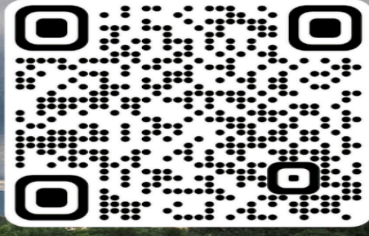



DAE HEP 2022
 HOSTED BY ISEK MOHALI
 DECEMBER 12TH TO 16TH

XXV DAE-BRNS HIGH ENERGY PHYSICS SYMPOSIUM 2022

TOPICS

- ASTROPARTICLE PHYSICS AND COSMOLOGY
- BEYOND THE STANDARD MODEL
- FORMAL THEORY
- FUTURE EXPERIMENTS AND DETECTOR DEVELOPMENT
- HEAVY IONS AND QCD
- HIGGS PHYSICS
- NEUTRINO PHYSICS
- QUARK AND LEPTON FLAVOUR PHYSICS
- SOCIETAL APPLICATIONS
- TOP QUARK AND EW PHYSICS

CME search in isobar (${}_{44}^{96}\text{Ru} + {}_{44}^{96}\text{Ru}$ and ${}_{40}^{96}\text{Zr} + {}_{40}^{96}\text{Zr}$) collisions at $\sqrt{s_{NN}} = 200$ GeV using SDM at RHIC

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Supported in part by



U.S. DEPARTMENT OF
ENERGY

Office of
Science

The STAR Collaboration

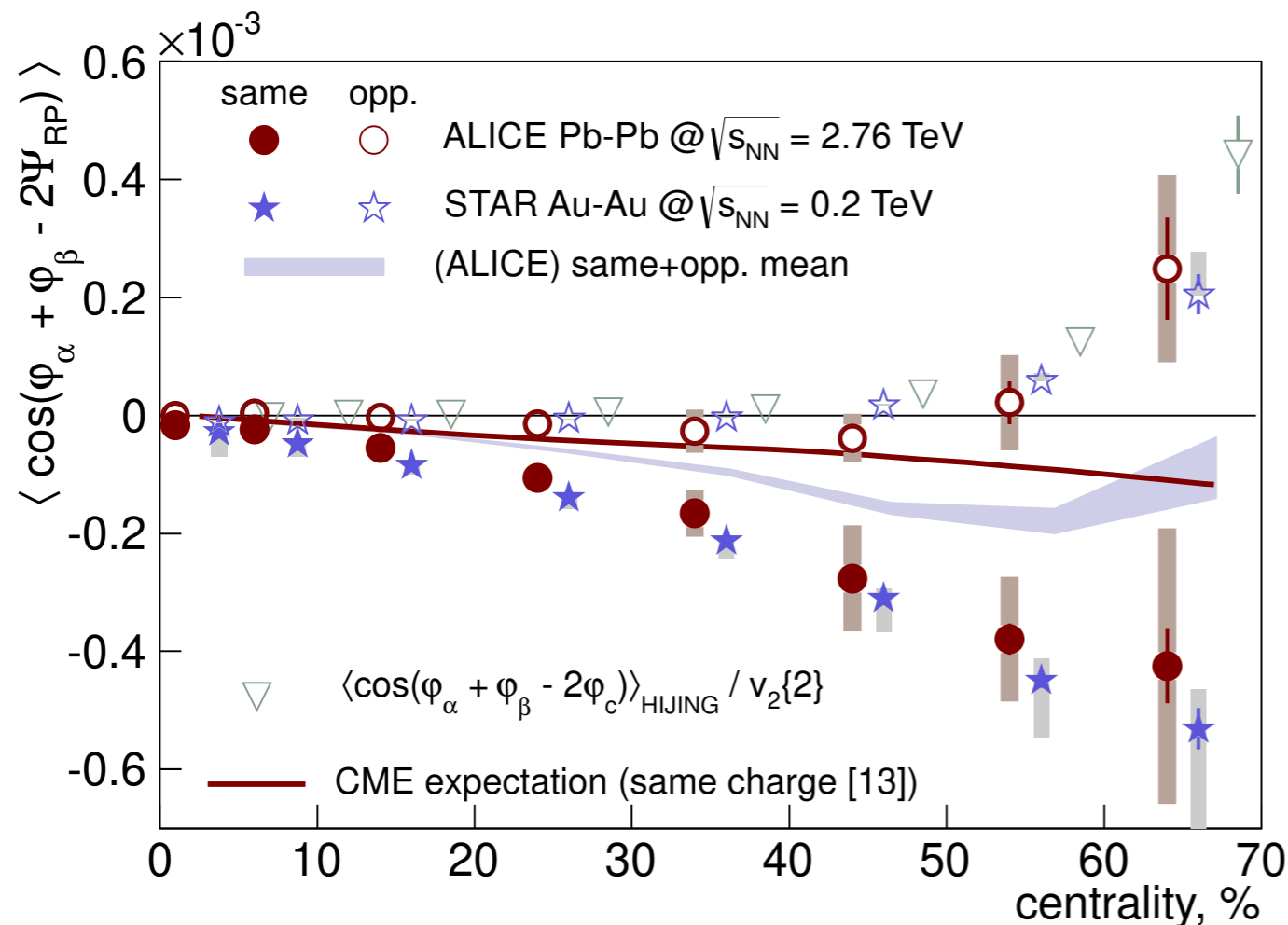
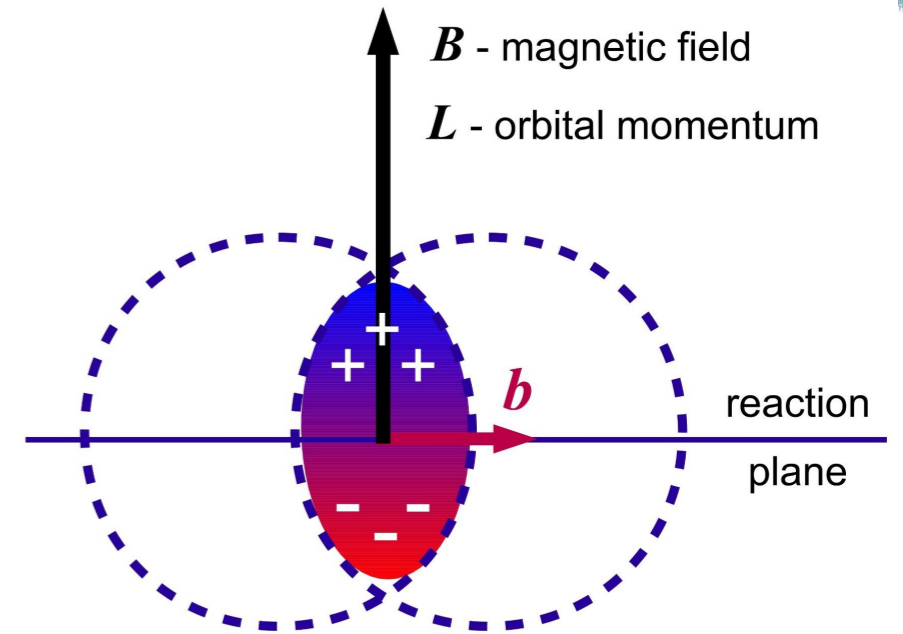
<https://drupal.star.bnl.gov/STAR/presentations>

Outline

- Introduction
- STAR Experiment
- Analysis Details
- Results
- Summary

Chiral Magnetic Effect?

- Strong magnetic field created by the fast-moving spectator protons causes the charge separation perpendicular to the reaction plane, known as the CME [1].



- The STAR at RHIC and the ALICE at the LHC have studied the CME by measuring the γ -correlator ($\gamma = \langle \cos(\phi_a + \phi_b - 2\Psi_{RP}) \rangle$) [2].

[1] K. Fukushima, D. E. Kharzeev and H. J. Warringa, Phys. Rev. D 78, 074033 (2008).

