



UNIVERSIDAD DE GRANADA





1

Standard Model and open problems

Mikael Chala

(Universidad de Granada)

11th IDPASC School, Olomouc; September 2, 2022

The rho parameter

$$\rho = \frac{m_W^2}{m_Z^2 c_W^2} = 1$$



 $T^{\pm} = T_1 \pm iT_2$

2

$$D_{\mu} \mathbf{H} = \left(\partial_{\mu} + igT_{I}W_{\mu}^{I} + ig'YB_{\mu}\right) \mathbf{H}$$
$$= \left[\partial_{\mu} + \frac{ig}{\sqrt{2}}\left(T^{+}W_{\mu}^{+} + T^{-}W_{\mu}^{-}\right) + \frac{ig}{c_{W}}\left(T_{3} - s_{W}^{2}Q\right)Z_{\mu} + ieQA_{\mu}\right] \mathbf{H}$$

$$(D_{\mu} \mathbf{H})^{\dagger} D^{\mu} \mathbf{H} \supset \underbrace{\frac{g^{2} v^{2}}{4}}_{W_{\mu}} W^{\mu} W^{\mu-} + \underbrace{\frac{g^{2} v^{2}}{8c_{W}^{2}}}_{m_{W}^{2}} Z_{\mu} Z^{\mu}$$
$$\underbrace{\frac{1}{2}m_{Z}^{2}}_{m_{Z}^{2}} Z_{\mu} Z^{\mu}$$



The T (rho-1) parameter



4

Proposed exercise: Show that if the Higgs boson is a colourless SU(2) triplet with Y=1, then the rho parameter differs from 1

Lepton and baryon numbers (accidentally) conserved: Neutrinos are massless; proton is stable

$$\mathcal{O}^{(5)} = (\overline{L_L^c} \tilde{\phi}^*) (\tilde{\phi}^\dagger L_L)$$

$$\begin{split} \varepsilon^{\alpha\beta\gamma}\varepsilon_{jk}\left[(d_{p}^{\alpha})^{T}Cu_{r}^{\beta}\right]\left[(q_{s}^{\gamma j})^{T}Cl_{t}^{k}\right]\\ \varepsilon^{\alpha\beta\gamma}\varepsilon_{jk}\left[(q_{p}^{\alpha j})^{T}Cq_{r}^{\beta k}\right]\left[(u_{s}^{\gamma})^{T}Ce_{t}\right]\\ \varepsilon^{\alpha\beta\gamma}\varepsilon_{jn}\varepsilon_{km}\left[(q_{p}^{\alpha j})^{T}Cq_{r}^{\beta k}\right]\left[(q_{s}^{\gamma m})^{T}Cl_{t}^{n}\right]\\ \varepsilon^{\alpha\beta\gamma}\left[(d_{p}^{\alpha})^{T}Cu_{r}^{\beta}\right]\left[(u_{s}^{\gamma})^{T}Ce_{t}\right] \end{split}$$

Lepton and baryon numbers are broken nonperturbatively:

$$\Delta L = \Delta B = 3$$

1601.03654

$$qq \to \overline{lll}\overline{q}\,\overline{q}\,\overline{q}\,\overline{q}\,\overline{q}\,\overline{q}\,\overline{q}$$

$$\Gamma \sim e^{-\frac{(4\pi)^2}{g^2}}$$

 $\Gamma \sim T^4$

At high temperature

Suppressed flavour-changing neutral currents

$$L_Y \to -\frac{v}{\sqrt{2}} \left(1 + \frac{h}{v}\right) \left(y_{ij}^u \overline{u_L^i} u_R^j + y_{ij}^d \overline{d_L^i} d_R^j + y_{ij}^l \overline{l_L^i} l_R^j + \text{h.c.}\right)$$

 $(\mathcal{U}_L^u)^{\dagger} y^u \mathcal{U}_R^u = \operatorname{diag}(y_u, y_c, y_t)$ $(\mathcal{U}_L^d)^{\dagger} y^d \mathcal{U}_R^d = \operatorname{diag}(y_d, y_s, y_b)$

$$c_Z^V \overline{\mathbf{u}_{\mathbf{L}}} \gamma_\mu \mathbf{u}_{\mathbf{L}} Z_\mu \to c_Z^V \overline{\mathbf{u}_{\mathbf{L}}} (\mathcal{U}_L^u)^{\dagger} \gamma_\mu \mathcal{U}_L^u \mathbf{u}_{\mathbf{L}} Z^\mu = c_Z^V \overline{\mathbf{u}_{\mathbf{L}}} \gamma_\mu \mathbf{u}_{\mathbf{L}} Z_\mu$$

