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Prague, 17 October 2022

The Tevatron



Highest energy proton-antiproton collider, CMS energy ≈ 2 TeV, 1986-2011

anti-proton
source



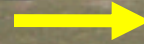
Recycler (anti- π @ 8 GeV)



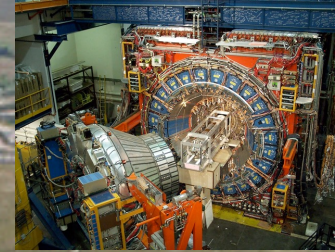
Main Injector (8-150 GeV)

Tevatron (150-1000 GeV)

protons



CDF



Dzero



anti-protons



Prague Scientists Check in at Fermilab

by Chad Boutin

The particles could travel, but the particle physicists couldn't.

Czech physicists now form a lively group at Fermilab's DZero. But for many Cold War decades, scientists in Prague were isolated from their Western colleagues.



"Fermilab was like Jupiter for us," said Ivan Wilhelm, Charles University rector and himself a nuclear physicist. "We knew it existed, but we couldn't reach it."

Researchers from three Prague institutions--Charles University, Czech Technical University, and the Academy of Sciences of the Czech Republic-- are now fully engaged with the community at DZero, contributing to the experiment and to the education of future scientists.

"The Prague group has been right up to speed with us from the outset," said John Womersley, spokesman for DZero. "It's as though there had never been a barrier between us at all."

Building a Collaboration

Once the Iron Curtain was torn down, senior physicist Vladimír Simák of the Academy of Sciences saw places like Fermilab as international destinations and sources of international collaboration, instead of as interplanetary dreams. Simák made contact with a number of Western colleagues at a 1996 workshop in Padua, Italy, which led to his collaboration on experiments at CERN, the European Laboratory for Particle Physics. It was in Geneva that he first heard about the planned upgrades to Fermilab's Tevatron. For a physicist with longstanding interest in probing the structure of the proton, the possibility of contributing to Collider Run II was a great temptation.



Vladislav Simak

Many contributions from Czech groups to Dzero detector components and physics:

- Muon HV system
- Silicon and luminosity detectors
- Computing
- Jet physics



The Dzero Detector

Tracking:

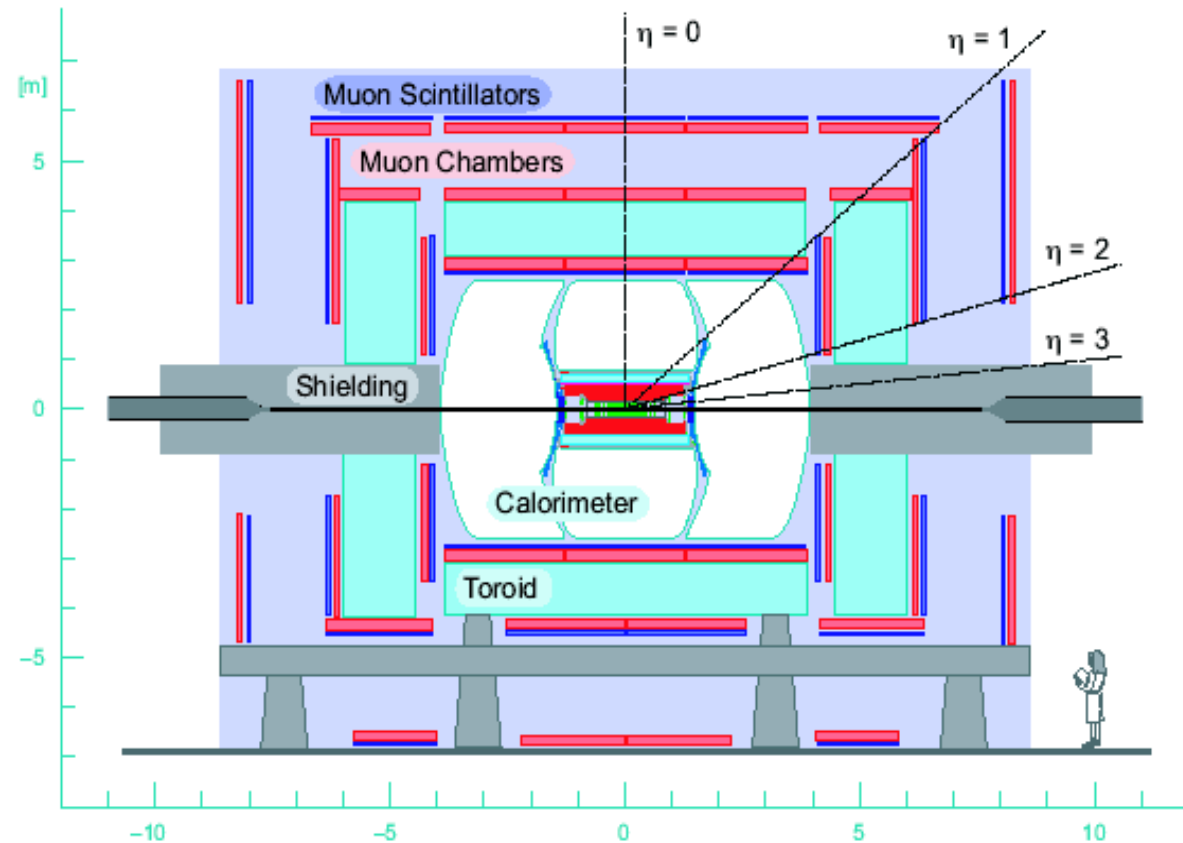
Silicon Microstrip Tracker
Central Fiber Tracker
2 T Solenoid

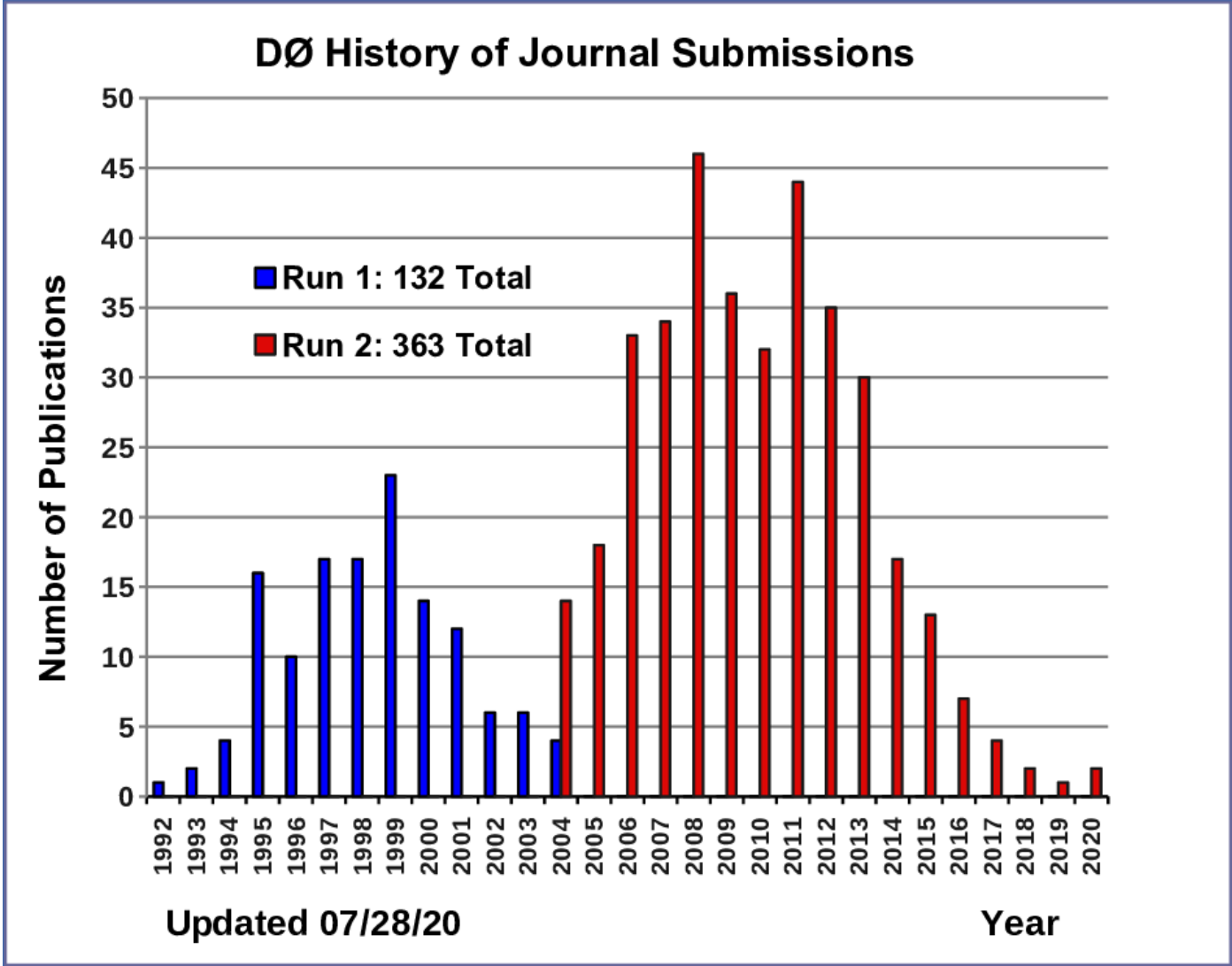
Calorimeter:

Liquid Argon Calorimeter
Scintillator Pre-shower

Muons:

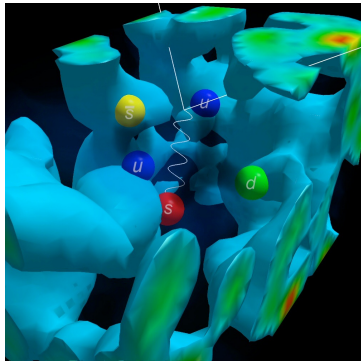
Drift Tubes
Scintillators
1.8 T Toroid





A rich physics programme

QCD



Top



B physics



New Phenomena

$$\begin{aligned} \mathcal{L}_{GWS} = & \sum_f (\bar{\Psi}_f (i\gamma^\mu \partial_\mu - m_f) \Psi_f - e Q_f \bar{\Psi}_f \gamma^\mu \Psi_f A_\mu) + \\ & + \frac{g}{\sqrt{2}} \sum_i (\bar{a}_i \gamma^\mu b_i W_\mu^+ + \bar{b}_i \gamma^\mu a_i W_\mu^-) + \frac{g}{2c_w} \sum_f \bar{\Psi}_f \gamma^\mu (I_f^3 - 2s_w^2 Q_f - I_f^3 \gamma_5) \Psi_f Z_\mu + \\ & - \frac{1}{4} |\partial_\mu A_\nu - \partial_\nu A_\mu - ie(W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 - \frac{1}{2} |\partial_\mu W_\nu^+ - \partial_\nu W_\mu^+ + \\ & - ie(W_\mu^+ A_\nu - W_\nu^+ A_\mu) + ig' c_w (W_\mu^+ Z_\nu - W_\nu^+ Z_\mu)|^2 + \\ & - \frac{1}{4} |\partial_\mu Z_\nu - \partial_\nu Z_\mu + ig' c_w (W_\mu^- W_\nu^+ - W_\mu^+ W_\nu^-)|^2 + \\ & - \frac{1}{2} M_\eta^2 \eta^2 - \frac{g M_\eta^2}{8 M_W} \eta^3 - \frac{g'^2 M_\eta^2}{32 M_W} \eta^4 + |M_W W_\mu^+ + \frac{g}{2} \eta W_\mu^+|^2 + \\ & + \frac{1}{2} |\partial_\mu \eta + i M_Z Z_\mu + \frac{ig}{2c_w} \eta Z_\mu|^2 - \sum_f \frac{g m_f}{2 M_W} \bar{\Psi}_f \Psi_f \eta \end{aligned}$$