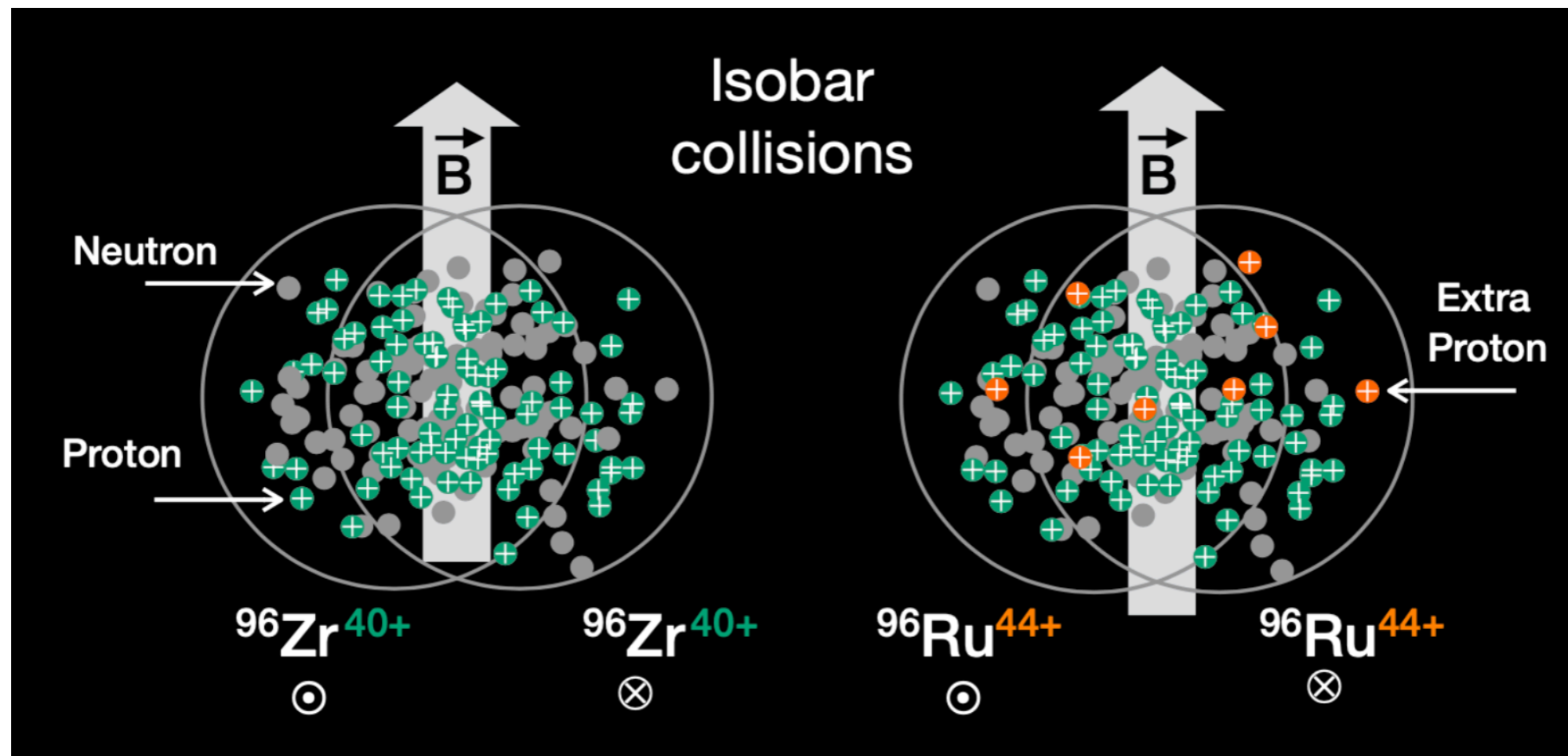
[2] S. Voloshin, Phys. Rev. C70, 057901 (2004).

B. Abelev et al., (ALICE Collaboration), PRL 110, 012301 (2013)

Introduction

Isobar Collisions

- Expected enhanced CME effect in Ru+Ru collisions than Zr+Zr collisions.
- The magnetic field is $\sim 10-18\%$ larger in Ru+Ru collisions



P. Tribedy, Free meson seminar, TIFR, Oct 7th, 2021

STAR Experiment

Two main detectors used in STAR for particle identification:

- **Time Projection Chamber (TPC)**
- **Time of Flight (TOF)**

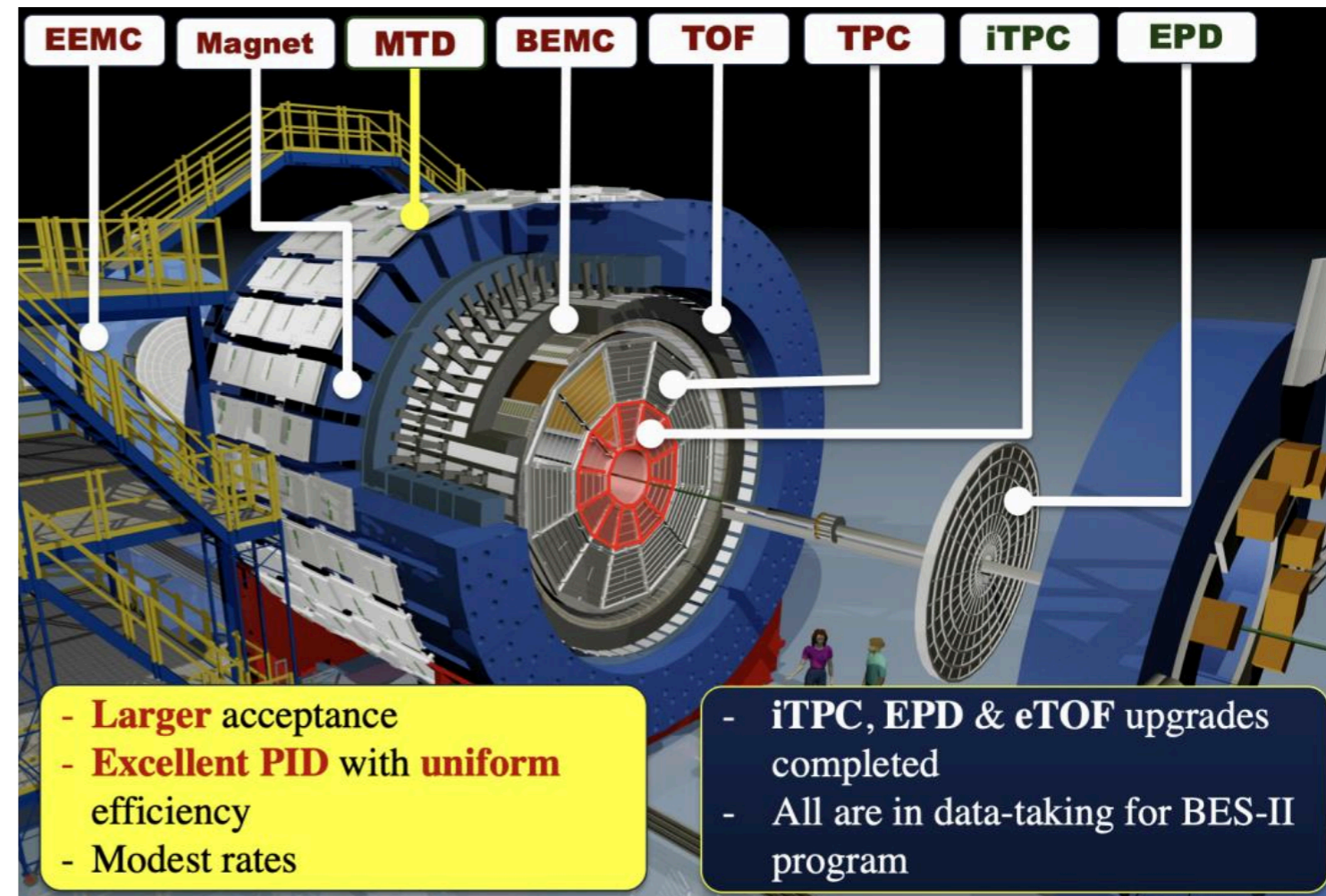
The main characteristics of the STAR:

- Large coverage i.e., $\phi(0,2\pi)$ and $\eta(-1,1)$
- Excellent particle identification at low p_T using TPC and at intermediate p_T using TOF

Data Set: Isobaric Collisions (**Ru+Ru & Zr+Zr**) at 200 GeV (~50% of the available data).

