

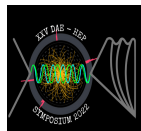
# Neutrino masses and mixing angles in the context of bilinear R-parity violating supersymmetry

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Work done in collaboration with A. Choudhury, S. Mitra and S. Mondal

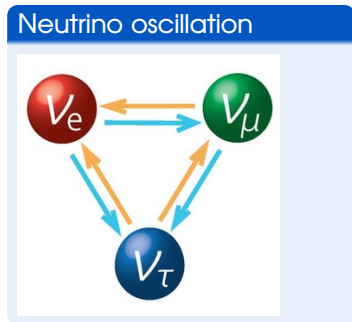
XXV DAE-BRNS HEP Symposium  
13 December, 2022



- Model definition
- Motivation of this model
- Neutrino mass generation
- Observables and constraints
- Parameters
- Analysis details
- Results and discussion

# Existence of neutrino mass

- Neutrino oscillation  $\rightarrow$  one of the most robust indications towards the existence of physics BSM
- Neutrino oscillation experiments  $\rightarrow$  Super-Kamiokande, KamLAND, RENO
- Accelerator experiments like T2K, K2K, MINOS etc.
- Neutrino oscillation  $\rightarrow$  tiny mass of neutrino and mixing
- Within Standard Model(SM) framework neutrino is massless  $\rightarrow$  no right handed neutrino
- No neutrino mass from RPC MSSM  $\rightarrow$  leads to RPV MSSM



# R-parity violating MSSM

## R-Parity violating superpotential

$$W_{\cancel{R_p}} = \varepsilon_i L_i H_u + \lambda_{ijk} L_i L_j \bar{e}_k + \lambda'_{ijk} L_i Q_j \bar{d}_k + \lambda''_{ijk} \bar{u}_i \bar{d}_j \bar{d}_k$$

## Motivation of considering only bilinear term

- We can generate neutrino mass by considering  $\lambda$  and  $\lambda'$  couplings
- They can generate neutrino mass at loop level only not at tree level
- Also their contribution to neutrino mass are Yukawa suppressed compared to contribution from bilinear term
- That's why we consider only bilinear term for our analysis

## Bilinear R-Parity violating Superpotential

$$W_{\cancel{R}_p} = \varepsilon_i L_i H_u$$

- $\varepsilon_i \rightarrow$  Bilinear R-parity violating coupling parameter
- $L_i$  is lepton multiplet
- $H_u$  is Higgs supermultiplet

## Lagrangian and corresponding soft term

$$\mathcal{L}_{\cancel{R}_p} = [\varepsilon_i (\tilde{H}_u^0 \nu_{iL} - \tilde{H}_u^+ l_{iL})] + h.c$$

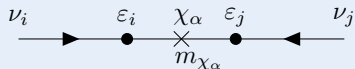
## Corresponding soft terms

$$\mathcal{L}_{soft} = B_i \tilde{L}_i H_u + h.c$$

$B_i$  is soft BRPV coupling

# Tree level contribution

## Tree level diagram

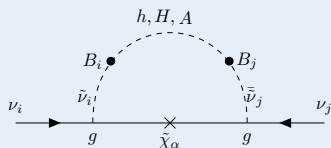


Tree level contribution to neutrino mass

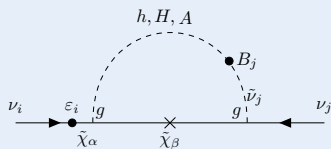
- Coupling between neutrino and up-type Higgsino
- $7 \times 7$  neutralino and neutrino mass matrix
- From the see-saw mechanism  $\rightarrow$  From high mass of neutralino neutrinos get some tiny mass
- **Only one neutrino becomes massive at tree level**

