30th conference on ultra-relativistic nucleus-nucleus collisions September 3-9, 2023 Houston, Texas, USA

# **Electromagnetic Probes**





# Raphaelle Bailhache Goethe University Frankfurt

## **Electromagnetic probes**



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### γ, γ<sup>\*</sup>

Emitted at all stages of the heavy-ion collisions with negligible final-state interactions contrary to hadronic probes !  $\rightarrow$  Undistorted information about the medium at the same of their emission





### **Electromagnetic probes Most interesting ones (Direct** $\gamma$ , $\gamma$ \*)



Electromagnetic probes





### **Electromagnetic probes** Most interesting ones (Direct $\gamma, \gamma^*$ )



### **Initial hard scattering**

- Test  $N_{\rm coll}$  scaling
- Constrain nuclear PDFs
- Candle for energy loss studies: γ-tagged jets..

### **Pre-equilibrium phase**

Mechanism of equilibration

### **Thermal radiation**

- Effective QGP temperature
- Constrain space-time evolution

### Chiral symmetry restoration with dileptons

- $\rho$  broadening
- Constrain mechanisms:

 $\rho$  –  $a_1$  mixing









Focus on soft (thermal) radiation

- Direct photons:  $low-p_T$
- Dileptons

Disclaimer: will not cover initial hard-scattering or UPC

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### Direct photons: low pt

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# **Direct photons**

Sources populate different  $p_{\rm T}$  ranges:

- Hard scattering: prompt photons
  - Direct production
  - Fragmentation photons
- **Pre-equilibrium**
- Thermal radiation from QGP and hot hadronic matter
- + possible jet-medium interaction



# **Direct photons**

Sources populate different  $p_{\rm T}$  ranges:

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  - Fragmentation photons
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### Thermal sources at low $p_{\rm T}$ :

inverse slope  $\propto$  effective fireball temperature  $T_{\rm eff}$ 

- Blueshifted due to radial flow
- Averaged over time
- $\rightarrow$  Need models to disentangle sources



## Direct photons: status before QM

**Direct**  $\gamma$  yield at low  $p_{\rm T}$  in A+A collisions above prompt hard-scattering  $\gamma$  expectation observed by:

- PHENIX with different methods at different energies  $\sqrt{s_{\rm NN}} = 39-200 \, {\rm GeV}$
- ALICE with different methods at  $\sqrt{s_{\rm NN}}$  = 2.76 TeV (ALICE results link)



PHENIX: Phys. Rev. C 107, 024914 (2023); arXiv:2203.17187 ALICE: Phys. Lett. B 754 (2016) 235-248; Ana Marin Hard Probes 2023

Carolina Arata #174 Vassu Doomra #655





## **Direct photons: status before QM**

**Extracted**  $T_{\text{eff}}$  from  $\gamma_{\text{non-prompt}} = \gamma_{\text{dir}} - \gamma_{\text{prompt}}^{\text{estimated}}$ 

- Increases with  $p_{\rm T}$  range used to fit
- Above deconfinement temperature
- No obvious variation of  $T_{\rm eff}$  with  $dN_{\rm ch}/dN_{\eta}|_{\eta=0}$ although do not exclude small increase

#### Carolina Arata #174 Vassu Doomra #655



ALICE: Phys. Lett. B 754 (2016) 235-248; Ana Marin Hard Probes 2023







## Interpretation of $T_{\rm eff}$

- Naive idea: higher  $p_{\rm T}$  , earlier emission, higher T
- But:
  - Bias due to radial flow:  $T_{\rm eff}^{\rm w/flow} > T_{\rm eff}^{\rm w/o \, flow}$ 
    - Locally: large for high  $p_{\rm T} \gamma_{\rm thermal}$  emitted at small T
    - Globally: integrated over space-time  $\rightarrow$  smaller
  - **Contributions from pre-equilibrium** (without well defined T) for  $p_{\rm T} \ge 2.5$ -3 GeV/c



#### **Jean-Francois Paquet #787**



 $T_{\rm min}$ 

C. Shen, U. W. Heinz, J-F Paquet, C. Gale, Phys. Rev. C 89 (2014) 044910 J-F Paquet, arXiv:2305.10669





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**Measure**  $v_2$  of direct  $\gamma$  in Au+Au collisions at  $\sqrt{s_{NN}} = 200$  GeV with 10 × larger data sample