Event and Track selection cuts:

- $-35 < V_z < 25$ cm
- $|\eta| < 1$
- $0.2 < p_t < 2.0 \text{ GeV}/c$
- $\text{DCA} < 3$ cm



STAR (Solenoidal Tracker at RHIC)

Analysis Details

Sliding Dumbbell Method

- The azimuthal plane in each event is scanned by sliding the dumbbell of $\Delta\phi = 90^\circ$ in steps of $\delta\phi = 1^\circ$ while calculating, Db_{+-} for each region to obtain maximum values of Db_{+-} (Db_{+-}^{max}) in each event.

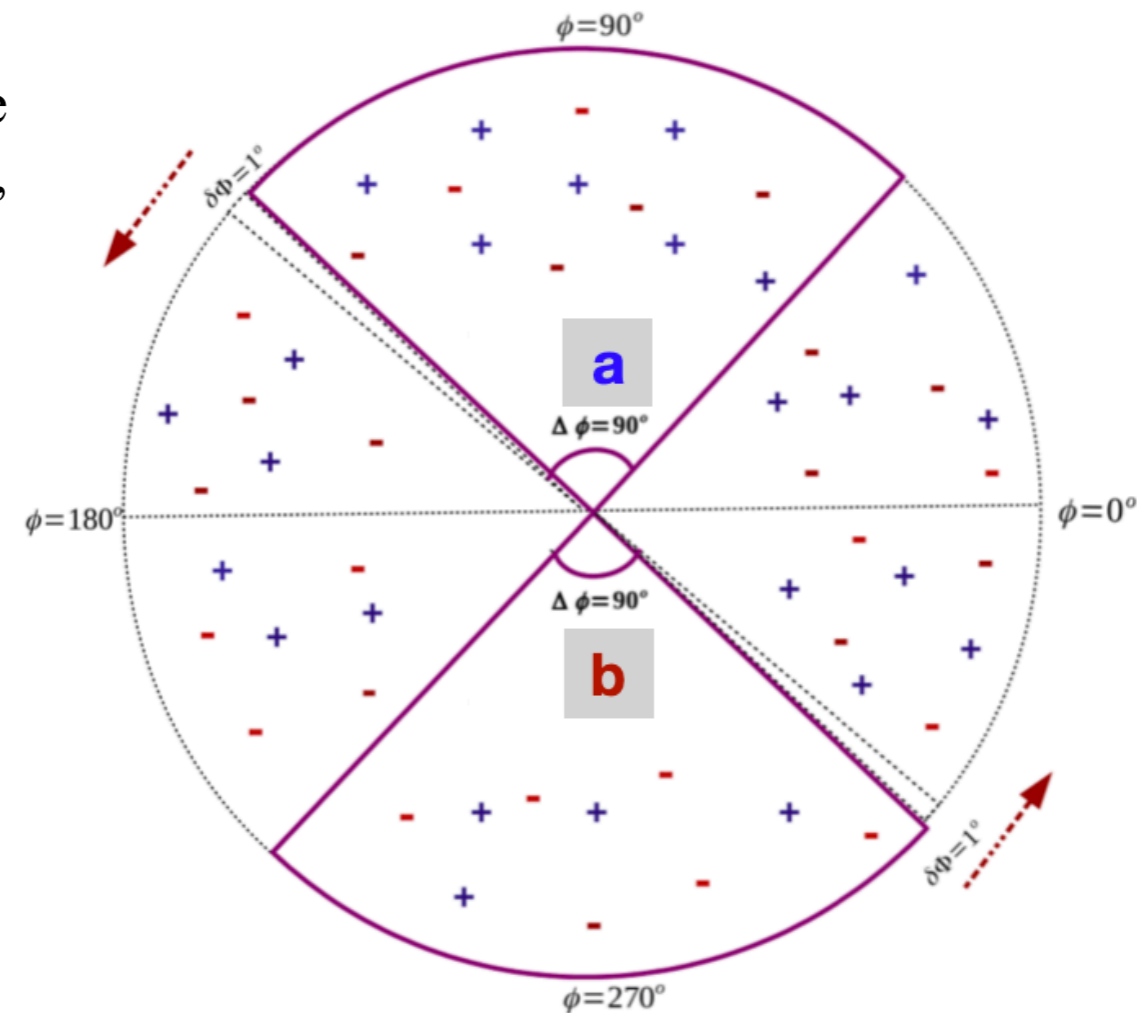
$$Db_{+-} = \frac{n_+^a}{(n_+^a + n_-^a)} + \frac{n_-^b}{(n_+^b + n_-^b)}$$

Where, n_+^a and n_-^a (n_+^b and n_-^b), the number of positive and negative charged particles on “a”(“b”) side of the dumbbell

- Fractional Charge separation (f_{DbCS}) across the dumbbell in each event is defined as :

