

# Darboux covariance: a hidden symmetry of perturbed Schwarzschild Black Holes

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June 7, 2022

M. Lenzi and C. F. Sopuerta, *Phys. Rev. D* **104**, 084053 (2021), M. Lenzi and C. F. Sopuerta, *Phys. Rev. D* **104**, 124068 (2021)

Einstein perturbed equations

$$\delta G_{\mu\nu} + \Lambda h_{\mu\nu} = 0$$

→

Master equation

$$(\square_2 - \Omega) \Psi(h, h', \dots) = 0$$

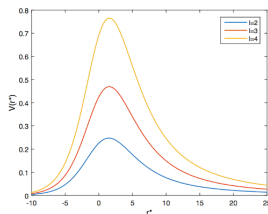
# Outline

- 1 Motivations
- 2 Perturbation theory of vacuum spherically symmetric spacetimes
- 3 Construction of master functions and equations
- 4 Darboux covariance and KdV hierarchy in Schwarzschild BH
- 5 Conclusions

# Motivations

# Motivations

- Perturbation theory in GR is a fundamental tool to describe physical systems and make predictions
- Scattering from BH
- Quasi-normal modes and gravitational waves (energy and angular momentum fluxes)



# Perturbation theory of vacuum spherically symmetric spacetimes

# Steps towards decoupling

- Perturbed Einstein equations at linear order

$$g_{\mu\nu} = \hat{g}_{\mu\nu} + h_{\mu\nu} \quad \longrightarrow \quad \begin{cases} \hat{G}_{\mu\nu} + \Lambda \hat{g}_{\mu\nu} = 0 \\ \delta G_{\mu\nu} + \Lambda h_{\mu\nu} = 0 \end{cases}$$

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- Metric splitting reflecting spherical symmetry

$$\hat{g}_{\mu\nu} = \begin{pmatrix} g_{ab} & 0 \\ 0 & r^2 \Omega_{AB} \end{pmatrix} \quad \longrightarrow \quad \begin{aligned} g_{ab} dx^a dx^b &= -f(r) dt^2 + dr^2/f(r) \\ \Omega_{AB} d\Theta^A d\Theta^B &= d\theta^2 + \sin^2 \theta d\varphi^2 \end{aligned}$$



# Steps towards decoupling

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- Harmonics expansion of perturbed metric

$$h_{\mu\nu} = \sum_{\ell, m} h_{\mu\nu}^{\ell m, \text{odd}} + h_{\mu\nu}^{\ell m, \text{even}}$$

# Harmonic decoupled Einstein equations

$$\delta G_{ab}^{\ell m}(x^c, \Theta^A) = \mathcal{E}_{ab}^{\ell m}(x^c) Y^{\ell m}(\Theta^A)$$

$$\delta G_{aA}^{\ell m}(x^b, \Theta^B) = \mathcal{E}_a^{\ell m}(x^b) Y_A^{\ell m}(\Theta^B) + \mathcal{O}_a^{\ell m}(x^b) X_A^{\ell m}(\Theta^B)$$

$$\delta G_{AB}^{\ell m}(x^a, \Theta^C) = \mathcal{E}_T^{\ell m}(x^a) T_{AB}^{\ell m}(\Theta^C) + \mathcal{E}_Y^{\ell m}(x^a) Y_{AB}^{\ell m}(\Theta^C) + \mathcal{O}_X^{\ell m}(x^a) X_{AB}^{\ell m}(\Theta^C)$$

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# Known master functions

## Odd parity

$$\Psi_{\text{RW}} = \frac{r^a}{r} \tilde{h}_a$$

$$\Psi_{\text{CPM}} = \frac{2r}{(\ell-1)(\ell+2)} \varepsilon^{ab} \left( \tilde{h}_{b:a} - \frac{2}{r} r_a \tilde{h}_b \right)$$

T. Regge and J. A. Wheeler, *Phys. Rev.* **108**, 1063–1069 (1957), C. T. Cunningham et al., *Astrophys. J.* **224**, 643–667 (1978)

## Known master functions

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## Even parity

$$\Psi_{\text{ZM}} = \frac{2r}{\ell(\ell+1)} \left\{ \tilde{K} + \frac{2}{\lambda} \left( r^a r^b \tilde{h}_{ab} - r r^a \tilde{K}_{:a} \right) \right\}$$