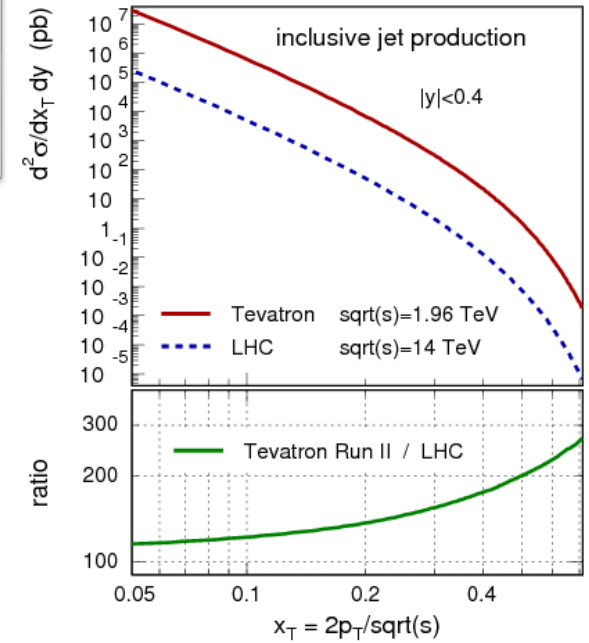
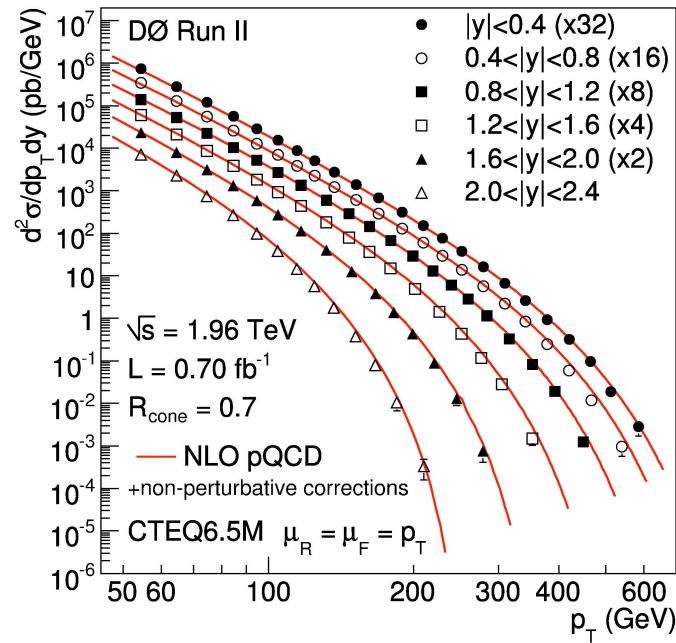
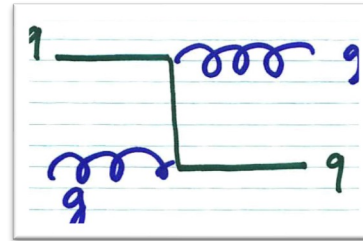
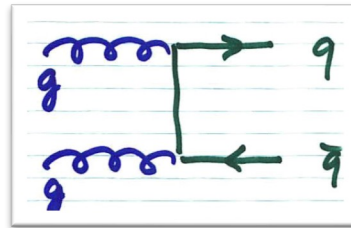
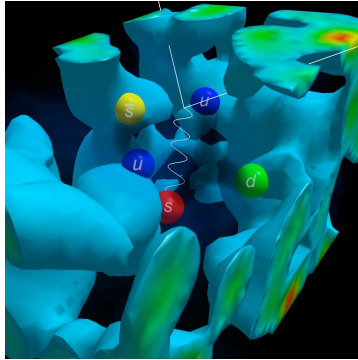
Electroweak



Higgs

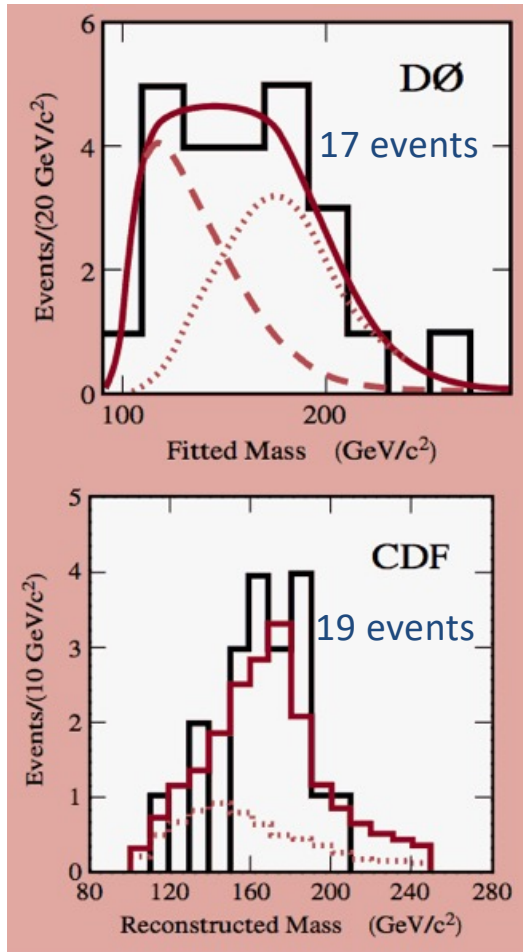
A rich physics programme

QCD



$$x_T = 2 \times 700 \text{ GeV} / 1.96 \text{ TeV} = 0.7$$

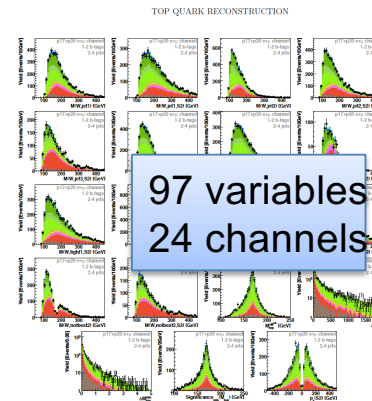
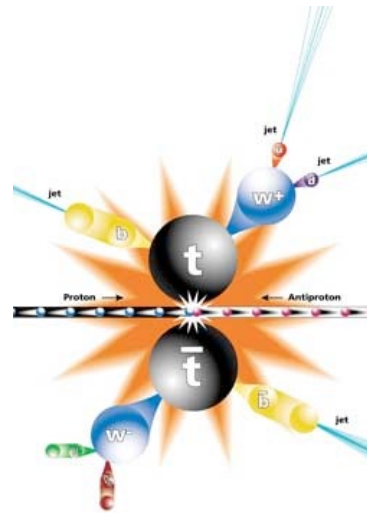
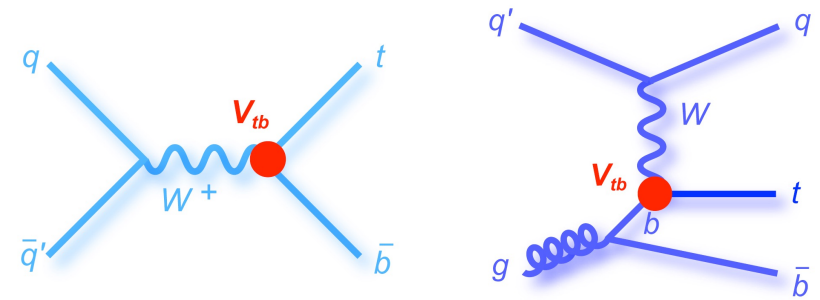
A rich physics programme