$$f_{DbCS} = Db_{+-}^{max} - 1$$

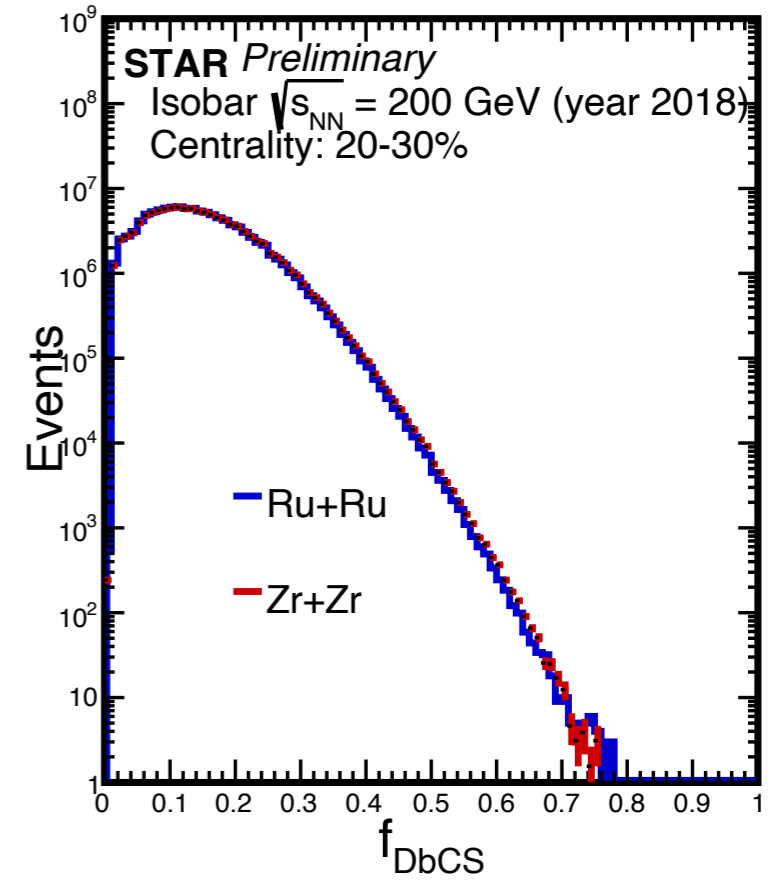
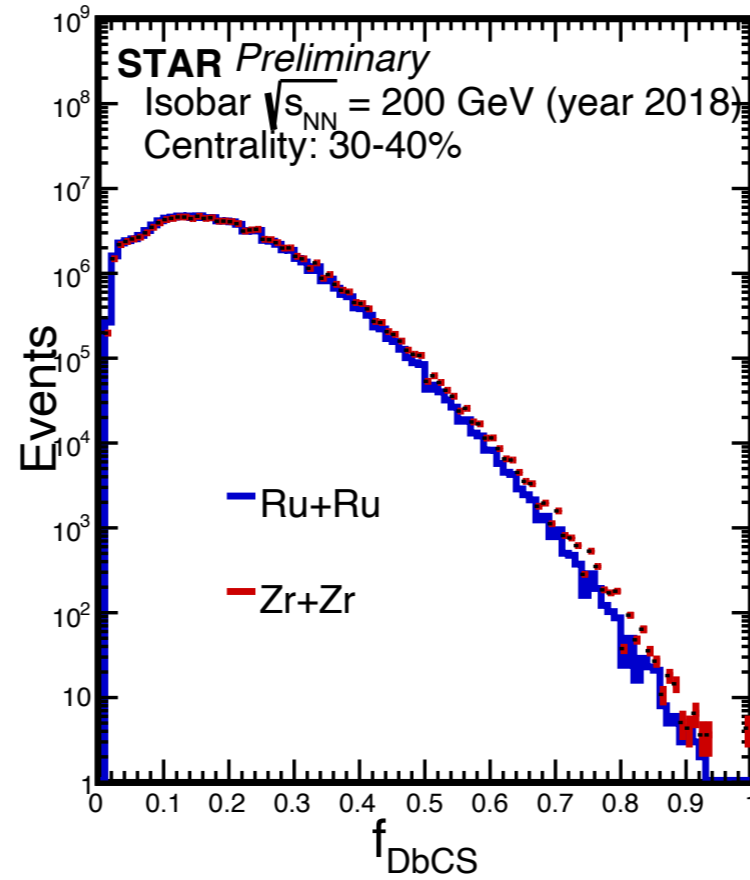
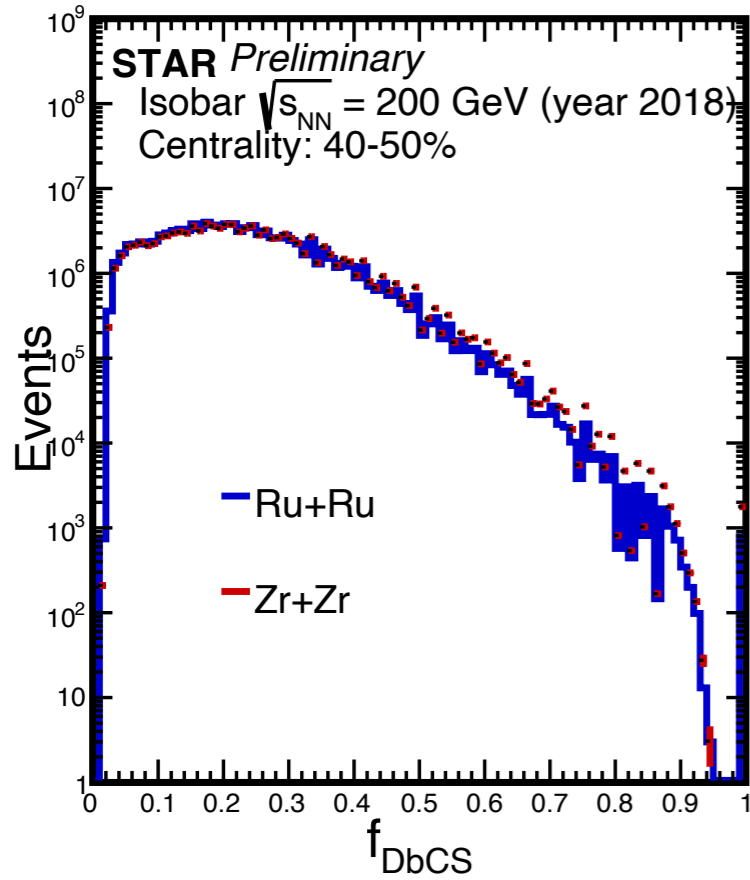
- f_{DbCS} distributions are obtained for different collision centralities and divided into 10-percentile bins.



Background Estimation

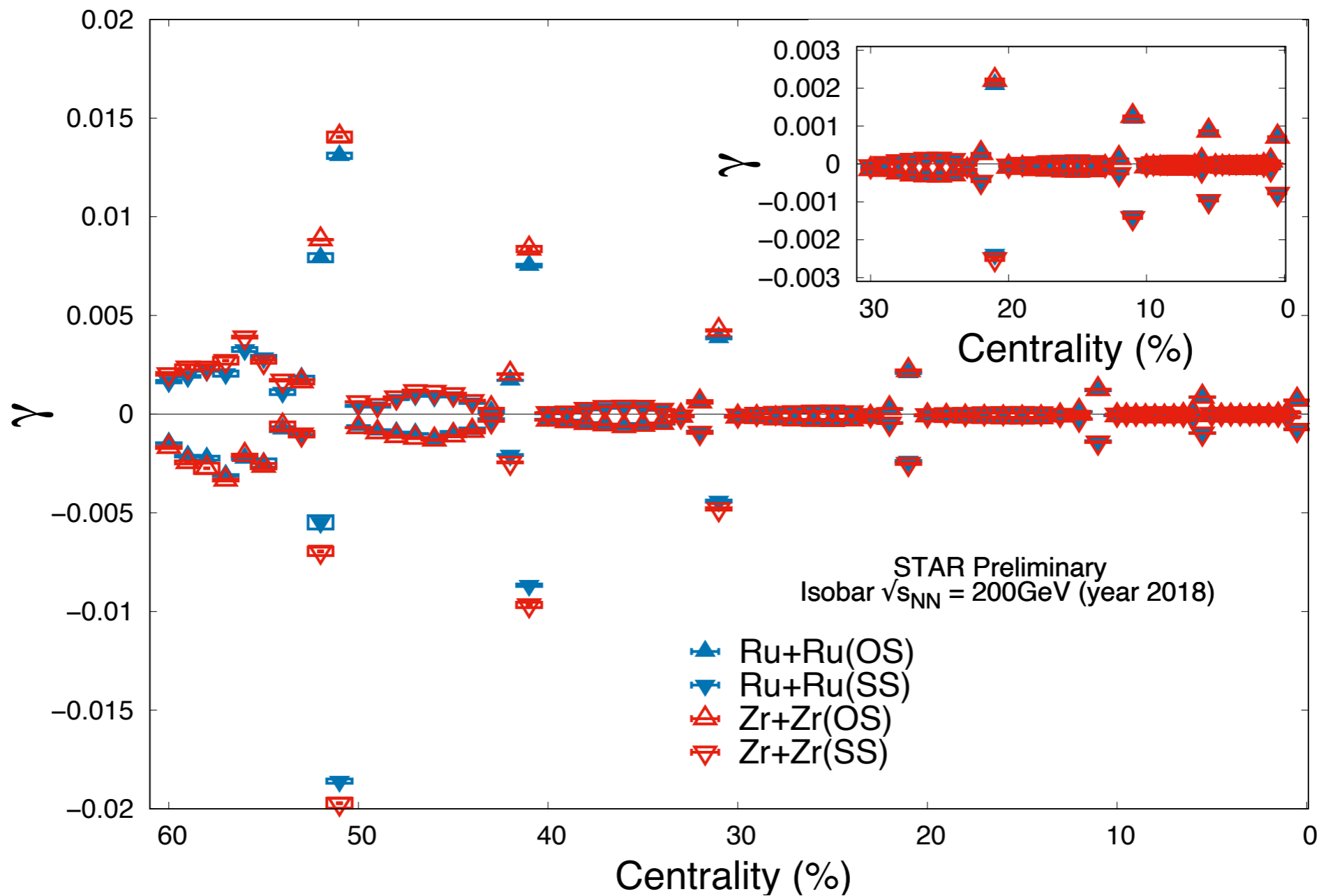
- **Charge Shuffle (ChS):** The charges of particles in each event are shuffled randomly to destroy the charge-dependent correlations amongst charged particles but keeping θ and ϕ of each particle unchanged in an event.
- **Correlated (Corr.) Background:** The shuffling of charges of particles in an event keeping the flow in, kills not only the CME-like correlations but also correlations amongst produced particles in an event. Correlations amongst particles that were destroyed during charge shuffling were recovered from the corresponding original events in a particular f_{DbCS} bin. This is termed as the correlated background.