FCNC amplitudes can not arise at tree level

Suppressed flavour-changing neutral currents

$$-\frac{g}{2\sqrt{2}}\left\{\left[\overline{u^{i}}\gamma^{\mu}(1-\gamma^{5})V_{ij}d^{j}+\overline{\nu^{i}}\gamma^{\mu}(1-\gamma^{5})l^{j}\right]W_{\mu}^{+}+\text{h.c.}\right\}$$

$$V = (\mathcal{U}_L^u)^{\dagger} \mathcal{U}_L^d$$

CKM matrix, it's unitary

FCNC amplitudes can arise at one loop, but suppressed:



For equal mass (GIM mechanism!)

Suppressed flavour-changing neutral currents

Proposed exercise: figure out an extension of the Standard Model that gives tree-level FCNCs.

CKM matrix for three generations:

$$V = (\mathcal{U}_L^u)^{\dagger} \mathcal{U}_L^d$$

CP violation!

$$\begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta} & c_{23}c_{13} \end{pmatrix}$$

All Shakarov conditions for baryogenesis are present within the SM:

- Baryon number violation
- C violation
- CP violation

Why the three are needed for generating ΔB ?

Let's see what happens in we have baryon number violation but C is conserved:

$$\begin{split} \Gamma(p^+ \to e^+ \gamma) \\ & \parallel \\ \Gamma(p_L^+ \to e_R^+ \gamma_L) + \Gamma(p_R^+ \to e_L^+ \gamma_R) \\ & \parallel \text{because C is conserved } \parallel \\ \Gamma(p_L^- \to e_R^- \gamma_L) \quad \Gamma(p_R^- \to e_L^- \gamma_R) \end{split}$$

Proposed exercise: Show that CP is also needed or else there exist processes that wash out the baryon asymmetry

However, it is well known that the amount of CPV within the SM it is not enough for baryogenesis

One can add new sources of CPV, but this (seemingly) conflicts with experiment; e.g.:



A way out: spontaneous CPV.

CPV must occur in the early Universe, namely at high T (it can be negligible at current experiments!)



Anomalies: classical symmetries broken at the quanutm level. Global anomalies are OK, but gauge anomalies are forbidden!

Among other conditions, for the fields of the SM we get:

$$Q_{\nu} + Q_{e} + N_{c}(Q_{u} + Q_{d}) = -1 + \frac{1}{3}N_{c} = 0 \quad \Rightarrow \quad N_{c} = 3 \quad (!!)$$

Leptons and quarks needed in every generation!

Gauge coupling evolution

 $\beta_g = 16\pi^2 \mu \frac{dg}{d\mu} \Rightarrow g(\mu)$

Renormalisation in a nutshell

All divergences are local, namely they can be cast in Lagrangian form

$$L_{QED}^{one-loop} = Z_{\psi} \, i \overline{\psi} \gamma^{\mu} D_{\mu} \psi, \quad Z_{\psi} = 1 + \mathcal{O}(1/\epsilon), \quad D = 4 - 2\epsilon$$
$$= (Z_{\psi} \overline{\psi} \gamma^{\mu} \partial_{\mu} \psi + \sqrt{Z_A} Z_e Z_{\psi} e \overline{\psi} \gamma^{\mu} \psi A_{\mu})$$
$$= 1$$
$$Z_e = Z_A^{-1/2} \Rightarrow \qquad \sum_{p} \sum_{k=p}^{k-p} \sum_{p} \sum_{k=p}^{k-p} \sum_{p} \sum_{k=p}^{k-p} \sum_{p} \sum_{p} \sum_{k=p}^{k-p} \sum_{p} \sum_$$

Gauge coupling running. Renormalisation in a nutshell

Anecdotal comment: This does not work for Yang-Mills out of the box, because quantization breaks gauge invariance (gauge fixing)

 \rightarrow Use the so-called background field method instead for QCD and SU(2)!