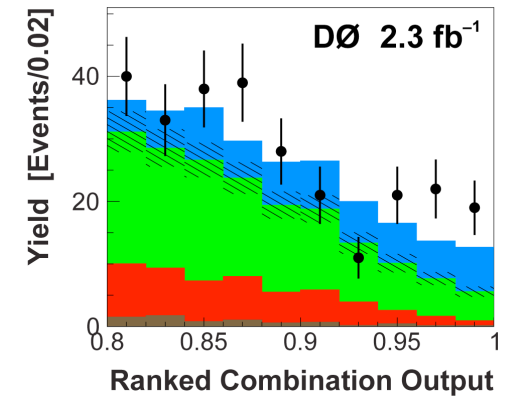
Top



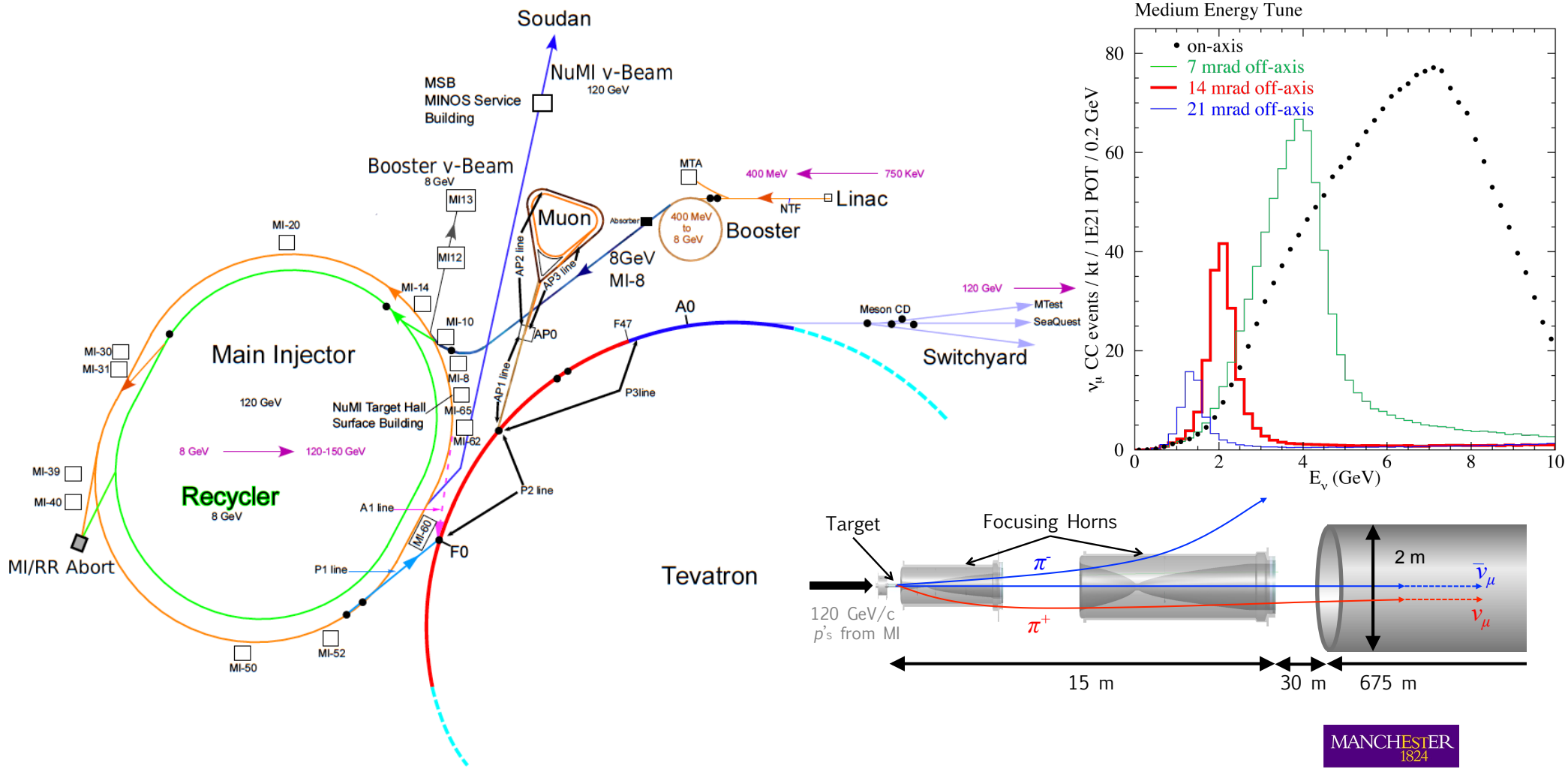
Electroweak production, single top



Signal Region



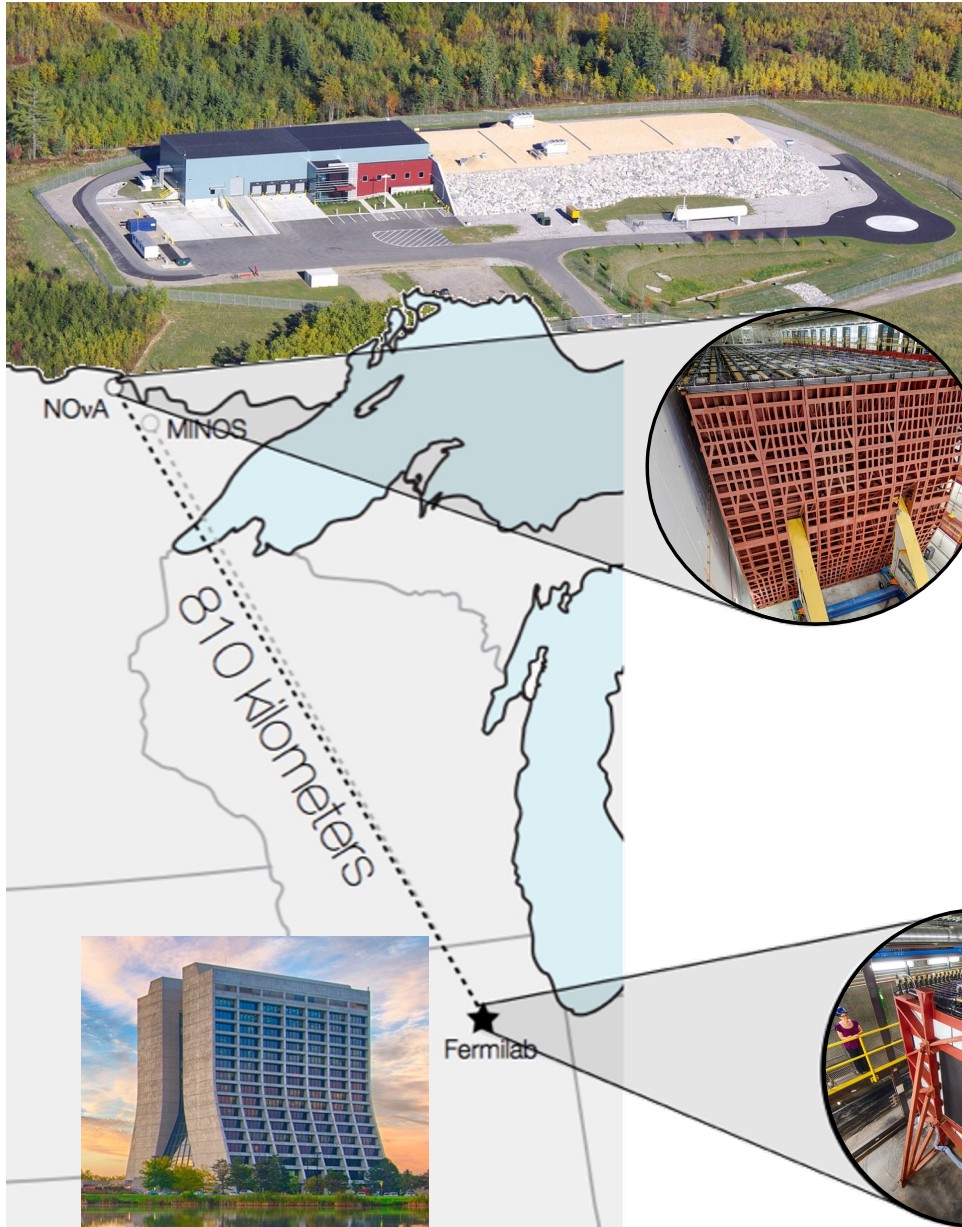
From the energy to the intensity frontier: neutrinos





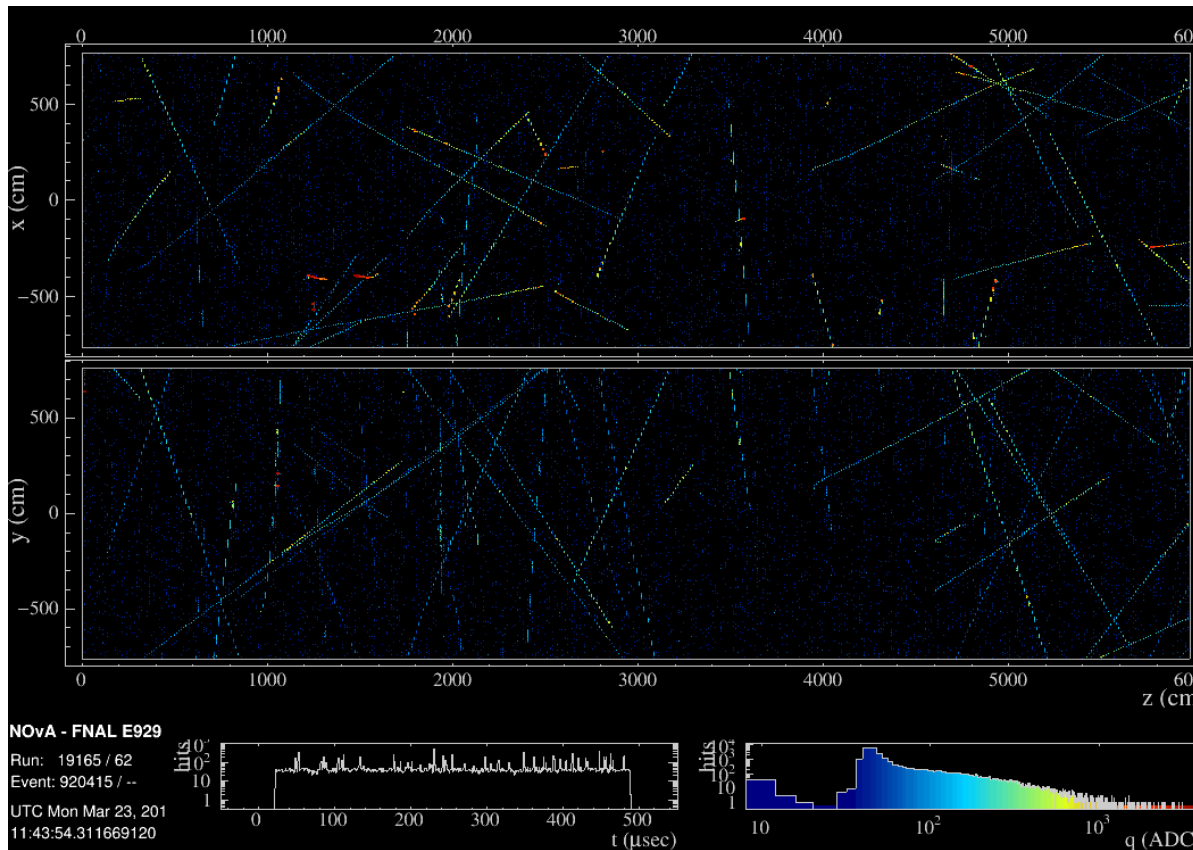
NOvA Experiment

- NuMI beam: ν_μ or $\bar{\nu}_\mu$
- 2 functionally identical, tracking calorimeter detectors
 - Near: 300 T underground
 - Far: 14 kT on the surface
 - Placed off-axis to produce a narrow-band spectrum
- 810 km baseline
 - Longest baseline of current experiments.

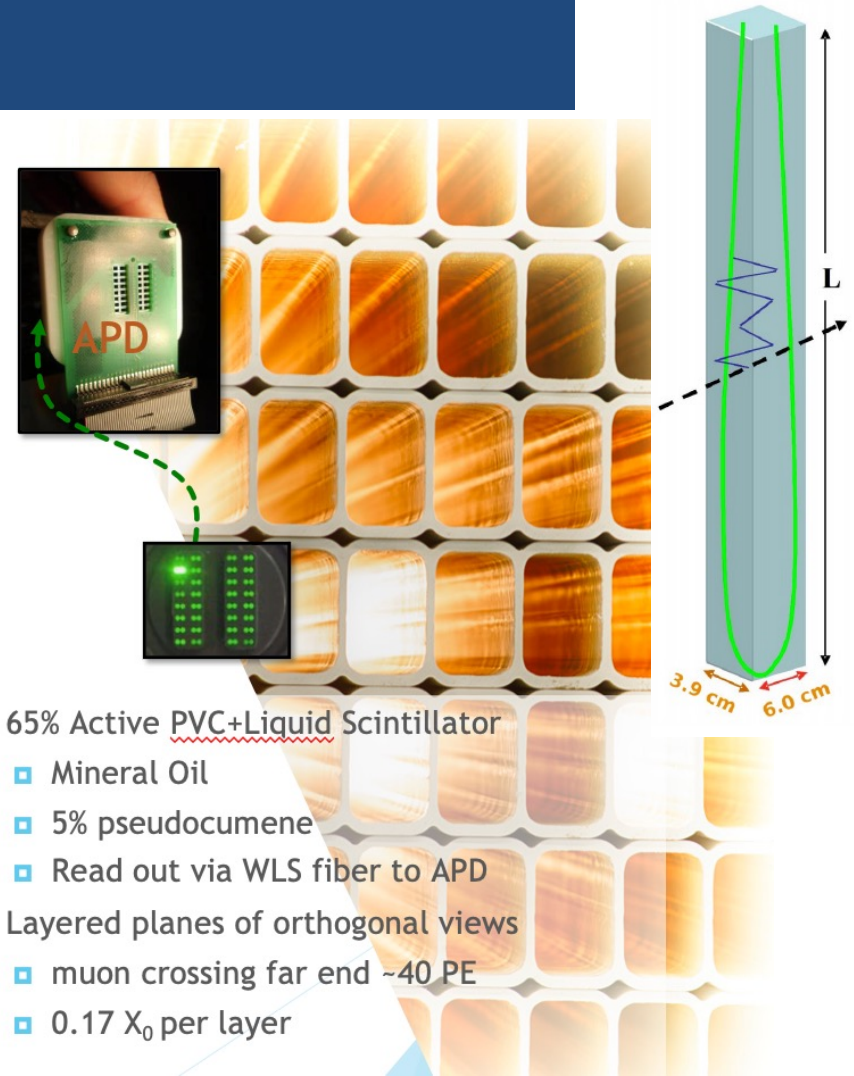




NOvA is on the surface..