 $Au - Au \sqrt{s_{NN}} = 200 \text{ GeV}$ 

$$v_2^{\rm dir} \sim v_2^{\gamma_{\rm decay}} \sim v_2^{\pi}$$

$$v_2^{\rm dir} \sim 0$$
  
at high  $p_{\rm T,ee}$ 

![](_page_11_Picture_10.jpeg)

![](_page_11_Picture_11.jpeg)

![](_page_12_Picture_0.jpeg)

Results in agreement with previous publication and with significant reduction of uncertainties

![](_page_12_Figure_2.jpeg)

Electromagnetic probes

Vassu Doomra #655

![](_page_12_Picture_7.jpeg)

![](_page_12_Picture_8.jpeg)

![](_page_12_Picture_9.jpeg)

## **Direct photon puzzle still there**

**Yield** 

![](_page_13_Figure_2.jpeg)

 $v_2^{\rm dir}$ 

#### Vassu Doomra #655

$$Au - Au \sqrt{s_{NN}} = 200 \text{ GeV}$$

### **Simultaneous description of** yield and $v_2$ challenging for theory

(Same trend at the LHC but ok within large uncertainties)

> Trigger some theoretical idea **J-A Sun Poster #236** but only showing  $v_2^{dir}$  !

![](_page_13_Picture_10.jpeg)

![](_page_13_Picture_11.jpeg)

![](_page_13_Figure_12.jpeg)

![](_page_13_Picture_13.jpeg)

![](_page_13_Picture_14.jpeg)

## News from ALICE: Pb – Pb 5.02 TeV

- Finalised  $\gamma_{dir}$  yield in central Pb—Pb collisions at 5.02 TeV using dielectrons
- Described by calculations including:
  - Prompt photons
  - Pre-equilibrium photons
  - Thermal photons

If anything: model tends to overestimate yield

Daiki Sekihata #171

![](_page_14_Figure_11.jpeg)

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# **RHIC and LHC energies**

### $dN_{\gamma dir}/dy$ at the LHC

consistent with universal scaling behaviour seen by PHENIX:

$$\frac{\mathrm{d}N_{\gamma_{\mathrm{dir}}}}{\mathrm{d}y} = C(p_{\mathrm{T}}) \times (\mathrm{d}N_{\mathrm{ch}}/\mathrm{d}\eta \mid_{\eta=0})^{\alpha}$$

with no obvious  $p_{\rm T}$  dependence of  $\alpha$  (not so trivial why)

![](_page_15_Picture_7.jpeg)

![](_page_15_Figure_9.jpeg)

![](_page_15_Figure_10.jpeg)

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# **RHIC and LHC energies**

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with no obvious  $p_{\rm T}$  dependence of  $\alpha$  (not so trivial why)

#### **But some issues remained opened:**

- Discrepancy between STAR and PHENIX results
- Same model (lines):
  - Underestimates PHENIX data from semi-peripheral to peripheral (increasing discrepancy)
- In agreement with LHC data (need to decrease uncertainties)

#### Daiki Sekihata #171

![](_page_16_Picture_12.jpeg)

![](_page_16_Figure_13.jpeg)

![](_page_16_Figure_14.jpeg)

![](_page_16_Picture_15.jpeg)

![](_page_16_Picture_16.jpeg)

# pp collisions at the LHC

**First measurement of direct photons** at low  $p_{\rm T}$  in small systems at the LHC

- Minimum bias pp collisions  $\sqrt{s} = 13$  TeV
  - Provide reference
  - Reproduced by both prompt only and lacksquareprompt + thermal radiation calculations
- High-multiplicity pp collisions  $\sqrt{s} = 13$  TeV
  - Search for onset of thermal radiation
  - Significant higher  $\gamma_{dir}$  yield
  - Call for predictions

![](_page_17_Figure_9.jpeg)

![](_page_17_Figure_14.jpeg)

### Minimum bias pp $\sqrt{s} = 13$ TeV

#### Minimum bias pp High-multiplicity pp

![](_page_17_Picture_17.jpeg)

![](_page_17_Picture_18.jpeg)

![](_page_17_Picture_19.jpeg)

Dileptons

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![](_page_18_Picture_4.jpeg)

### Virtual photons

![](_page_19_Figure_1.jpeg)

### Carry mass ( $m_{ee}$ ):

• Can serve as an approximate clock

### $\rightarrow$ Separate thermal radiation from different stages

### Schematic view of dielectron invariant mass spectrum

![](_page_19_Figure_8.jpeg)

### Virtual photons

![](_page_20_Figure_1.jpeg)