Proposed exercise: Compute the beta function of QED and demonstrate that it grows at high energies

Gauge coupling running



Gauge coupling running: Let's talk about SUSY



Gauge coupling running: Let's talk about SUSY



So, at very high energies we could have a single group...



$$\begin{aligned} \mathbf{SU}(5) \\ \mathbf{\overline{5}}_{F} = \begin{pmatrix} d_{1}^{c} \\ d_{2}^{c} \\ d_{3}^{c} \\ e^{-} \\ -\nu_{e} \end{pmatrix} \quad \mathbf{10}_{F} = \begin{pmatrix} 0 & u_{3}^{c} & -u_{2}^{c} & u_{1} & d_{1} \\ -u_{3}^{c} & 0 & u_{1}^{c} & u_{2} & d_{2} \\ u_{2}^{c} & -u_{1}^{c} & 0 & u_{3} & d_{3} \\ \hline -u_{1} & -u_{2} & -u_{3} & 0 & e^{+} \\ -d_{1} & -d_{2} & -d_{3} & -e^{+} & 0 \end{pmatrix} \end{aligned}$$

SO(10) $16_F = (Q, u^c, d^c, L, \nu^c, e^c)$

Still, 3 families like in the SM

Georgi–Jarlskog mass relation

$$m_{\rm b} = m_{\tau}, \quad m_{\mu} = 3m_{\rm s}, \quad m_{\rm d} = 3m_{\rm e}$$

Relations that hold at the GUT scale

One group to unify them all...

$\underbrace{171}_{SU(19)} \rightarrow \underbrace{4Q + 4u^c + 5d^c + 5L + 4e^c + Q^c + u + 2d + 2L^c + e + (\text{more vector fermions})}_{SU(3)_C \times SU(2)_L \times U(1)_Y}$

Ekstedt, Fonseca and Malinsky '20

Problems:

- * hard computations
- * gravity effects not always under control
- * decoupling all scalars but the Higgs
- * proton decay!



Mass hierarchy



Hierarchical CKM



 $V_{\rm CKM} \sim 1$

Hierarchical CKM



Complementary to the GIM mechanism!

$$\sum V_{ki}V_{kj}^{*}F(m_{u^{k}}) = V_{ui}V_{uj}^{*}F(m_{u}) + V_{ci}V_{cj}^{*}F(m_{c}) + V_{ti}V_{tj}^{*}F(m_{t})$$
$$\sim (V_{ui}V_{uj}^{*} + V_{ci}V_{cj}^{*})F(0) + V_{ti}V_{tj}^{*}F(m_{t})$$
$$\sim V_{ti}V_{tj}^{*}[F(m_{t}) - F(0)]$$

Some (many?) people claim this hierarchy of masses **cries** for an explanation

One possibility: GUTs, but hard to make them work properly

Other possibility: Frogatt-Nielsen-like models:

$$L_Y = y_t \overline{q_L^3} H t_R + y_u \overline{q_L^1} H u_R + y_u \overline{q_L^1} H \frac{S}{f} u_R$$

[e.g.: Q(qL1) = Q(uR) = -1, Q(S) = 2, Q(rest) = 0]

$$y_t \sim y_u \Rightarrow m_t \gg m_u$$

You predict the hierarchy?

What's the argument?

If yt and yu are taken randomly from a **flat distribution**, the probability of yt/yu being O(1) is large, while yt/yu being O(10000) is small

Why assuming a flat distribution on the first place? Because in the absence of other knowledge, flat distribution = maximal ignorance

OK, do the same reasoning but for $\log(yt)/\log(yu)$ instead. What's the result in this case?

What's the argument?

If yt and yu are taken randomly from a **flat**

towardsdatascience.com/stop-using-uniform-priors-47473bdd0b8a

ununa or the minucinee or outhers, try using t unstribution us a prior, rus

mentioned at the start of this section, a normal prior is almost always better than a uniform prior.

In Closing

Whenever you think about using a uniform prior, remember: just don't.

