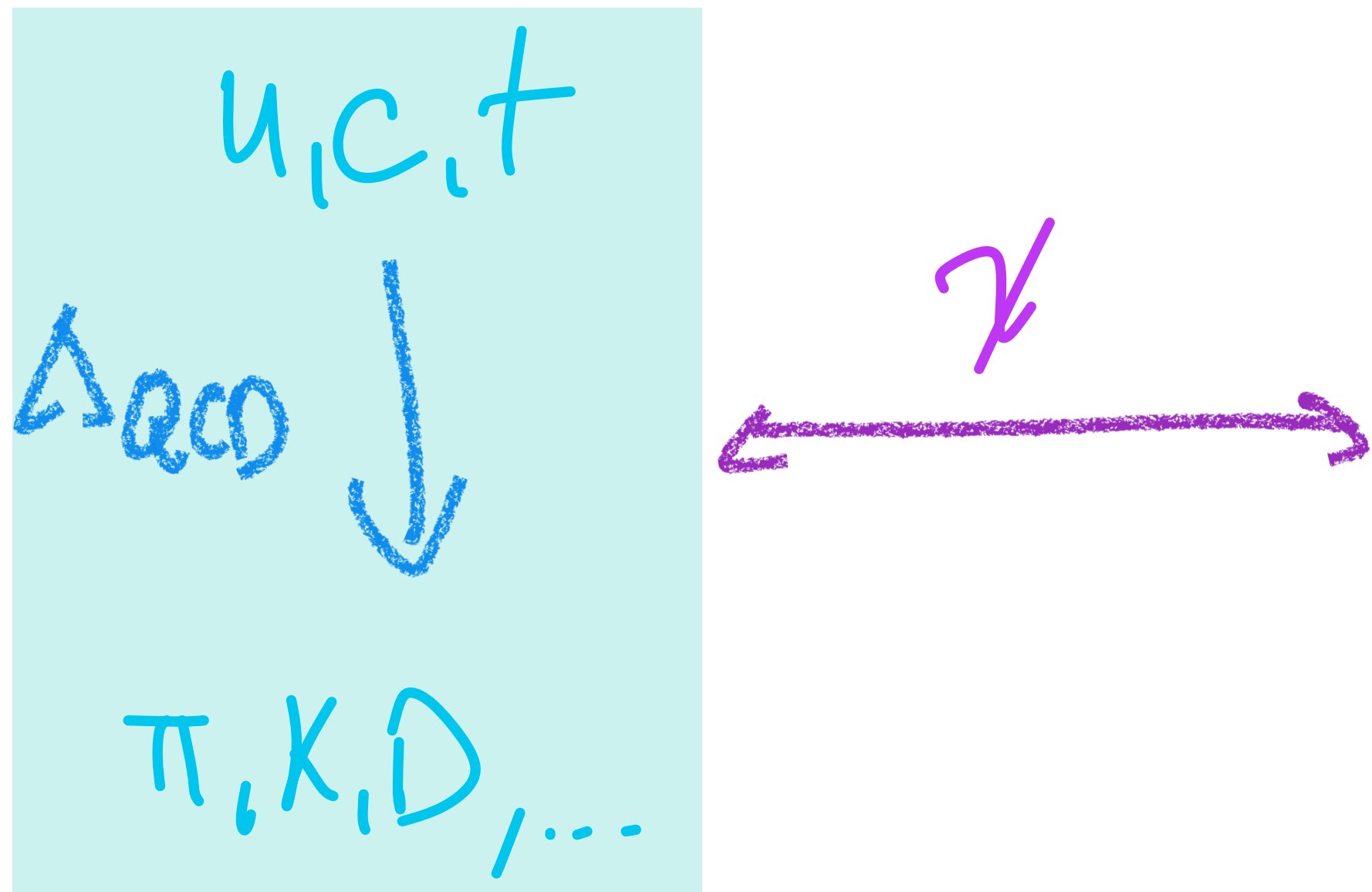
Schwaller, Bai, '14



The SM couplings are fixed by the quantum numbers of χ , bifundamental of both strong gauge groups

A QCD-LIKE DARK SECTOR

Schwaller, Bai, '14



The SM couplings are fixed by the quantum numbers of χ , bifundamental of both strong gauge groups

A QCD-LIKE DARK SECTOR

Schwaller, Bai, '14

When $m_Q \rightarrow 0$, $m_\chi \rightarrow \infty$, $SU(3)_{DL} \otimes SU(3)_{RD} \rightarrow SU(3)_{DV}$ by $\langle \bar{Q}_\alpha Q_\beta \rangle \sim \delta_{\alpha\beta} \Lambda_{DQCD}^3$
delivering 8 pNGB

Dark Pions

Dark Quark content

$$\pi_D^{(1,2)}$$

$$\bar{Q}_2 Q_1$$

$$\pi_D^{(1,3)}$$

$$\bar{Q}_3 Q_1$$

$$\pi_D^{(2,3)}$$

$$\bar{Q}_3 Q_2$$

$$\pi_{D3}$$

$$\frac{1}{\sqrt{2}} [\bar{Q}_1 Q_1 - \bar{Q}_2 Q_2]$$

$$\pi_{D8}$$

$$\frac{1}{\sqrt{6}} [\bar{Q}_1 Q_1 + \bar{Q}_2 Q_2 - 2\bar{Q}_3 Q_3]$$

A QCD-LIKE DARK SECTOR

Schwaller, Bai, '14

Depending of the quantum numbers of the heavy scalar we can have different low energy EFTs

$$SU(3)_C \times SU(3)_D \times SU(2)_L \times U(1)_Y$$

$$\chi \sim (3, \bar{3}, 1, 1/3)$$

Schwaller, Renner '18

$$\mathcal{L}_{int} \supset -K_{\alpha i} \bar{d}_{Ri} Q_{L\alpha} \chi + h.c$$

$$\chi \sim (3, \bar{3}, 1, -2/3)$$

AC, Scherb, Schwaller '21

$$\mathcal{L}_{int} \supset -K_{\alpha i} \bar{u}_{Ri} Q_{L\alpha} \chi + h.c$$

A QCD-LIKE DARK SECTOR

Schwaller, Bai, '14

Depending of the quantum numbers of the heavy scalar we can have different low energy EFTs

$$SU(3)_C \times SU(3)_D \times SU(2)_L \times U(1)_Y$$

$$\chi \sim (3, \bar{3}, 1, 1/3)$$

Schwaller, Renner '18

$$\mathcal{L}_{\text{eff}} \supset \frac{f_D^2}{m_\chi^2} K_{\alpha i} K_{\beta j}^* \partial_\mu \pi_0^{(\alpha, \beta)} \bar{d}_{Ri} \gamma^M d_{Rj}$$

$$\chi \sim (3, \bar{3}, 1, -2/3)$$

AC, Scherb, Schwaller '21

$$\mathcal{L}_{\text{eff}} \supset \frac{f_D^2}{m_\chi^2} K_{\alpha i} K_{\beta j}^* \partial_\mu \pi_0^{(\alpha, \beta)} \bar{u}_{Ri} \gamma^M u_{Rj}$$