F. J. Zerilli, Phys. Rev. D **2**, 2141–2160 (1970), V. Moncrief, Ann. Phys. (N.Y.) **88**, 323 (1974)

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$$\tilde{h}_a = h_a - \frac{1}{2} h_{2:a} + \frac{r_a}{r} h_2,$$

$$\tilde{h}_{ab} = h_{ab} - \kappa_{a:b} - \kappa_{b:a},$$

$$\lambda(r) = (\ell+2)(\ell-1) - \Lambda r^2 - 3(f-1)$$

$$\tilde{K} = K + \frac{\ell(\ell+1)}{2} G - 2 \frac{r^a}{r} \kappa_a$$

# Known master equations

## Master equation

$$(\square_2 - \Omega_{\text{even/odd}}) \Psi_{\text{even/odd}} = 0$$



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$$(\square_2 - \Omega_{\text{even/odd}}) \Psi_{\text{even/odd}} = 0$$

- In Schwarzschild background

$$\Omega_{\text{odd}}(r) = \frac{\ell(\ell+1)}{r^2} - \frac{3r_s}{r^3}$$

$$\Omega_{\text{even}}(r) = \frac{1}{\lambda^2} \left[ \frac{(\ell-1)^2(\ell+2)^2}{r^2} \left( \ell(\ell+1) + \frac{3r_s}{r} \right) + \frac{9r_s^2}{r^4} \left( (\ell-1)(\ell+2) + \frac{r_s}{r} \right) \right]$$

- Tortoise coordinate  $dx/dr = 1/f(r)$

$$\left( -\frac{\partial^2}{\partial t^2} + \frac{\partial^2}{\partial x^2} - V_{\text{even/odd}} \right) \Psi_{\text{even/odd}} = 0, \quad V_{\text{even/odd}} = f \Omega_{\text{even/odd}}$$

Construction of master functions and equations

# Assumptions

- 1 Linear in the metric perturbations and first-order derivatives

$$\begin{aligned}\Psi_{\text{odd}}^{\ell m} &= C_0^\ell h_0^{\ell m} + C_1^\ell h_1^{\ell m} + C_2^\ell h_2^{\ell m} \\ &+ C_3^\ell \dot{h}_0^{\ell m} + C_4^\ell h_0'^{\ell m} + C_5^\ell \dot{h}_1^{\ell m} \\ &+ C_6^\ell h_1'^{\ell m} + C_7^\ell \dot{h}_2^{\ell m} + C_8^\ell h_2'^{\ell m}\end{aligned}$$

- 2 Time independent coefficients

$$C_i^\ell = C_i^\ell(r)$$

- 3 Satisfies a wave equation and the potential is left arbitrary

$$(\square_2 - \Omega) \Psi = 0$$

- 4 Arbitrary perturbative gauge

# Procedure

## Odd parity

$$K_1 \left[ \Omega^{\text{odd}}(r) - \Omega_*^{\text{odd}}(r) \right] = 0$$

- Standard branch

$$\Omega^{\text{odd}}(r) = \Omega_*^{\text{odd}}(r)$$

- Darboux branch

$$K_1 = 0$$

## Even parity

$$K_{13} \left[ \Omega^{\text{even}}(r) - \Omega_*^{\text{even}}(r) \right] = 0$$

- Standard branch

$$\Omega^{\text{even}}(r) = \Omega_*^{\text{even}}(r)$$

- Darboux branch

$$K_{13} = 0$$

# Odd parity sector: standard branch

$$(\square_2 - \Omega_{\text{odd}}) S \Psi_{\text{odd}} = 0$$

- A single potential selected  $\Omega^{\text{odd}}(r) = \Omega_*^{\text{odd}}(r)$

$$\Omega_*^{\text{odd}}(r) = \Lambda + \frac{\ell(\ell + 1) + 3(f - 1)}{r^2} \quad (= \Omega_{\text{RW}} \quad \text{in Sch})$$