### Carry mass ( $m_{ee}$ ):

- Can serve as an approximate clock  $\rightarrow$  Separate thermal radiation from different stages
- Invariant mass not affected by radial flow
  - $\rightarrow$  Access to QGP properties without blue shift Inverse slope  $\rightarrow$  Access to early averaged temperature

### Schematic view of dielectron invariant mass spectrum

![](_page_20_Figure_9.jpeg)

<u>R. Rapp and H. Van Hees, Phys. Lett. B753 (2016) 586</u>

### Virtual photons

![](_page_21_Figure_1.jpeg)

### Carry mass ( $m_{ee}$ ):

- Can serve as an approximate clock
  - $\rightarrow$  Separate thermal radiation from different stages
- Invariant mass not affected by radial flow
  - $\rightarrow$  Access to QGP properties without blue shift Inverse slope  $\rightarrow$  Access to early averaged temperature
- Radiation from hot-hadronic matter Sensitive to in-medium spectral function of  $\rho$  meson Related to chiral symmetry restoration, lifetime of the fireball

### Schematic view of dielectron invariant mass spectrum

![](_page_21_Figure_11.jpeg)

![](_page_21_Picture_12.jpeg)

## **Results over a wide range of energies**

### **Probe the phase diagram** from (high $\mu_{\rm B}/{ m low}~T$ ) to small (low $\mu_{\rm B}/{ m high}~T$ )

New results from:

- HADES: Ag-Ag  $\sqrt{s_{\rm NN}}$  = 2.42 and 2.55 GeV + pp reference for Ag-Ag  $\sqrt{s_{\rm NN}}$  = 2.55 GeV
- STAR:

Au-Au  $\sqrt{s_{\rm NN}}$  = 7.7, 14.6, 19.6 GeV (BES-II COL)

• ALICE:

$$Pb-Pb\sqrt{s_{NN}} = 5.02 \text{ TeV}$$

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![](_page_22_Figure_10.jpeg)

Odyniec, G. (2022) In: Blaschke, D., Redlich, K., Sasaki, C., Turko, L. Lecture Notes in Physics, vol 999. Springer, Charm, ISBM 978-3-030-95490-1

![](_page_22_Picture_13.jpeg)

![](_page_22_Picture_14.jpeg)

# HADES: Ag – Ag at very high $\mu_{\rm R}$

- Clear excess of  $e^+e^-$  pairs over:
  - Cocktail of hadronic decays at freeze-out
  - + initial NN contributions pp analysis for 2.55 GeV ongoing

**Karina Scharmann Poster #201** 

#### **Niklas Schild Poster #683**

 $\mu_{
m B}pprox\,$  800 MeV

Ag-Ag  $\sqrt{s_{\rm NN}} = 2.42$ , **2.55** GeV

![](_page_23_Figure_10.jpeg)

Au-Au at  $\sqrt{s_{NN}} = 2.42 \text{ GeV}$ : <u>HADES, Nature Phys, 15 (2019) 10, 1040</u>  $\pi^- p \rightarrow ne^+ e^-$  at  $\sqrt{s_{\pi^- p}} = 1.49 \text{ GeV}$ : <u>HADES, arXiv:2205.15914</u> 24

![](_page_23_Picture_14.jpeg)

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**Karina Scharmann Poster #201** 

#### **Niklas Schild Poster #683**

![](_page_24_Figure_8.jpeg)

Ag-Ag  $\sqrt{s_{\rm NN}} = 2.42$ , **2.55** GeV

![](_page_24_Figure_10.jpeg)

Au-Au at  $\sqrt{s_{NN}} = 2.42 \text{ GeV}$ : <u>HADES, Nature Phys, 15 (2019) 10, 1040</u>  $\pi^- p \rightarrow ne^+ e^-$  at  $\sqrt{s_{\pi^- p}} = 1.49 \text{ GeV}$ : <u>HADES, arXiv:2205.15914</u> 25

![](_page_24_Picture_14.jpeg)

# HADES: Ag – Ag at very high $\mu_R$

- Clear excess of  $e^+e^-$  pairs over:
  - Cocktail of hadronic decays at freeze-out
  - + initial NN contributions pp analysis for 2.55 GeV ongoing

**Karina Scharmann Poster #201** 

•  $v_2$  of e<sup>+</sup>e<sup>-</sup> pairs  $\approx$  0 in excess region

At SIS18 energies,  $v_2 < 0$  for  $\pi^{0,\pm}$  due to spectator shadowing Differential studies as a function of mass,  $p_{\rm T}$ , y and centrality Comparison with transport model **Renan Hirayama Poster #133** 

 $\rightarrow$  Confirm penetrating nature

![](_page_25_Figure_10.jpeg)