# Loop Contributions

## Loop diagrams



a)  $BB$  loop generated neutrino mass



b)  $\epsilon B$  loop generated neutrino mass

- Sneutrino mix with different Higgs  
→ produce splitting of sneutrino masses

- $BB$  loop is the dominant one

- $m_1$  and  $m_2$  become massive at loop level

- $[m_\nu]_{ij}^{(\epsilon\epsilon)} \propto \cos^2 \beta \propto \frac{1}{\tan^2 \beta}$

- $[m_\nu]_{ij}^{(BB)} \propto \frac{1}{\cos^2 \beta} \propto \tan^2 \beta$

- $[m_\nu]_{ij}^{(\epsilon B)} \propto \frac{1}{\cos \beta} \propto \tan \beta$

# Observables

- Two mass square splitting values ( $\Delta m_{21}^2$  and  $\Delta m_{31}^2$ )
- Three mixing angles ( $\theta_{12}$ ,  $\theta_{13}$ ,  $\theta_{23}$ )

## Neutrino Observables

Observable	best fit value $\pm 1\sigma$
$\Delta m_{21}^2$ ( $10^{-5} \text{eV}^2$ )	$7.50^{+0.22}_{-0.20}$
$ \Delta m_{31}^2 $ ( $10^{-3} \text{eV}^2$ ) (NH)	$2.55^{+0.02}_{-0.03}$
$ \Delta m_{31}^2 $ ( $10^{-3} \text{eV}^2$ ) (IH)	$2.45^{+0.02}_{-0.03}$
$\theta_{12}/^\circ$	$34.3 \pm 1.0$
$\theta_{13}/^\circ$ (NH)	$8.53^{+0.13}_{-0.12}$
$\theta_{13}/^\circ$ (IH)	$8.58^{+0.12}_{-0.14}$
$\theta_{23}/^\circ$ (NH)	$49.26 \pm 0.79$
$\theta_{23}/^\circ$ (IH)	$49.46^{+0.60}_{-0.97}$



## Constraints from Higgs

- Higgs Mass: From experiment mass close to 125 GeV. Considered  $\pm 3$  GeV as theoretical uncertainty.

Phys. Rev. Lett. 114 191803 (2015)

- Higgs coupling strength data at  $\sqrt{s} = 13$  TeV collected with CMS experiment  $\rightarrow$  Higgs coupling to Z, W, b, t,  $\mu$ ,  $\tau$ , and  $\gamma$  particle.

CMS-PAS-HIG-19-005, 2020

## Constraints from flavor physics

- rare  $b$ -hadron decays as  $\mathcal{B}(B \rightarrow X_s + \gamma)$  and  $\mathcal{B}(B_s^0 \rightarrow \mu^+ + \mu^-)$

Eur. Phys. J. C 81 226 (2021) and Phys. Rev. Lett. 128 041801 (2022)

- Total 15 observables

# Parameters

- Using these observables and constraints  $\rightarrow$  try to constrain the parameter space of BRPV
- Also considered the recent collider constraints also
- Considered minimal set of parameters

## List of fixed parameters

$$M_1 = 300 \text{ GeV}$$

$$M_2 = 1.1 \text{ TeV}$$

$$M_3 = 3 \text{ TeV}$$

$$M_A = 3 \text{ TeV}$$

$$M_{\tilde{q}} = 3 \text{ TeV}$$

$$M_{\tilde{l}} = 2 \text{ TeV}$$

$$A_t = -3.5 \text{ TeV}$$

- Now we have 2 MSSM parameters ( $\mu$  and  $\tan \beta$ )
- 3 BRPV coupling parameters ( $\varepsilon_1, \varepsilon_2, \varepsilon_3$ )
- 3 soft coupling parameters ( $B_1, B_2, B_3$ )
- 3 vev parameters ( $v_1, v_2, v_3$ )
- So we have total 11 free parameters