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- Master function is a linear combination of RW and CPM

$${}_S\Psi_{\text{odd}}(t, r) = K_1 \Psi_{\text{RW}}(t, r) - \frac{(\ell+2)(\ell-1)}{4} K_0 \Psi_{\text{CPM}}(t, r)$$

$$t^a \Psi_{\text{CPM};a} = 2 \Psi_{\text{RW}}$$

# Odd parity sector: Darboux branch

$$(\square_2 - \Omega_{\text{odd}})_D \Psi_{\text{odd}} = 0$$

- Potential dependent master functions

$${}_D \Psi_{\text{odd}}(t, r) = -\frac{(\ell+2)(\ell-1)}{4} \left( K_0 + \hat{K}_7 \Xi(r) \right) \Psi_{\text{CPM}}(t, r) + K_7 \Phi_{\text{ON}}(t, r)$$

with

$$\hat{K}_7 = \frac{2K_7}{(\ell+2)(\ell-1)}$$

$$\Xi(r) = \Xi(r, \Omega(r))$$

$$\Phi_{\text{ON}}(t, r) = \varepsilon^{ab} \tilde{h}_{a:b}$$

# Odd parity sector: Darboux branch

$$(\square_2 - \Omega_{\text{odd}})_D \Psi_{\text{odd}} = 0$$

## ■ Family of potentials

$$\left( \frac{\delta V_{,x}}{\delta V} \right)_{,x} + 2 \left( \frac{V_{,x}^{\text{odd}}}{\delta V} \right)_{,x} - \delta V = 0$$

where

$$\delta V = V - V^{\text{odd}}$$

$$V^{\text{odd}} = f \Omega_*^{\text{odd}}$$



# Even parity sector: standard branch

$$(\square_2 - \Omega_{\text{even}}) S \Psi_{\text{even}} = 0$$

- A single potential selected  $\Omega^{\text{even}}(r) = \Omega_*^{\text{even}}(r)$

$$\begin{aligned} \Omega_*^{\text{even}} &= \frac{\lambda^3(r) - 2\Lambda r^2[\lambda(r) - (\ell + 2)(\ell - 1)]^2 + 2(\ell + 2)^2(\ell - 1)^2(\ell^2 + \ell + 1)}{3r^2\lambda^2(r)} \\ &= (\Omega_Z \quad \text{Sch}) \end{aligned}$$

## Even parity sector: standard branch

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- Master function is a linear combination of ZM and a new one

$${}_S \Psi_{\text{even}}(t, r) = \frac{\ell(\ell + 1)}{4} K_2 \Psi_{\text{ZM}}(t, r) - K_{13} \Psi_{\text{NE}}(t, r)$$

$$\Psi_{\text{NE}}(t, r) = \frac{1}{\lambda(r)} t^a \left( r \tilde{K}_{:a} - \tilde{h}_{ab} r^b \right)$$

$$t^a \Psi_{\text{ZM}:a} = 2 \Psi_{\text{NE}}$$

# Even parity sector: Darboux branch

$$(\square_2 - \Omega_{\text{even}}) {}_D\Psi_{\text{even}} = 0$$

- Potential dependent master functions

$${}_D\Psi_{\text{even}}(t, r) = (K_2 + K_{24} \Sigma(r)) \Psi_{\text{ZM}}(t, r) - \frac{K_{24}}{2} \tilde{K}(t, r)$$

with

$$\Sigma(r) = \Sigma(r, \Omega(r))$$

# Even parity sector: Darboux branch

$$(\square_2 - \Omega_{\text{even}})_D \Psi_{\text{even}} = 0$$

- Family of potentials

$$\left( \frac{\delta V_{,x}}{\delta V} \right)_{,x} + 2 \left( \frac{V_{,x}}{\delta V} \right)_{,x} - \delta V = 0$$

where

$$\delta V = V - V^{\text{even}}$$

$$V^{\text{even}} = f \Omega_*$$

# Darboux covariance and KdV hierarchy in Schwarzschild BH

# Full space of master equations

$$(-\partial_t^2 + L_V) \Psi = (-\partial_t^2 + \partial_x^2 - V) \Psi = 0$$

## Standard branch

$${}_S\Psi_{\text{odd}}^{\text{even}} = \mathcal{C}_1 \Psi_{\text{CPM}}^{\text{ZM}} + \mathcal{C}_2 \Psi_{\text{RW}}^{\text{NE}}$$

$$V = V_{\text{RW}}^{\text{Z}}$$

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## Darboux branch

$${}_D \Psi_{\text{odd}}^{\text{even}} = \mathcal{C}_1 \Psi_{\text{CPM}}^{\text{ZM}} + \mathcal{C}_2 \left( \Sigma_{\text{odd}}^{\text{even}} \Psi_{\text{CPM}}^{\text{ZM}} + \Phi_{\text{odd}}^{\text{even}} \right)$$

$$\left( \frac{\delta V_{,x}}{\delta V} \right)_{,x} + 2 \left( \frac{V_{,x}^{\text{even/odd}}}{\delta V} \right)_{,x} - \delta V = 0$$