![](_page_25_Figure_11.jpeg)

Au-Au at  $\sqrt{s_{NN}} = 2.42 \text{ GeV}$ : <u>HADES, Nature Phys, 15 (2019) 10, 1040</u>  $\pi^- p \rightarrow ne^+ e^- \text{ at } \sqrt{s_{\pi^- p}} = 1.49 \text{ GeV}$ : <u>HADES, arXiv:2205.15914</u> 26

![](_page_25_Figure_14.jpeg)

![](_page_25_Picture_15.jpeg)

# HADES: Ag – Ag at very high $\mu_R$

### **Excess = Data - Cocktail - measured initial NN contributions**

has an exponential shape:

$$\frac{\mathrm{d}N}{\mathrm{d}M_{\mathrm{ee}}} \propto M_{\mathrm{ee}}^{3/2} \exp(-\frac{M_{\mathrm{ee}}}{T_{\mathrm{fit}}})$$

 $\rightarrow$  Extract integrated fireball temperature  $T_{\rm fit} \approx$  70 MeV for  $0.2 < M_{\rm ee} < 0.9 ~{\rm GeV/c^2}$ now differentially in centrality classes

#### Excess reproduced by hadronic thermal rates folded with coarse-grained medium evolution from transport

#### **Niklas Schild Poster #683**

![](_page_26_Picture_9.jpeg)

$$\mu_{
m B}pprox$$
 800

Ag-Ag  $\sqrt{s_{\rm NN}}$  = **2.42**, 2.55 GeV

![](_page_26_Figure_12.jpeg)

![](_page_26_Picture_14.jpeg)

### ) MeV

# **STAR:** Au – Au at intermediate $\mu_{\rm B}$

observed over cocktail of hadronic decays at freeze-out (without  $\rho$ ) + Drell-Yan

![](_page_27_Figure_3.jpeg)

Electromagnetic probes

Yiding Han #301 **Chenliang Jin Poster #683** 

![](_page_27_Picture_7.jpeg)

**Excess of**  $e^+e^-$  pairs at low  $m_{e^+e^-}$ 

![](_page_27_Picture_9.jpeg)

![](_page_27_Picture_10.jpeg)

![](_page_27_Picture_11.jpeg)

# **STAR:** Au – Au at intermediate $\mu_{\rm R}$

observed over cocktail of hadronic decays at freeze-out (without  $\rho$ ) + Drell-Yan

![](_page_28_Figure_3.jpeg)

Electromagnetic probes

Yiding Han #301 **Chenliang Jin Poster #683** 

 $\mu_{
m B}pprox\,$  200–500 MeV

**Excess of**  $e^+e^-$  pairs at low  $m_{e^+e^-}$ 

![](_page_28_Picture_10.jpeg)

![](_page_28_Picture_11.jpeg)

### **STAR: excess at low mass**

**Excess yield at low**  $m_{ee}$  / (d $N_{\pi^0}$ /dy)

• For  $\sqrt{s_{\rm NN}} \ge 17.3$  GeV: described (over larger mass range) by calculations including thermal production of  $\rho$ with in-medium broadening spectral function (+ QGP)

### In BES-II region (NEW):

Baryon density increases,  $T_{\rm ch}$  decreases  $\rightarrow$  Probe the role of baryons and temperature effects

Hint for a decrease with  $\sqrt{s_{\rm NN}}$ , need to reduce uncertainties (future experiments CBM, NA60+)

![](_page_29_Picture_8.jpeg)

#### Yiding Han #301 **Chenliang Jin Poster #683**

 $\mu_{
m B}pprox\,$  200–500 MeV

![](_page_29_Figure_11.jpeg)

STAR BES-I: PRC 107, L061901 (2023) R. Rapp and H. van Hees, Phys. Lett. B 753, 586 (2016)

![](_page_29_Figure_14.jpeg)

![](_page_29_Picture_15.jpeg)

# **STAR: higher mass**

**Excess / (** $dN_{ch}/d\eta$ **) at higher**  $\sqrt{s_{NN}}$  (> 19 GeV BES-I)

**Compared to new calculations for QGP thermal radiation** with:

- Production rates up to NLO at finite  $\mu_{\rm B}$
- Integrated over space-time with a realistic hydro
- $\rightarrow$  Quite good agreement with the data

**C. Gale #626** 

 $\mu_{
m B}pprox\,$  0–200 MeV

Au-Au 
$$\sqrt{s_{\rm NN}}$$
 = 19, 27, 39, 62.4, 200 G

![](_page_30_Figure_11.jpeg)