Results



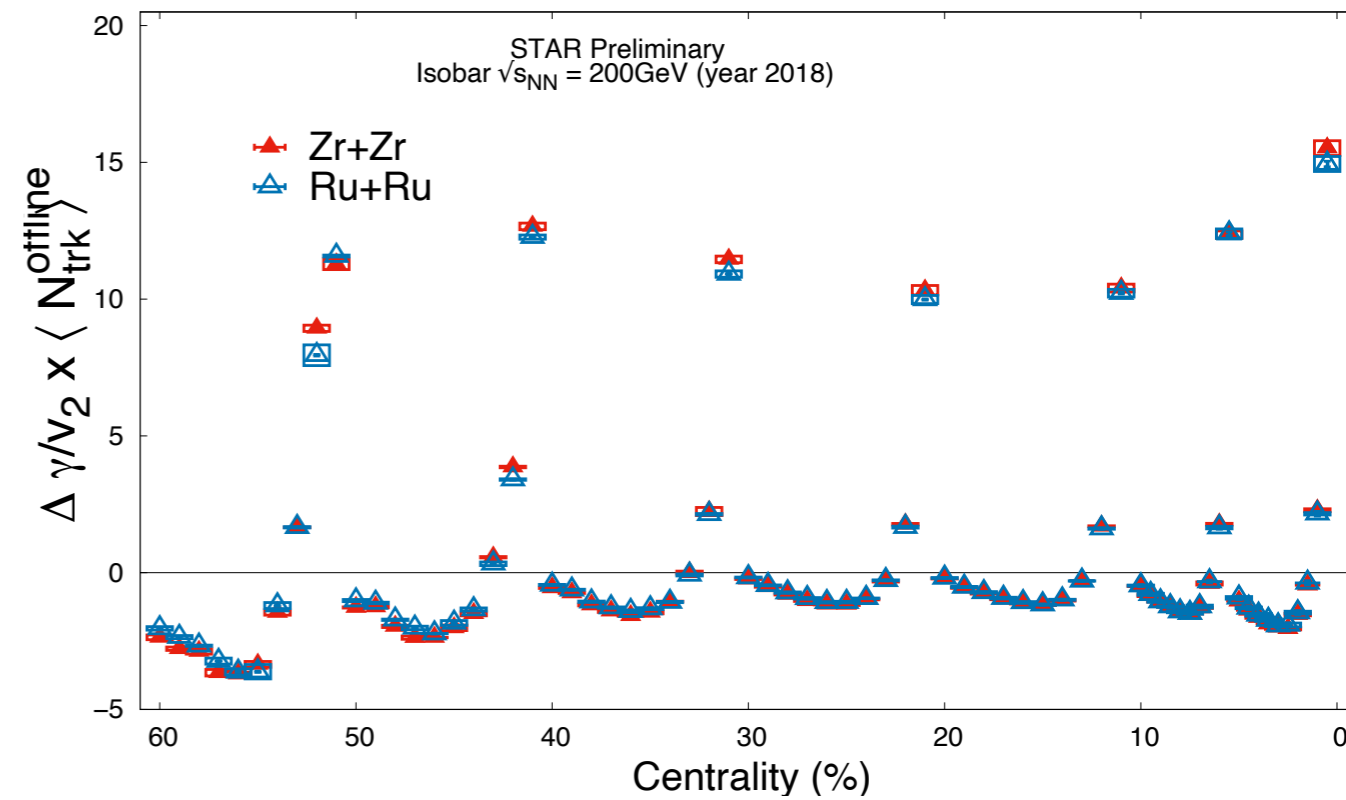
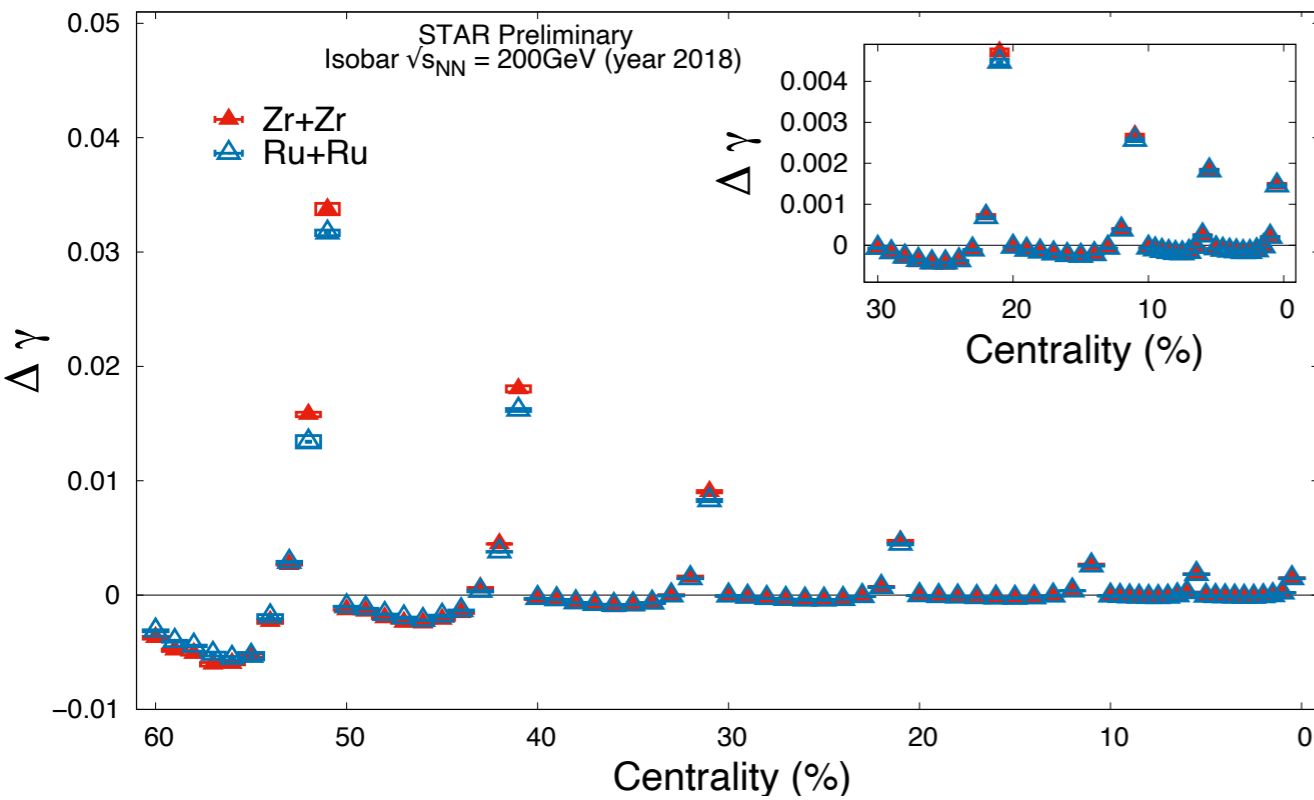
- Charge separation (f_{DbCS}) distributions extend towards higher f_{DbCS} values with decreasing collision centrality.