Higgs mass: the wrong calculation



$$\int_{0}^{\Lambda} \dots \Rightarrow \, \delta m_{H}^{2} = \left[\frac{1}{4}(9g^{2} + 2g'^{2}) - 6y_{t}^{2} + 6\lambda\right] \frac{\Lambda^{2}}{32\pi^{2}}$$

Stability under quantum corrections

Masses and mixing parameters (including the Higgs mass!) are stable under **Standard Model** quantum corrections

$$\beta_{\mu^{2}} = \left[2 \operatorname{Tr}(y^{e} y^{e^{\dagger}}) + 6 \operatorname{Tr}(y^{u} y^{u^{\dagger}}) + 6 \operatorname{Tr}(y^{d} y^{d^{\dagger}}) - \frac{3}{2}g_{1}^{2} - \frac{9}{2}g_{2}^{2} - 12\lambda \right] \mu^{2}$$

$$\beta_{y^{u}} = \left\{ \frac{3}{2}y^{u} y^{u^{\dagger}} - \frac{3}{2}y^{d} y^{d^{\dagger}} + 3 \left[\operatorname{Tr}(y^{u} y^{u^{\dagger}}) + \operatorname{Tr}(y^{d} y^{d^{\dagger}}) \right] + \operatorname{Tr}(y^{e} y^{e^{\dagger}}) - \frac{17}{12}g_{1}^{2} - \frac{9}{4}g_{2}^{2} - 8g_{3}^{2} \right\} y^{u}$$

The actual hierarchy problem



Fine tuning subjective (like for CKM). However, prediction of charm quark indication that naturalness might work...³⁸

Most clear solution: SUSY again



Difficulty: susy must be broken

One more solution: composite Higgs models

Cannot resolve the loop to high energies because I see the Higgs structure... T



The constituents of the Higgs make also other composite particles: most importantly **vector-like quarks (VLQs)**

VLQs should be much heavier than the Higgs, otherwise we would have seen them already

Solution? Copy the structure of QCD

QCD

 $SU(2)_L \times SU(2)_R \rightarrow SU(2)_{L+R}$

3 Goldstones, the pions, much lighter than other hadrons

CHMs

$G \to H$

4 Goldstones, the Higgs degrees of freedom, much lighter than VLQs

Who can be G and H?

Solution? Copy the structure of QCD

QCD

 $SU(2)_L \times SU(2)_R \rightarrow SU(2)_{L+R}$

3 Goldstones, the pions, much lighter than other hadrons

CHMs

$$G \to H$$

4 Goldstones, the Higgs degrees of freedom, much lighter than VLQs

Who can be G and H?

 $SO(5) \rightarrow SO(4)$ $SO(6) \rightarrow SO(5)$

. . .

Difficulty: *G* must be broken

Some more fundamental problem: dark matter (DM)

One interesting hypothesis: DM is formed by **neutral weakly-interacting non-relativistic** particles



Some more fundamental problem: dark matter (DM)

One interesting hypothesis: DM is formed by **neutral weakly-interacting non-relativistic** particles



44

Composite Higgs models also work! e.g. $SO(6) \rightarrow SO(5)$

$$L = \frac{1}{2} (\partial_{\mu} h)^{2} + \frac{1}{2} (\partial_{\mu} S)^{2}$$
$$- \left\{ \frac{1}{2} \mu^{2} h^{2} + \frac{1}{4} \lambda_{h} h^{4} + \frac{1}{2} \mu_{S}^{2} S^{2} + \frac{1}{3} \kappa S^{3} + \frac{1}{4} \lambda_{S} S^{4} + \frac{1}{3} \kappa_{hS} h^{2} S + \frac{1}{4} \lambda_{hS} h^{2} S^{2} \right\}$$

A comparison with R-parity in SUSY (motivation relies *also* on dark matter)



Signals and status of SUSY and CHMs:



Signals and status of SUSY and CHMs:



Neutrino masses

They can arise at dimension-5. How can this be completed in the UV? $\mathcal{O}^{(5)} = (\overline{L_L^c} \tilde{\phi}^*) (\tilde{\phi}^\dagger L_L)$



Neutrino masses

Searches at the LHC



Neutrino masses



- Other (potential) problems of the SM:
 - * g-2, flavour anomalies, ...
 - * Why three generations?* Why charge is quantised?

* How to make computations without using fields? What do we learn from that?