## Range of input parameters for scanning

From literature study we came up with some exhaustive range of each parameter such as

$$\mu: 1 \text{ to } 3 \text{ TeV}$$

$$\tan \beta: 1 \text{ to } 60$$

$$\varepsilon_i (i = 1, 2, 3): -1.0 \text{ to } 1.0 \text{ GeV}$$

$$B_i (i = 1, 2, 3): 0.1 \text{ GeV to } 10 \text{ TeV}$$

$$v_i (i = 1, 2, 3): 10^{-8} \text{ to } 0.1 \text{ GeV}$$

# Analysis details

- Generated model by using SARAH-4.14.5
- Output masses and mixing matrix generated by SPheno-4.0.4
- For scanning we use Markov Chain Monte Carlo(MCMC) based likelihood analysis
- Publicly available code which is a Python implementation of the ensemble sampler

Publications of the Astronomical Society of the Pacific, 125 306 (2013)

- We use a flat prior on all the parameters
- We use 500 walkers and 400 steps for each walker

# Analysis details

- We find the maximum likelihood function  $L \propto \exp(-\mathcal{L})$

- Log likelihood  $\mathcal{L} = \frac{\chi^2}{2} = \frac{1}{2} \sum_{i=1}^{n_{\text{obs}}} \left[ \frac{\Gamma_i^{\text{obs}} - \Gamma_i^{\text{th}}}{\sigma_i} \right]^2$

- $\Gamma_i^{\text{obs}}$  represents the set of  $n_{\text{obs}}$  observed data points
- The corresponding errors  $\sigma_i$  for each data
- $\Gamma_i^{\text{th}}$  is the calculated value of each observable using our theoretical model
- Maximum likelihood means we find the **minimum**  $\chi^2$  here
- Degrees of freedom(D.O.F) = 15 independent observables - 11 free parameters = 4

# Results and discussion for NH scenario

## Parameters at best-fit point

$$\nu_1 = 3.37 \times 10^{-4} \text{ GeV}$$

$$\nu_2 = 4.00 \times 10^{-4} \text{ GeV}$$

$$\nu_3 = 9.66 \times 10^{-4} \text{ GeV}$$

$$\varepsilon_1 = -4.55 \times 10^{-3} \text{ GeV}$$

$$\varepsilon_2 = -8.86 \times 10^{-3} \text{ GeV}$$

$$\varepsilon_3 = -3.06 \times 10^{-2} \text{ GeV}$$

$$B_1 = 546.42 \text{ GeV}$$

$$B_2 = 268.79 \text{ GeV}$$

$$B_3 = 1932.49 \text{ GeV}$$

$$\mu = 1248.78 \text{ GeV}$$

$$\tan \beta = 12.48$$

- We have got  $\chi_{min}^2 = 3.47$  and  $\frac{\chi_{min}^2}{D.O.F} = 0.87$
- Neutrino masses at the best-fit point are  $m_{\nu_1} = 3.41 \times 10^{-6}$  eV,  $m_{\nu_2} = 8.67 \times 10^{-3}$  eV and  $m_{\nu_3} = 5.05 \times 10^{-2}$  eV
- $\sum m_\nu = 0.059$  eV  $\rightarrow$  satisfies upper limit on the sum of neutrino masses coming from the cosmological data  $\rightarrow \sum m_\nu < 0.12$  eV