Results



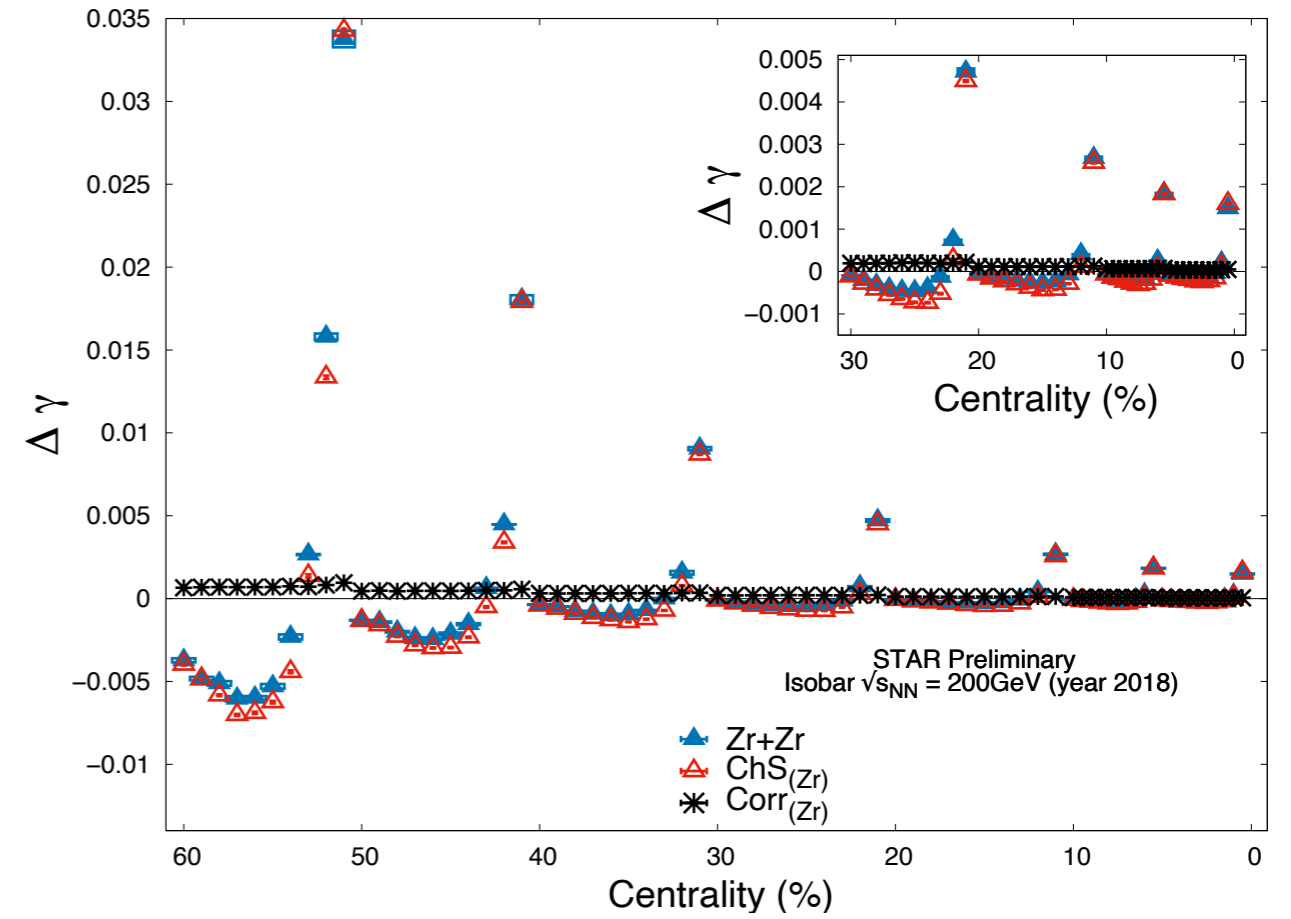
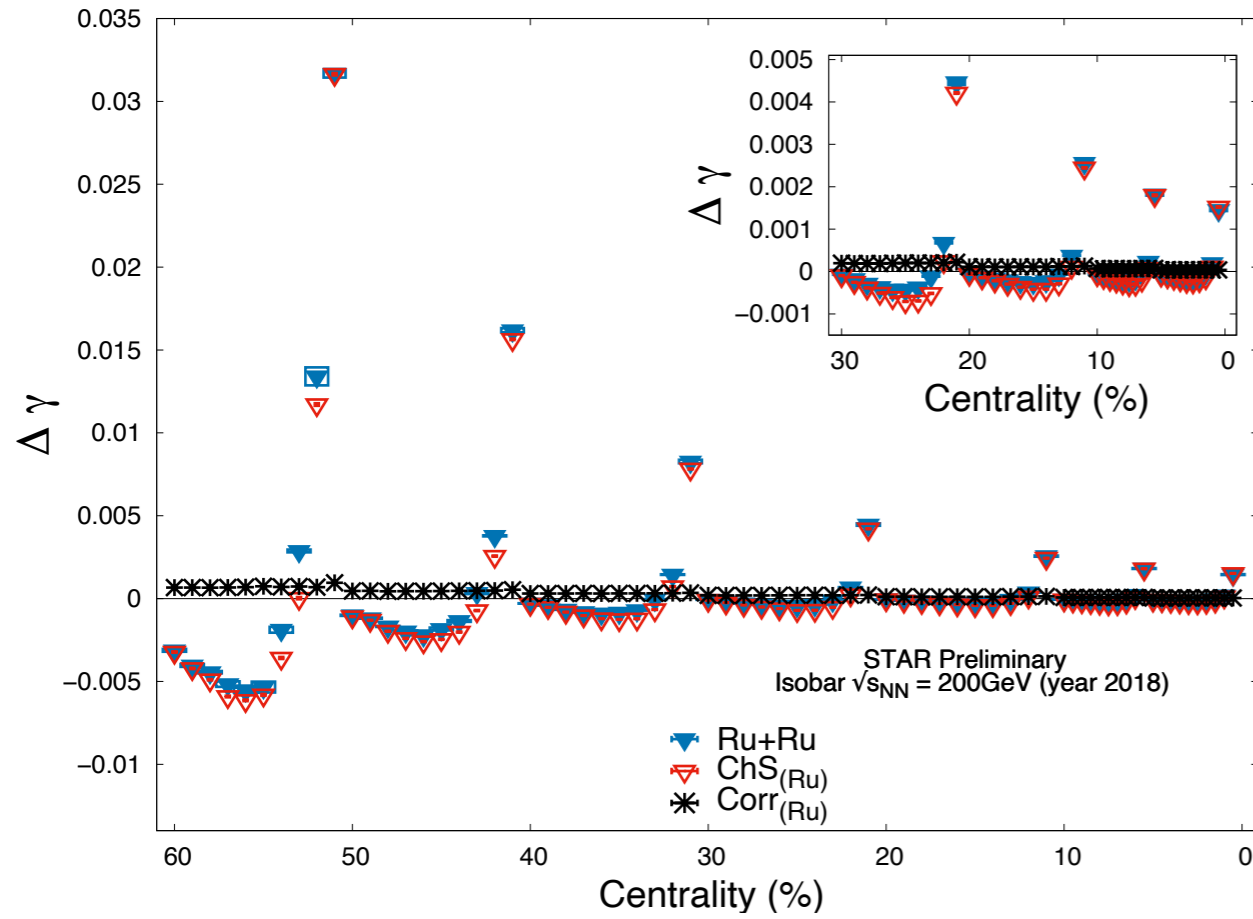
- For top f_{DbCS} bins (0-20%), $\gamma_{OS} > 0$ and $\gamma_{SS} < 0$.