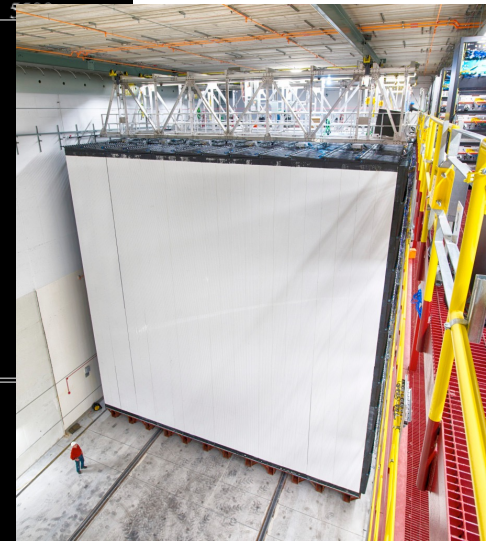
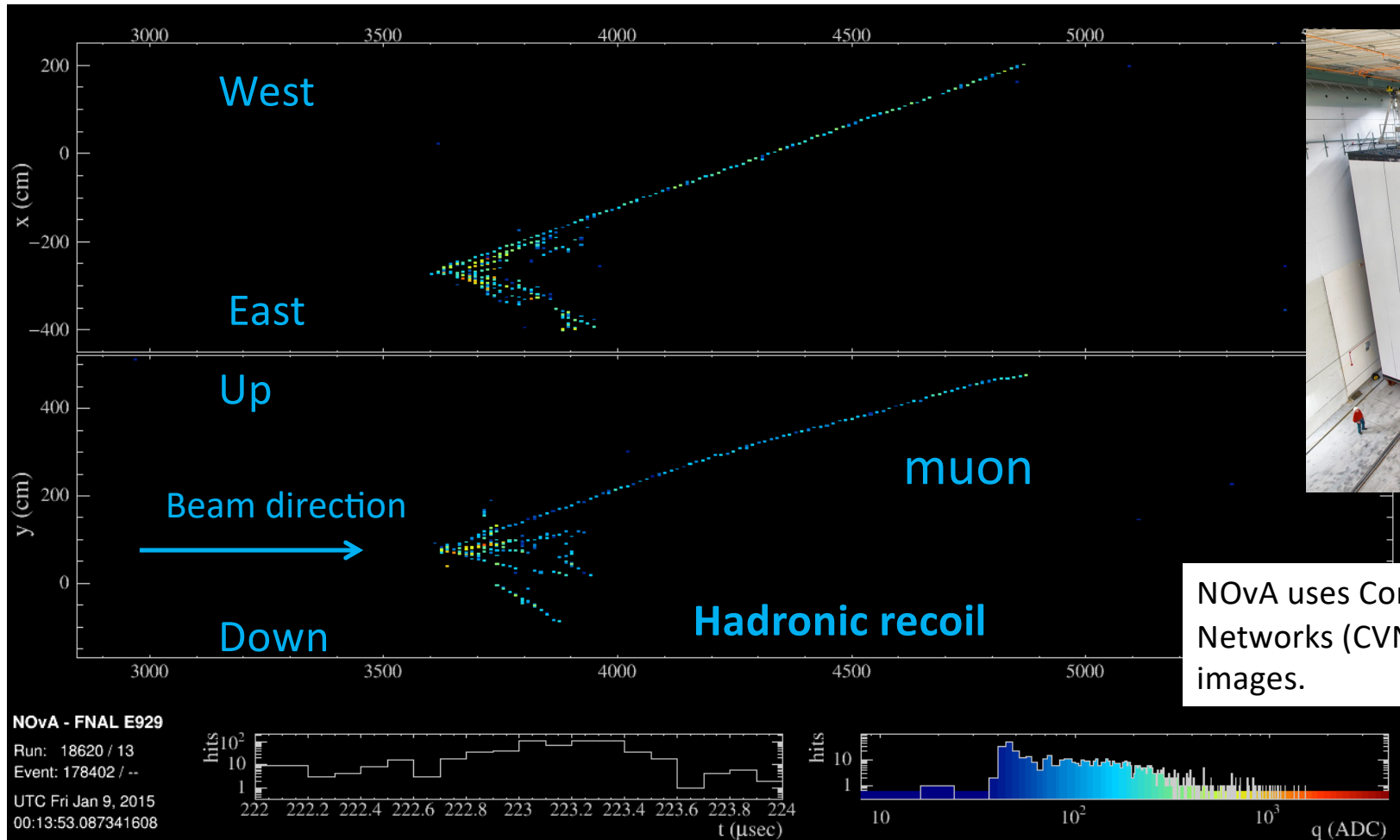


- 14 kt Far Detector
- Equivalent Near Detector



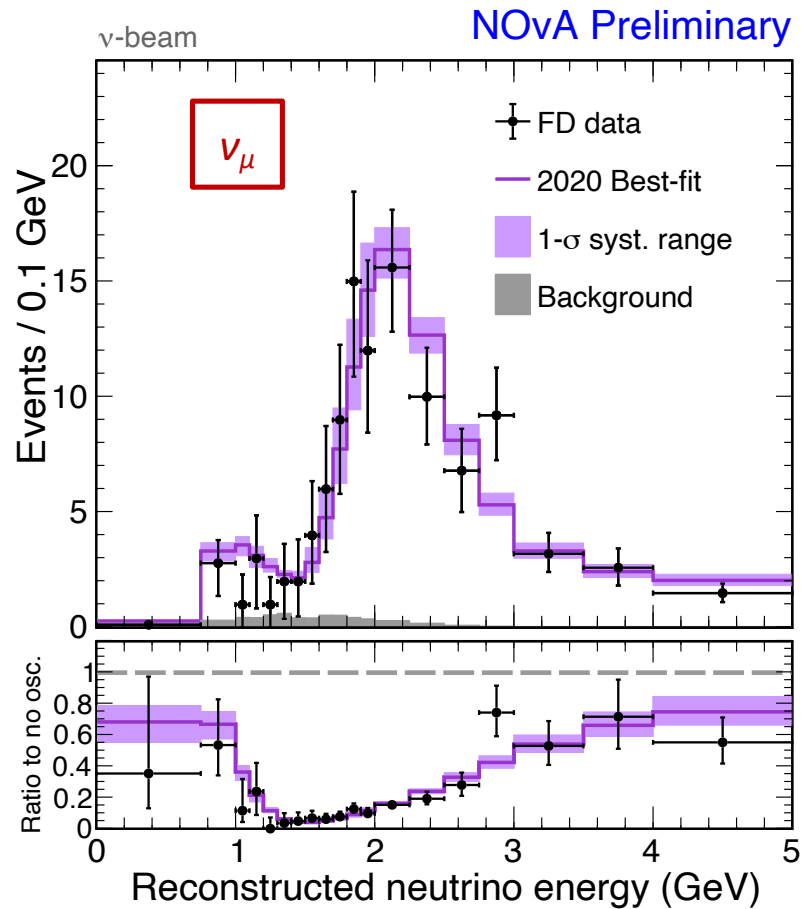
- 65% Active PVC+Liquid Scintillator
 - Mineral Oil
 - 5% pseudocumene
 - Read out via WLS fiber to APD
- Layered planes of orthogonal views
 - muon crossing far end ~40 PE
 - 0.17 X_0 per layer

NOvA Detector

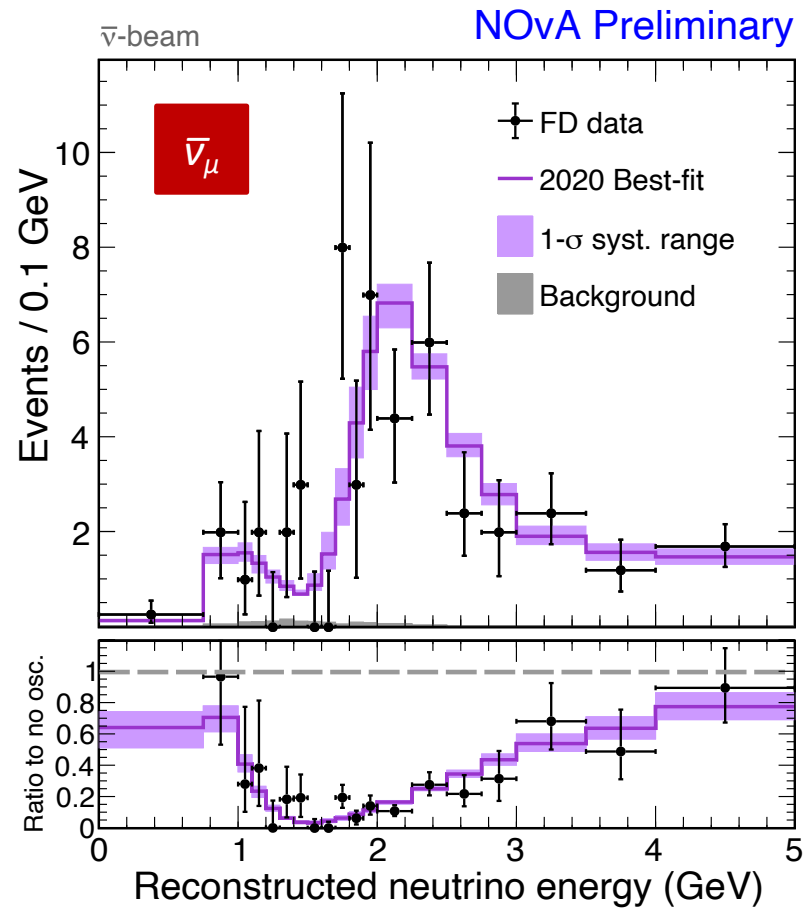


NOvA uses Convolutional Neural Networks (CVNs) to reconstruct images.

ν_μ and $\bar{\nu}_\mu$ disappearance at the NOvA Far Detector



211 events, 8.2 background



105 events, 2.1 background

To give you a “flavour”

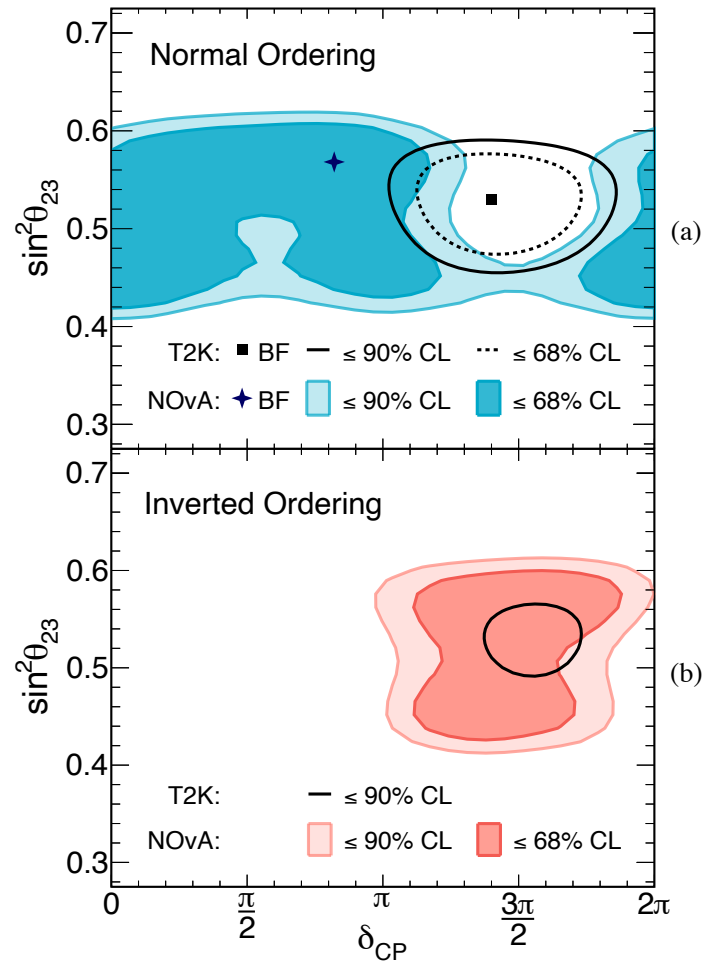
$$\begin{aligned}
 P(\nu_\mu \rightarrow \nu_e) \approx & \sin^2 \theta_{23} \sin^2 2\theta_{13} \frac{\sin^2(\Delta_{31} - aL)}{(\Delta_{31} - aL)^2} \Delta_{31}^2 \\
 & + \sin 2\theta_{23} \sin 2\theta_{13} \sin 2\theta_{12} \frac{\sin(\Delta_{31} - aL)}{(\Delta_{31} - aL)} \Delta_{31} \frac{\sin(aL)}{aL} \Delta_{21} \cos(\Delta_{31} - \delta) \\
 & + \cos^2 \theta_{23} \sin^2 2\theta_{12} \frac{\sin(aL)}{aL} \Delta_{21}^2
 \end{aligned}$$

$$\begin{aligned}
 a &= \frac{G_F N_e}{\sqrt{2}} \\
 \Delta_{ij} &= \frac{\Delta m_{ij}^2 L}{4E}
 \end{aligned}$$