![](_page_30_Picture_14.jpeg)

![](_page_30_Picture_15.jpeg)

![](_page_30_Picture_16.jpeg)

![](_page_30_Picture_17.jpeg)

# **STAR: higher mass**

**Excess / (** $dN_{ch}/d\eta$ **) at higher \sqrt{s\_{NN}} (> 19 GeV BES-I**)

**Compared to new calculations for QGP thermal radiation** with:

- Production rates up to NLO at finite  $\mu_{\rm B}$
- Integrated over space-time with a realistic hydro

#### $\rightarrow$ Quite good agreement with the data

#### **Confirm theoretically:**

- Role of thermal QGP dileptons as thermometers
- Discriminating power of dilepton polarisation (e.g.  $\mu_{\rm R}$ )

No pre-equilibrium contributions yet in the calculations

 $\rightarrow$  Need more precise data

**C. Gale #626** 

 $\mu_{
m B}pprox\,$  0–200 MeV

Au – Au 
$$\sqrt{s_{\rm NN}}$$
 = 19, 27, 39, 62.4, 200 G

![](_page_31_Figure_16.jpeg)

![](_page_31_Picture_19.jpeg)

![](_page_31_Picture_20.jpeg)

![](_page_31_Picture_21.jpeg)

![](_page_31_Picture_22.jpeg)

# **ALICE:** Pb – Pb at $\mu_{\rm B} \approx 0$

### **Finalised dielectron yield**

compared to background cocktail from hadronic decays

- Large heavy-flavour (HF) background not easy to estimate:  $\rightarrow$  two versions:
  - Vacuum expectations (pp  $\times \langle N_{coll} \rangle$ )
  - Medium effects (measured  $R^{c,b \rightarrow e^{\pm}}_{\Delta \Delta}$ , EPS09 nPDF)
- No clear excess over the background  $\rightarrow$  Hint at low  $m_{ee}$  (0.18 <  $m_{ee}$  < 0.5 GeV/c<sup>2</sup>)

#### Daiki Sekihata #171

![](_page_32_Picture_10.jpeg)

 $Pb-Pb\sqrt{s_{NN}} = 5.02 \text{ TeV}$ 

#### ALICE, arXiv:2308.16704

![](_page_32_Figure_13.jpeg)

![](_page_32_Picture_15.jpeg)

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![](_page_33_Picture_10.jpeg)

 $Pb-Pb\sqrt{s_{NN}} = 5.02 \text{ TeV}$ 

#### ALICE, arXiv:2308.16704

![](_page_33_Figure_13.jpeg)

![](_page_33_Picture_15.jpeg)

### ALICE: excess at low mass

### Excess = Dielectron yield - background cocktail (w/o $\rho$ )

Compared to calculations for thermal radiation

- from hadronic phase (in-medium  $\rho$ )
- from QGP

Tension in  $0.5 < m_{\rm ee} < 0.8~{\rm GeV/c^2}$  (~ 3 $\sigma$ )

### $\rightarrow$ Need to reduce uncertainties

#### Daiki Sekihata #171

 $\mu_{
m B}pprox$  O

Pb-Pb
$$\sqrt{s_{\rm NN}}$$
 = 5.02 TeV

#### ALICE, arXiv:2308.16704

![](_page_34_Figure_13.jpeg)

![](_page_34_Figure_14.jpeg)

35

![](_page_35_Picture_0.jpeg)

### Much more data (up to a factor 100 for Pb-Pb)

Daiki Sekihata #171

ALICE: Jerome Jung Poster #665 PHENIX: Roli Esha Poster #217

![](_page_35_Figure_6.jpeg)

![](_page_35_Picture_7.jpeg)

![](_page_35_Picture_9.jpeg)

![](_page_35_Figure_10.jpeg)

![](_page_35_Picture_11.jpeg)

![](_page_35_Picture_12.jpeg)

![](_page_36_Picture_0.jpeg)

![](_page_36_Figure_2.jpeg)

Key tool to handle the heavy-flavour background at the LHC !