Results



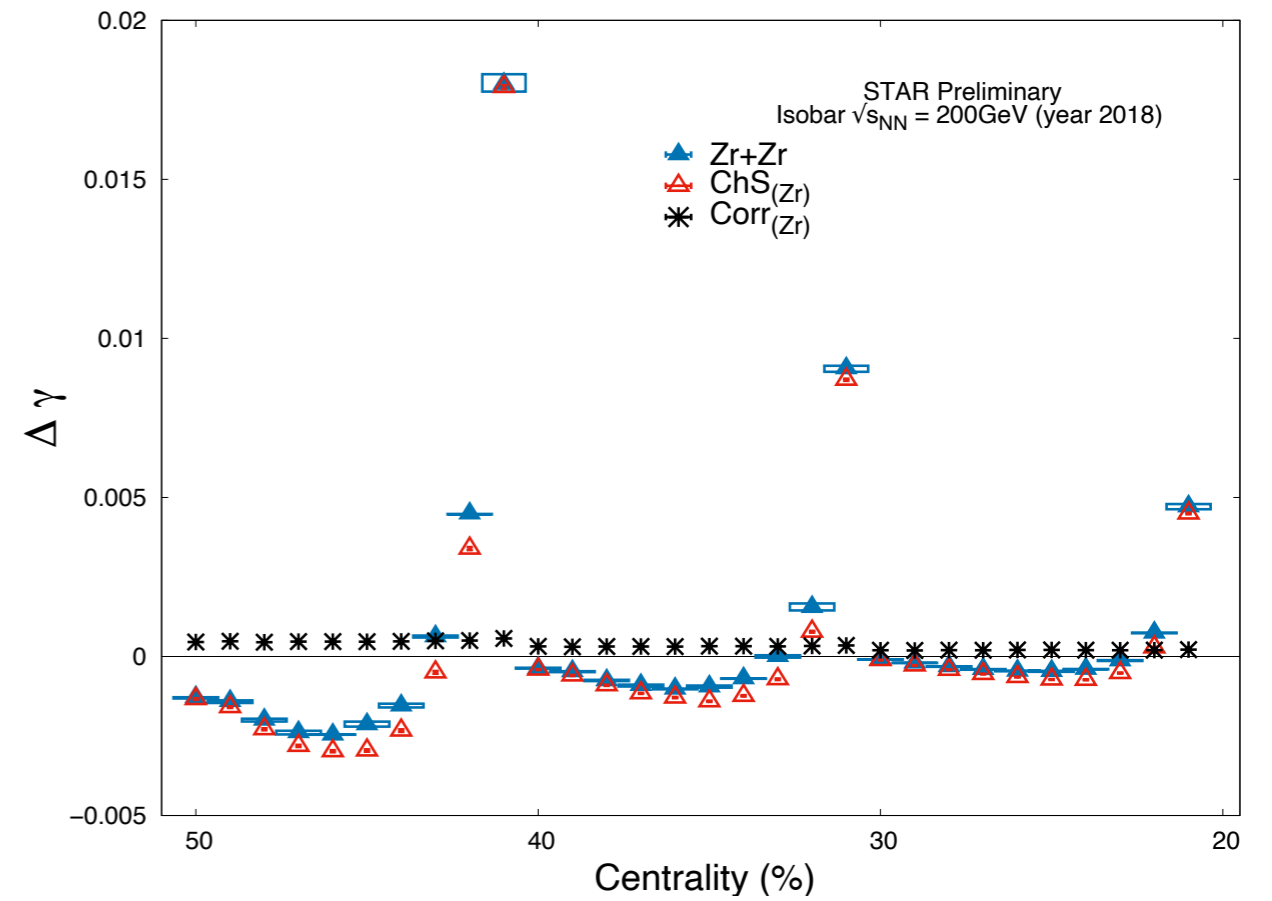
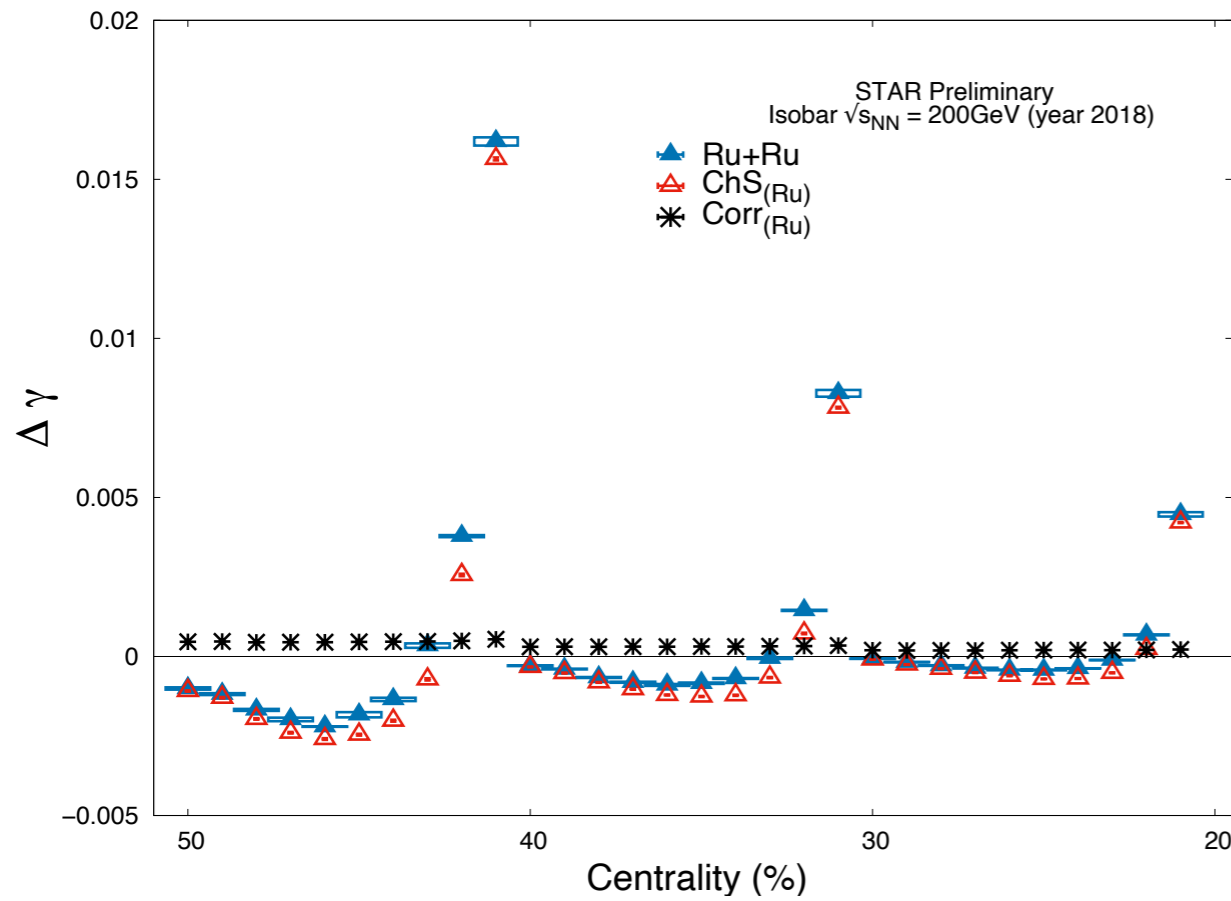
- $\Delta\gamma$ ($= \gamma_{OS} - \gamma_{SS}$) is positive for the top 20%(30%) f_{DbCS} .
- $\Delta\gamma$ is smaller for Ru than those of Zr.
- $\Delta\gamma$ scaled with $\langle N_{trk}^{offline} \rangle/v_2$ (Left figure).

Results



- $\Delta\gamma$ for Ru+Ru and Zr+Zr collisions are compared with their respective backgrounds for each f_{DbCS} bin.