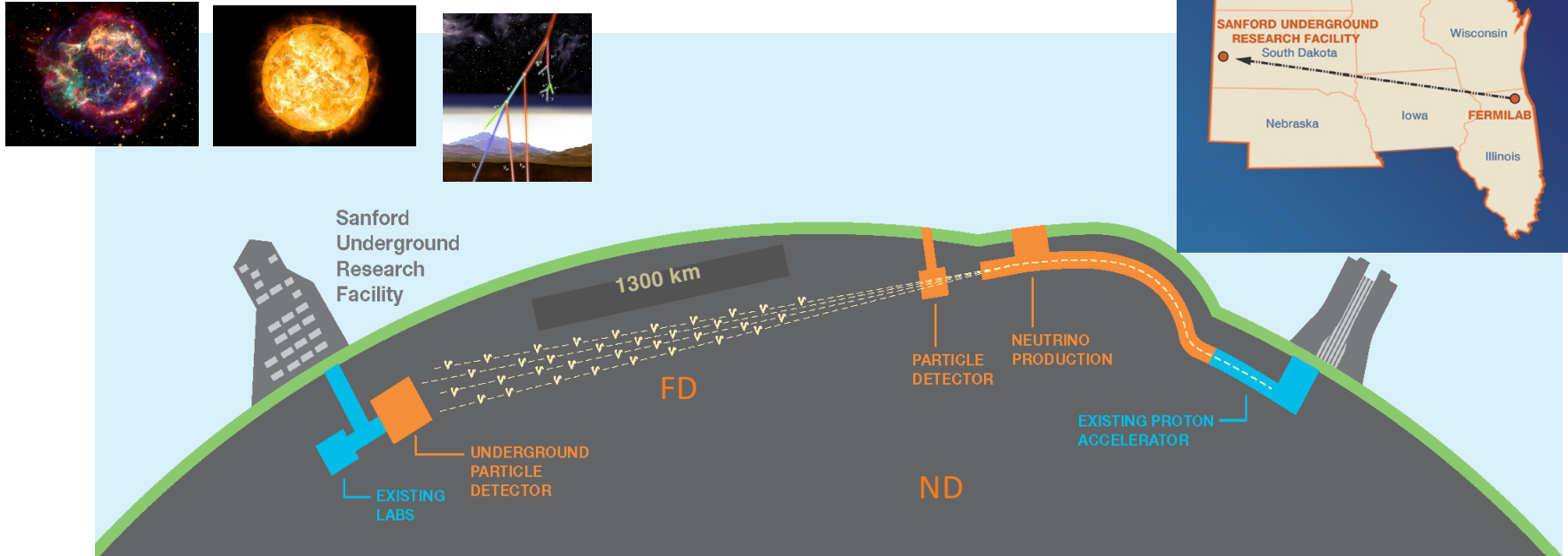
- Electron-neutrino appearance is sensitive to the CP phase.
- Mass ordering effect depends on electron density N_e .
- Simultaneous determination of mass ordering and CP phase only possible with long-baselines by fitting the energy dependence of the flux modulation.

NOvA and T2K

- Current data are inconclusive – expect some improvements with further running.
- Need next-generation experiments to discover CPV and resolve mass ordering.



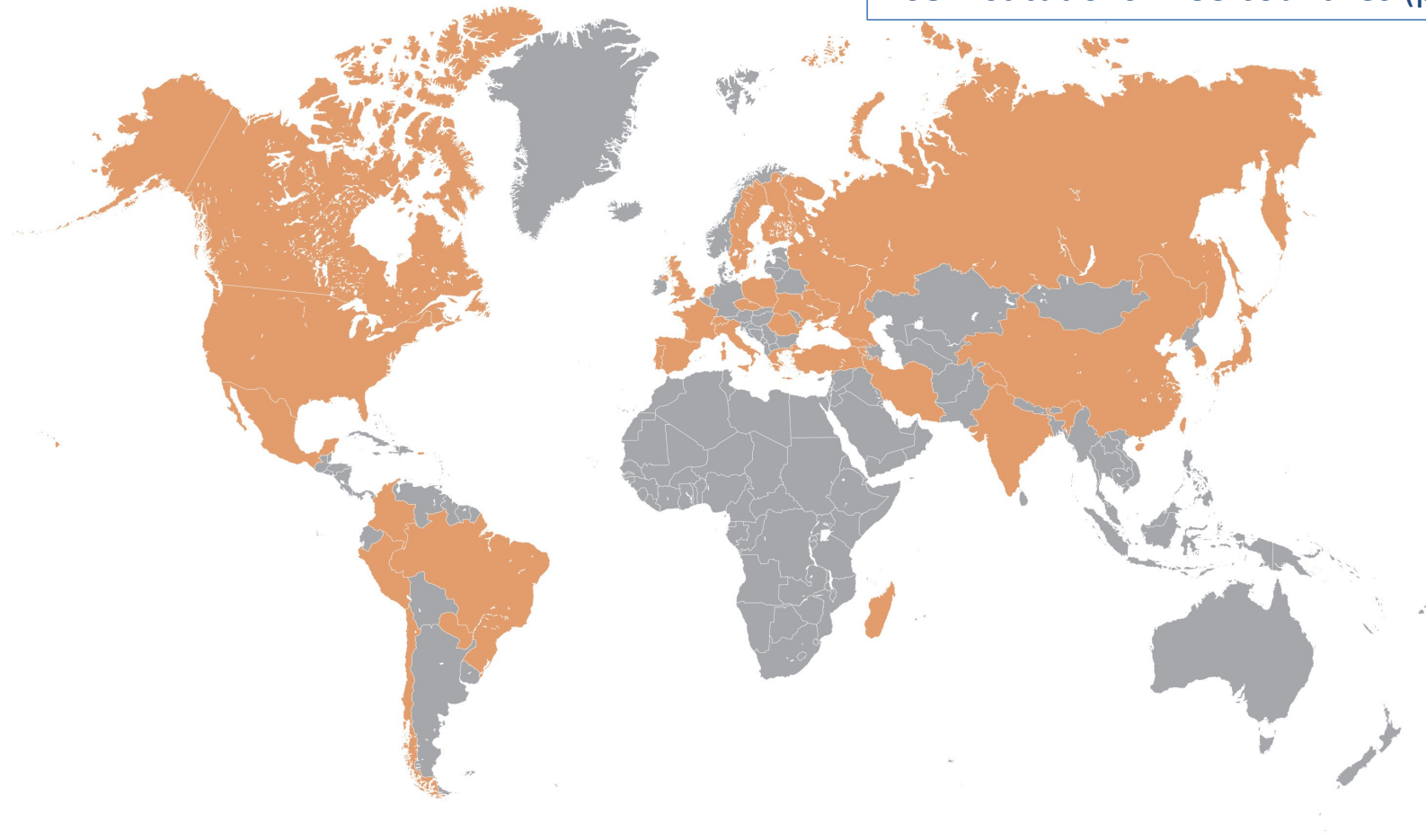
DUNE in a Nutshell



1. A high-power, wide-band **neutrino beam** (\sim GeV energy range).
2. A \approx 70 kt **liquid-argon Far Detector** in South Dakota, located 1478 m underground in a former gold mine.
3. A **Near Detector** located approximately 575 m from the neutrino source at Fermilab.

DUNE – a global collaboration

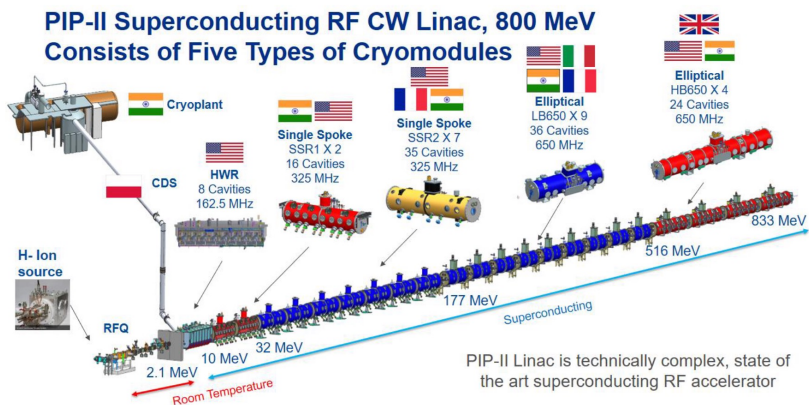
1317 collaborators from
208 institutions in 33 countries (plus CERN)



Proton Improvement Plan (PIP-II)



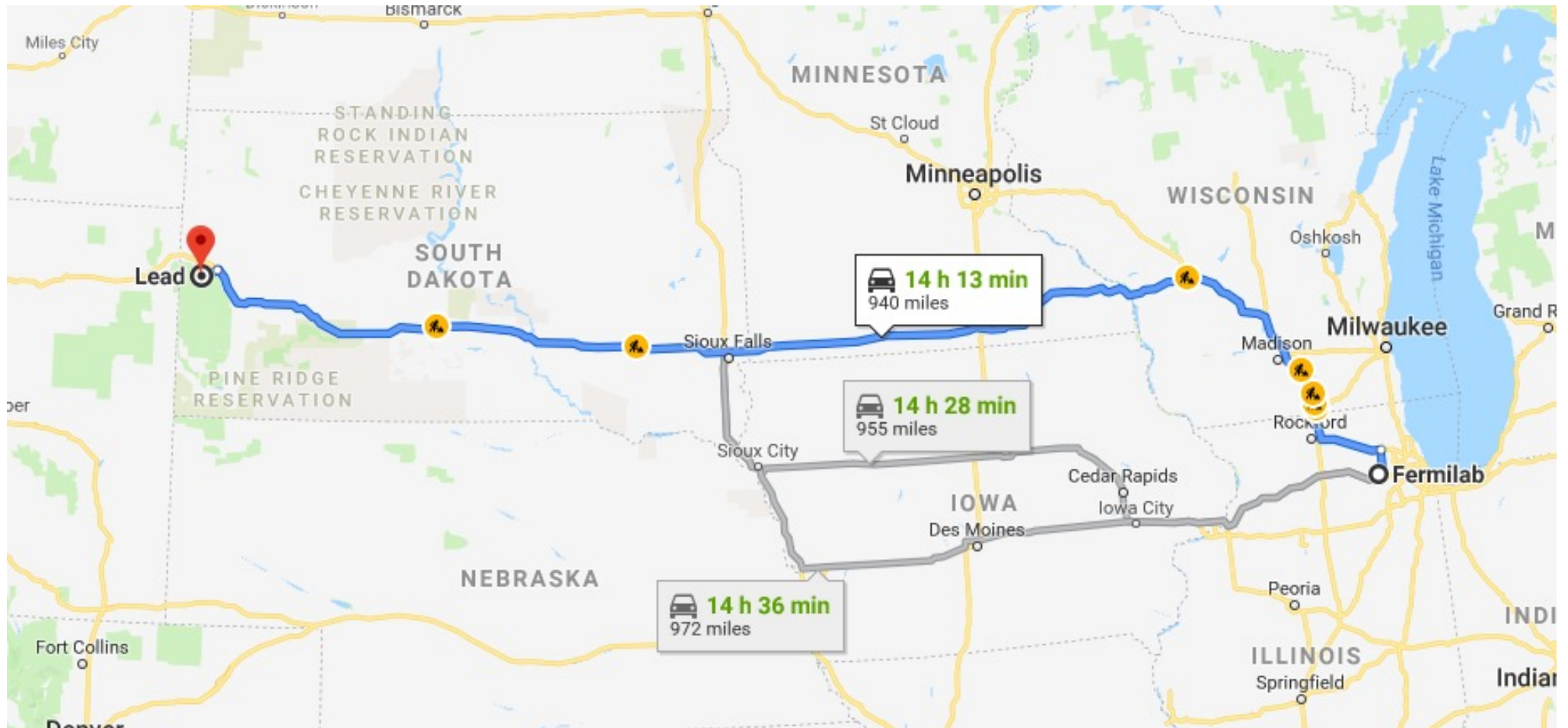
PIP-II Superconducting RF CW Linac, 800 MeV
Consists of Five Types of Cryomodules