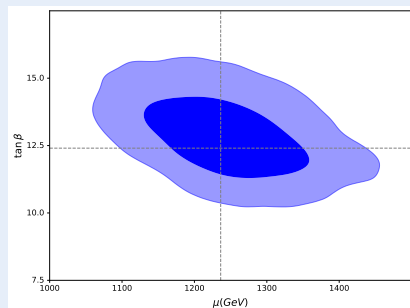
# Results and discussion

## Contribution to $\chi^2_{min}$ from each observable

	Observable	Best-fit value	$\chi^2$ contribution	Total contribution
Neutrino observables	$\Delta m_{21}^2$	$7.51 \times 10^{-5}$	$2.1 \times 10^{-3}$	$8.1 \times 10^{-2}$
	$\Delta m_{31}^2$	$2.55 \times 10^{-3}$	$2.45 \times 10^{-2}$	
	$\theta_{13}$	8.54	$7.29 \times 10^{-3}$	
	$\theta_{12}$	34.40	$1.06 \times 10^{-2}$	
	$\theta_{23}$	49.10	$3.43 \times 10^{-2}$	
Higgs Mass	$m_h$	124.70	$9.5 \times 10^{-3}$	$9.5 \times 10^{-3}$
Flavor Observables	$B \rightarrow X_s + \gamma$	$3.14 \times 10^{-4}$	1.43	1.50
	$B_s \rightarrow \mu^+ + \mu^-$	$3.21 \times 10^{-9}$	$6.9 \times 10^{-2}$	
Higgs coupling	$k_z$	1.0	0.32	1.87
	$k_W$	1.0	0.61	
	$k_b$	1.001856	0.43	
	$k_\tau$	1.001856	0.26	
	$k_\mu$	1.001856	$8.8 \times 10^{-3}$	
	$k_t$	0.9999882	$8.2 \times 10^{-3}$	
	$k_\gamma$	1.074519	0.21	

# Results and discussion

## $\mu - \tan \beta$ contour plot



The dashed line corresponds to the best-fit point

- Both are tightly constrained

- $[m_\nu]_{ij}^{(\varepsilon\varepsilon)} \propto \frac{1}{\tan^2 \beta}$
- $[m_\nu]_{ij}^{(BB)} \propto \tan^2 \beta$
- $[m_\nu]_{ij}^{(\varepsilon B)}$   $\propto \tan \beta$
- Without loop contributions  
→ no mixing of neutrinos
- From theory  $\tan \beta$  should not be very large or small
- $2\sigma$  allowed region for  $\mu \rightarrow$   
1078-1423 GeV
- $2\sigma$  allowed region for  $\tan \beta$   
→ 10.5-15



# Results and discussion

- $2\sigma$  region are:

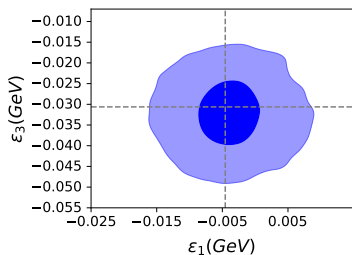
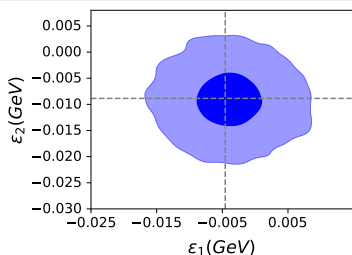
$\varepsilon_1$ :  $-1.7 \times 10^{-2}$  to  $4.0 \times 10^{-3}$  GeV

$\varepsilon_2$ :  $-2.1 \times 10^{-2}$  to  $3.2 \times 10^{-3}$  GeV

$\varepsilon_3$ :  $-4.9 \times 10^{-2}$  to  $-1.5 \times 10^{-2}$  GeV

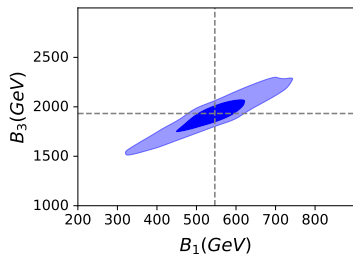
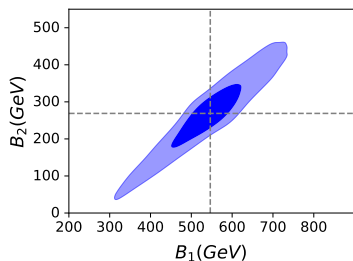
- $\varepsilon_1 - \varepsilon_2 - \varepsilon_3$  parameter space is tightly constrained
- $\varepsilon_3$  is loosely constrained as compared to other two  $\varepsilon$