Results



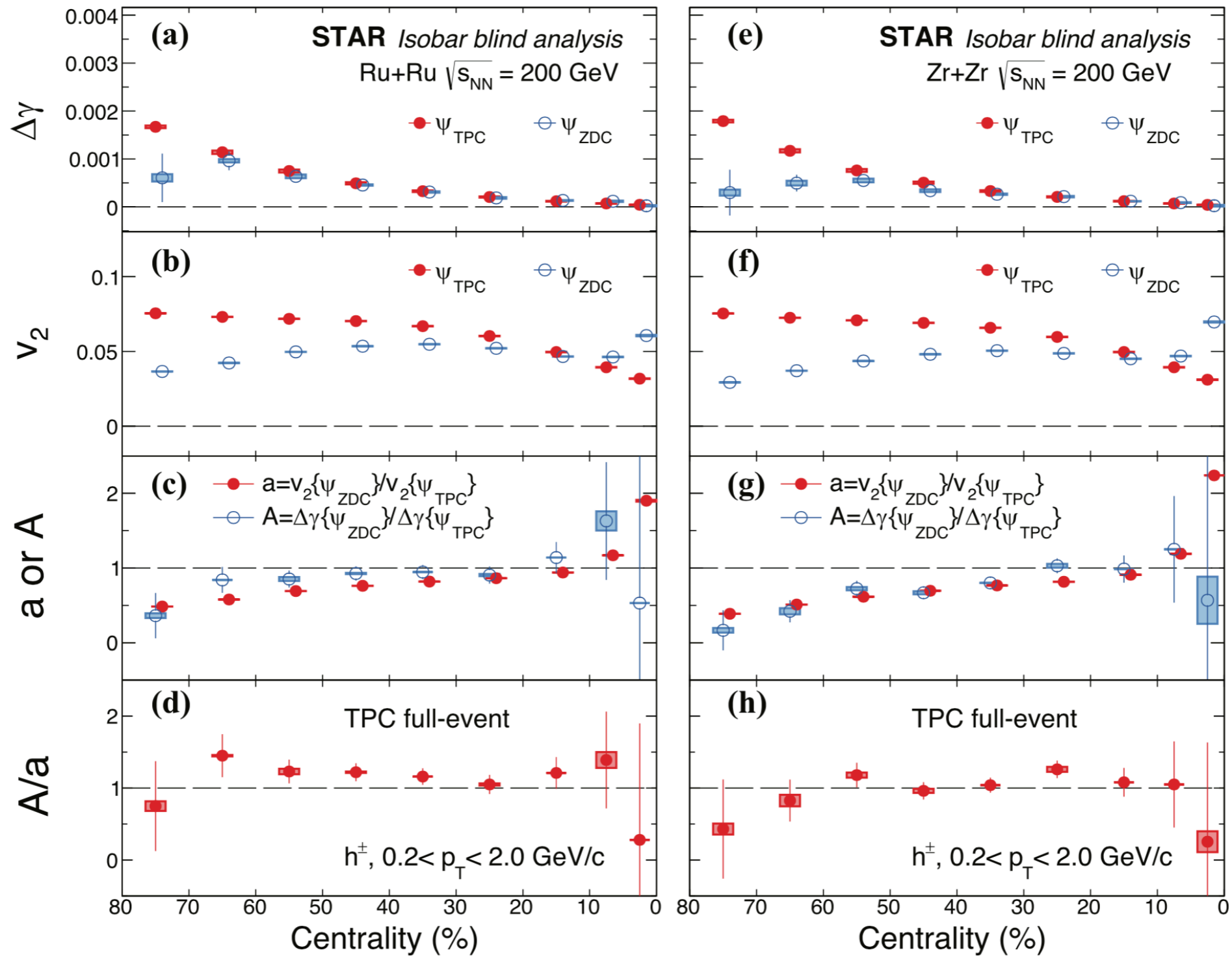
- The data points for the top 20% f_{DbCS} bins look higher than the total background (ChS+Corr) for the 30-50% collision centralities.

Summary

- ◆ The charge separation (f_{D_bCS}) distribution extends towards higher f_{D_bCS} values with decreasing collision centrality.
- ◆ It is seen that $\gamma_{OS} > 0$ and $\gamma_{SS} < 0$ for the top 20%(30%) f_{D_bCS} bins for 0-40% (40-60%) centralities.
- ◆ It can be seen that $\Delta\gamma$ are smaller for Ru than those of Zr for the top 10% (top 20%) f_{D_bCS} bins for 20-40% (40-60%) centralities. However, the difference between them decreases if $\Delta\gamma$ is scaled with $\langle N_{trk}^{offline} \rangle / v_2$.
- ◆ $\Delta\gamma$ for Ru+Ru and Zr+Zr are compared with their respective backgrounds (i.e., Charge shuffled (ChS) and Correlated (Corr)) for 0-60% collision centralities.
- ◆ The data points for the top 20% f_{D_bCS} bins look higher than the total background (ChS+Corr) for both Ru and Zr for the 30-50% collision centralities. We are analyzing the full available data set to get a detailed comparison.

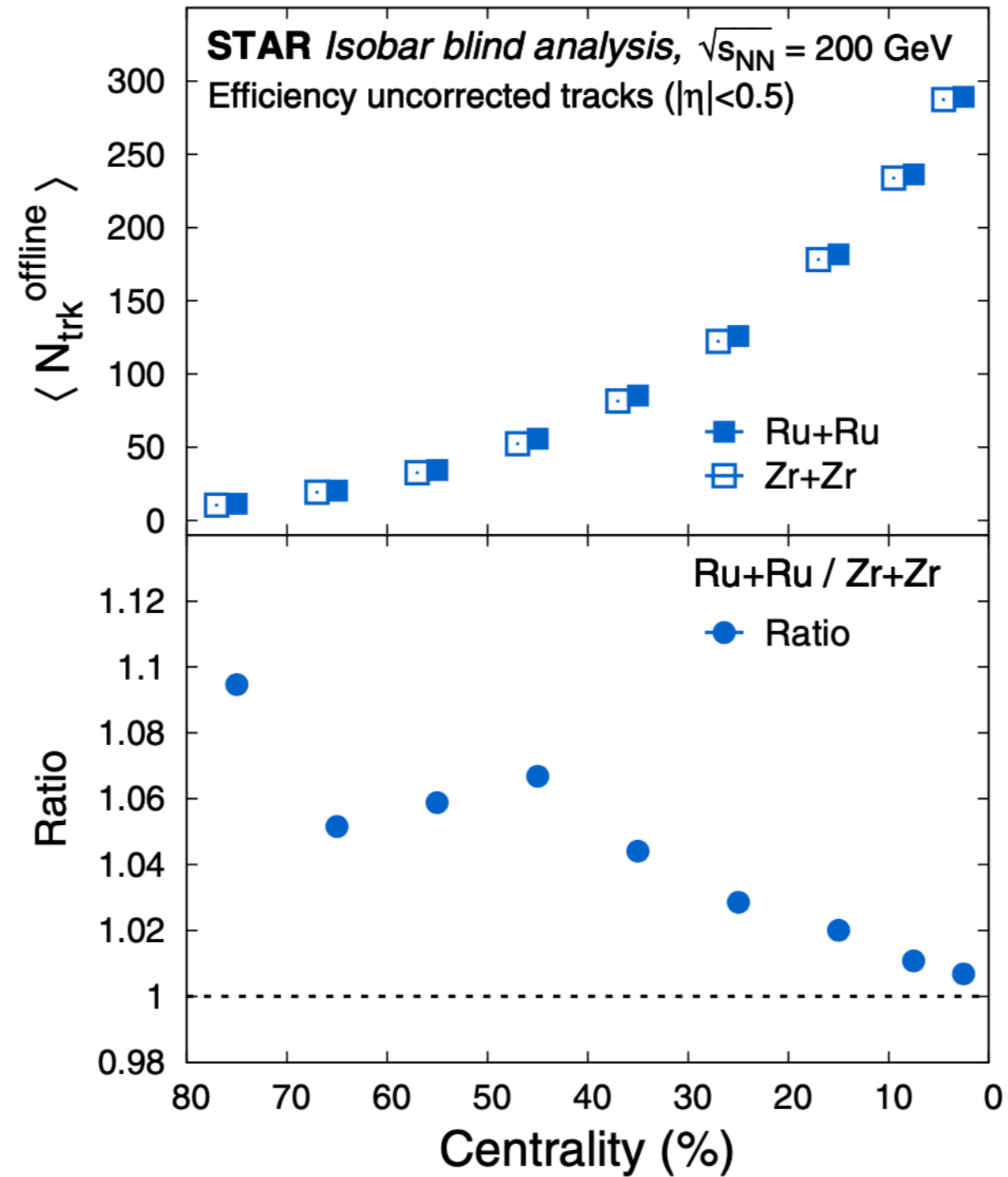
Thank you for your attention!!

Backup



M. S. Abdallah *et al.*, Phys. Rev. C 105, 014901 (2022).

Backup



M. S. Abdallah *et al.*, Phys. Rev. C 105, 014901 (2022).