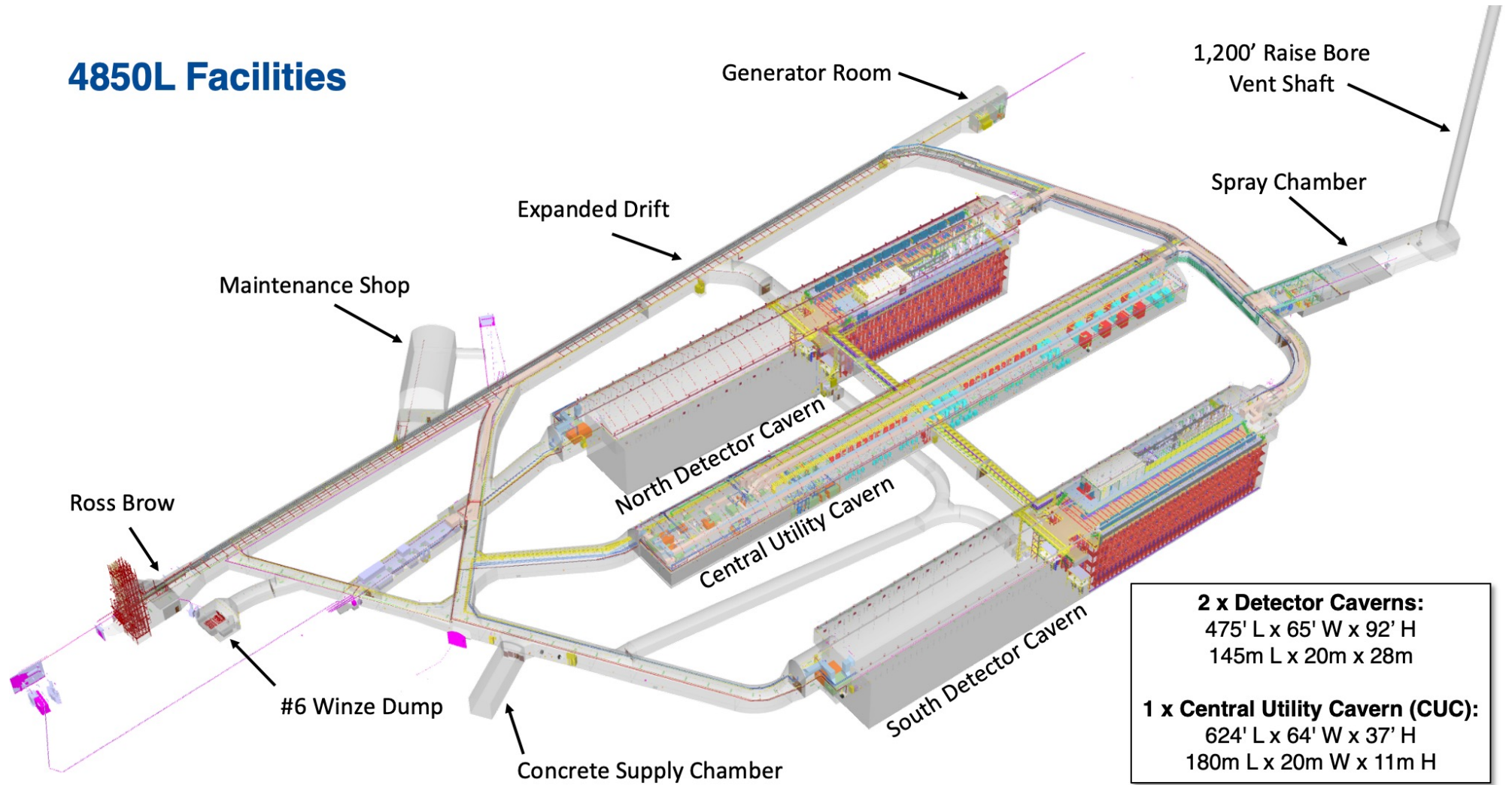
PIP-II Linac is technically complex, state of the art superconducting RF accelerator



$$L/E = 500 \text{ km/GeV} \Rightarrow L = 1300 \text{ km}$$



4850L Facilities



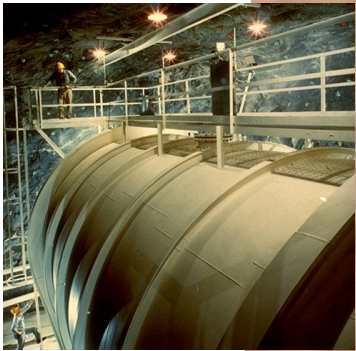
12 Feb 2022

Central Utility Cavern Pilot Drift Breakthrough



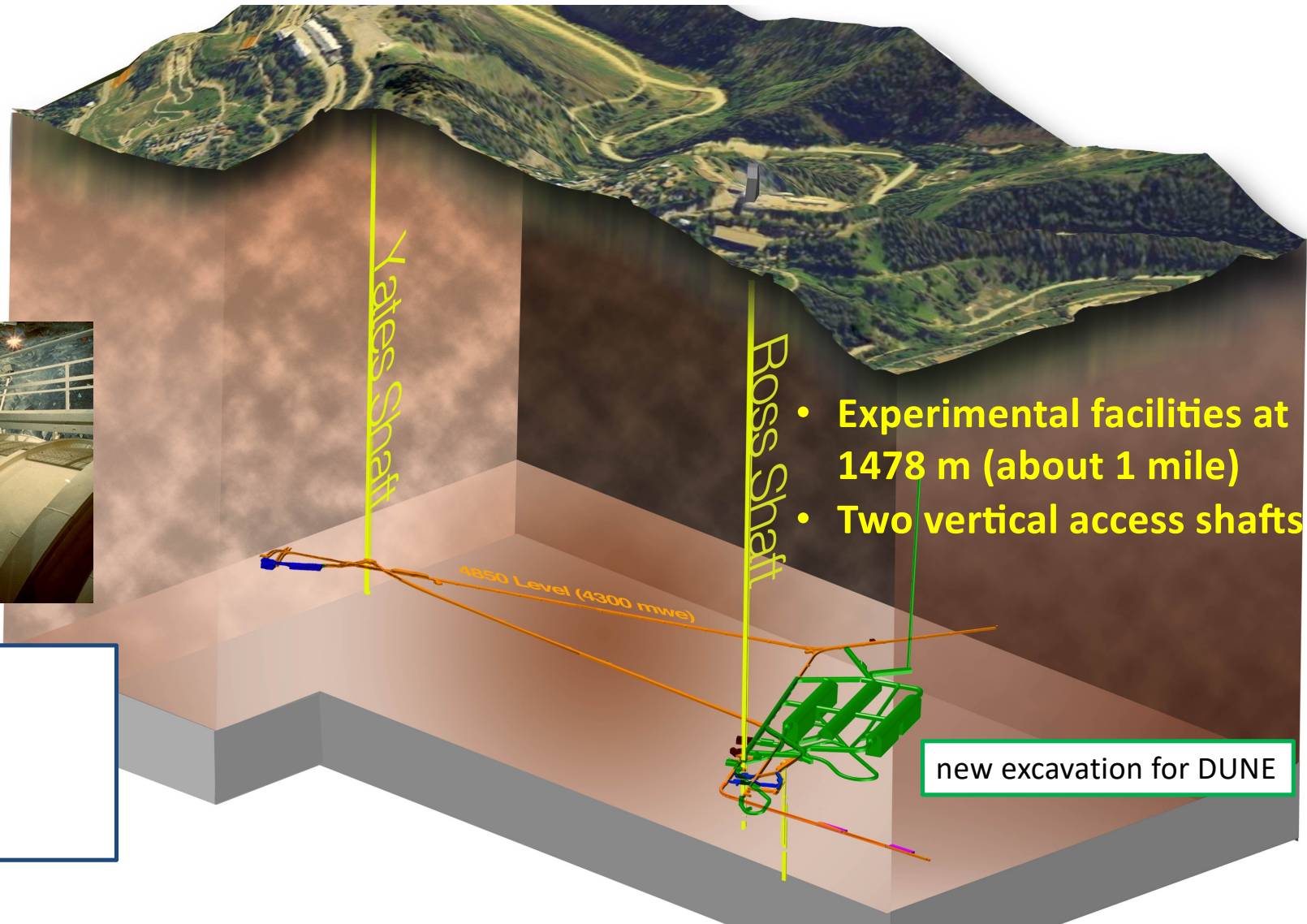
Sanford Underground Research Facility (SURF)





Davis Campus:

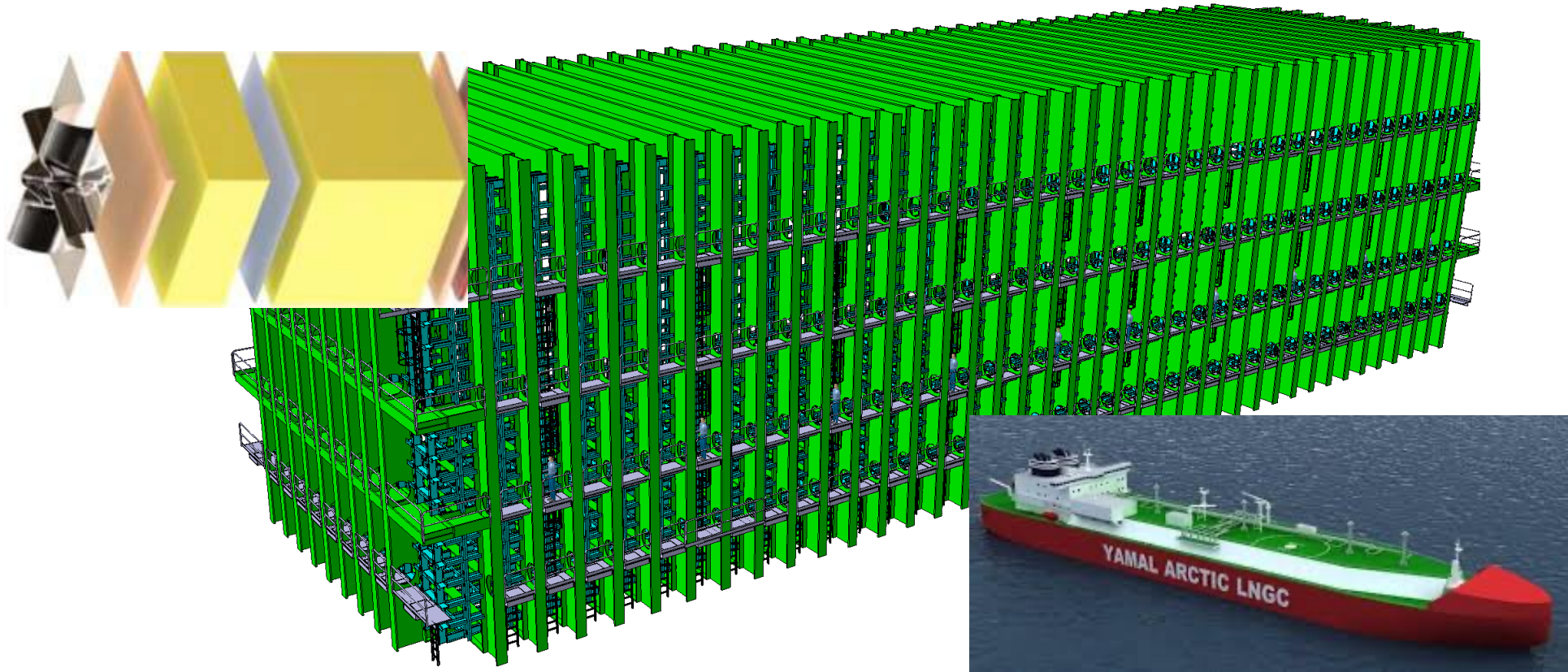
- LUX
- Majorana
- ...
- LZ





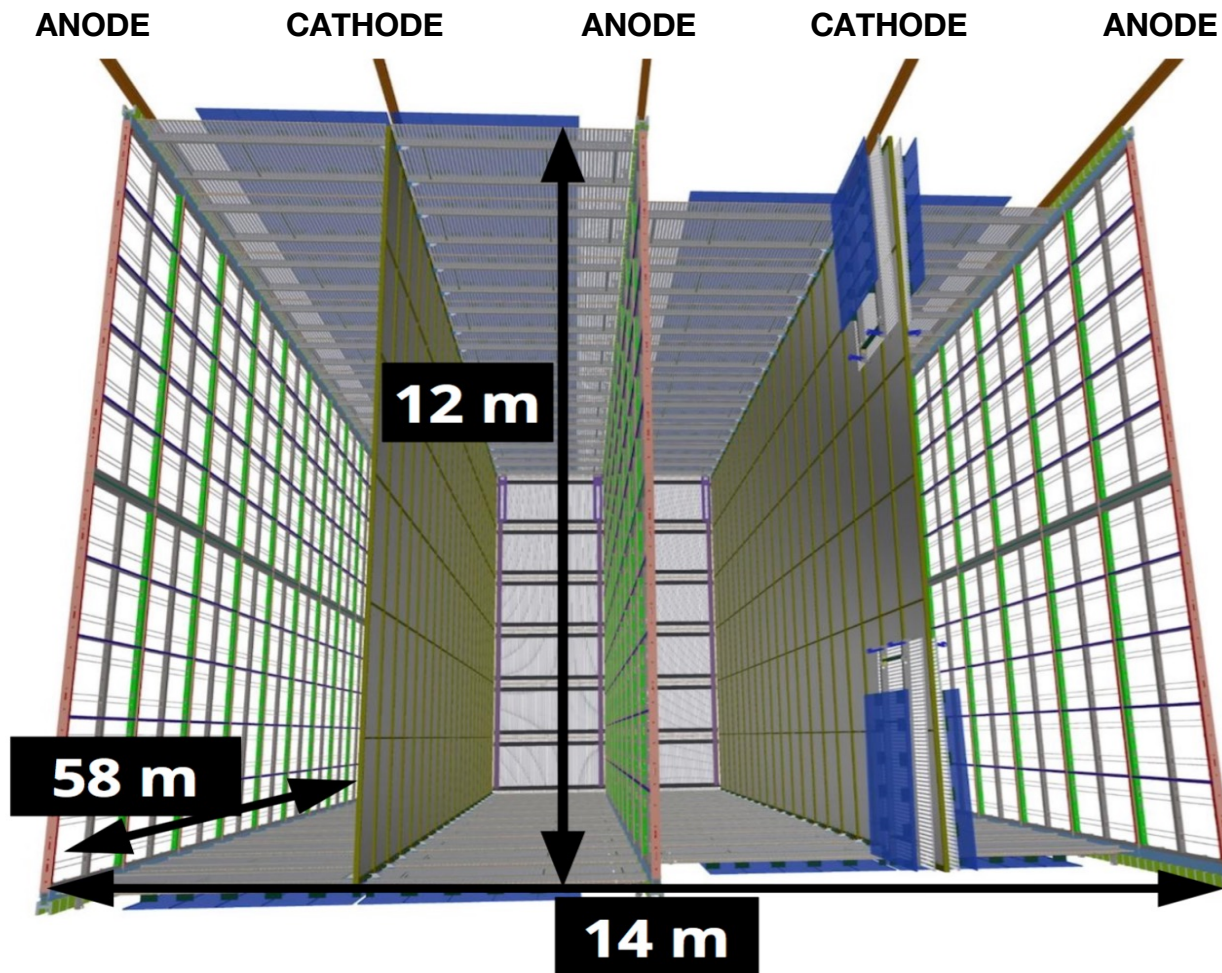
Four cryostats filled with liquid argon

Each of the four cryostats contains 17,000 tons of liquid argon at 89 K (-184°C or -299°F)

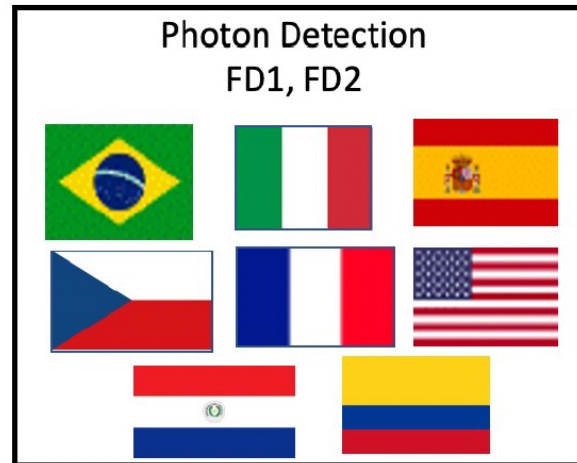


External Dimensions: 19 m x 18 m x 66 m

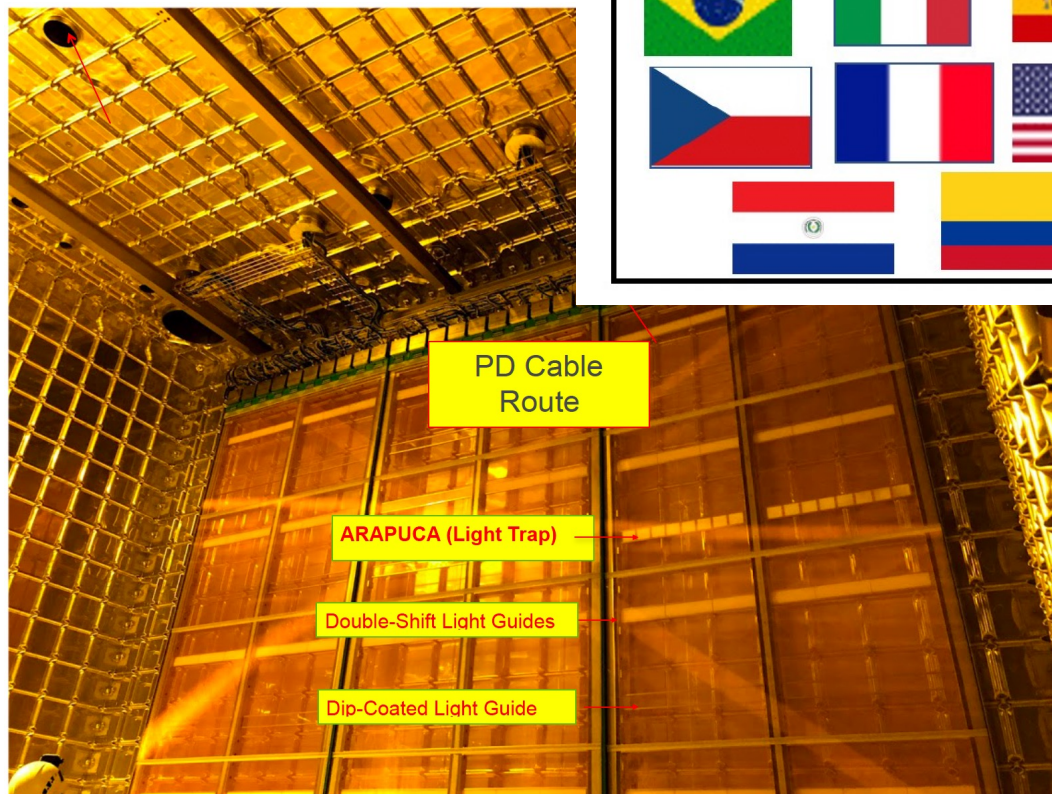
Horizontal Drift Detector (FD Module 1)