$\varepsilon_1 - \varepsilon_2 - \varepsilon_3$  contour plot



# Results and discussion

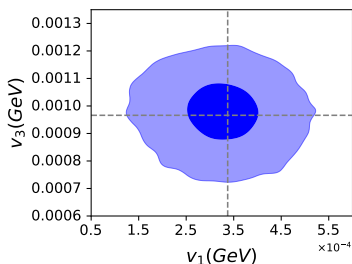
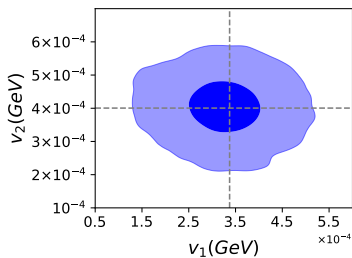
$B_1 - B_2 - B_3$  contour plot



- $2\sigma$  region are:
  - $B_1$ : 310 - 730 GeV
  - $B_2$ : 35 - 460 GeV
  - $B_3$ : 1530 - 2285 GeV
- $B_1$  and  $B_2$  have almost same allowed parameter space
- $B_3$  has large allowed parameter space
- They all are highly correlated and for large value of  $B_1$ ,  $B_2$  and  $B_3$  are also large

# Results and discussion

$\nu_1 - \nu_2 - \nu_3$  contour plot



- $2\sigma$  region are:

$$\nu_1: 1.2 \times 10^{-4} - 5.2 \times 10^{-3} \text{ GeV}$$

$$\nu_2: 2.1 \times 10^{-4} - 5.9 \times 10^{-4} \text{ GeV}$$

$$\nu_3: 7.2 \times 10^{-4} - 1.2 \times 10^{-3} \text{ GeV}$$

- $\nu_3$  parameter space is loosely constrained as compared to  $\nu_1$  and  $\nu_2$

# Results for IH scenario

The MCMC run for this scenario is still going on

## Contribution to $\chi^2_{min}$ from each observable

	Observable	Best-fit value	$\chi^2$ contribution	Total contribution
Neutrino observables	$\Delta m_{21}^2$	$7.49 \times 10^{-5}$	$6.82 \times 10^{-4}$	$4.85 \times 10^{-3}$
	$\Delta m_{31}^2$	$2.45 \times 10^{-3}$	$2.07 \times 10^{-3}$	
	$\theta_{13}$	8.56	$8.23 \times 10^{-3}$	
	$\theta_{12}$	34.42	$1.03 \times 10^{-2}$	
	$\theta_{23}$	49.33	$2.72 \times 10^{-2}$	
Higgs Mass	$m_{h_1}$	123.98	0.11	0.11
Flavor Observables	$B \rightarrow X_s + \gamma$	$3.16 \times 10^{-4}$	1.15	1.22
	$B_s \rightarrow \mu^+ + \mu^-$	$3.21 \times 10^{-9}$	0.07	
Higgs coupling	$k_z$	1.0	0.32	1.89
	$k_{W_2}$	1.0	0.61	
	$k_b$	1.001822	0.43	
	$k_\tau$	1.001822	0.26	
	$k_\mu$	1.001822	$8.84 \times 10^{-3}$	
	$k_t$	0.9999762	$8.30 \times 10^{-3}$	
	$k_\gamma$	1.077167	0.23	
Total $\chi^2_{min} = 3.28$				
$\frac{\chi^2_{min}}{D.O.F} = 0.82$				

# Conclusion

- We have considered neutrino observables along with recent higgs data and flavor physics data
- We have considered minimal set of parameter including 9 BRPV parameters along with 2 MSSM parameters( $\mu$  and  $\tan \beta$ )
- Scanned the parameter space using MCMC
- We obtained that the BRPV model can explain neutrino and other experimental data.
- We have also shown the allowed  $1\sigma$  and  $2\sigma$  region for each parameter space along with their correlation
- But the allowed parameter space is tightly constrained

**THANK YOU**