ALICE: Jerome Jung Poster #665 PHENIX: Roli Esha Poster #217

![](_page_36_Figure_9.jpeg)

### **Future measurements**

**Different experiments at different accelerator facilities** 

- From high  $\mu_{\rm B}$  to vanishing  $\mu_{\rm B}$ HADES, STAR, CBM, NA60+, ALICE 2&3
- To answer open questions:
  - Mechanism of Chiral symmetry restoration ( $\rho a_1$  mixing)
  - Dilepton  $v_2$  (Input for direct photon puzzle)
  - Dilepton polarisation (discriminating variable)
  - Equilibrium mechanisms
  - QGP EoS ( $T_{eff}$ )
  - First order phase transition at large  $\mu_{
    m B}$  ( $T_{
    m eff}$  in IMR vs  $\sqrt{s_{
    m NN}}$ )
  - Electric conductivity

NA60+: Giacomo Alocco #761 **CBM: Claudia Hoehne #687** ALICE 3: Isabella Sanna #317 **Poster Sebastian Scheid #158** 

![](_page_37_Figure_14.jpeg)

Courtesy of Thomas Ullrich

![](_page_37_Picture_17.jpeg)

![](_page_37_Picture_18.jpeg)

![](_page_37_Picture_19.jpeg)

### **Future measurements**

**Different experiments at different accelerator facilities** 

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#### Azumi Sakai #721

- Mechanism of Chiral symmetry restoration ( $\rho a_1$  mixing)
- Dilepton  $v_2$  (Input for direct photon puzzle)
- Dilepton polarisation (discriminating variable) Florian Seck #741
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  m B}$  ( $T_{
  m eff}$  in IMR vs  $\sqrt{s_{
  m NN}}$ )
- Electric conductivity Toru Nishimura #785

NA60+: Giacomo Alocco #761 **CBM: Claudia Hoehne #687** ALICE 3: Isabella Sanna #317 **Poster Sebastian Scheid #158** 

![](_page_38_Figure_15.jpeg)

Courtesy of Thomas Ullrich

![](_page_38_Picture_18.jpeg)

![](_page_38_Picture_19.jpeg)

![](_page_38_Picture_20.jpeg)

## Summary

- Direct photons: low- $p_{\rm T}$ 

  - Some issues remain opened (direct photon puzzle...)  $\rightarrow$  Common effort between theorists and experimentalists to solve them
- Dileptons
  - Measurements at very different  $\mu_{\rm B}/T$
  - Large potential of dilepton measurements shown by theory at this QM  $\bullet$
  - Huge experimental efforts to make such measurements possible.....starting now !

Many thanks to H. Appelshäuser, H. Busching, V. Doomra, R. Esha, T. Galatyuk, C. Hoehne, C. Gale, D. Gabor, Y. Han, J. Jung, A. Marin, T. Nishimura, J-F Paquet, P. Plaschke, R. Rapp, K. Reygers, A. Sakai, M. Sas, N. Schild, S. Scheid, F. Seck, D. Sekihata, J-a Sun

• Uncertainties of measurements improving in time at RHIC and at the LHC (effort still on going)

![](_page_39_Picture_15.jpeg)

40

![](_page_40_Picture_0.jpeg)

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![](_page_40_Picture_3.jpeg)

# **High-** $p_{T}$ **isolated photons**

### **Prompt photons** with fragmentation contribution suppressed

- Central Pb—Pb collisions
- $p_{\rm T} > 20$  GeV/c:  $R_{\rm AA} = 1$  $\rightarrow$  Verify  $N_{coll}$  scaling  $\rightarrow$  Calibrated reference for  $\gamma$ -h studies
- $p_{\rm T} < 20$  GeV/c: Cold nuclear matter effects expected May be overestimated by JETPHOX  $\rightarrow$  Constrain nPDFs

**Carolina Arata #174** 

![](_page_41_Picture_8.jpeg)

![](_page_41_Figure_9.jpeg)

Scale uncertainty  $p_{\tau}^{\gamma}/2 < \mu < 2p_{\tau}^{\gamma}$ 

![](_page_41_Picture_11.jpeg)

![](_page_41_Picture_12.jpeg)

# **High-** $p_{T}$ isolated photons

### **Prompt photons** with fragmentation contribution suppressed

Central Pb—Pb collisions

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- $p_{\rm T} < 20$  GeV/c: Cold nuclear matter effects expected May be overestimated by JETPHOX  $\rightarrow$  Constrain nPDFs

### Peripheral Pb—Pb collisions

•  $R_{AA} < 1$ : centrality selection bias of Glauber model Agreement with model by C.Loizides & A. Morsch C. Loizides & A. Morsch, arXiv:1705.08856 Use prompt photons as centrality estimators

(In peripheral AA and in small systems) **PHENIX Daniel Firak #654** 

Raphaelle Bailhache, University of Frankfurt

Electromagnetic probes

![](_page_42_Picture_10.jpeg)

**Carolina Arata #174** 

![](_page_42_Figure_12.jpeg)

