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# A Combination of Perturbative and Non-perturbative Kahler Moduli Stabilization Can Connect String Theory to Inflation

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In recent years, cosmological experiments like PLANCK-2018 [1,2] and BICEP/KECK [3] have shown the efficacy of single field slow-roll inflaton potential in explaining various experimental parameters regarding LSS, CMBR anisotropy and polarization data with significant precession. Therefore, obtaining a low energy effective inflationary theory consistent with such a class of potentials from superstring theory has been a subject of major efforts, although it is seriously at tension with swampland conjecture [4] and trans-Planckian censorship conjecture (TCC) [5]. In this paper, we have proposed that, such a connection is in principle possible in a pleasant way, if we stabilize all Kahler moduli by incorporating several perturbative and non-perturbative quantum corrections in the Kahler potential and super-potential respectively, a suitable uplifting mechanism and a novel canonical normalization technique. Our framework is based on  $10d$  type-IIB superstring theory compactified on a  $T^6/Z_N$  Calabi-Yau (CY) orientifold, equipped with three magnetized non-intersecting and intersecting  $D7$  branes,  $O7$  planes and the non-trivial quantised  $RR$  and  $NS$  closed 3-form world volume fluxes threading the 4-cycles of  $CY$ -volume. The perturbative corrections arising from  $\alpha'^3$  expansion in LVS [6], multi-graviton scattering upto one-loop with log-correction [7] and non-perturbative corrections related to  $E3$ -instanton [8] and gaugino condensation [9] break the supersymmetric no-scale structure giving an  $F$ -term  $AdS_4$  potential which is dynamically uplifted by  $D$ -term potential originating from  $U(1)$  charges of  $D7$  branes in gravitational sector thereby providing the inflaton potential after normalization. All the parameters of the derived  $dS_4$  potential are carefully tuned to maintain the inflationary plateau region. Cosmological parameters are obtained by  $k$ -space analysis of cosmological perturbations by dynamical horizon exit method [10] and found [11] to be consistent with PLANCK and BICEP/KECK constraints viz.,  $n_s = 0.9652 - 0.9662$ ,  $r = 5.8 \times 10^{-4} - 6.2 \times 10^{-4}$ ,  $N = 55.0 - 56.7$ ,  $n_t = (-7.28 \times 10^{-5}) - (-7.76 \times 10^{-5})$  at  $k = 0.001 - 0.009 \text{ Mpc}^{-1}$ .

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