# Standard view on DTs

- Isospectrality of  $V_{RW}$  and  $V_Z$

S. Chandrasekhar, Proc. Roy. Soc. Lond. A **369**, 425–433 (1980), K. Glampedakis et al., Phys. Rev. D **96**, 024036 (2017)



# Standard view on DTs

- Isospectrality of  $V_{RW}$  and  $V_Z$

DT in frequency domain  $\Psi(t, r) = e^{i\omega t} \psi(x; \omega)$

$$L_V \psi(x; \omega) = -\omega^2 \psi(x; \omega)$$

$$L_V \psi_o = -\omega_o^2 \psi_o \longrightarrow g(x) = -(\ln \psi_o)_{,x} \longrightarrow \begin{cases} L_v \phi = -\omega^2 \phi \\ \phi = W[\psi, \psi_o] / \psi_o \end{cases}$$

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- Darboux transformation between RW and ZM

$$\begin{aligned} \psi_o &= \frac{\lambda(r)}{2} e^{-i\omega_* x} \\ \omega_* &= -i\alpha = -i \frac{(\ell+2)!}{6M(\ell-1)!} \end{aligned} \longrightarrow \begin{aligned} G(x) &= \frac{6M f(r)}{\lambda(r) r^2} \\ V_{RW}^Z &= \pm G_{,x} + \alpha G + G^2 \end{aligned}$$

# Darboux transformation

- Darboux transformation between  $(v, \varphi)$  and  $(V, \Psi)$

$$(\square_2 - v) \varphi = 0 \quad \longrightarrow \quad \begin{cases} \Psi = \varphi_{,x} + g \varphi \\ V = v + 2g_{,x} \\ g_{,x} - g^2 + v = \mathcal{C} \end{cases} \quad \longrightarrow \quad (\square_2 - V) \Psi = 0$$

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$$\longrightarrow \quad \left( \frac{\delta V_{,x}}{\delta V} \right)_{,x} + 2 \left( \frac{v_{,x}}{\delta V} \right)_{,x} - \delta V = 0$$

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$$\rightarrow \quad \left( \frac{\delta V_{,x}}{\delta V} \right)_{,x} + 2 \left( \frac{v_{,x}}{\delta V} \right)_{,x} - \delta V = 0$$

- Darboux covariance of perturbations of spherically-symmetric BHs

$$g_{\text{odd}}^{\text{even}} = \frac{1}{2} \int dx \left( V - V_{\text{RW}}^Z \right)$$

$$g^{\text{RW} \rightarrow Z} = \frac{1}{2} \int dx \left( V_Z - V_{\text{RW}} \right)$$

# Inverse scattering methods in BH perturbation theory

- Scattering through a potential barrier

$$\psi \rightarrow \begin{cases} e^{i\omega x} & \text{for } x \rightarrow -\infty, \\ a(\omega)e^{i\omega x} + b(\omega)e^{-i\omega x} & \text{for } x \rightarrow +\infty, \end{cases}$$

# Inverse scattering methods in BH perturbation theory

- Scattering through a potential barrier

$$\psi \rightarrow \begin{cases} e^{i\omega x} & \text{for } x \rightarrow -\infty, \\ a(\omega)e^{i\omega x} + b(\omega)e^{-i\omega x} & \text{for } x \rightarrow +\infty, \end{cases}$$

- Bogoliubov coefficients completely determine the problem
  - Reflection and transmission coefficients

$$R(\omega) = \frac{b(\omega)}{a(\omega)}, \quad T(\omega) = \frac{1}{a(\omega)}$$

- QNMs  $\omega_i$  are the zeros of  $a(\omega)$

$$a(\omega_i) = 0$$

# KdV hierarchy in BH perturbation theory

- The logarithm of  $a(\omega)$  can be expanded in inverse powers

$$\ln[a(\omega)] = \sum_{n=0}^{\infty} \frac{I_n}{(2i\omega)^n} \quad \Longrightarrow \quad I_n = \int_{-\infty}^{+\infty} dx P_n(V, V', \dots)$$

- $I_n$  are the KdV integrals associated to the infinite symmetries of the KdV equation



# KdV hierarchy in BH perturbation theory

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The KdV integrals  $I_n$  are equal for all the Darboux related potentials

## Conclusions

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- Computational advances (QNMs, scattering data..)
- Extensions to modified theories of gravity, Kerr BH...

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Thank you!