![](_page_42_Picture_14.jpeg)

![](_page_42_Picture_15.jpeg)

## **Direct photons: status before QM**

**Direct**  $\gamma$  yield at low  $p_{\rm T}$  in A+A collisions above prompt hard-scattering  $\gamma$  expectation observed by:

- PHENIX with different methods at different energies  $\sqrt{s_{\rm NN}} = 39-200 \, {\rm GeV}$
- ALICE with different methods at  $\sqrt{s_{\rm NN}}$  = 2.76 TeV

Electromagnetic probes

 $Pb-Pb\sqrt{s_{NN}} = 2.76 \text{ TeV}$ 

![](_page_43_Figure_7.jpeg)

ALI-PREL-538511

PHENIX: Phys. Rev. C 107, 024914 (2023); arXiv:2203.17187 ALICE: Phys. Lett. B 754 (2016) 235-248; Ana Marin Hard Probes 2023

![](_page_43_Figure_10.jpeg)

![](_page_43_Picture_11.jpeg)

## **Direct photons: status before QM**

**Extracted**  $T_{\text{eff}}$  from  $\gamma_{\text{non-prompt}} = \gamma_{\text{dir}} - \gamma_{\text{prompt}}^{\text{estimated}}$ 

- increases with  $p_{\rm T}$  range used to fit
- Above deconfinement temperature
- No obvious variation of  $T_{\rm eff}$  with  $dN_{\rm ch}/dN_{\eta}|_{\eta=0}$ Although Do not exclude small increase

![](_page_44_Figure_7.jpeg)

PHENIX: Phys. Rev. C 107, 024914 (2023); arXiv:2203.17187 ALICE: Phys. Lett. B 754 (2016) 235-248; Ana Marin Hard Probes 2023

![](_page_44_Picture_10.jpeg)

### Interpretation of $T_{\rm eff}$

- Naive idea: higher  $p_{\rm T}$  , earlier emission, higher TAnalytic expression with simple symmetric hydro (Gubser) solutions

 $T_{\rm eff} \approx \frac{T_0}{1 + \frac{5}{2} \frac{T_0}{p}}$  Initial maximum *T* of plasma  $p_{\rm T}$  where  $T_{\rm eff}$  is fitted for  $\eta = 0$ 

- But bias due to radial flow:  $T_{\rm eff}^{\rm w/flow} > T_{\rm eff}^{\rm w/o\,flow}$ 
  - Locally: large for high  $p_{\rm T} \gamma_{\rm thermal}$  emitted at small T
  - Globally: integrated over space-time  $\rightarrow$  smaller

And further effects under study e.g.  $\gamma$  from pre-equilibrium contribution for  $p_{\rm T} \ge 2.5$  GeV/c

![](_page_45_Figure_7.jpeg)

#### **Jean-Francois Paquet #787**

![](_page_45_Figure_9.jpeg)

![](_page_45_Figure_10.jpeg)

C. Shen, U. W. Heinz, J-F Paquet, C. Gale, Phys. Rev. C 89 (2014) 044910 J-F Paquet, arXiv:2305.10669

![](_page_45_Picture_13.jpeg)

![](_page_45_Picture_14.jpeg)

![](_page_45_Picture_15.jpeg)

# **RHIC and LHC energies**

![](_page_46_Figure_1.jpeg)

Electromagnetic probes

#### Daiki Sekihata #171

![](_page_46_Figure_5.jpeg)

ALICE, arXiv:2308.16704

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# **RHIC and LHC energies**

### $dN_{\gamma dir}/dy$ at the LHC consistent with universal scaling behaviour seen by PHENIX:

$$\frac{\mathrm{d}N_{\gamma_{\mathrm{dir}}}}{\mathrm{d}y} = C(p_{\mathrm{T}}) \times (\mathrm{d}N_{\mathrm{ch}}/\mathrm{d}\eta \mid_{\eta=0})^{\alpha}$$

with no obvious  $p_{\rm T}$  dependence of  $\alpha$  (not so trivial why)

PHENIX Au—Au data at 200 GeV:  $\alpha = 1.11 \pm 0.02$  (stat)  $\pm_{0.08}^{0.09}$  (syst)

To be taken with care: LHC Pb—Pb data at 2.76 TeV and 5.02 TeV:  $\alpha = 1.07 \pm 0.25$  (syst + stat)

#### Daiki Sekihata #171

![](_page_47_Picture_9.jpeg)

![](_page_47_Figure_10.jpeg)

![](_page_47_Figure_12.jpeg)