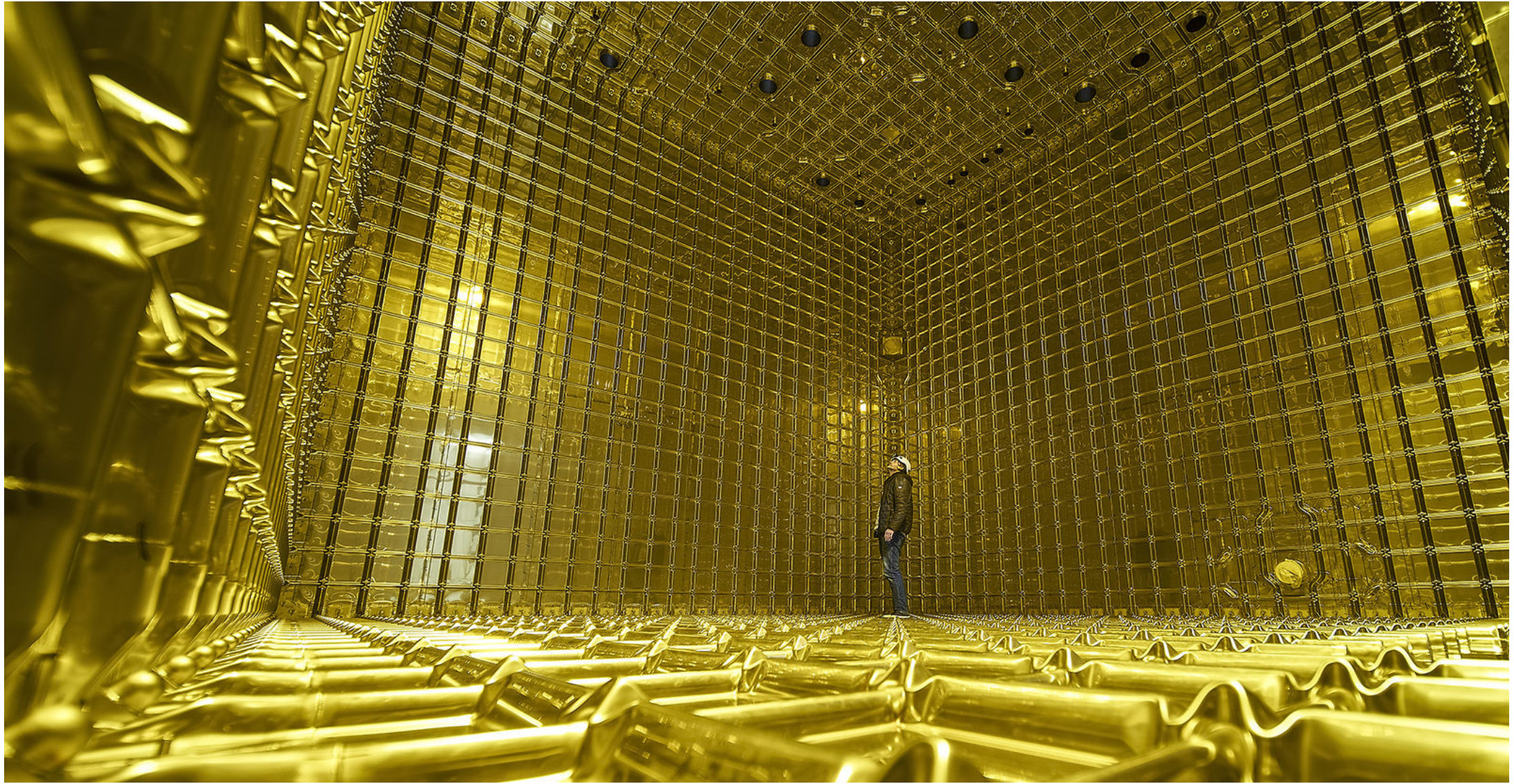


Photon Detection



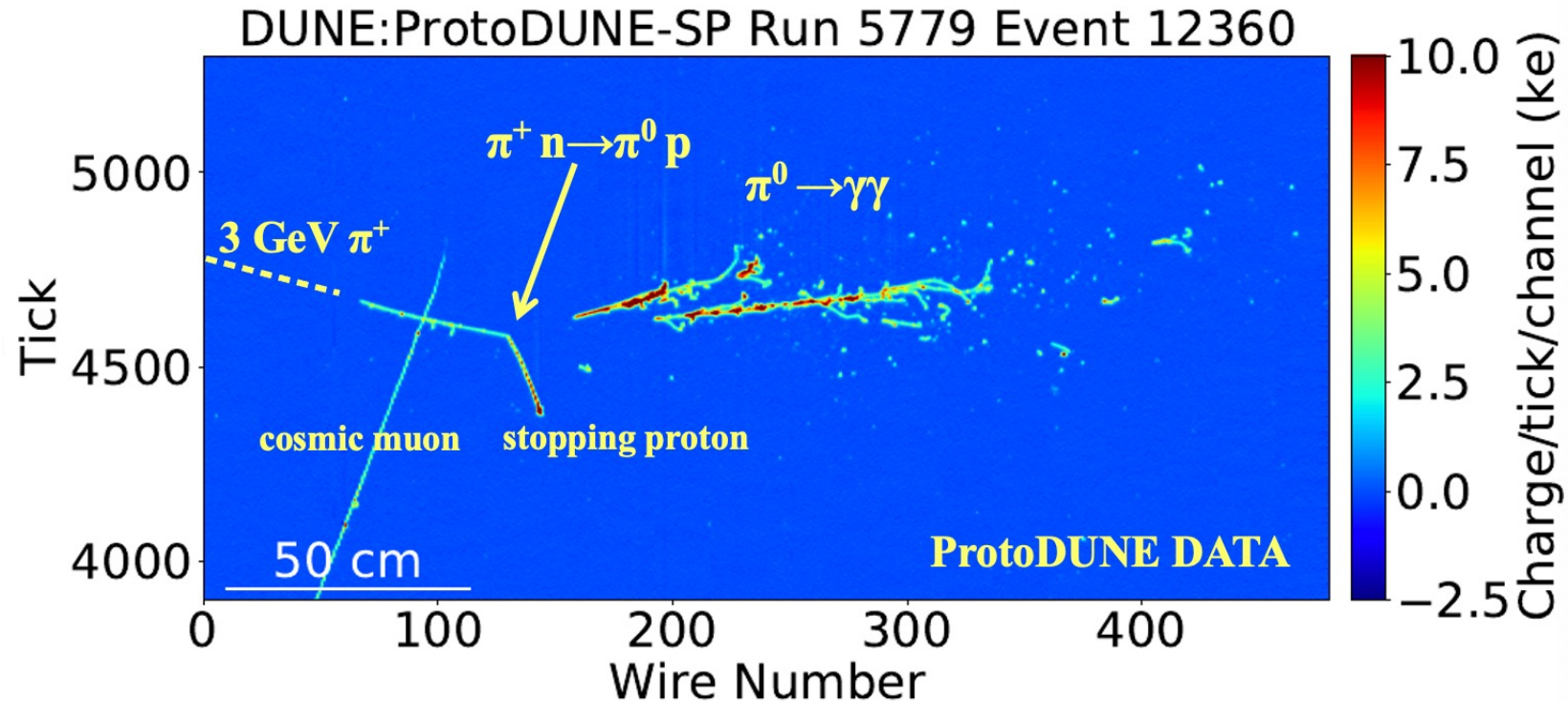
- Argon scintillation light is very abundant (40k photons/MeV)
- Arapuca technology with SiPMs for read-out
- Provides timing and event reconstruction
- Complementary to charge readout





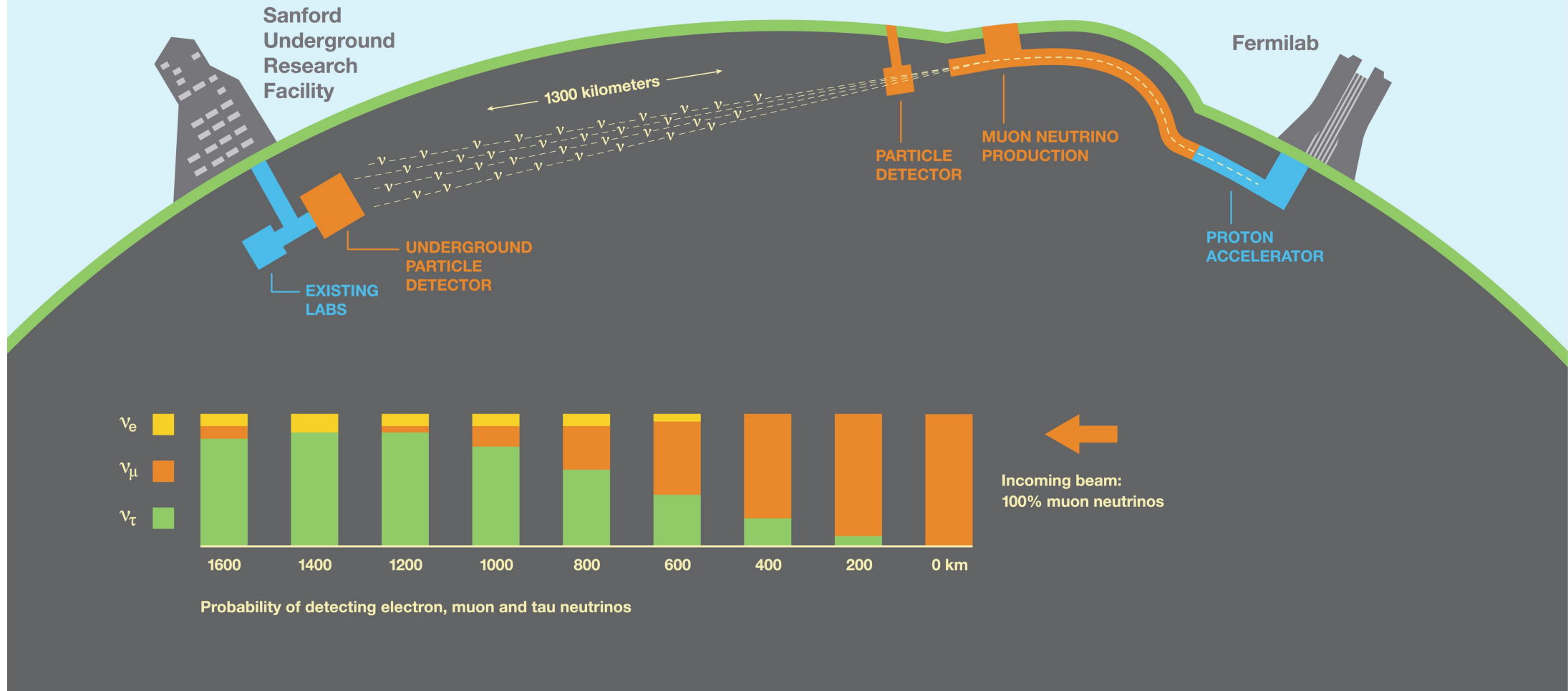


A ProtoDUNE-HD Data Event

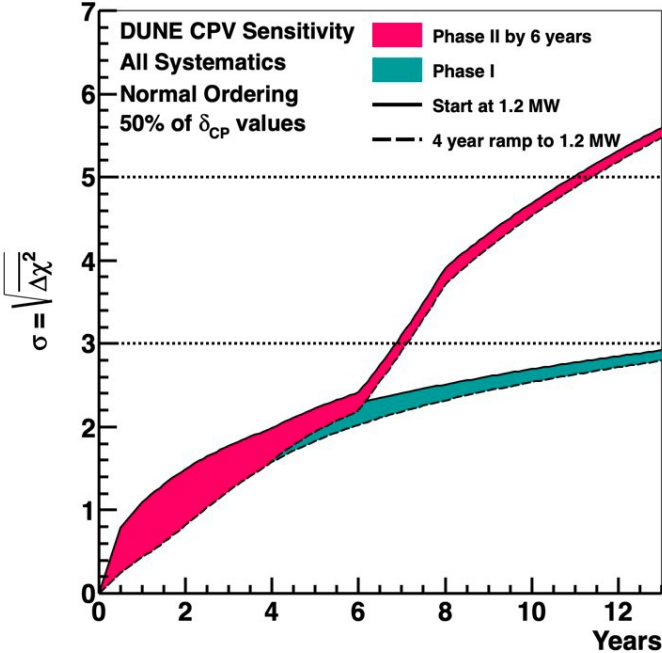


Reconstruction of events performed by PANDORA framework with the use of Grid computing resources, both areas UK-led.

Deep Underground Neutrino Experiment

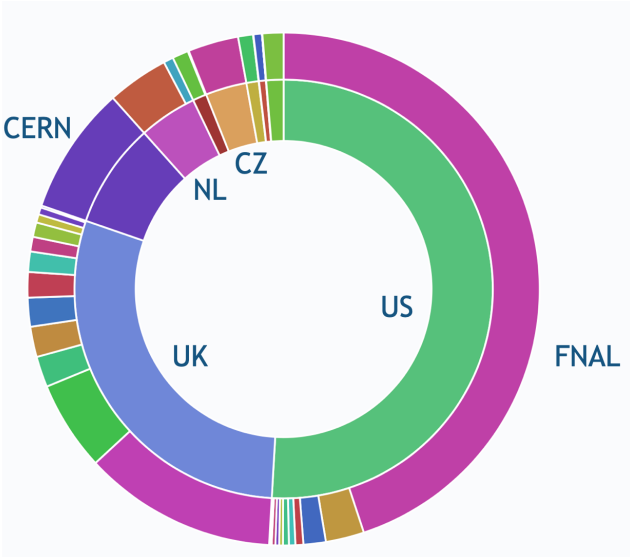


DUNE: Sensitivity to CP Violation



Phase II of DUNE will add Modules 3 and 4

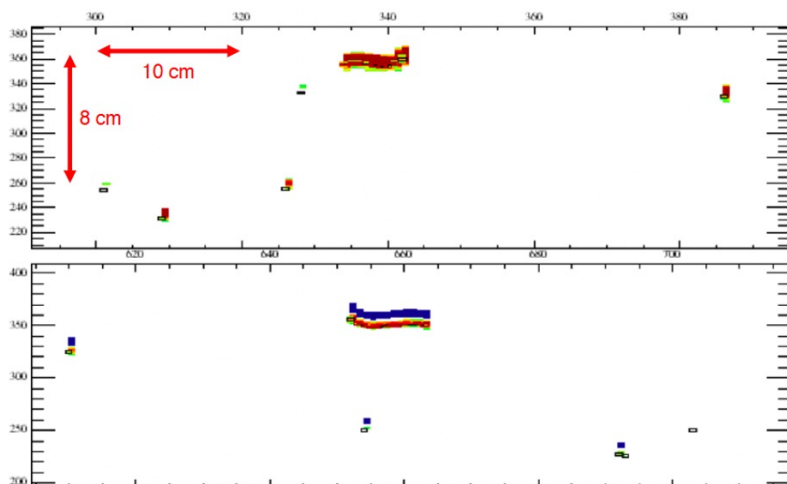
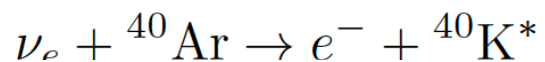
DUNE production jobs 2021



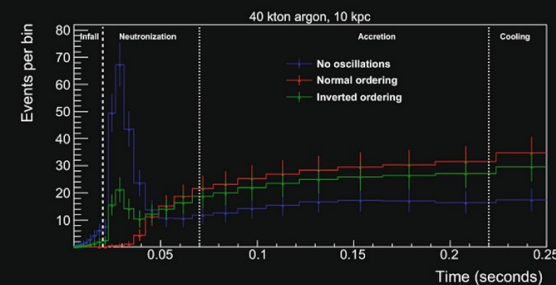
Czech institutes are major contributor to computing for NOvA and DUNE.

Neutrino Observatories

- Rich non-accelerator physics programme studying neutrinos from a supernova, solar, atmospheric neutrinos..
- Measurement at early times tests mass ordering and Detector requirements different from beam physics.



CC, 20 MeV



Expected event rates as a function of time for the electron-capture model for 40 kton of argon during early stages of the event – the neutronization burst and early accretion phases, for which self-induced effects are unlikely to be important. Shown are: the event rate for the unrealistic case of no flavor transitions (blue) and the event rates including the effect of matter transitions for the normal (red) and inverted (green) orderings. Error bars are statistical, in unequal time bins.

From the DUNE Collaboration: Supernova neutrino burst detection with the Deep Underground Neutrino Experiment, Eur. Phys. J. C 81, 423 (2021)

Fermilab – from the Tevatron to DUNE

- Czech particle physicists have made (and continue to make!) important contributions to the science programme at Fermilab.
- The Tevatron's rich physics programme included the discovery of the top quark and world-leading measurements of the W boson mass. It also made important contributions to Higgs boson physics.
- NOvA and DUNE provide pioneering measurements at the intensity frontier, searching for CPV with neutrinos. DUNE will also be an observatory for supernova neutrinos.