![](_page_47_Picture_13.jpeg)

![](_page_47_Picture_14.jpeg)

### **STAR: excess BES-**

![](_page_48_Picture_1.jpeg)

![](_page_48_Figure_2.jpeg)

#### Yiding Han #301 **Chenliang Jin Poster #683**

Au-Au  $\sqrt{s_{\rm NN}}$  = 19, 27, 39, 62.4, 200 GeV

STAR BES-I: PRC 107, L061901 (2023) R. Rapp and H. van Hees, Phys. Lett. B 753, 586 (2016)

![](_page_48_Picture_9.jpeg)

![](_page_48_Picture_10.jpeg)

## **ALICE: higher mass**

#### **Dielectron yield in the intermediate mass**

 $(1.2 < m_{ee} < 2.6 \text{ GeV/c}^2)$ 

Predicted thermal contribution from QGP:

- Expanding fireball model
- Transport model

Small compared to heavy-flavour decay background

#### $\rightarrow$ Need an other approach

#### Daiki Sekihata #171

![](_page_49_Figure_11.jpeg)

Pb-Pb 
$$\sqrt{s_{\rm NN}}$$
 = 5.02 TeV

#### ALICE, arXiv:2308.16704

![](_page_49_Figure_14.jpeg)

![](_page_49_Picture_16.jpeg)

# **STAR: higher mass**

**Excess / (** $dN_{ch}/d\eta$ **) at a bit higher**  $\sqrt{s_{NN}}$  (> 19 GeV BES-I)

**Compared to new calculations for QGP thermal radiation** with:

- Production rates up to NLO at finite  $\mu_{\rm B}$
- Integrated over space-time with a realistic hydro
- $\rightarrow$  Quite good agreement with the data

#### **Confirm:**

- Role of thermal QGP dileptons as thermometers
- Discriminating power of dilepton polarisation (e.g.  $\mu_{\rm B}$ )

No pre-equilibrium contributions yet in the calculations

**C. Gale #626** 

 $\mu_{
m B}pprox\,$  0–200 MeV

Au-Au
$$\sqrt{s_{\rm NN}}$$
 = 19, 27, 39, 62.4, 200 G

![](_page_50_Figure_15.jpeg)

![](_page_50_Figure_16.jpeg)

STAR BES-I: PRC 107, L061901 (2023)

![](_page_50_Picture_19.jpeg)

![](_page_50_Figure_20.jpeg)

![](_page_50_Picture_21.jpeg)

![](_page_50_Picture_22.jpeg)

## **ALICE: finalised results in Pb—Pb**

Excess = Dielectron yield - background cocktail

- Significance of excess:  $1.8\sigma$  ( $1.5\sigma$ ) in  $0.18 < m_{ee} < 0.5$  GeV/ $c^2$
- Consistent with calculations for thermal radiation
  - from hadronic phase  $(\rho)$
  - from QGP

Parton-Hadron-String Dynamics (PHSD) model = transport model

ALICE, arXiv:XXX

$$Pb-Pb\sqrt{s_{NN}} = 5.02 \text{ TeV}$$

![](_page_51_Figure_12.jpeg)

![](_page_51_Picture_14.jpeg)

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## Feasibility studies: CBM dielectrons

After 3 years, 5 days/energy, 100 kHz IR

- Low mass ( < 1 GeV/c<sup>2</sup>) dominated by thermal rho, reconstructed with precision of 1.5-4.5%
- $\rightarrow$  Allow for fireball lifetime measurement
- Intermediate mass (> 1 GeV/ $c^2$ ): accessible, statistics not yet sufficient to extract physics

![](_page_52_Figure_5.jpeg)

### from partonic to hadronic fireballs

Raphaelle Bailhache, University of Frankfurt

Electromagnetic probes

Christian Pauly, Hard Probes 2023

![](_page_52_Picture_12.jpeg)

![](_page_52_Picture_13.jpeg)

![](_page_52_Picture_14.jpeg)

# The little big bang

#### QGP

#### **Initial state**

### hard scatterings, pre-equilibrium

**KoMPoST** 

#### pre-equilibrium Initial conditions dynamics

Relativistic viscous hydro (MUSIC), equation of state (Hadron resonance gas, lattice QCD) Electromagnetic probes

IP-Glasma, Trento ansatz quark suppression... Raphaelle Bailhache, University of Frankfurt

#### hadronic decays

![](_page_53_Picture_9.jpeg)

### hot hadronic phase

### Hydrodynamic phase

### Hadronic transport

SMASH...

![](_page_53_Picture_14.jpeg)